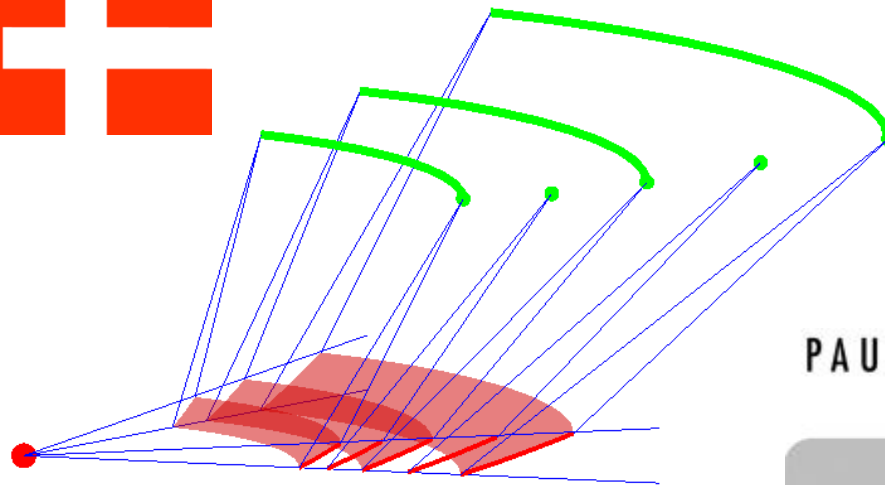
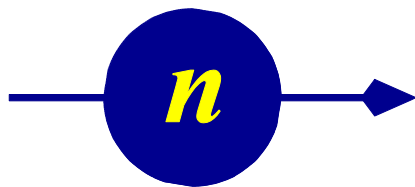


# CAMEA

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Lefmann, Niels Bech Christensen, Henrik M. Ronnow



*McStas*



PAUL SCHERRER INSTITUT



Technical University of Denmark



EUROPEAN  
SPALLATION  
SOURCE

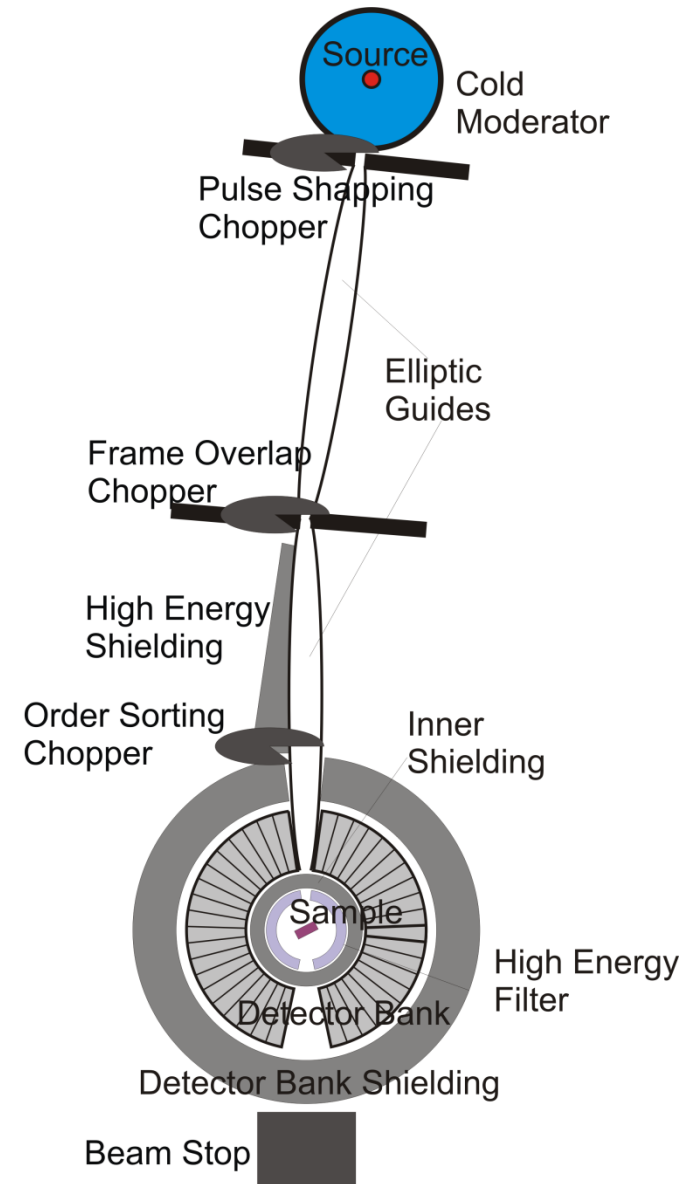


ÉCOLE POLYTECHNIQUE  
FÉDÉRALE DE LAUSANNE

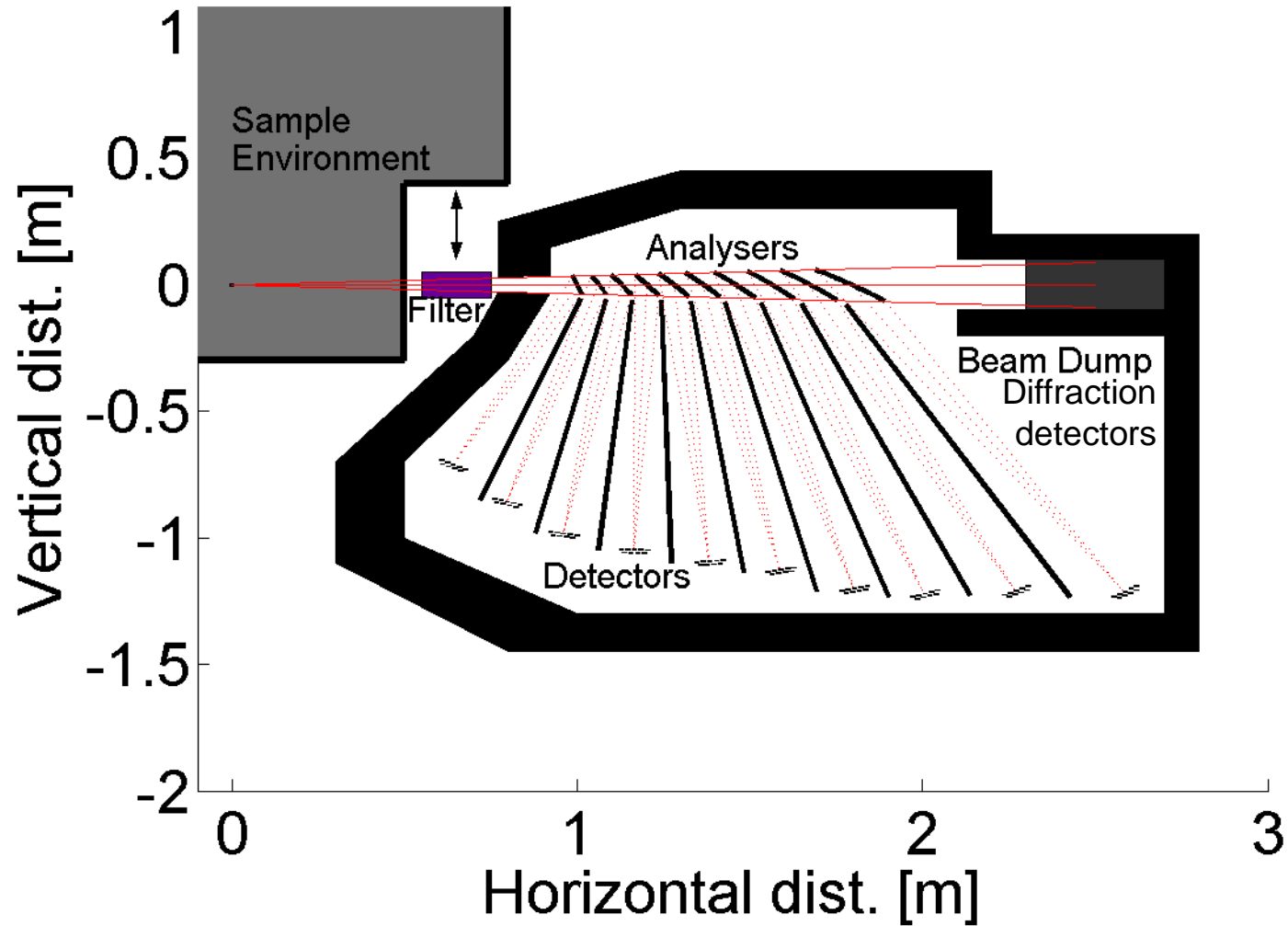
# Continuous Angle Multiple Energy Analysis

- Inverse time of flight
- Vertically scattering analyzers.
- Multiple analyzers behind each other select several energies.
- $E_f = 2\text{--}32\text{meV}$
- Energy resolution as cold TAS and medium cold TOF
- Can use full ESS long pulse shaping  $\Rightarrow$  better res
- Optimized for small sample size  $1\text{cm}^3$  down to  $1\text{mm}^3$
- High field (25T), high pressure (100kbar)
- Prototype tested at PSI-MARS

## CAMEA



# Multiple energies



# Dimensions ✓

$E_{\text{Analyser}}$ (meV)	2.5	2.8	3.1	3.5	4.0	4.5	5.0	5.5	6.5	8.0
$D_{\text{Sample-Analyser}}$ (m)	1.00	1.06	1.13	1.20	1.28	1.37	1.46	1.56	1.67	1.79
$D_{\text{Analyser-Detector}}$ (m)	0.80	0.90	1.00	1.05	1.10	1.15	1.25	1.30	1.35	1.45

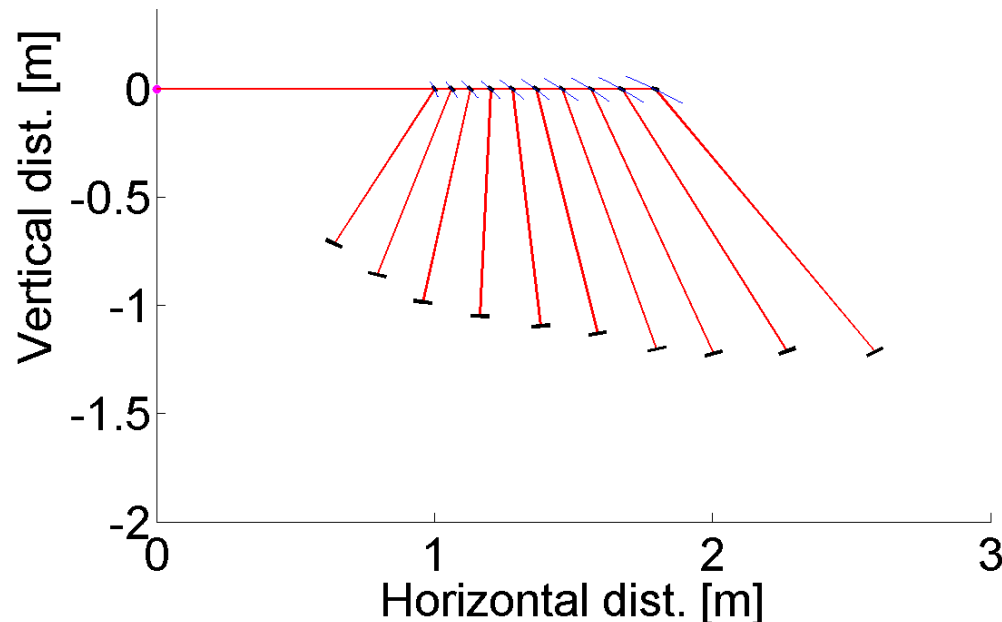
Scattering angle:  $3^\circ$ - $135^\circ$

Mosaicity: 60 minutes

Sample size:  $\frac{1}{2} \times \frac{1}{2} \times \frac{1}{2} \text{ cm}^3$

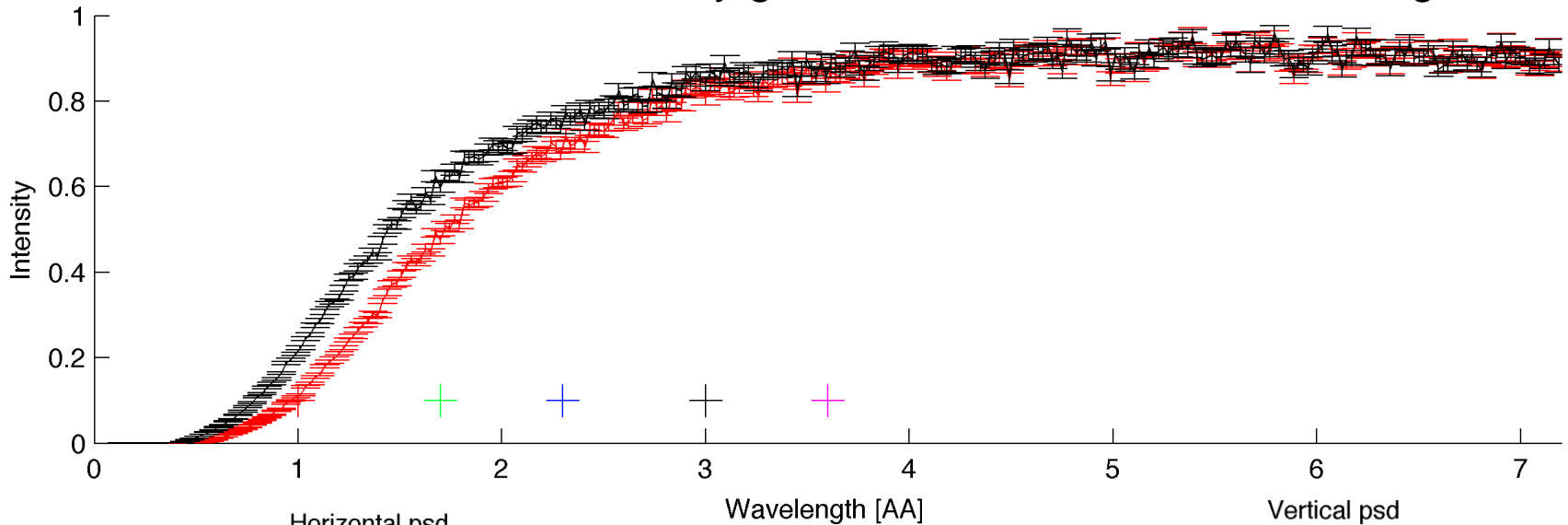
Detector width:  $\frac{1}{2}$  inch

Pixel size:  $\frac{1}{2} \text{ cm}$

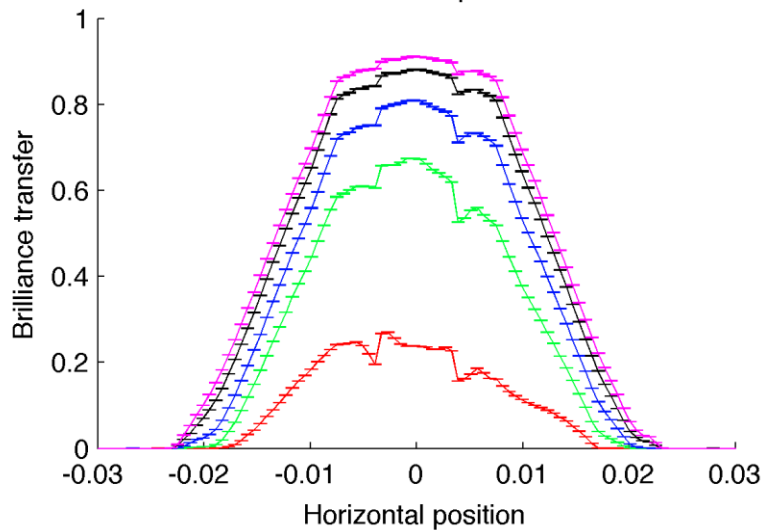


# Guide ✓

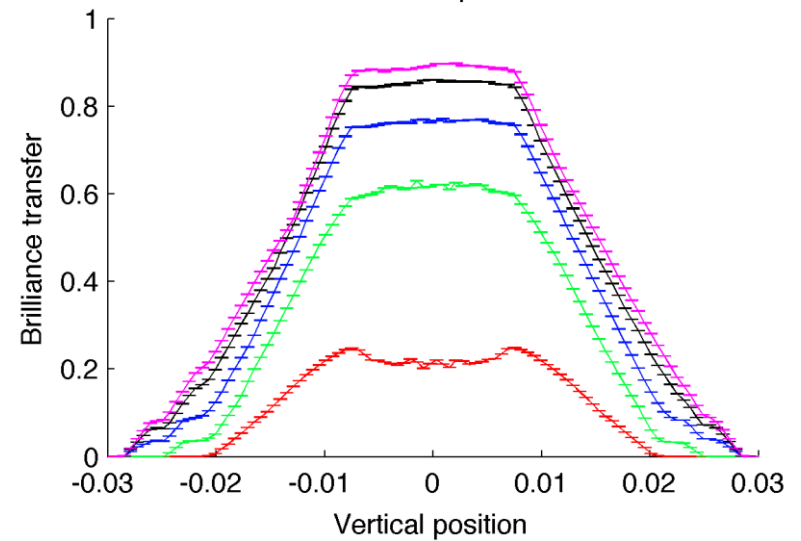
Brilliance transfer of Guide very good above 2 Å even if mirrors degrade



Horizontal psd



Vertical psd



# Choppers ✓

2 Pulse shaping  
Choppers

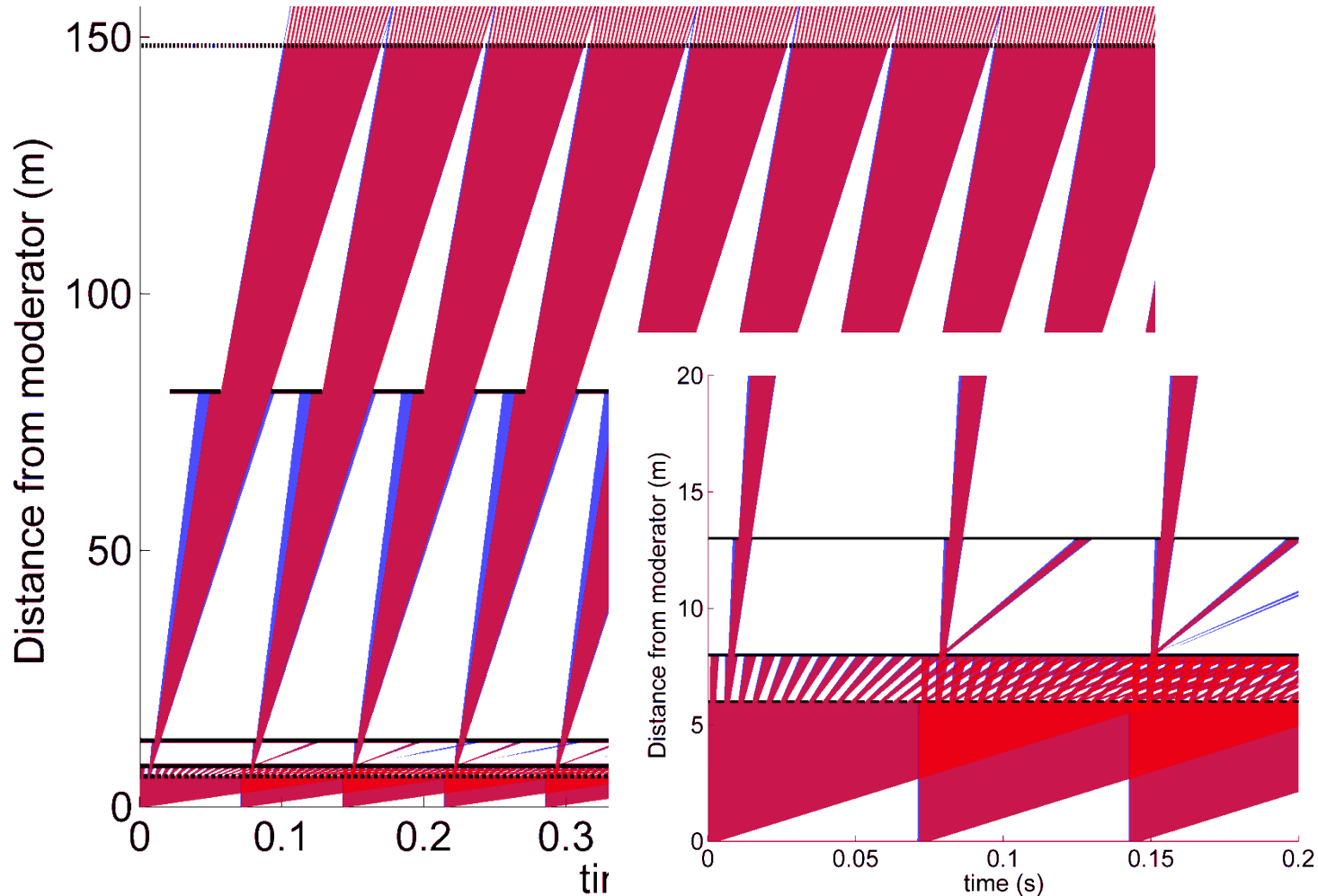
2 Frame overlap  
Choppers

1 Tail removal  
Chopper

2 Order sorting  
Choppers

Selects:

- Bandwidth range of  $E_i$
- Pulse-length resolution



# $E_i$ -band and resolution tuning ✓

Bandwidth chopper

e.g.  $\Delta\lambda = 1.8 \text{ \AA}$ , and

$E_i = 4.2 \text{ meV} - 12 \text{ meV}$

Only desired energy range hit sample

Pulse length chopper:

Full ESS long pulse

$\Rightarrow$   $\sim 4\%$  resolution

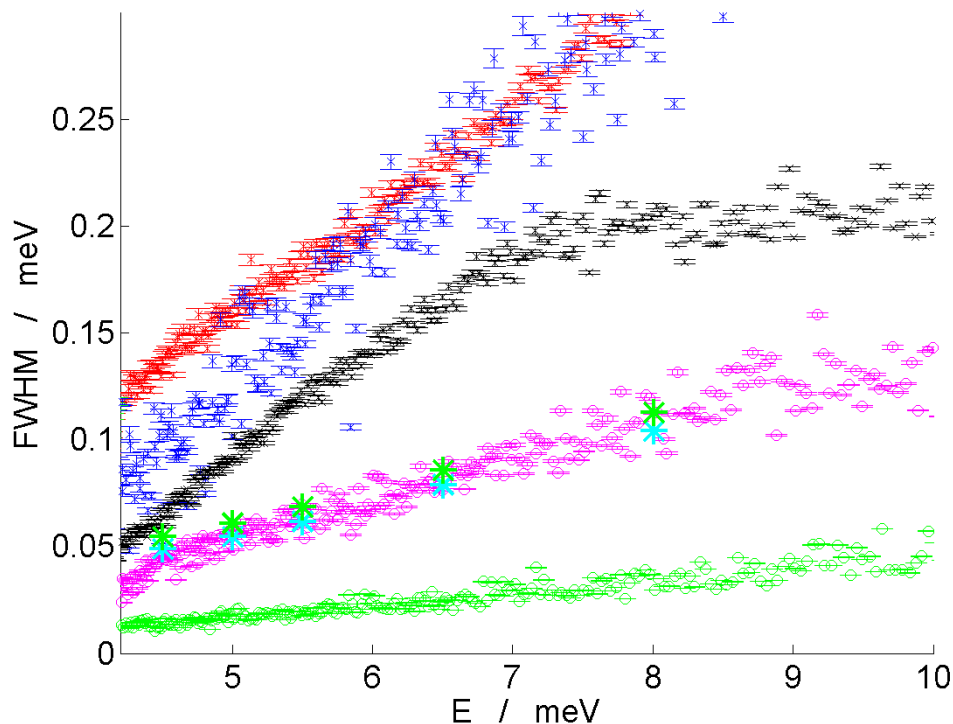
Pulse length:  $/3.3$  neutrons

$\Rightarrow$   $\sim 1.2\%$  resolution

Analysers:  $\sim 1.2\%$  resolution

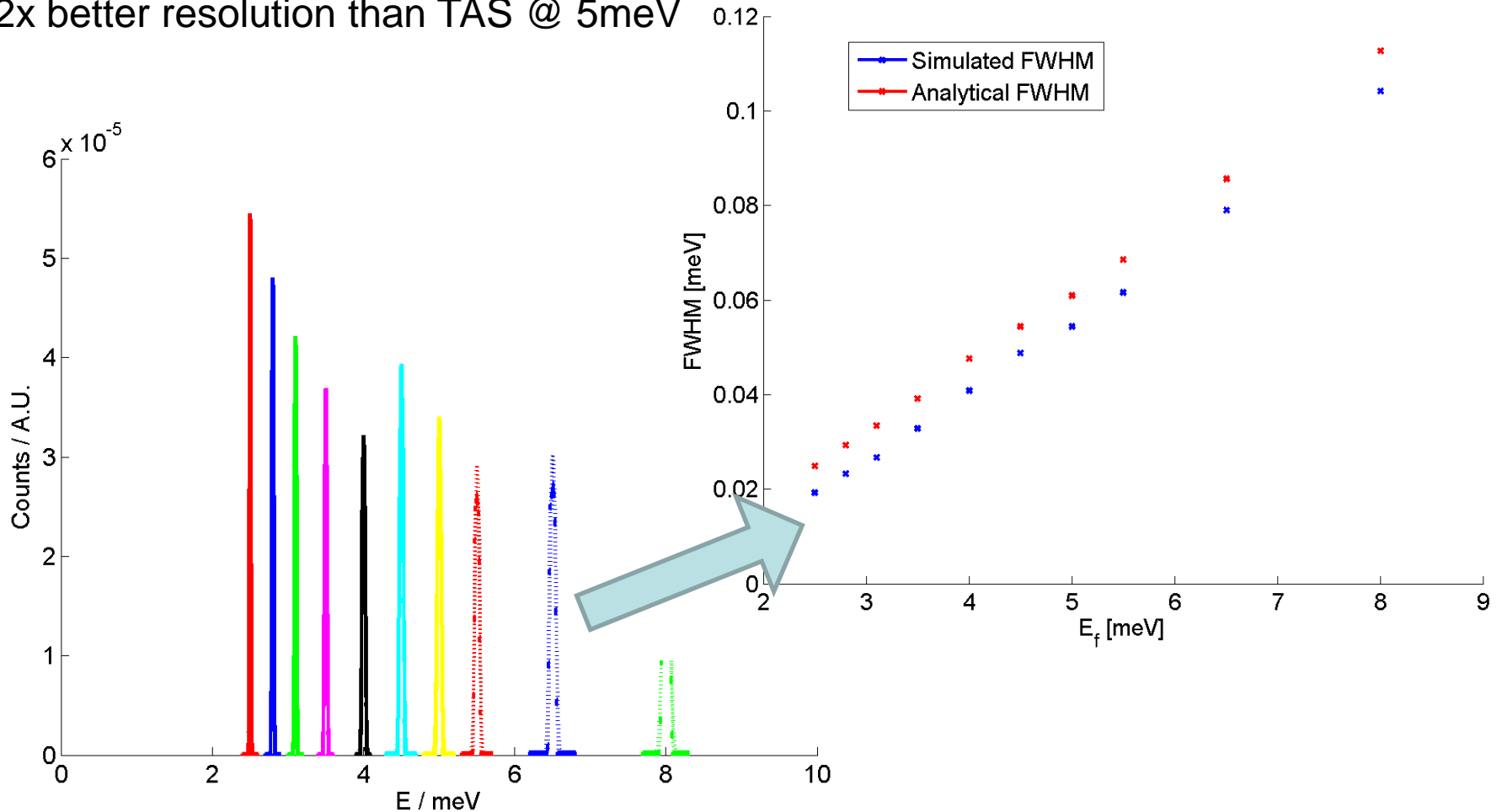
Total  $\Rightarrow$   $\sim 1.6\%$  resolution

$\sim 30\%$  better than focused TAS

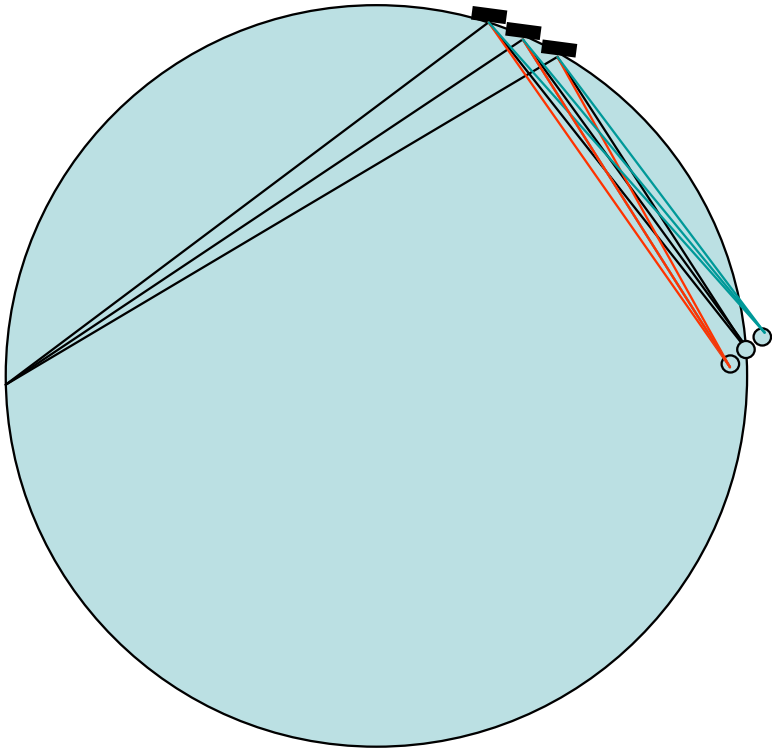


# Analyser energy resolution ✓

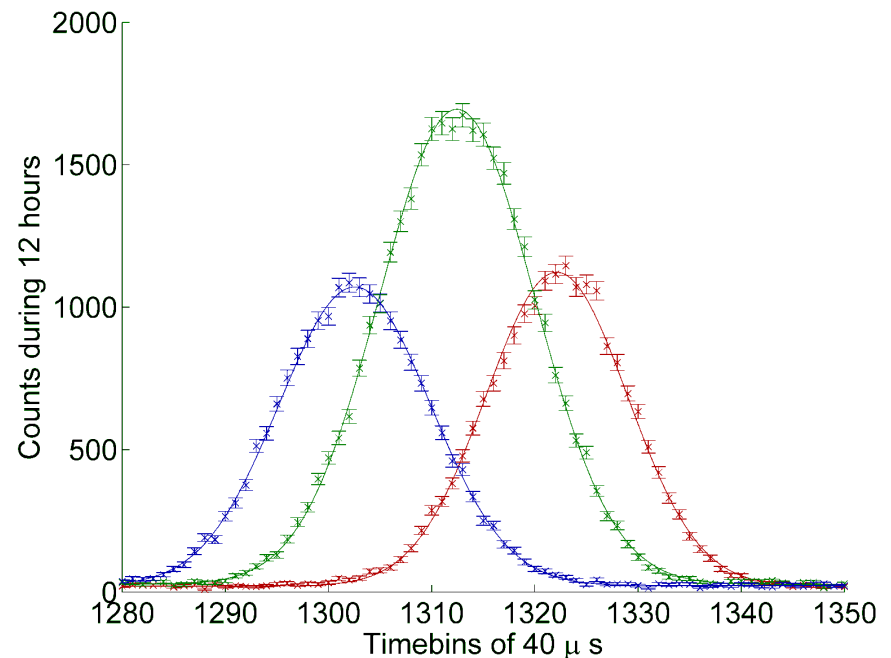
Simulated and calculated energy-resolutions of secondary spectrometer shows ~1.2%  
~2x better resolution than TAS @ 5meV



# Better $E_f$ resolution at no flux cost !



- Distance collimation  $\Rightarrow$  better resolution than standard TAS
- Multiple energies for each analyser
- Adjacent detectors record different energies  $\Rightarrow$  2.7x neutrons



# Momentum coverage ✓

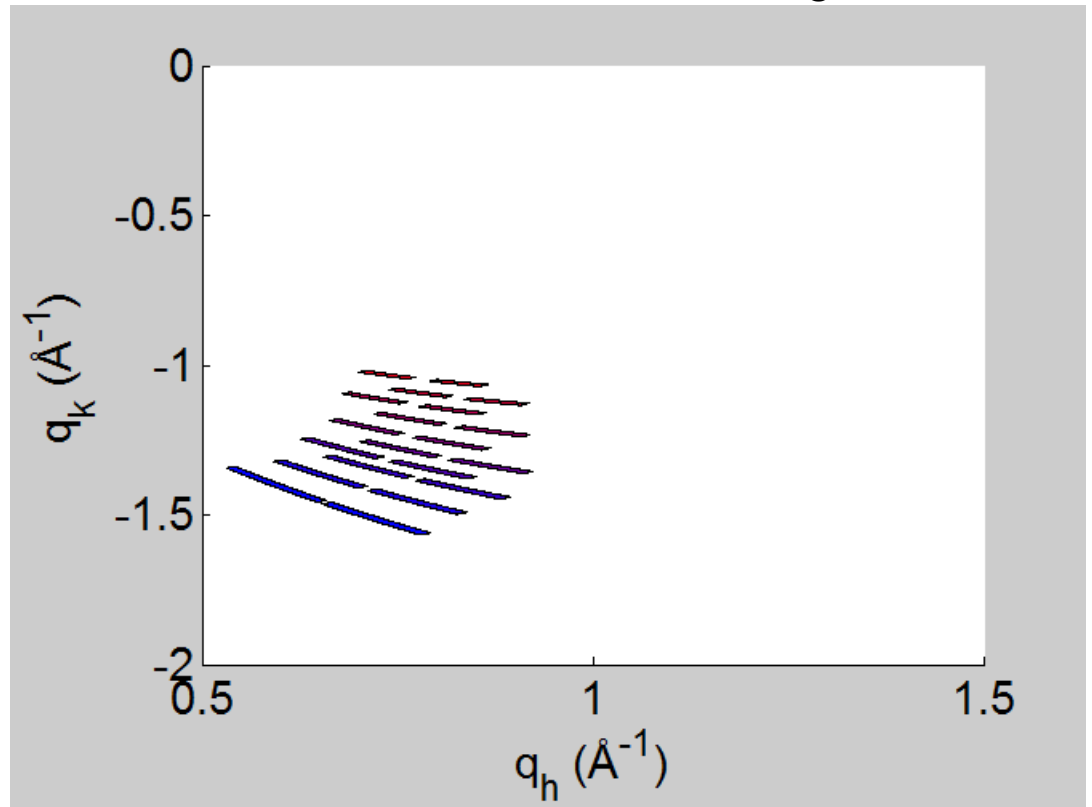
Each analyser covers  
a line in Q-plane  
(with small gaps)

Sample rotation  
covers reciprocal plane

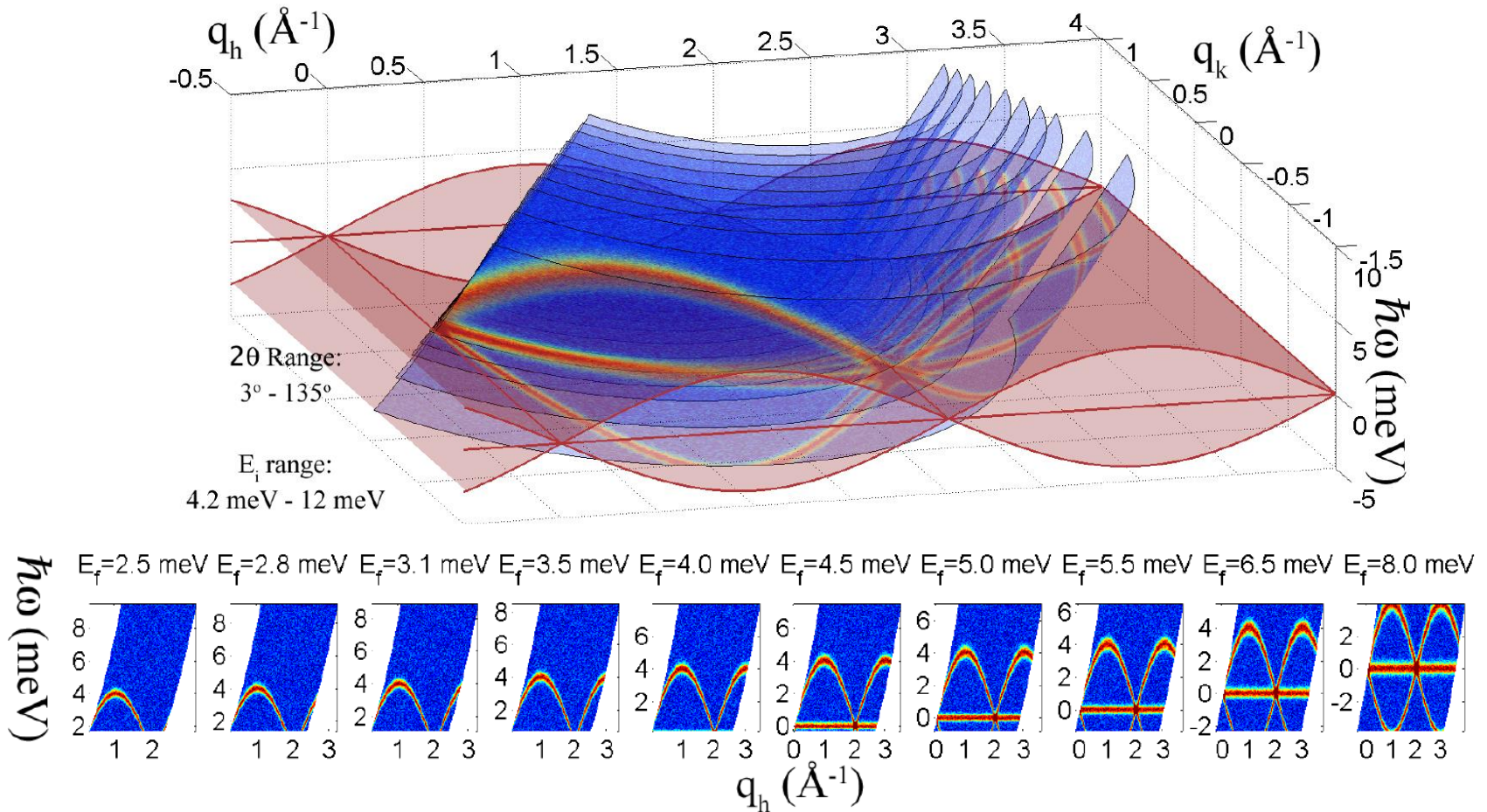
Small gaps are covered  
by adjacent energies

For completely homogeneous  
coverage, analyser tank can be  
rotated

Coverage at  $\Delta E=0$



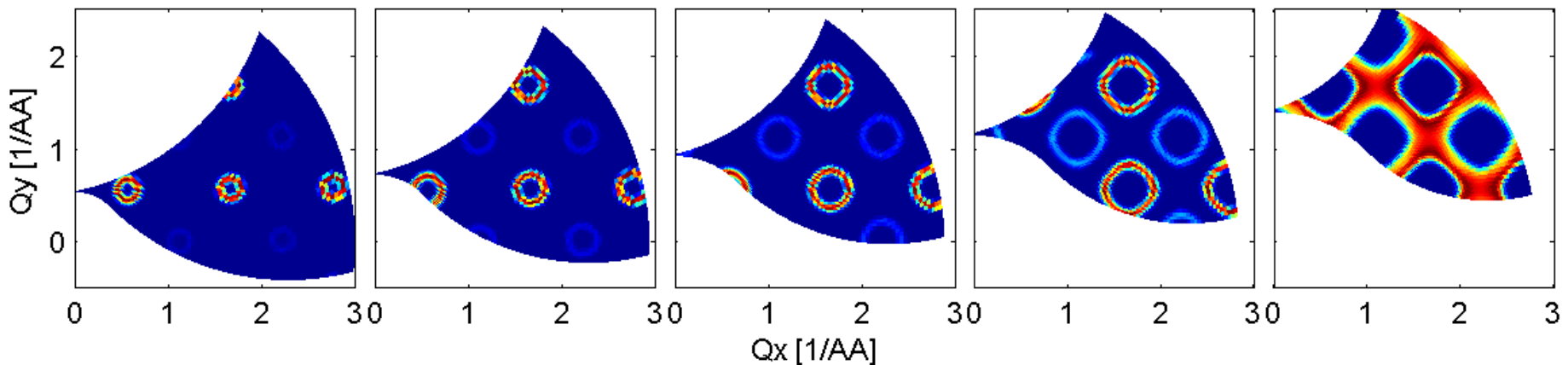
# Example



- Single acquisition gives N Q-E manifolds
- Often sufficient to determine dispersion changes
- Fast parametric studies, or sufficient stats for very small samples

# CAMEA: best possible “in-plane spectrometer”

- Direct TOF give 3D parabolic manifold of 4D  $S(q,w)$ 
  - Must analyse off-symmetry data
  - Or perform full “Horace” scan
- CAMEA: focus on flat 2D Q-plane and energy
  - Allows measuring Q-planes in 3D systems
    - ⇒ larger fraction of measured neutrons analysed
  - Several complete Q-Energy manifolds in one acquisition



# Analyzer Solid Angle

Instrument	Facility	Analyzer	Solid Angle (steradians)	$\pm 1.4^\circ$ Solid Angle (steradians)	$\pm 1.4^\circ$ Gain per analyzer
CAMEA		PG (002) or (004)	0.13 x 10	0.13 x 10	-
OSIRIS	ISIS	PG (002) or (004)	1.09	0.12	1.08
Iris	ISIS	PG (002) or (004)	0.36	0.11	1.18
PRISMA	ISIS	PG (002)	0.021 @ 5 meV	0.0147	8.8
MACS	NIST	PG (002)	0.15	0.075	1.7
Flatcone	ILL	Si(111)	0.05	0.05	2.6

- Same coverage as other indirect time of flight.
- Higher coverage than multiplexing TAS
- Multiply these numbers by the 10 analyser energies

# Flux ✓

CAMEA  $\pm 1^\circ$  vertical,  $\pm 0.75^\circ$  horizontal,  $1.4 \times 10^9 \text{ ncm}^{-1} \text{ s}^{-1}$  centered on  $3 \text{ \AA}$

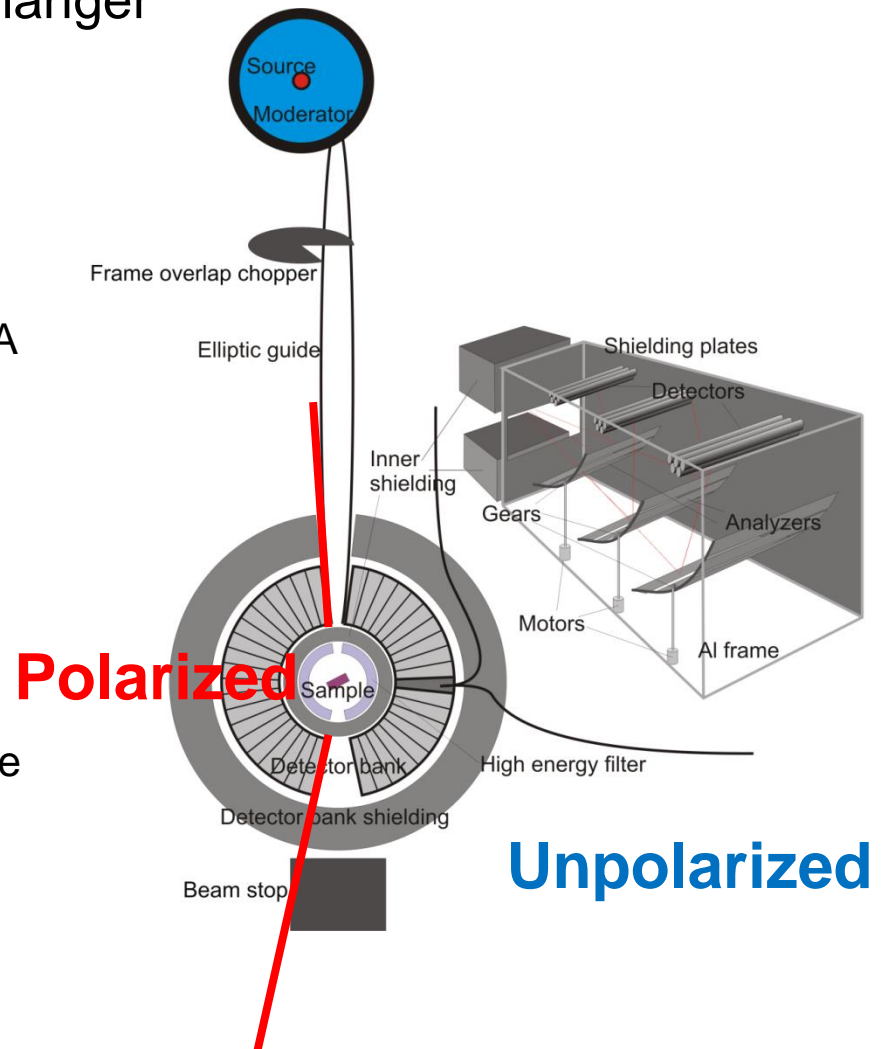
Instrument	Facility	Monochromator	Flux	CAMEA Gain	Energy Range
			N per $\text{cm}^2$ per s		(meV)
IN14	ILL	PG(002)	$1.6 \times 10^7$	88	0.1-17
PANDA	FRM-II	PG(002)	$1.9 \times 10^7$	74	0.1-20
MACS	NIST	PG(002)	$5 \times 10^8$	2.8	2.3-14*
THALES	ILL	PG(002)	$3.5 \times 10^8$	4	0.1-20
OSIRIS	ISIS	Time-of-Flight	$3.24 \times 10^7 @$ 180uA	43	-3 to 4
IRIS	ISIS	Time-of-Flight	$1.2 \times 10^7$ @180uA	117	-3.5 to 4
IN20 (polarized)	ILL	<u>Heusler</u>	$1.05 \times 10^7$	>30??	2-90

THERMAL Spectrometer

Wide divergence – Simulated Flux

# Polarization analysis ✓

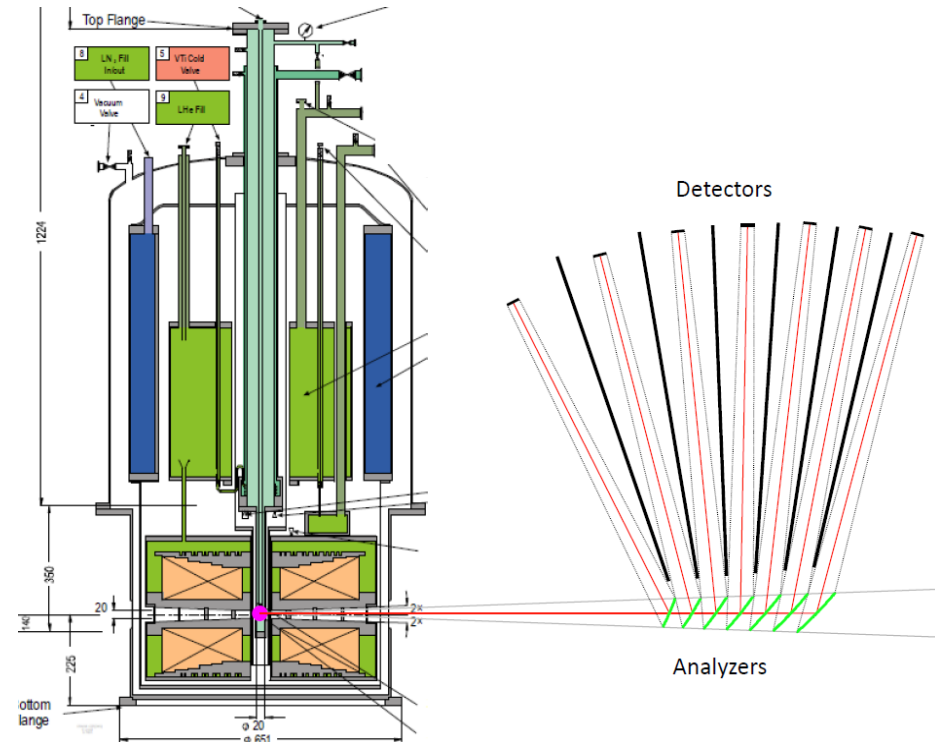
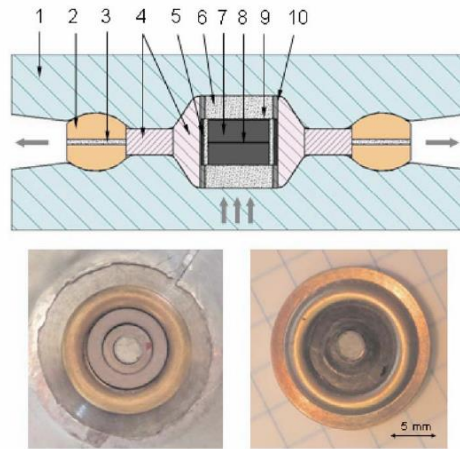
- Supermirror polarizers in a guide changer
- Option 1: PASTIS- wide angle He-3 spin cell
  - consistent with sample space of CAMEA
  - OK for pressure cells
  - Not OK for magnets
- Option 2: Supermirror array
  - Easier/cheaper than direct TOF because only  $\pm 2$  degree vertically
  - D7 proven concept
  - Cost based on D7 quote 2M Euros



# Experimental capabilities

- Planar Q-E maps by sample rotation
- Small samples  $1\text{mm}^3$  to  $1\text{cm}^3$
- Single acquisition scans  $\Rightarrow$  parametric studies
  - Magnetic field, Pressure, Temperature, etc.
- Extreme conditions
- In situ studies
- Time resolved studies:  $20\mu\text{s}$  resolution

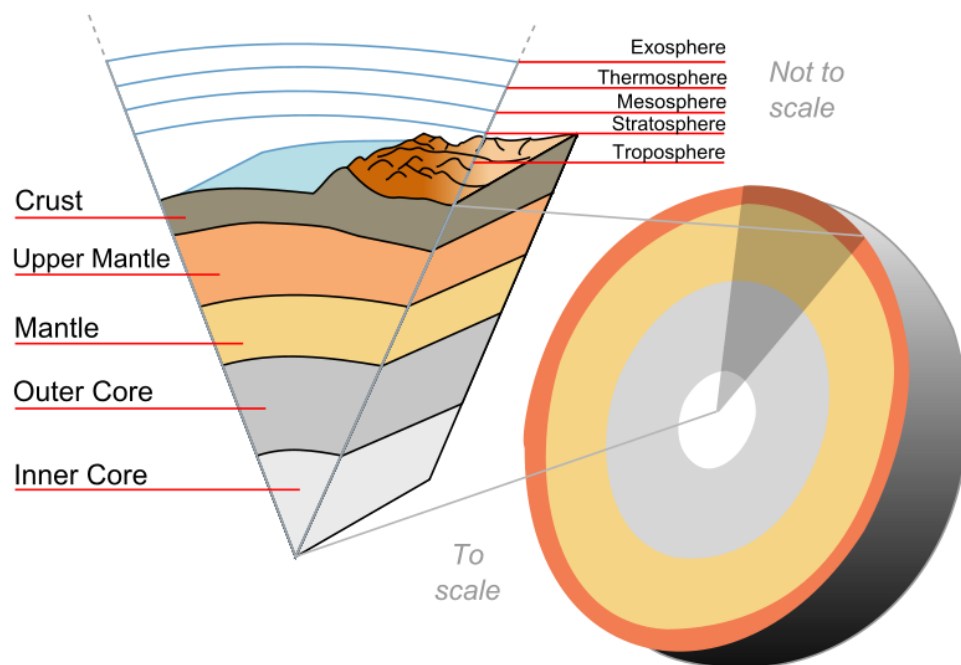
# Ideal for extreme environments



- 16+2T exists today, 25 T split coil HTSC likely in 2020
  - Quantum Magnetism, Superconductivity, Magnetoelectrics, Magnetocalorics
  - 18T is x2 over existing mapping instruments
- 100kbar, 0.3K-300K and 300-2000K possible
  - Quantum phase transitions, Planetary sciences

# Planetary science: High pressure and High Temperature

- Study of the structure of planets
- Upper mantle studies 30 GPa
- **High Temperatures** and **30 GPa** maximum **1-5 mm<sup>3</sup>**
- **QENS H-diffusion** water dynamics, Sound velocities
- PE cells used for geo-science
- Sintered DAC for NS emerging
- Pressure determination: **1% wavelength resolution** mode and **diffraction detector** ✓



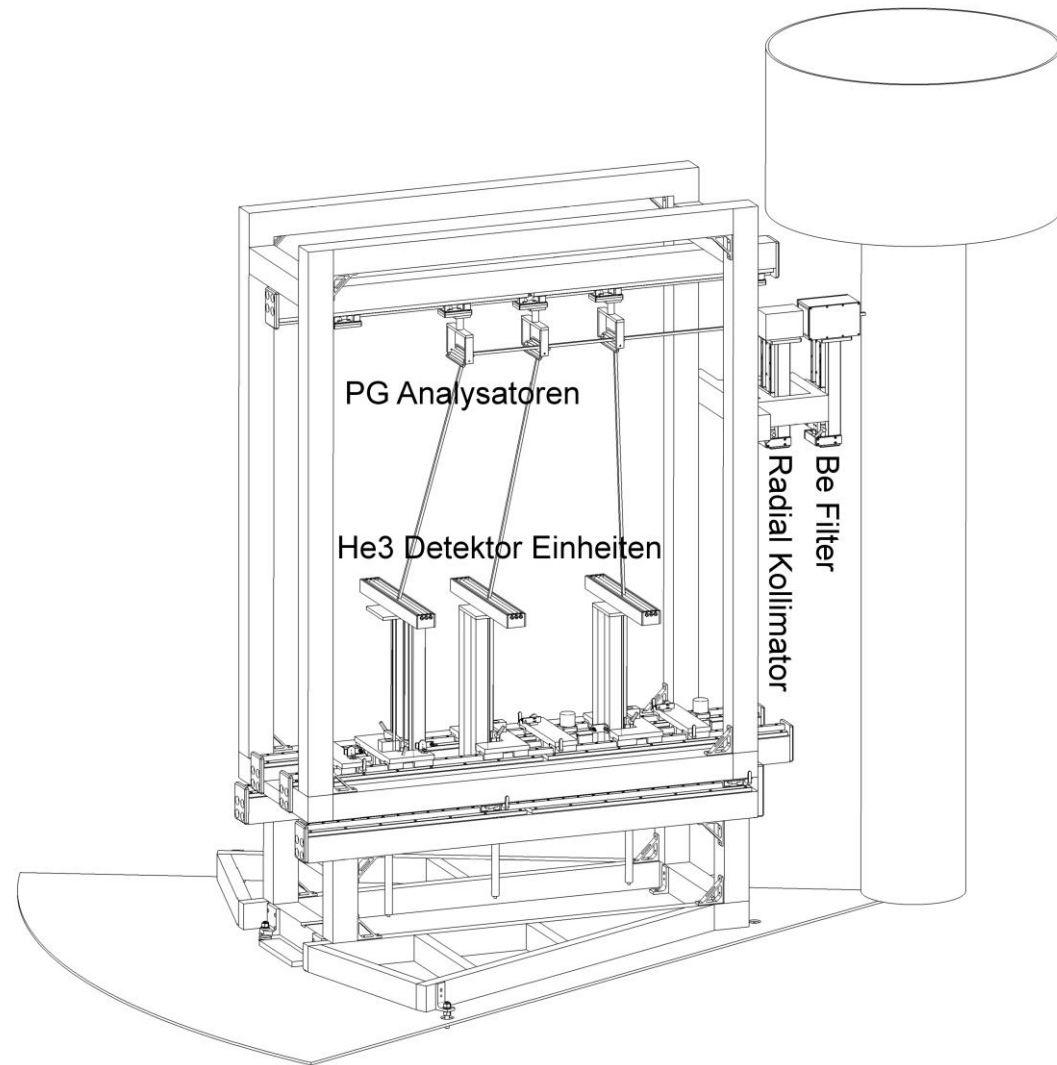
# Scientific case



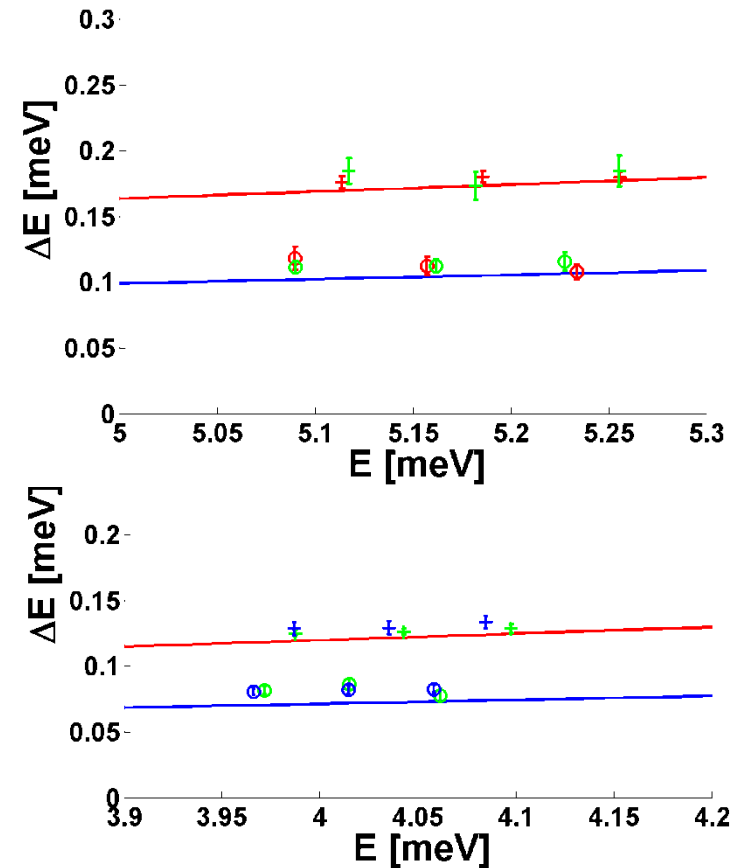
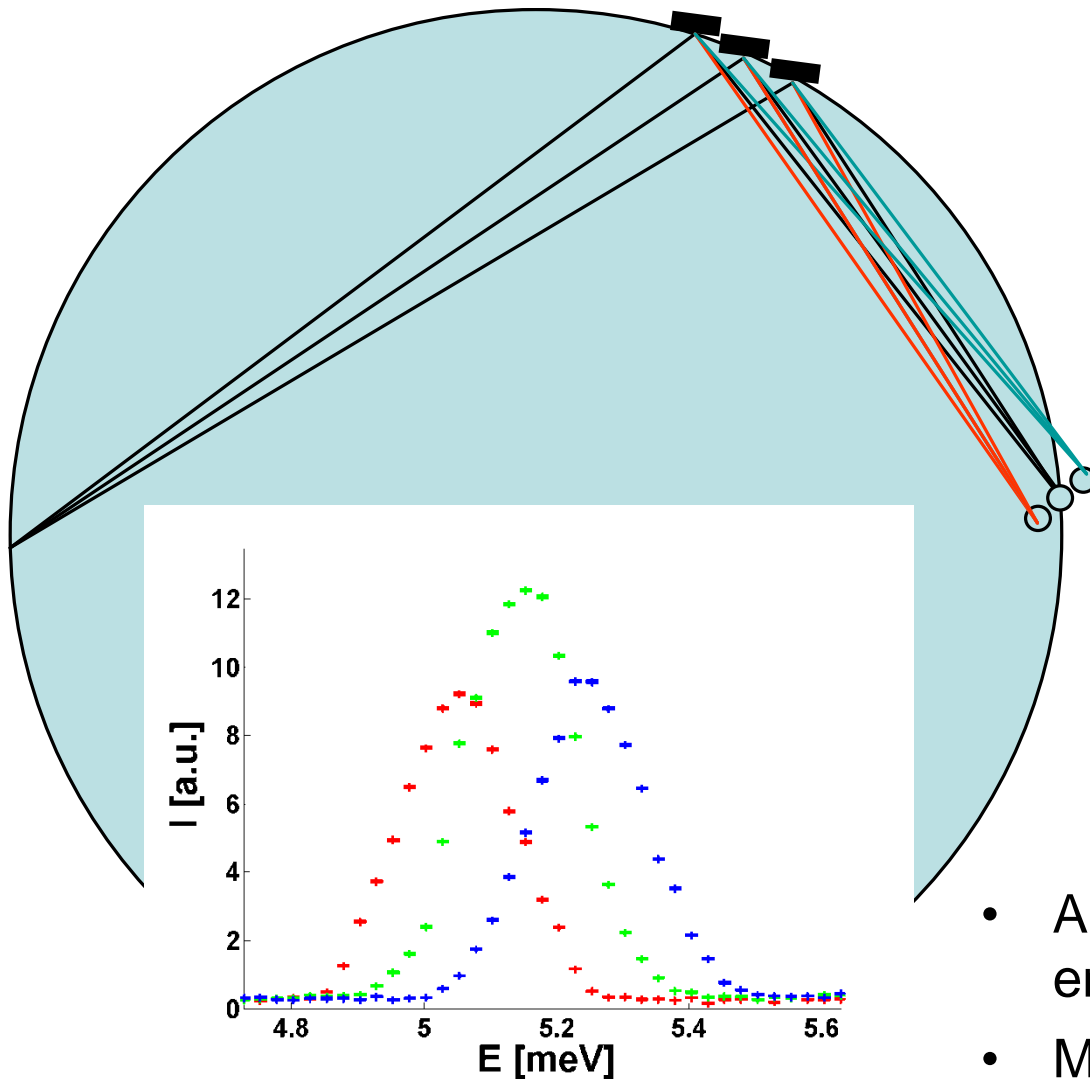
Report On Instrument Concept and Scientific Case  
PG Freeman et al. 2013

# Prototype at PSI

- Proving that CAMEA works
- Validation of calculations (McStas, analytical)
- Trying different geometries and solutions
- Getting experiences in building
- Checking background conditions, and searching for unknown effects



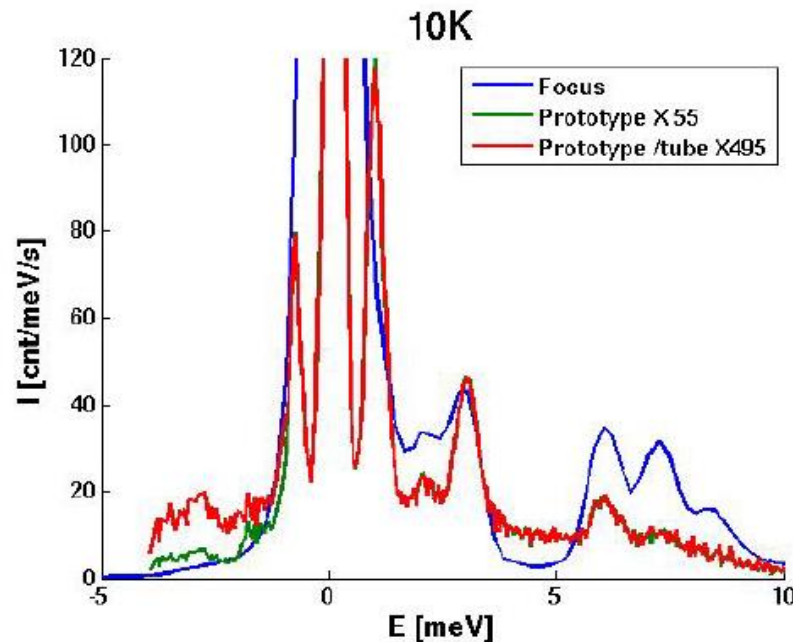
# Resolution and energy multiplication works as calculated



- Adjacent detectors record different energies  $\Rightarrow$  2.7x neutrons
- Mosaicity no effect on resolution

# Real experiments

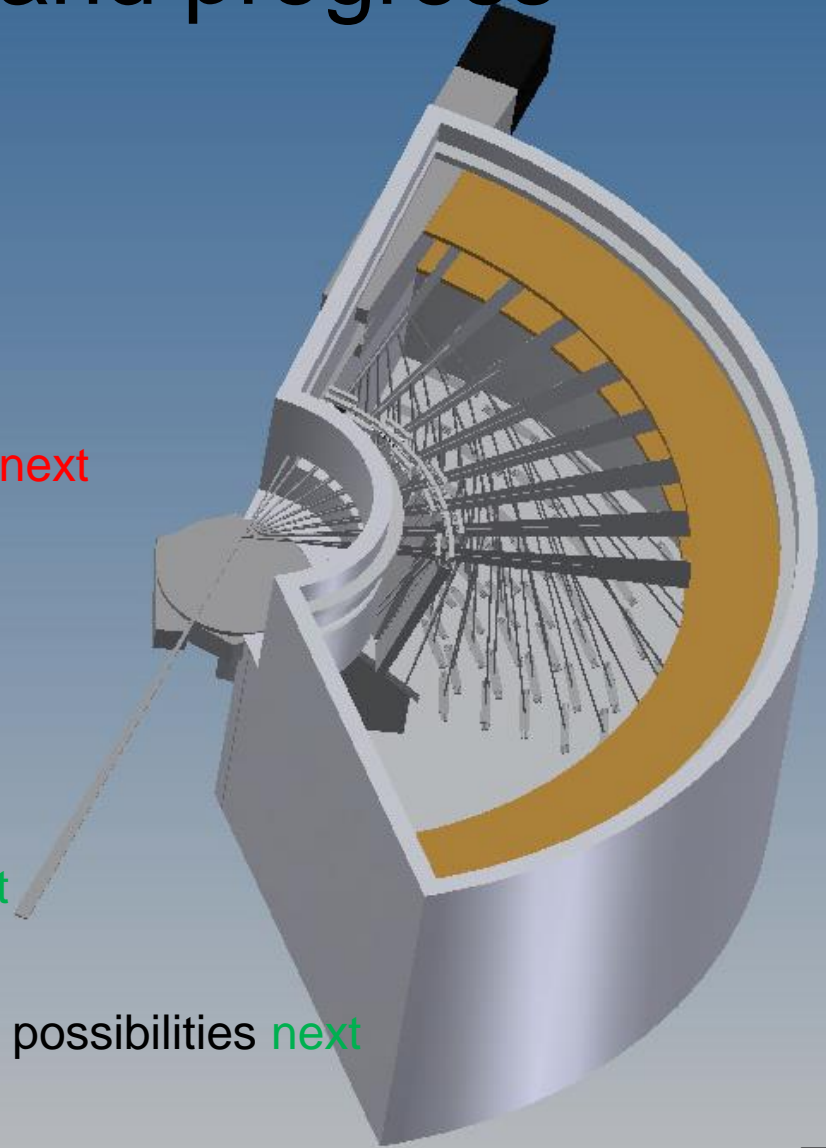
- $\text{LiHoF}_4$  crystal field excitations
- Comparison with FOCUS direct TOF



- Prototype has better resolution
- Prototype has >100 times less graphite than CAMEA
- MARS has “slightly” less flux than optimized guide at ESS

# Project plan and progress

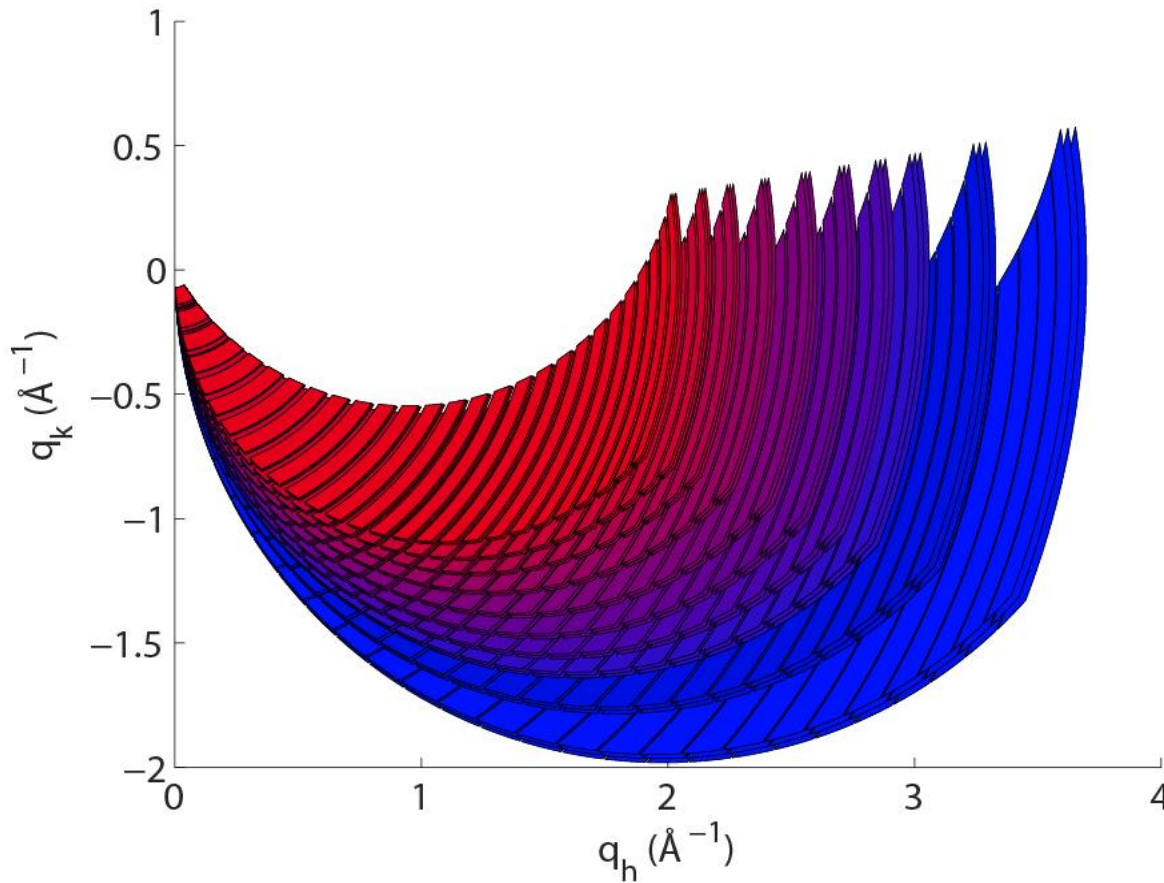
- Concept and science case ✓
- Calculations and Simulations:
  - Kinetics and  $(q, \omega)$  coverage ✓
  - Flux and resolution possibilities ✓
  - Guide and chopper system ✓
  - background and spurious scattering **next**
  - Virtual experiments **next**
  - Geometry optimizations **ongoing**
- Prototyping
  - Demonstration on RITA II ✓
  - Prototype on MARS at PSI in ✓
  - Pilot-project on RITA-II at PSI in **next**
- Sample environment
  - Magnet, P-cell 18T, 100kbar ✓; 2020 possibilities **next**



-  ⇒ Instrument proposal October 2013 ✓

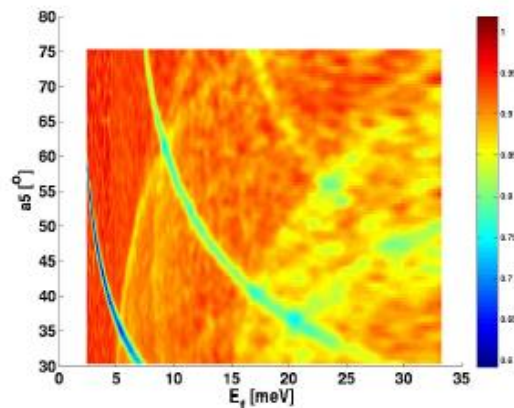
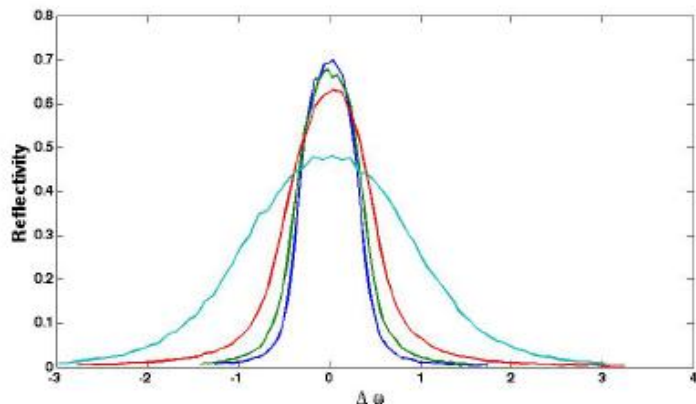
Thank y

# Momentum coverage ✓

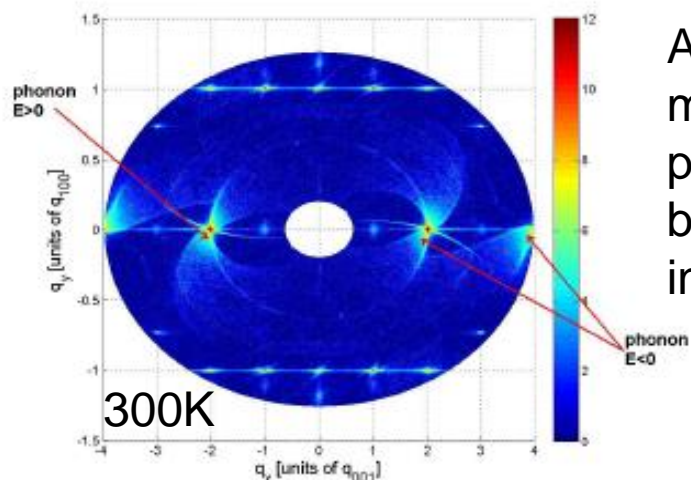
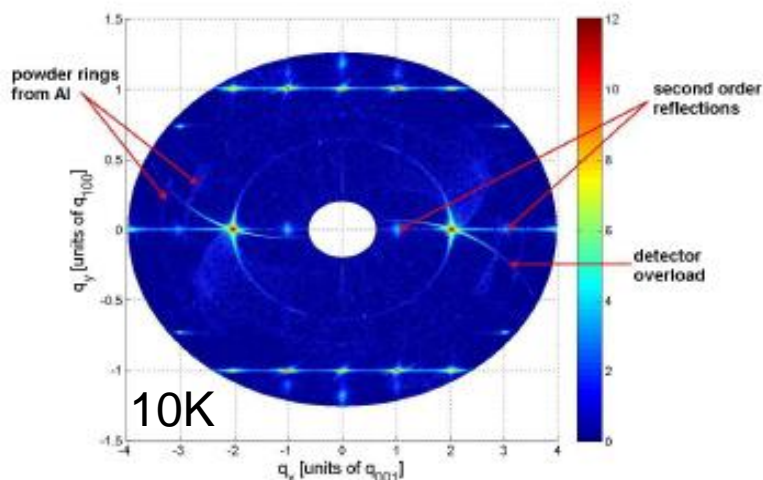


# Graphite thoroughly characterized

## Reflectivity and transmission



## Inelastic contamination



Analysers mounted on pulse-tube cooled base plate in vacuum tank

# Background

- Background suppression so far successful still being optimized
- Important: so far we understand all sources of bck.

