

# ESS Target Station Optimization: new moderator concept



# Planning and timeline

## Main milestones:

**Q1 2013: Technical Design Review (TDR) published**  
Tight schedule, “best practice” design for feasibility and costing

**Q3 2014: Target design optimization complete, **lay-out frozen****  
Plenty of time left (~ 1 year) to meet accelerator schedule

- lower costs at equal performance
- increase performance at equal costs

**Q1 2016: Main manufacturing contracts placed**

**Q4 2018: Target Station installed, commissioning without beam (Risk: delay with target hall)**

**Q3 2019: First beam on target, **commissioning with beam starts****

**Q2 2020: Commissioning complete, routine operation**

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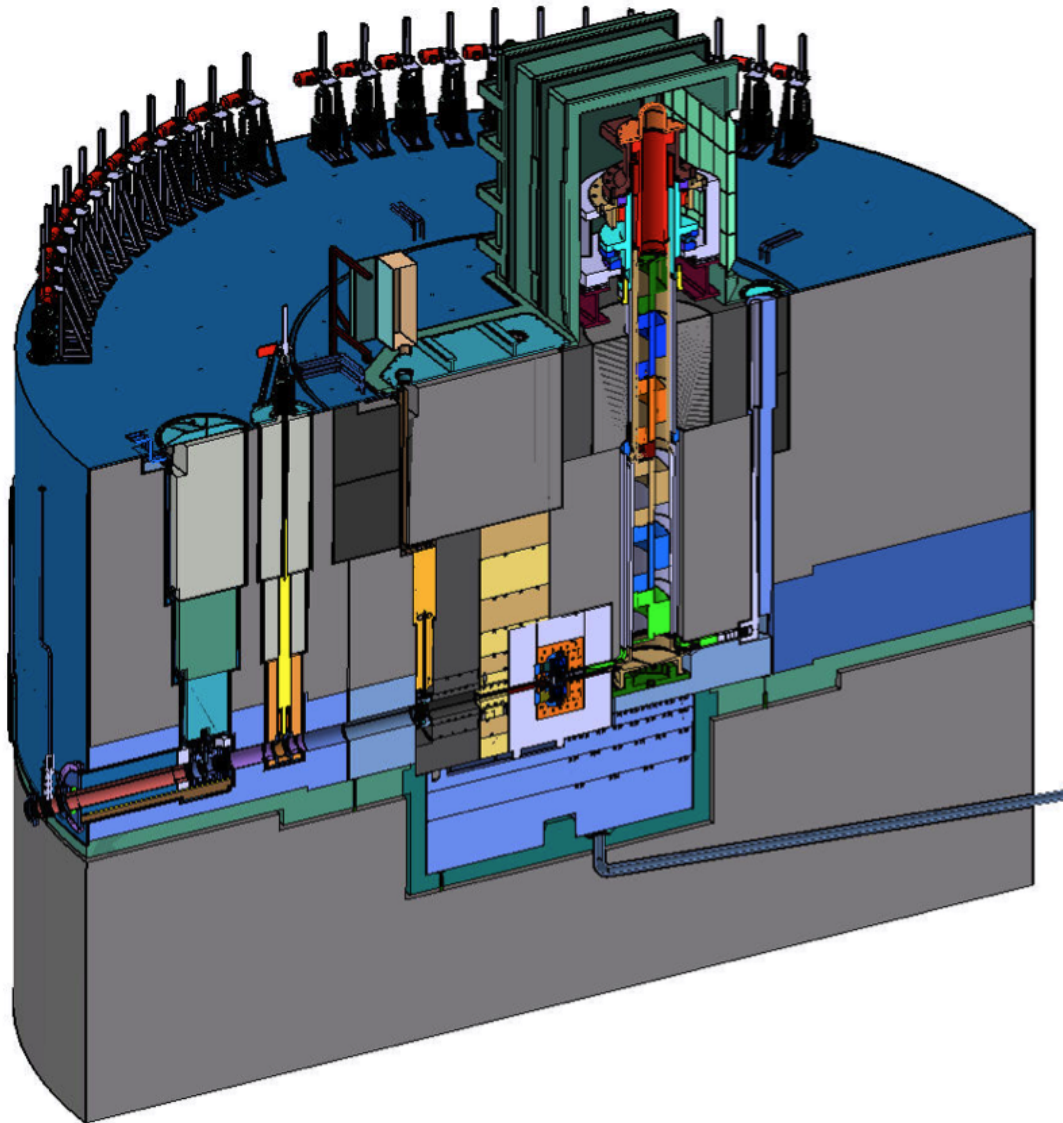
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EUROPEAN  
SPALLATION  
SOURCE

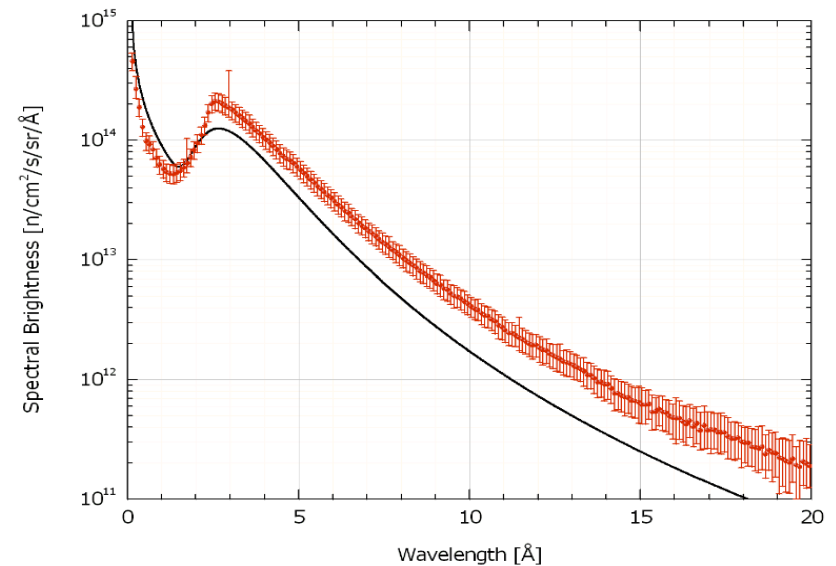
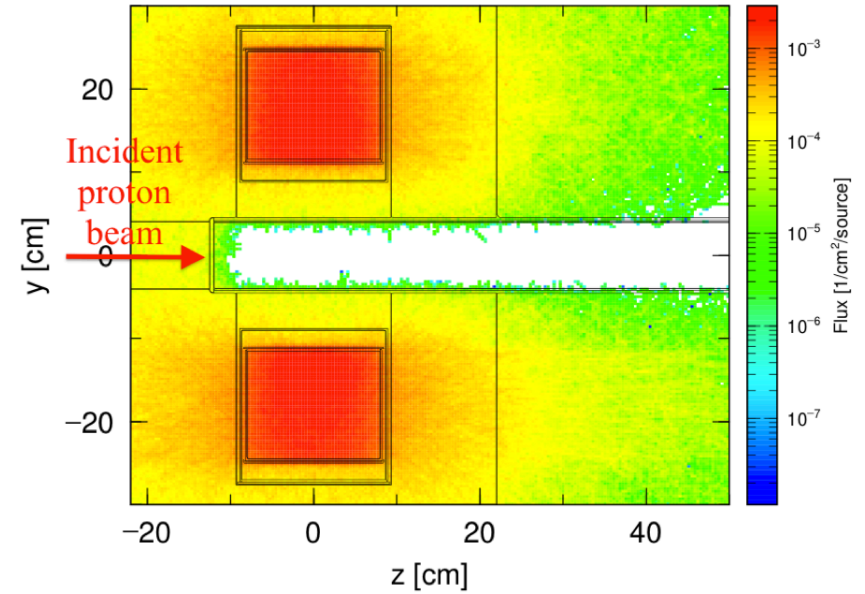
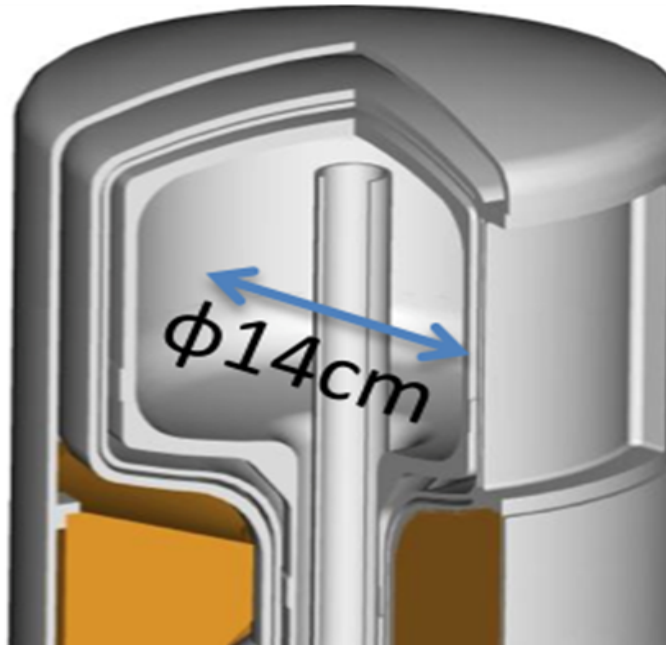
# Target station: the monolith



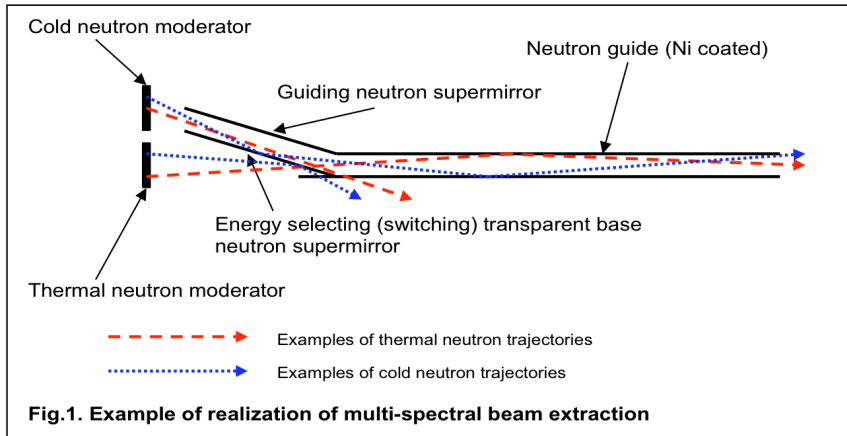
# High performance cold moderators

- Volume moderator: para-H<sub>2</sub>**
- implemented at J-PARC
  - flux gain  $\sim 2$  vs. box mods. without filter

reflector

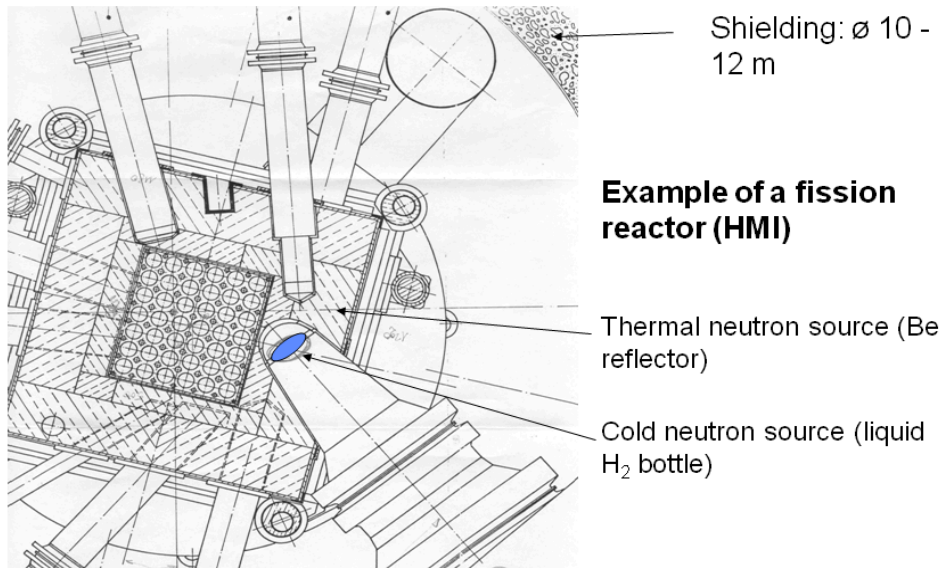


# Bi-spectral beam extraction

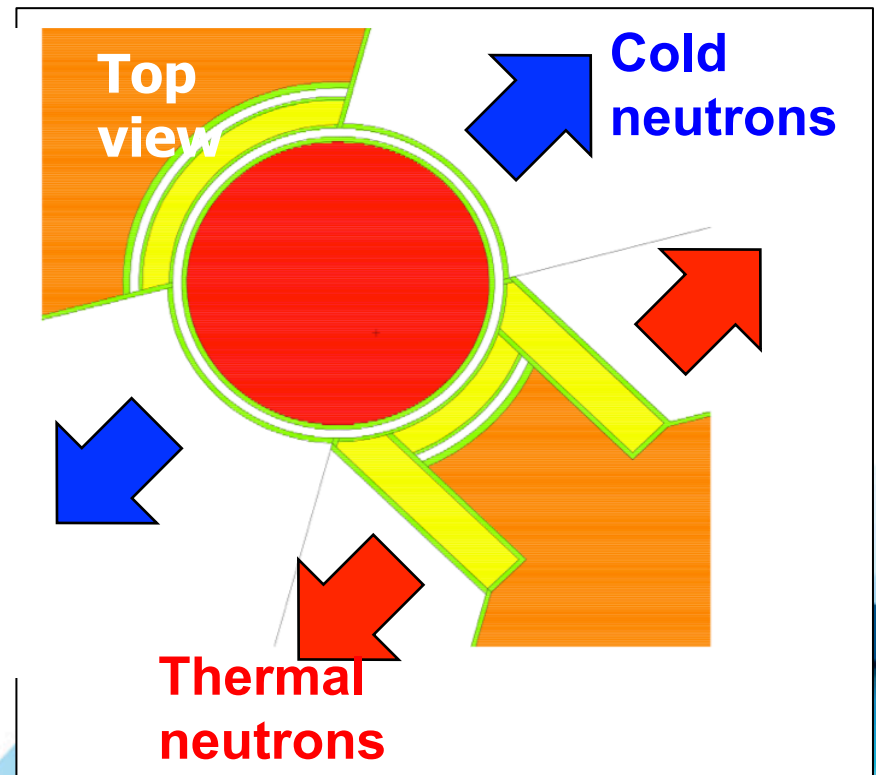


(F. Mezei, M. Russina, 2002)

**HZ Berlin: thermal source = Be reflector**

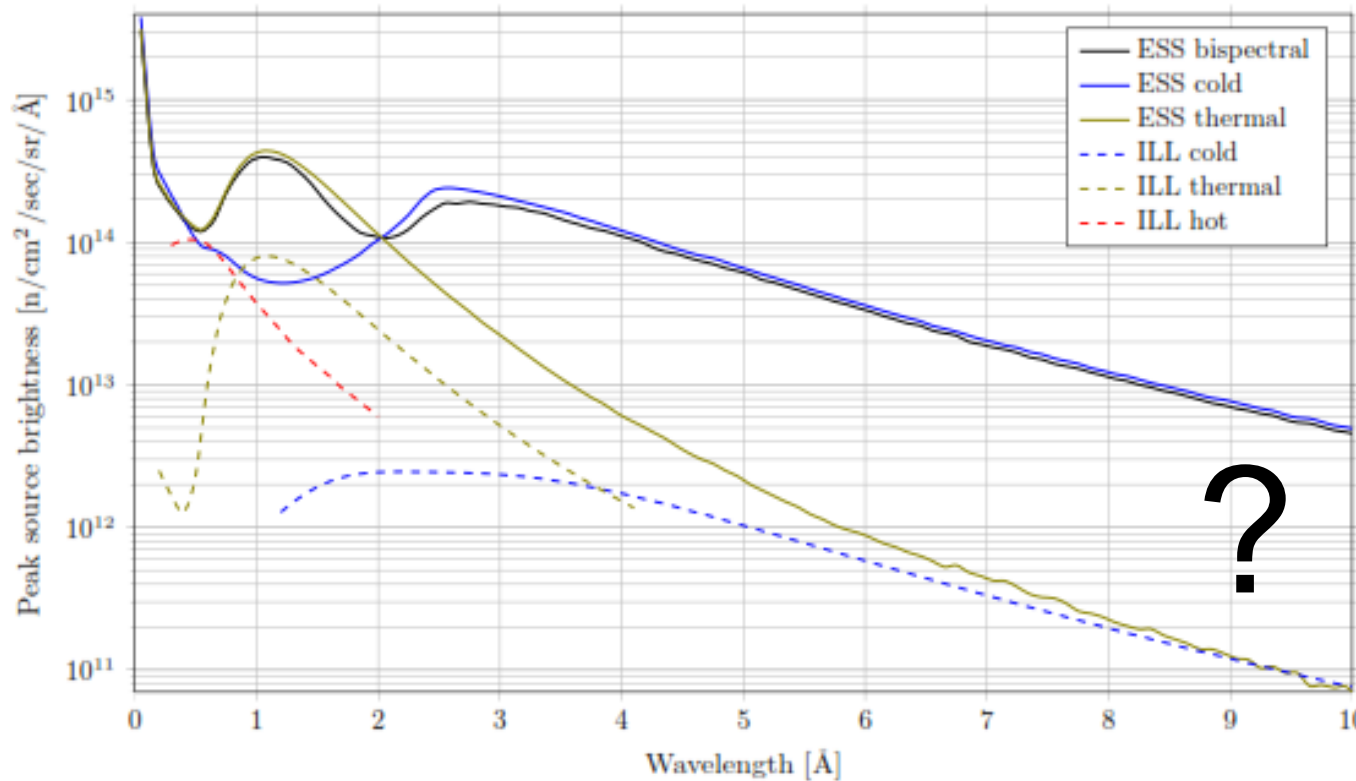


**Wide neutron energy dynamic range available on all beam-lines**

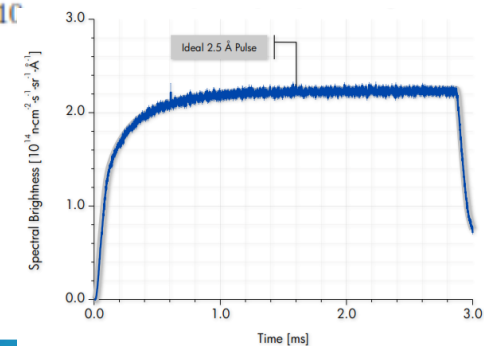
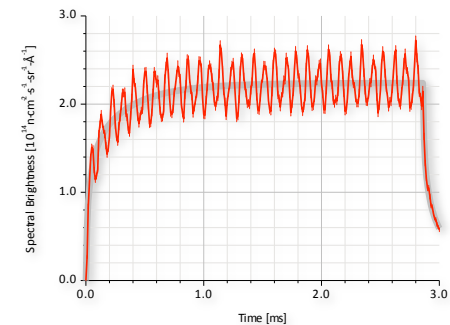


# Neutronic performance

Current baseline: ESS peak flux vs. ILL yellow book



Proton beam  
rastering



For long enough pulses:  
performance scales with peak flux



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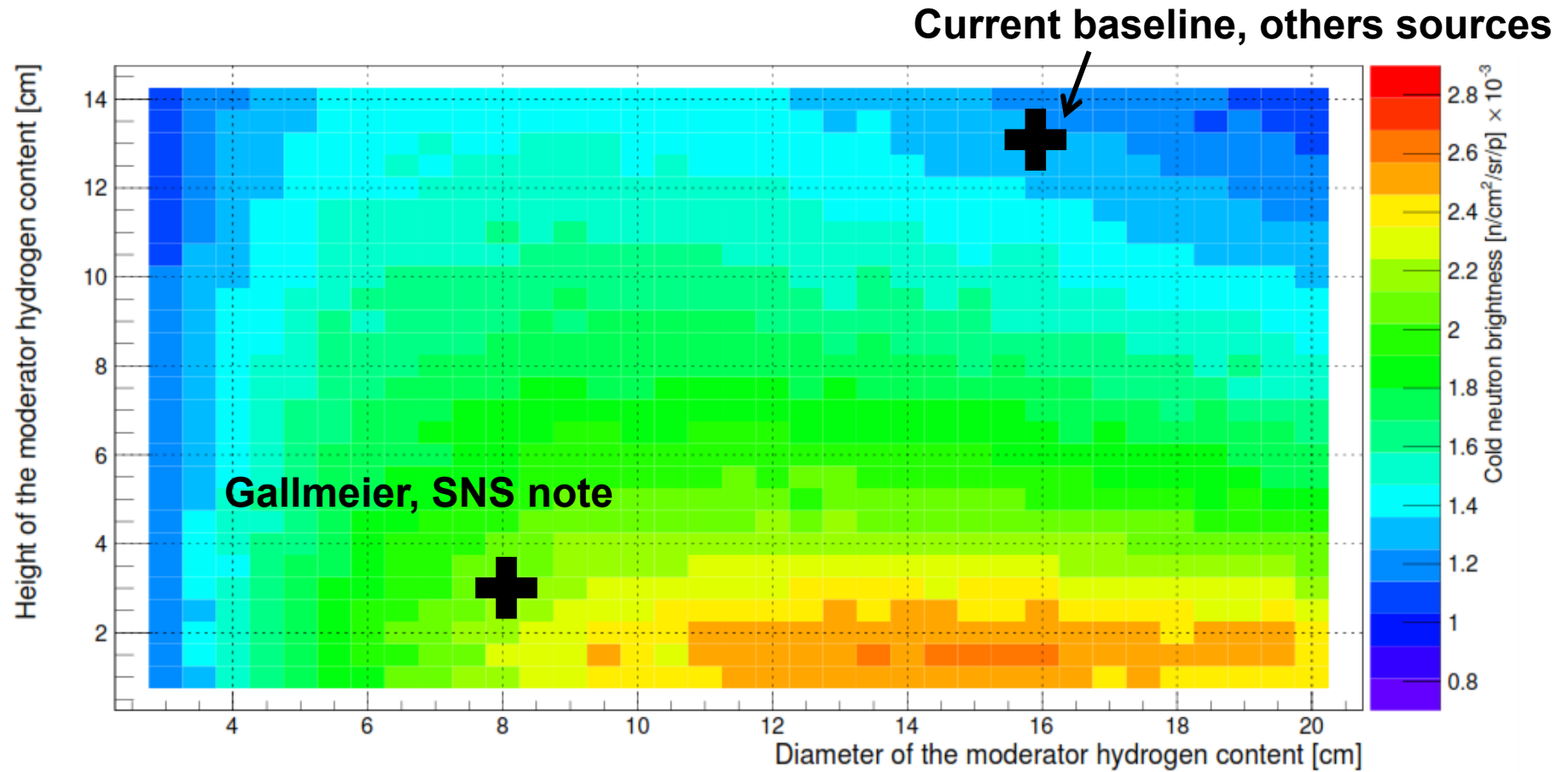
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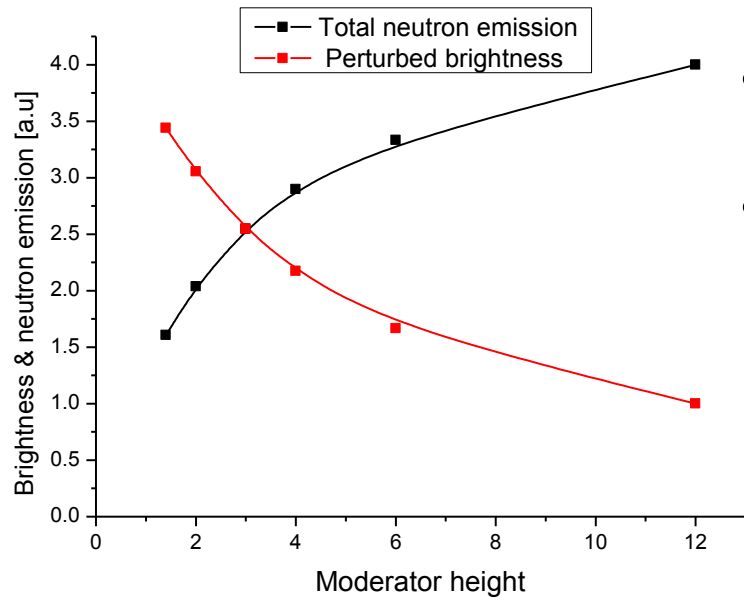
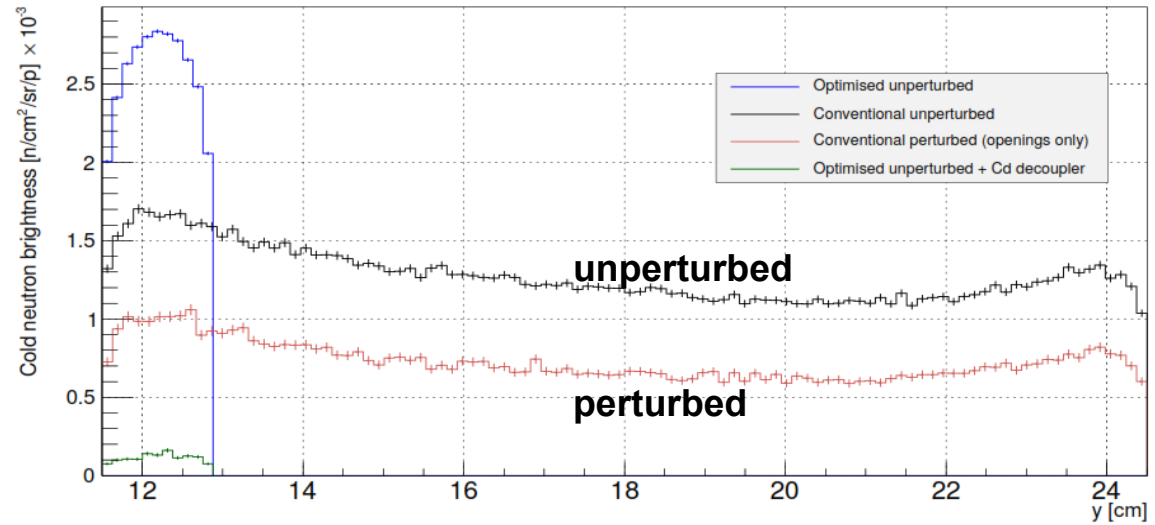
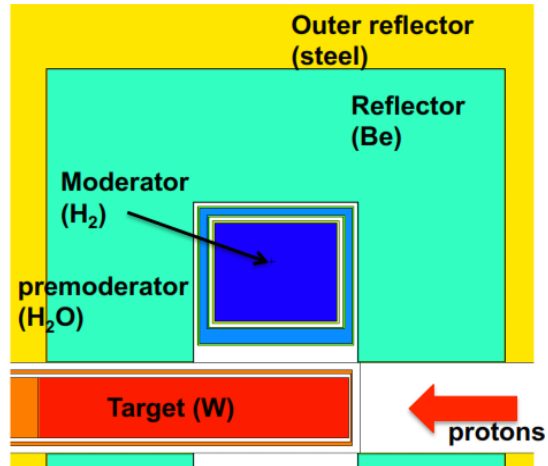


# Unperturbed moderator flux



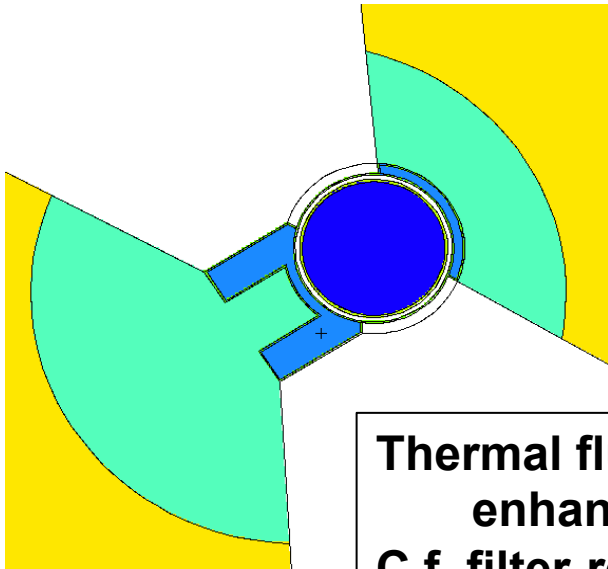
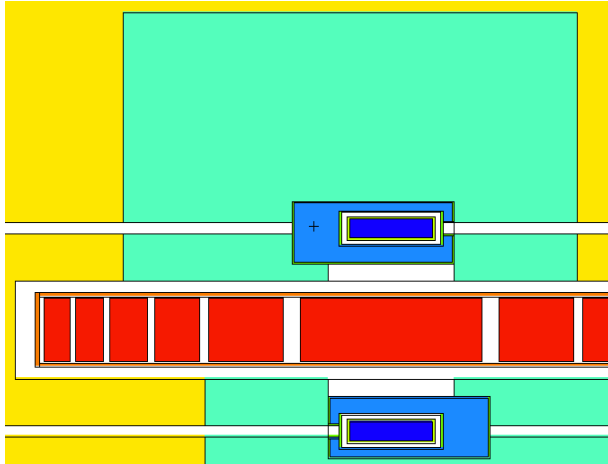
**Moderator size for highest unperturbed brightness:**  
 $\phi=15$  cm,  $h=1.4$  cm !!!

# Unperturbed moderator flux

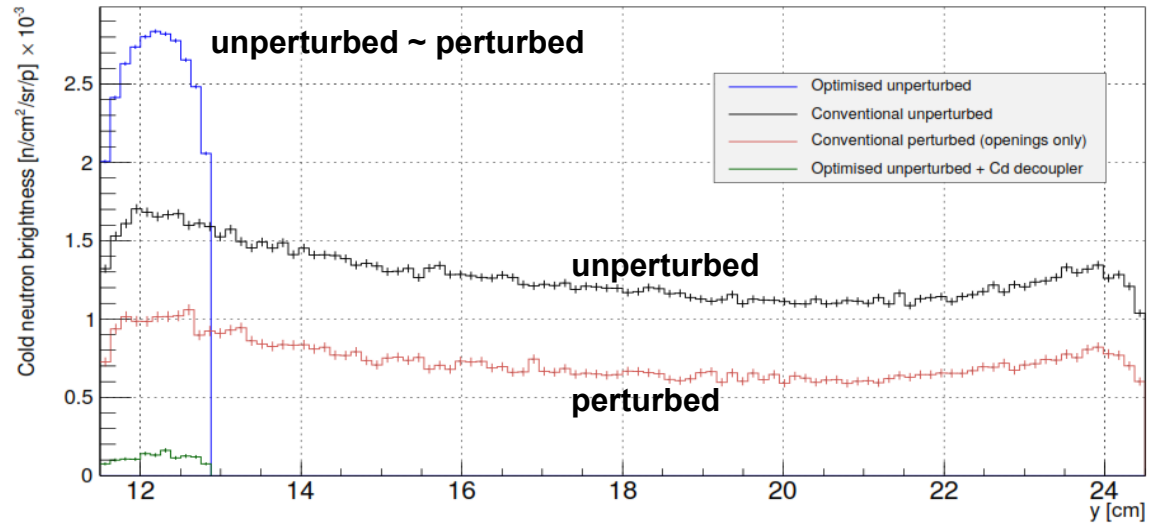


- Trend **even stronger** for the perturbed flux
- Also applies to the thermal flux from water moderator / Be reflector

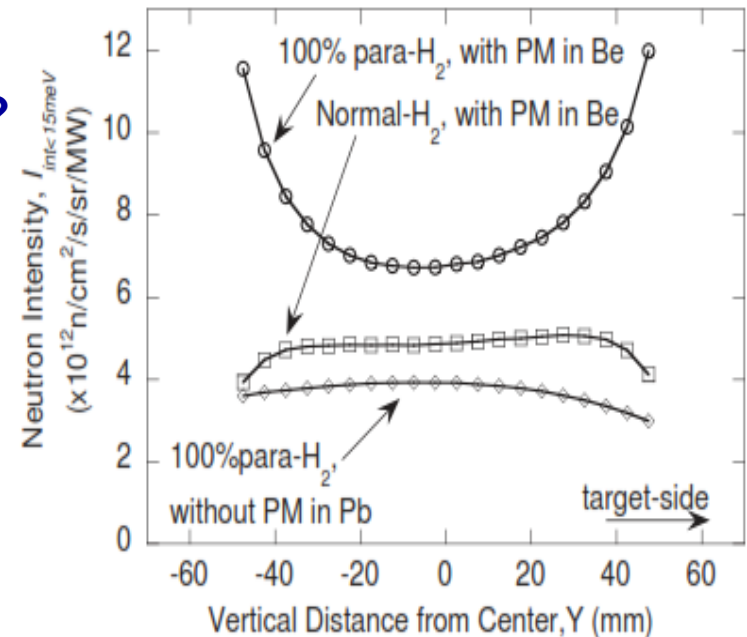
# Perturbed moderator flux



**Thermal flux: similarly enhanced  
C.f. filter-reflector !**



**Physics?**



# Optimized brightness moderator option

**How to use the high brightness of 12 cm x 1.5 cm cold  
or (12 cm + 12 cm) x 1.5 cm bi-spectral  
moderators?**

**Conventional wisdom: moderator size > guide entrance**

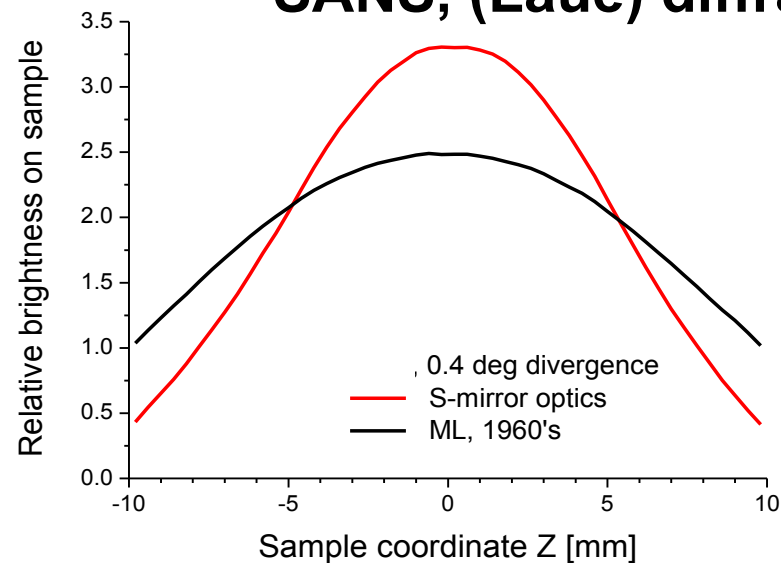
**Demonstrative basic examples for proof of principle**

- **Vertical and horizontal beam propagation: ~ independent**
- **Vertical dimension: ~ no influence on chopper action**
- **Global optimization important!!**

# Optimized brightness moderator option

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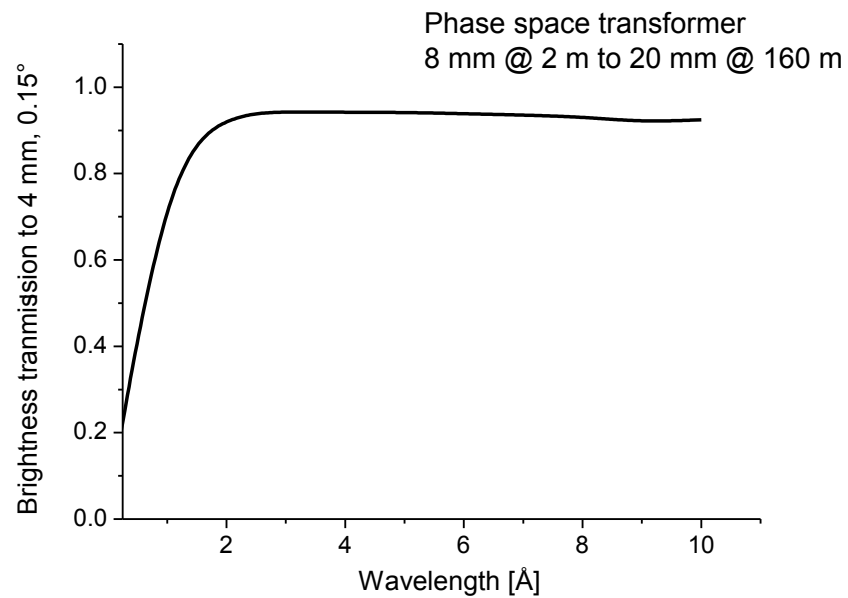
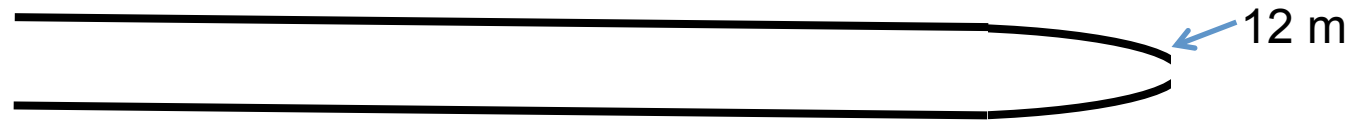
- **Direct view : (non-specular) reflectometry**
- **Conventional approach :  $\leq 1$  cm guide entrance height  
SANS, (Laue) diffraction, Laue diffraction ...**



# Conventional small entrance guides

**Demonstrative example:**

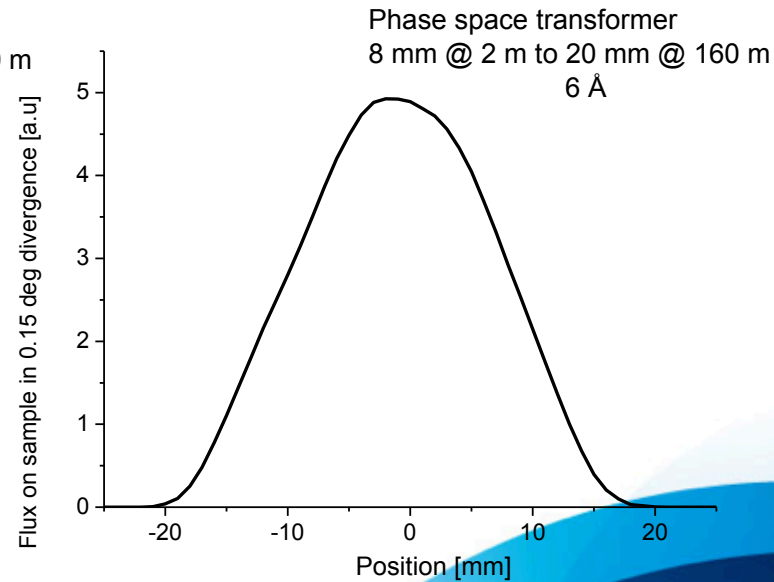
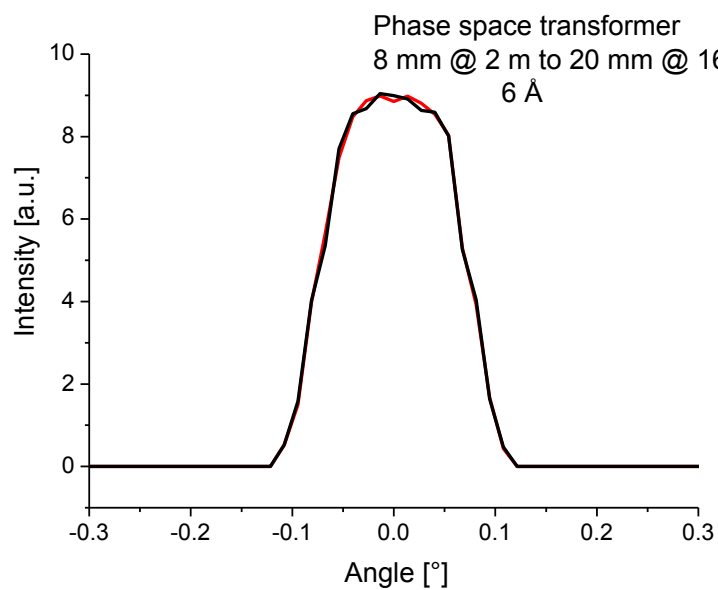
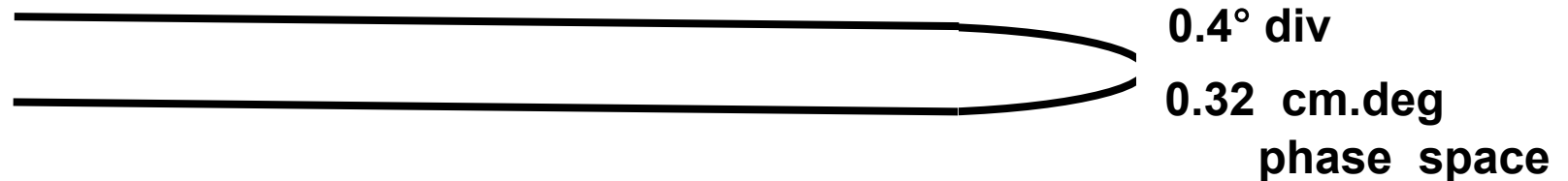
**160 m phase space transformer, Ni coated**



# Conventional small entrance guides

## Demonstrative example:

### 160 m phase space transformer, Ni coated

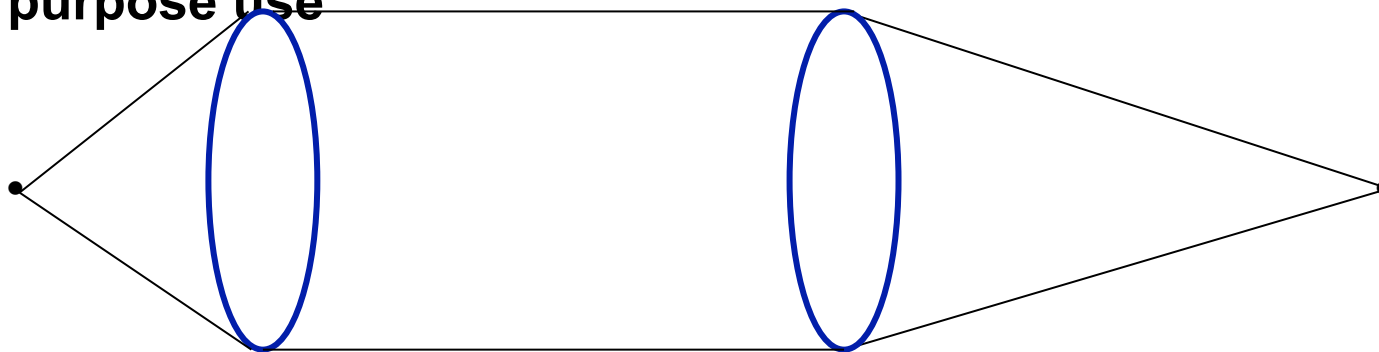




# Focusing optics

**How to use the high brightness of 12 cm x 1.5 cm cold  
or (12 cm + 12 cm) x 1.5 cm bi-spectral  
moderators?**

- **Liouville principle: high quality optics allows for general purpose use**

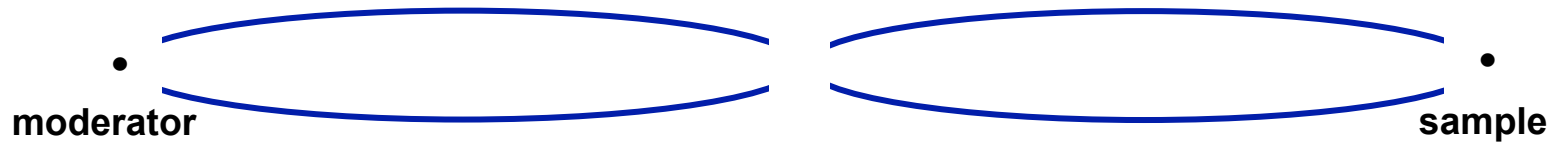


Lens combination

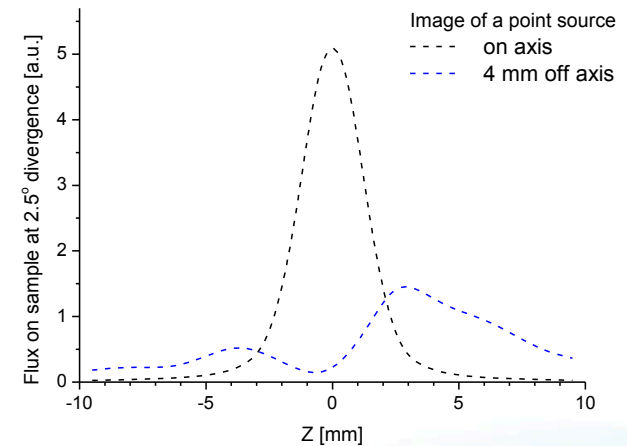
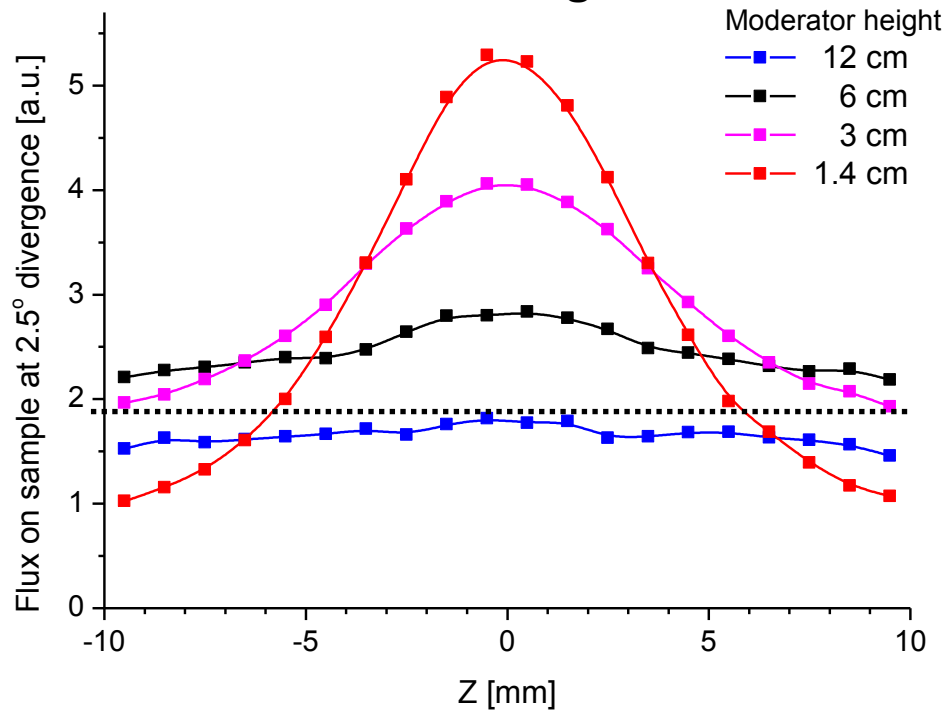
**Neutron optics: low aperture, limited quality imaging  
Supermirrors help, lot of room for progress**

# Focusing optics: large guide entrance

- Using best proposed / tested supermirror optics (cf. Stahn)

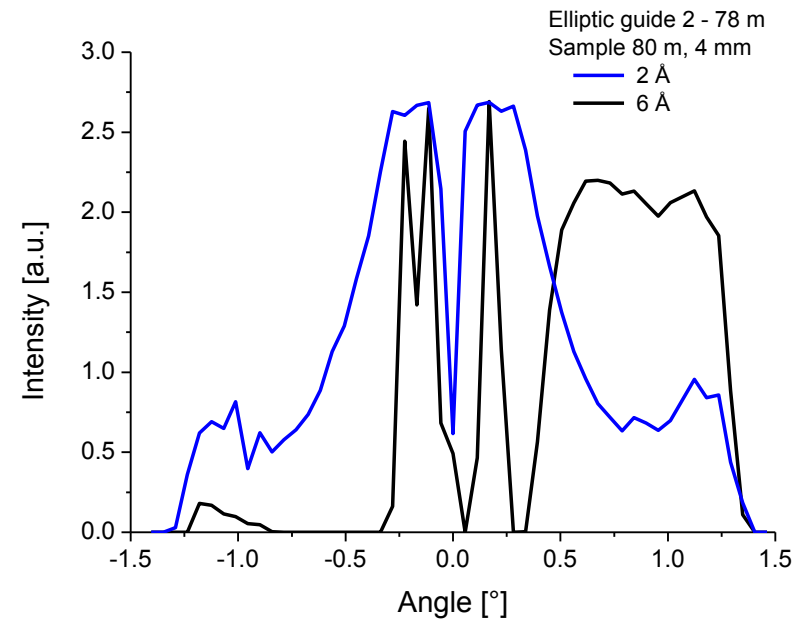
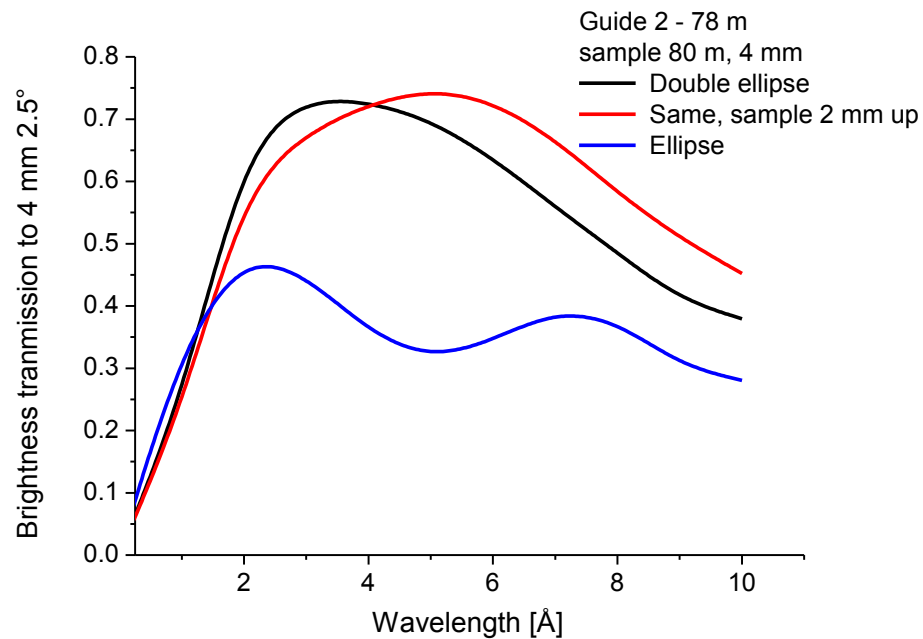


**Example: 2.5° beam divergence at 80 m  
4 Å wavelength**



# Focusing optics: large guide entrance

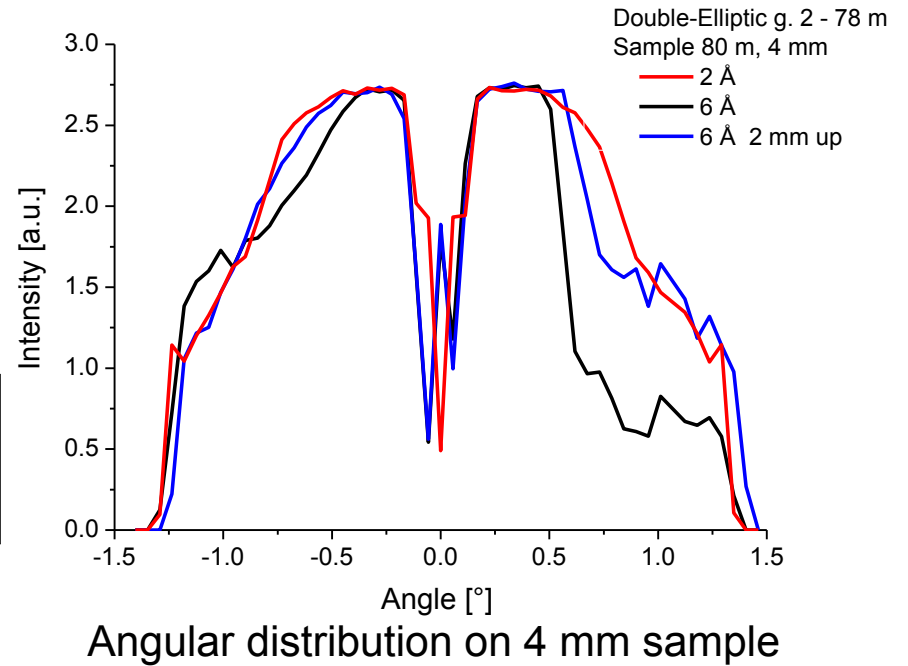
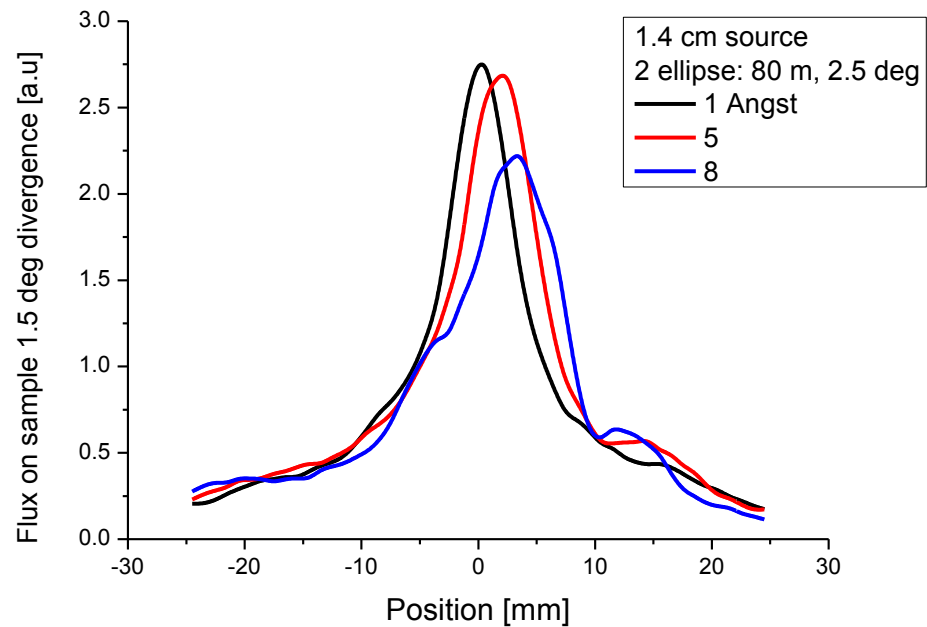
## Ellipse vs. double ellipse



Angular distribution on 4 mm sample

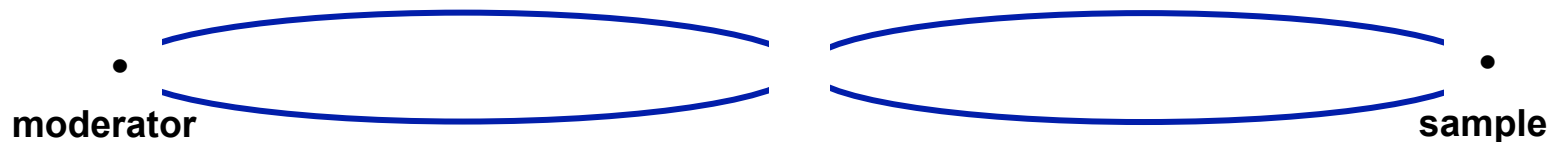
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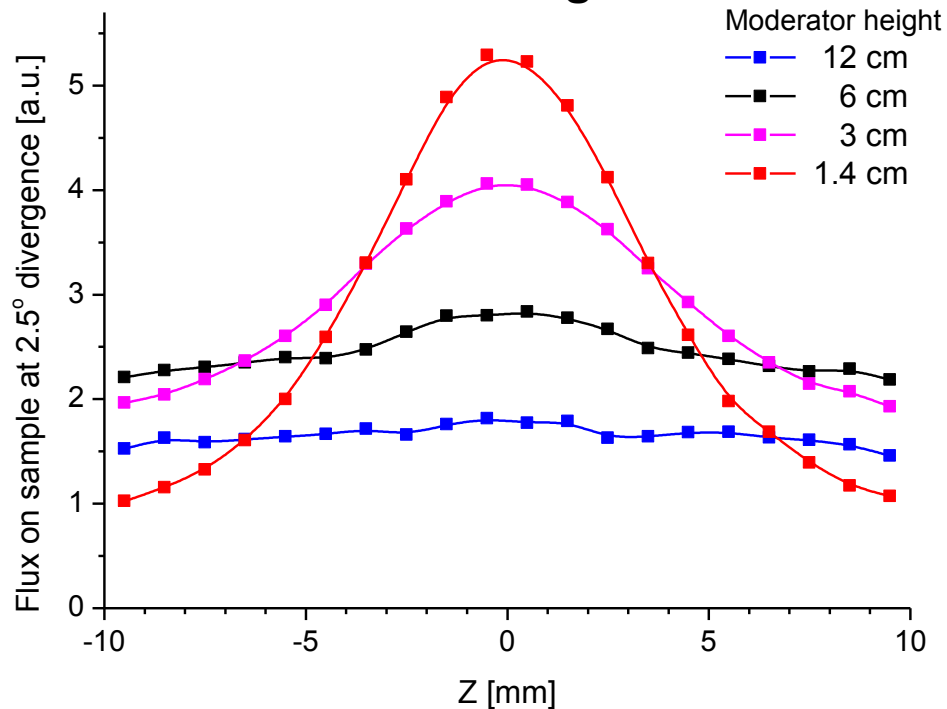


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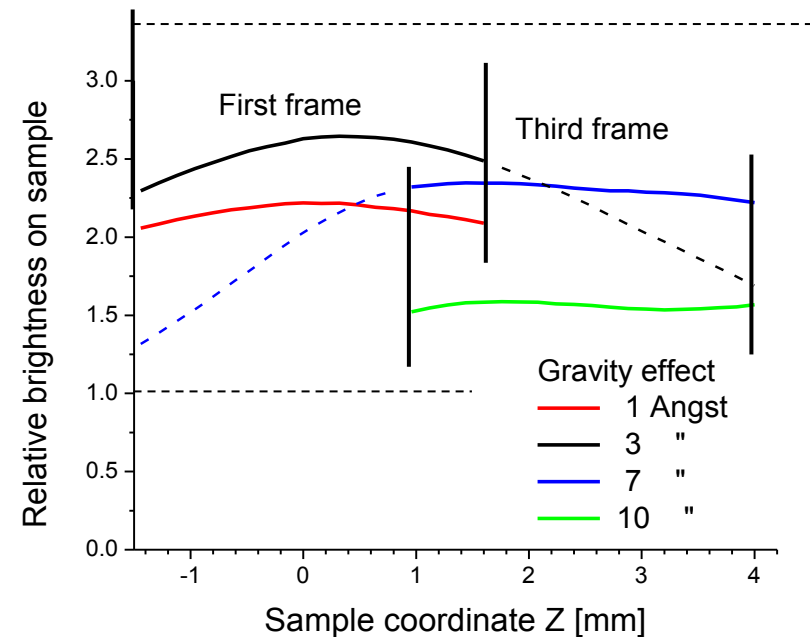
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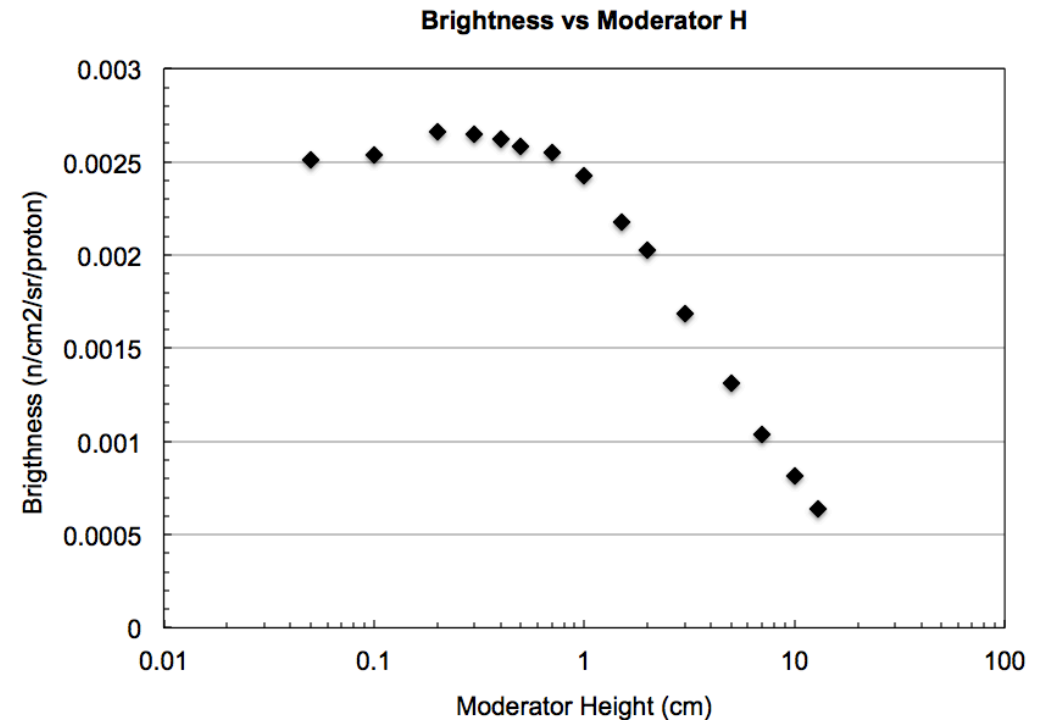


Effect of gravity & mitigation

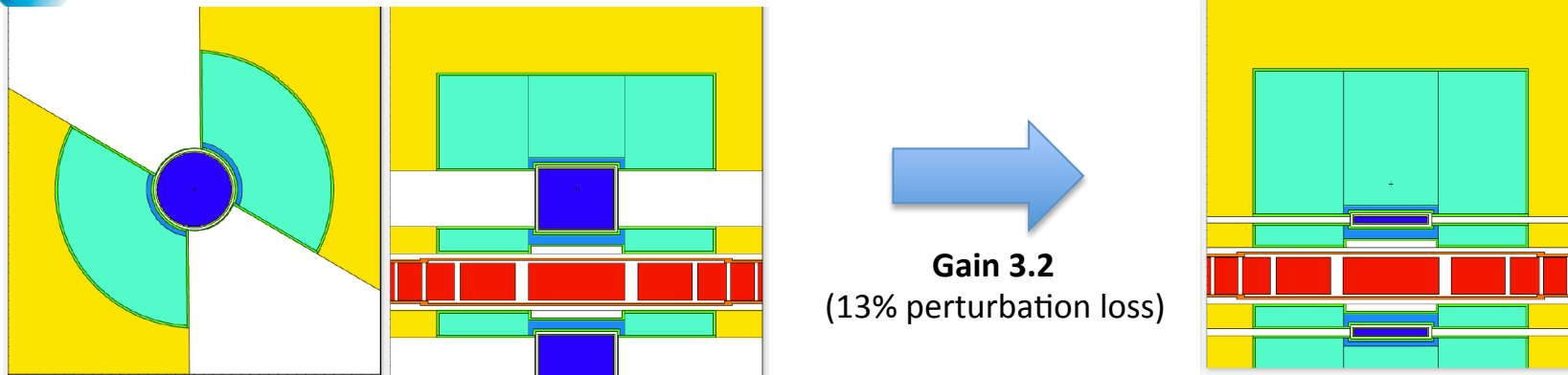


# Optimal perturbed moderator thickness

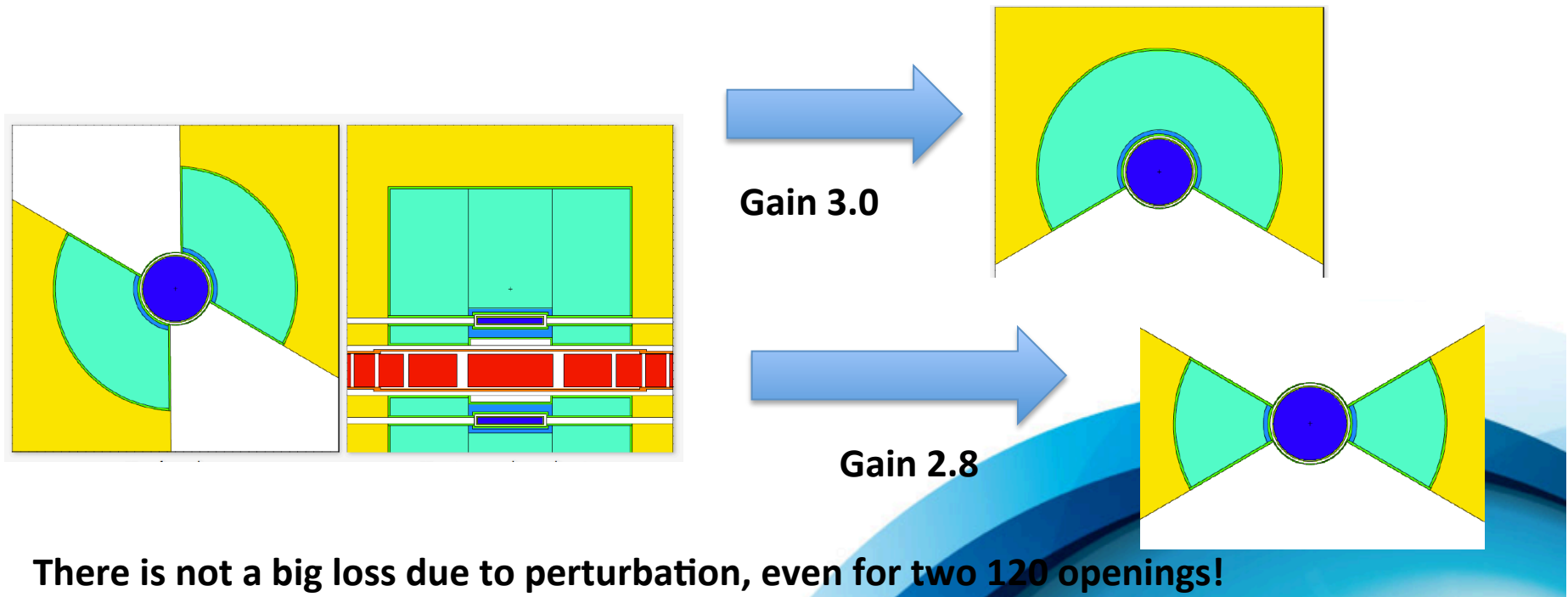
- Brightness for a moderator with fixed diameter of 16 cm, viewing a surface of 12 cm X (moderator height).
- Cole neutrons < 5 meV
- Maximum below 1 cm height.
- Calculations performed so far consider 1.5 cm height.



# From volume to flat moderator



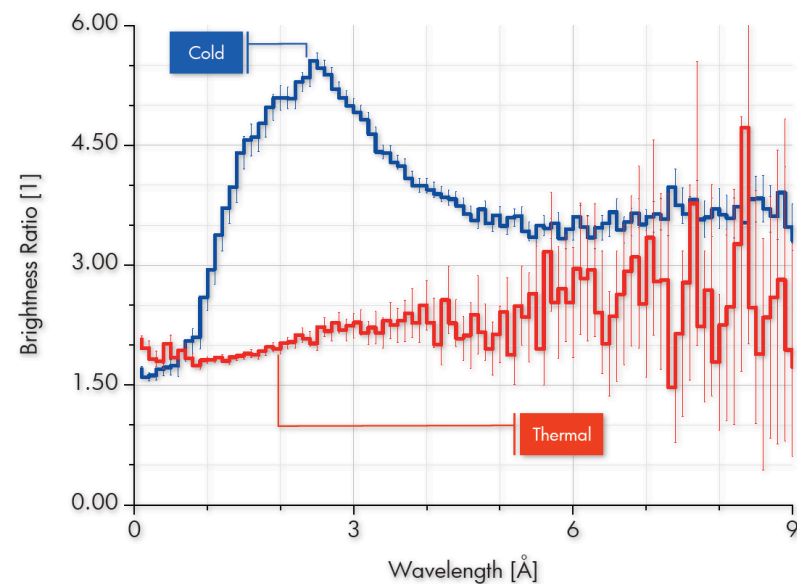
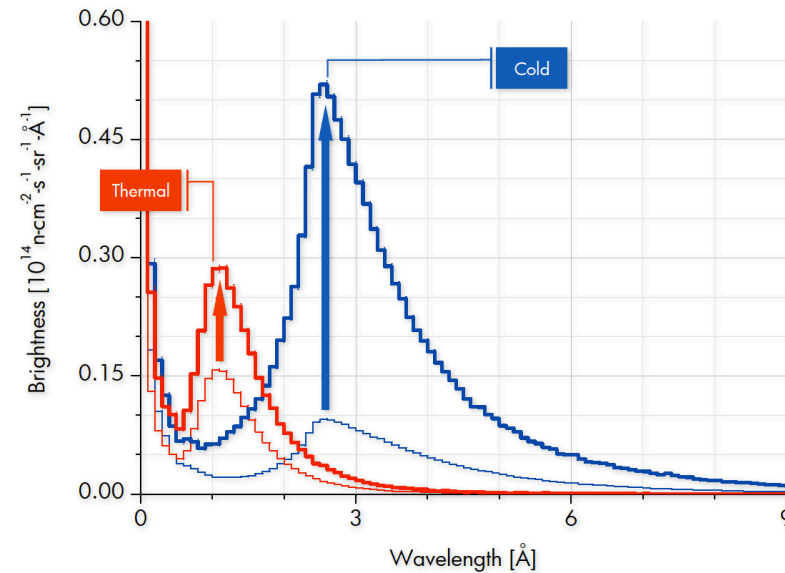
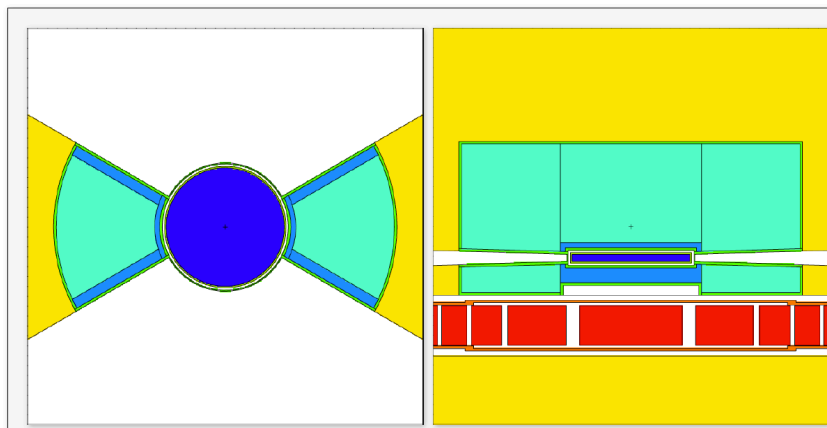
## Considering different opening options:



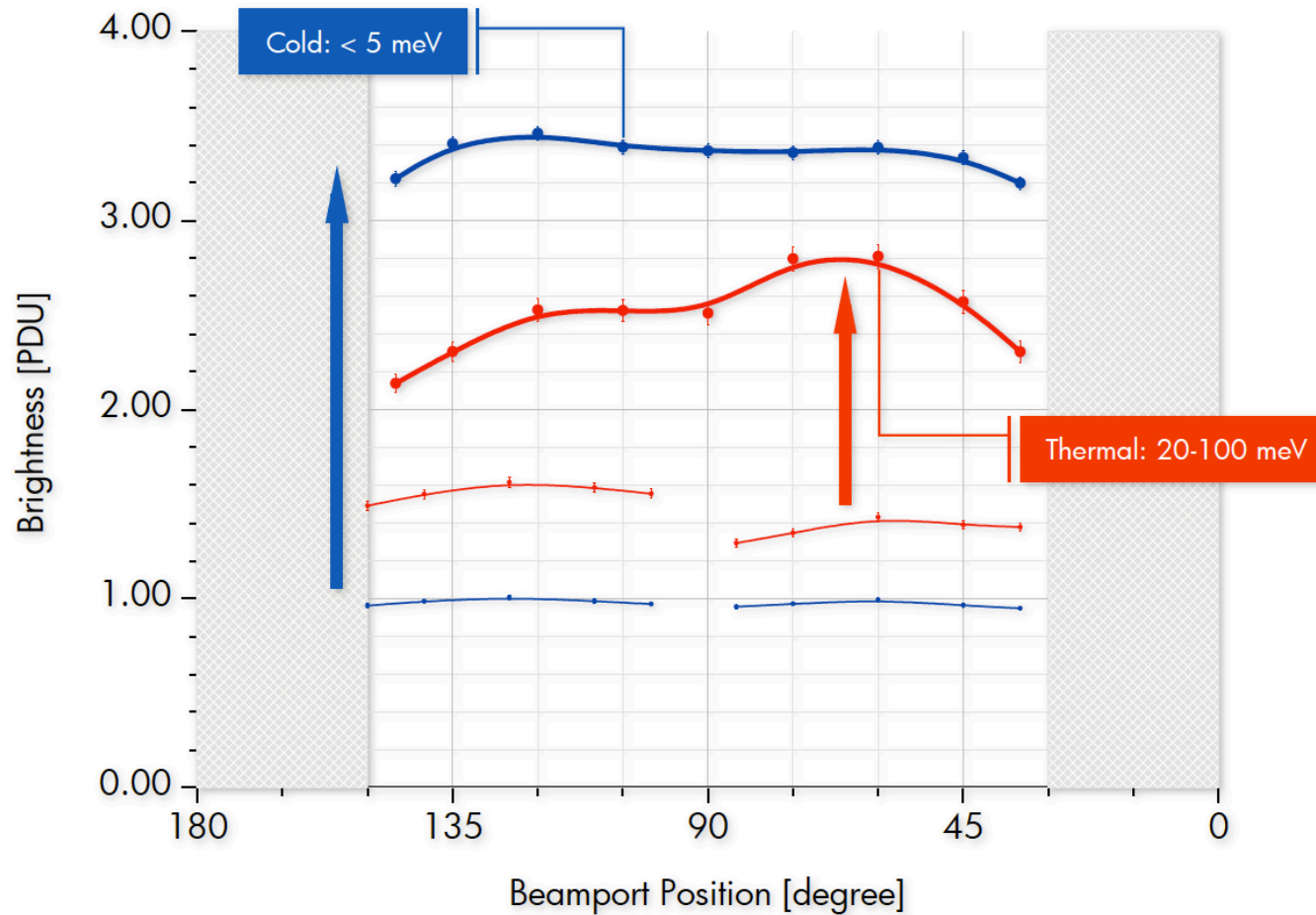


# Gain in brightness

- There is an increase for thermal and cold brightness.
- Thermal flux increase by about factor of 2.
- Cold flux increase by a factor of 5 (from 2 to 3 Angstrom) and factor 3.5 (above 4 Angstrom)

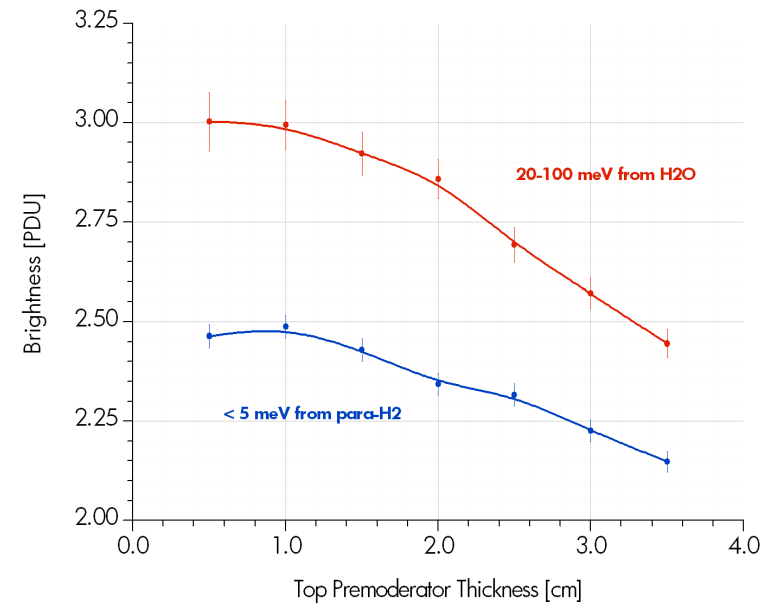
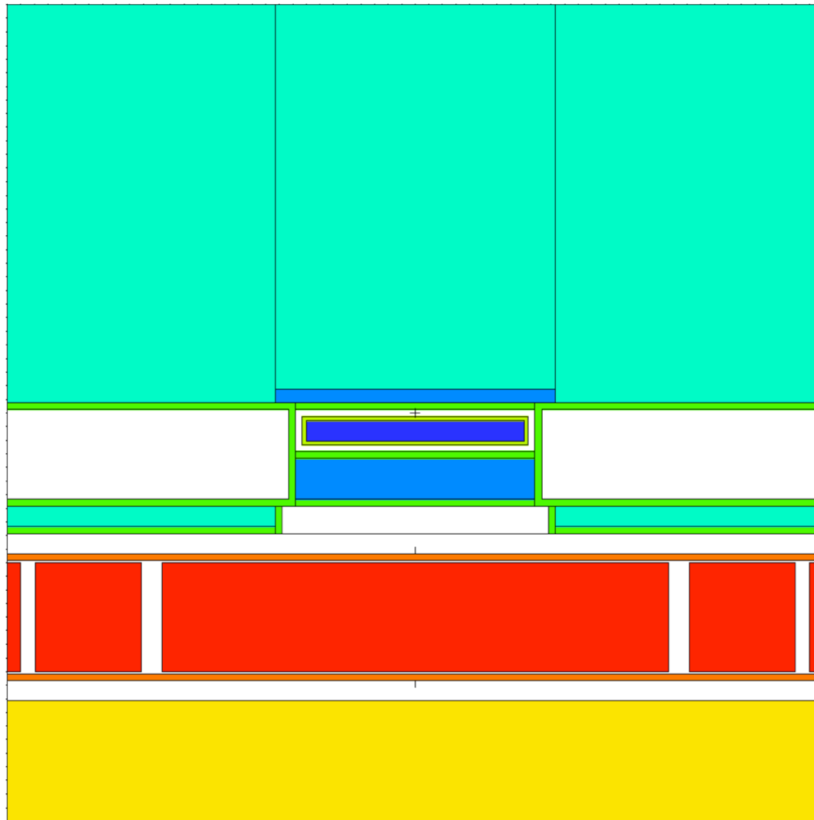


# Angular distribution



# Thermal vs. cold neutrons

- Find good balance between thermal and cold brightness.
- Gap in Be reflector is a crucial parameter.



# Optimized brightness moderator option

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or (12 cm + 12 cm) x 1.5 cm bi-spectral moderators?**

- **Proof of principle demonstrative examples:**

It is possible to transfer much of the brightness gain to the samples for low face space volume applications

**low beam divergence and/or small samples**

*Losses would cumulate for a small moderator in 2D!*

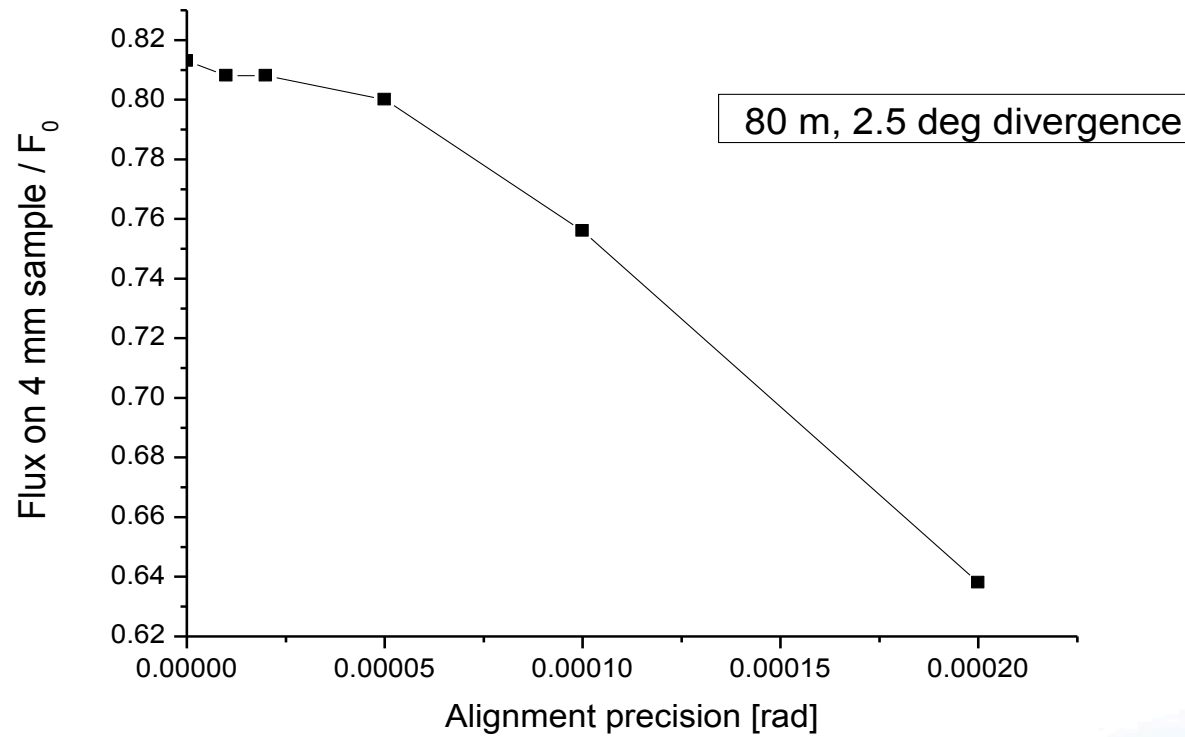
**Comparison to ILL: x 50 -100**

- **Optimization, implementation for the facility:**

**Collaborative effort** needed between instrument scientists, target and simulators: One “flat”, one “thicker” moderator? 2 flat moderators? Single “360° moderator”? Best “flat” moderator & beam extraction / delivery combination? Best bi-spectral combination?

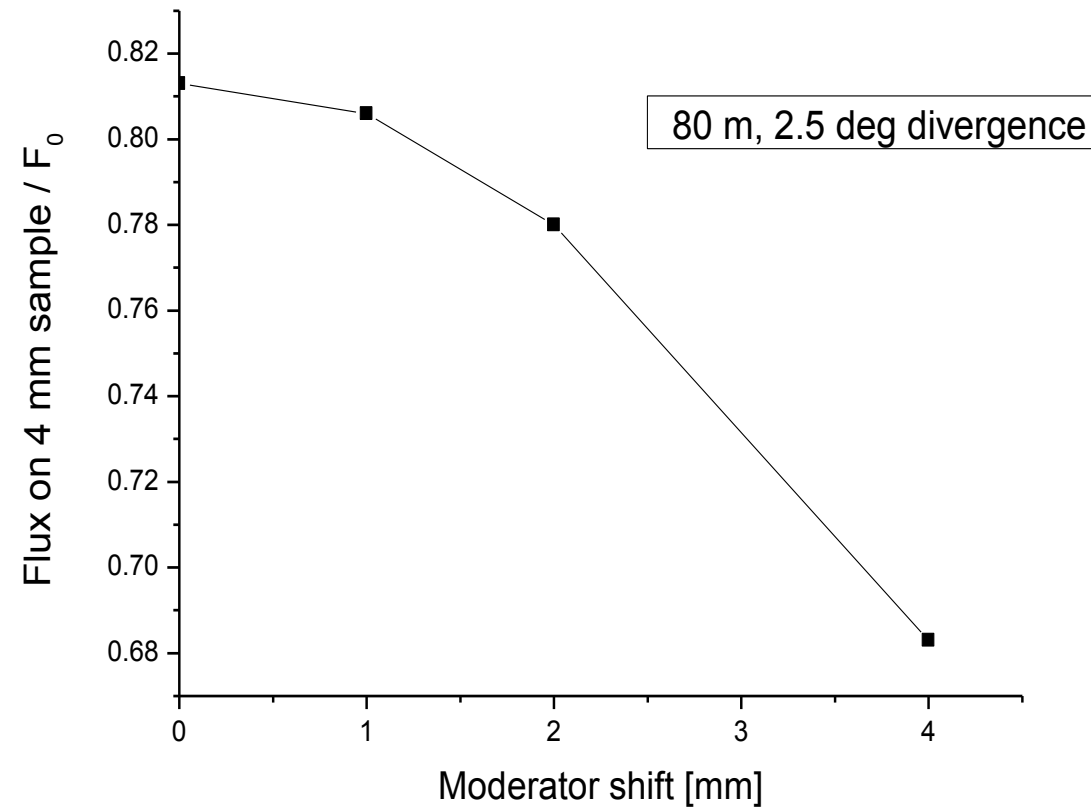
**Thank you!**

# Impact of plate misalignment



**Needs pre-aligned section >1 m**

# Impact of moderator elevation



Engineering homework