



DETECTOR BACKGROUNDS

**Fast Neutron and Gamma-ray Sensitivity
of Helium-3 and Boron-10 Detectors
and Effect of Window and Cosmic Neutrons**

Francesco Piscitelli
on behalf of the ESS Detector Group

IKON18
Lund
2020/02/25

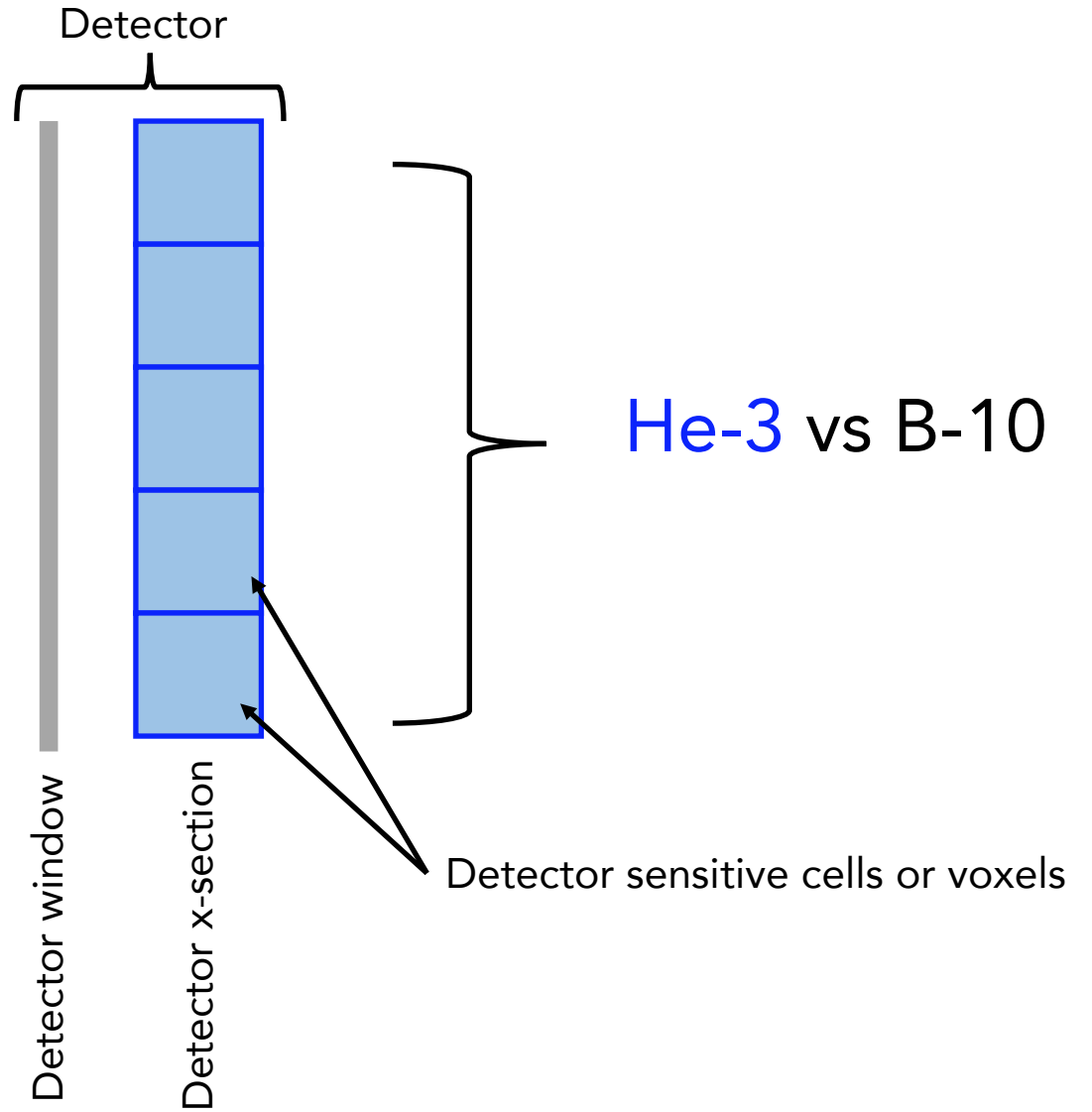
DETECTOR BACKGROUND

- Cosmic neutrons
- Gamma-rays
- Fast Neutrons
- Scattered Neutrons

} He-3 vs B-10

DETECTOR BACKGROUND

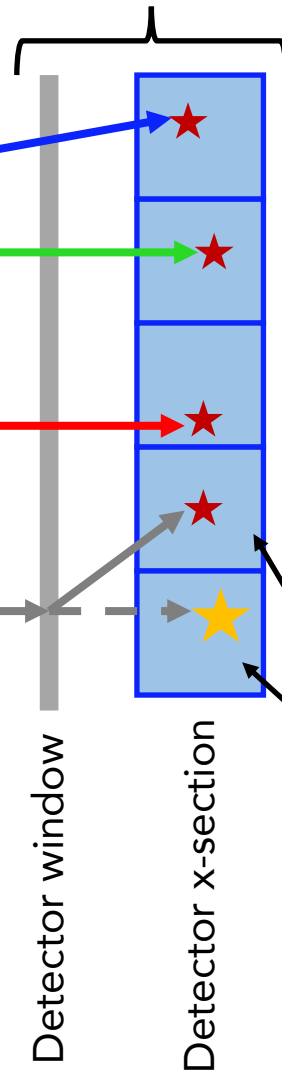
- Cosmic neutrons
- Gamma-rays
- Fast Neutrons
- Scattered Neutrons



DETECTOR BACKGROUND

- Cosmic neutrons
- Gamma-rays
- Fast Neutrons
- Scattered Neutrons

Detector



LEGEND

- ★ Good events
- ★ Unwanted events

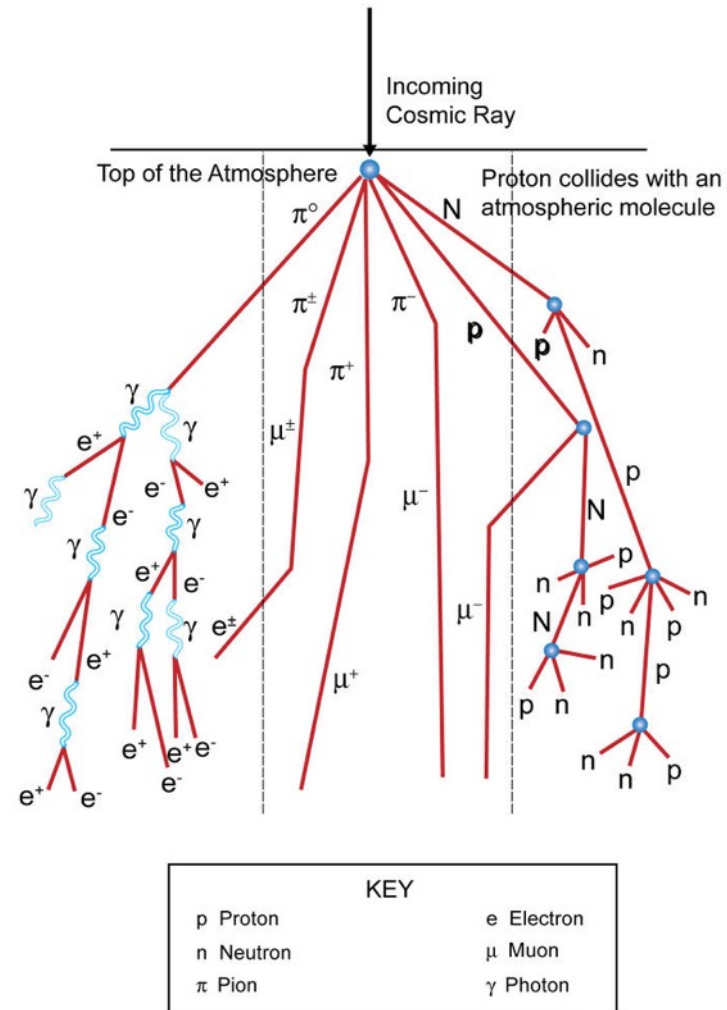
He-3 vs B-10

Detector sensitive cells or voxels

COSMIC NEUTRON BACKGROUND

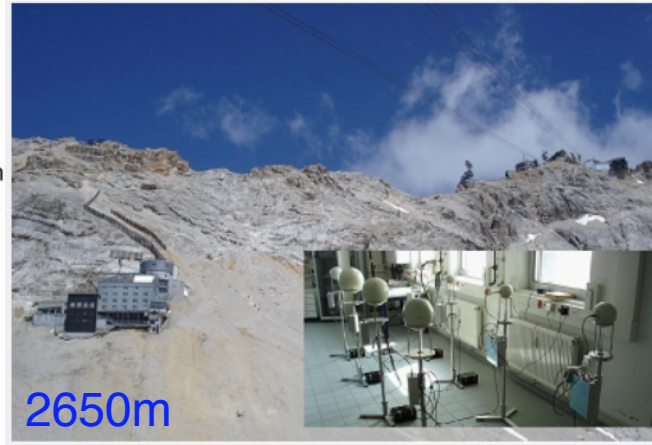
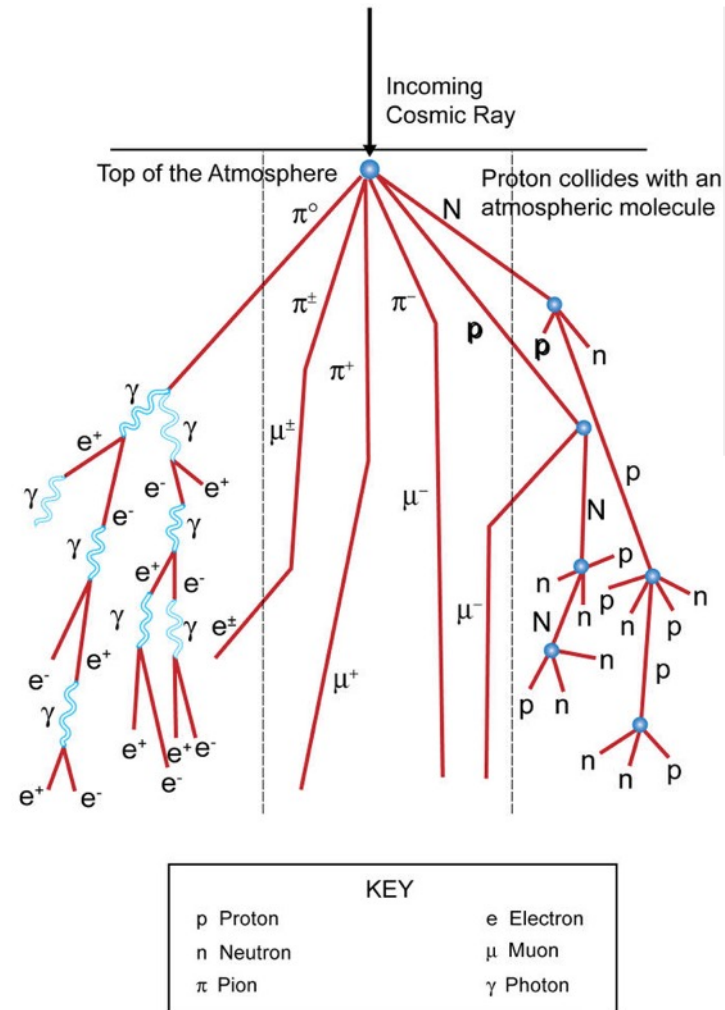
BACKGROUND: cosmic neutrons

Neutrons are created by cosmic ray spallation in the high atmosphere.
Energies from 10^{-9} to 10^3 MeV

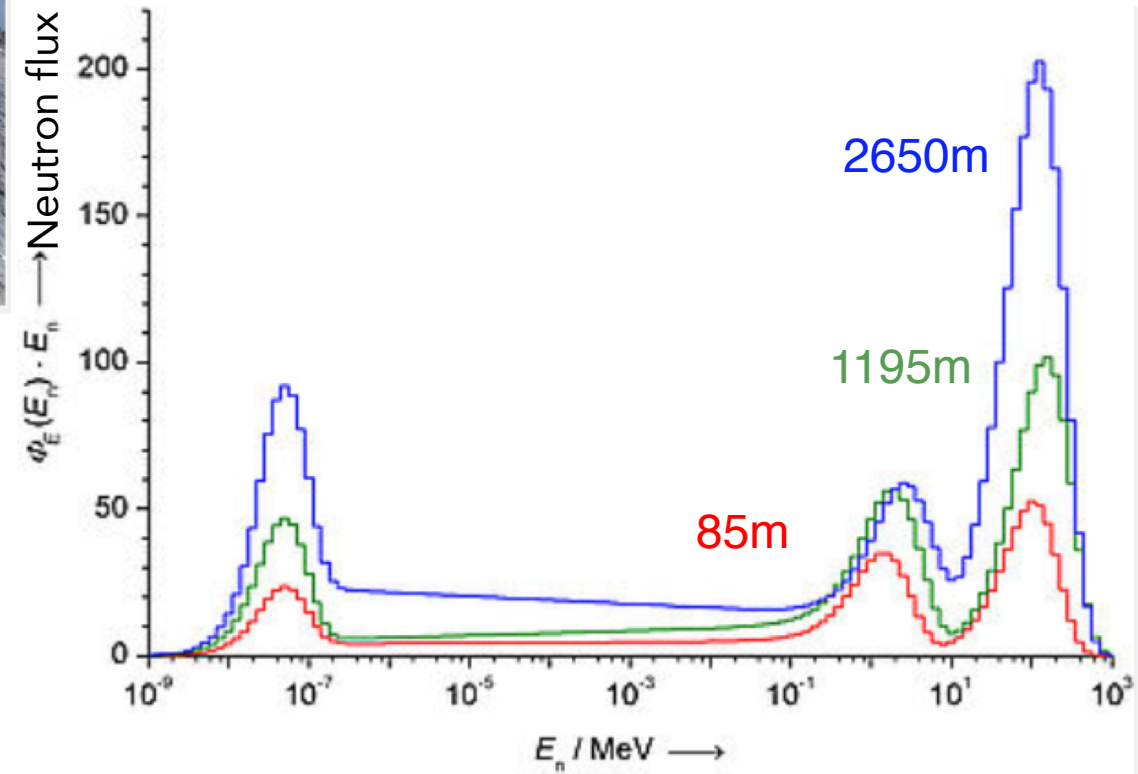


BACKGROUND: cosmic neutrons

Neutrons are created by cosmic ray spallation in the high atmosphere.
Energies from 10^{-9} to 10^3 MeV



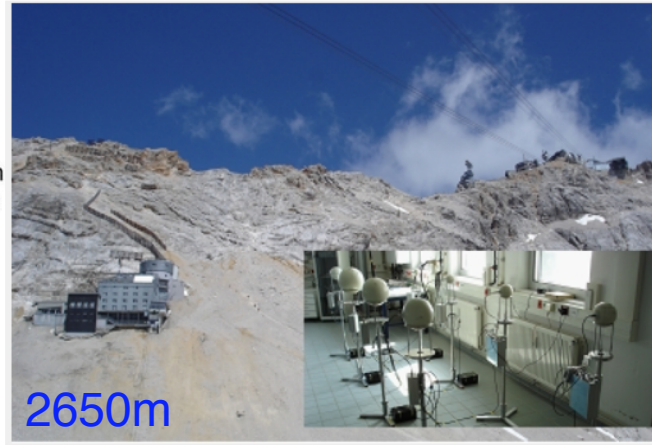
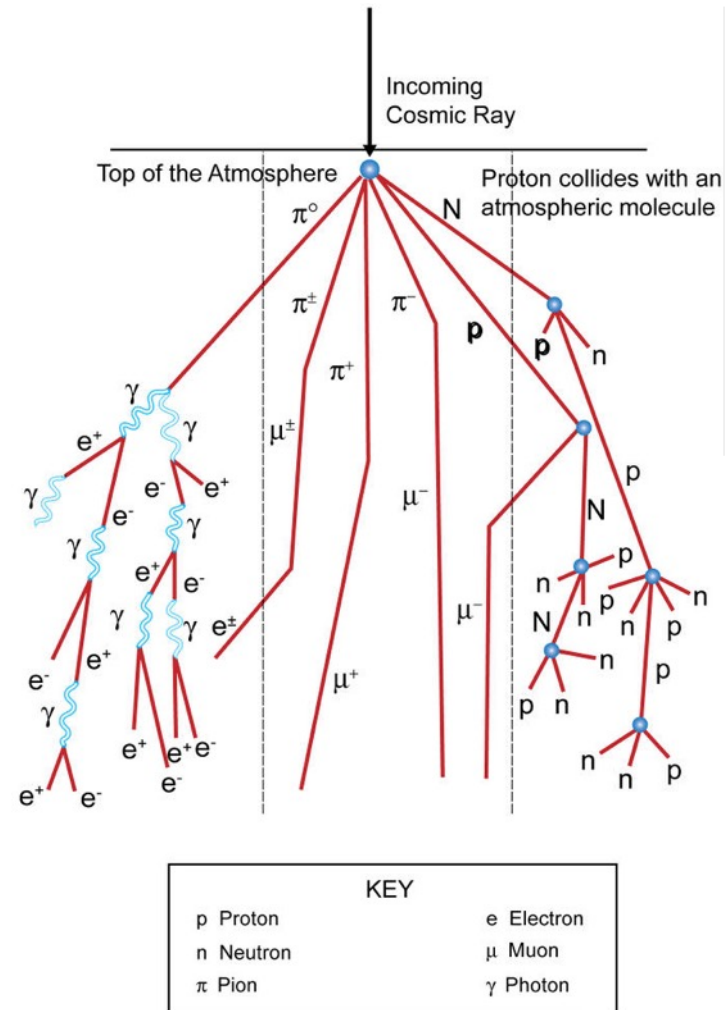
Measured with Bonner spheres in the whole energy spectrum



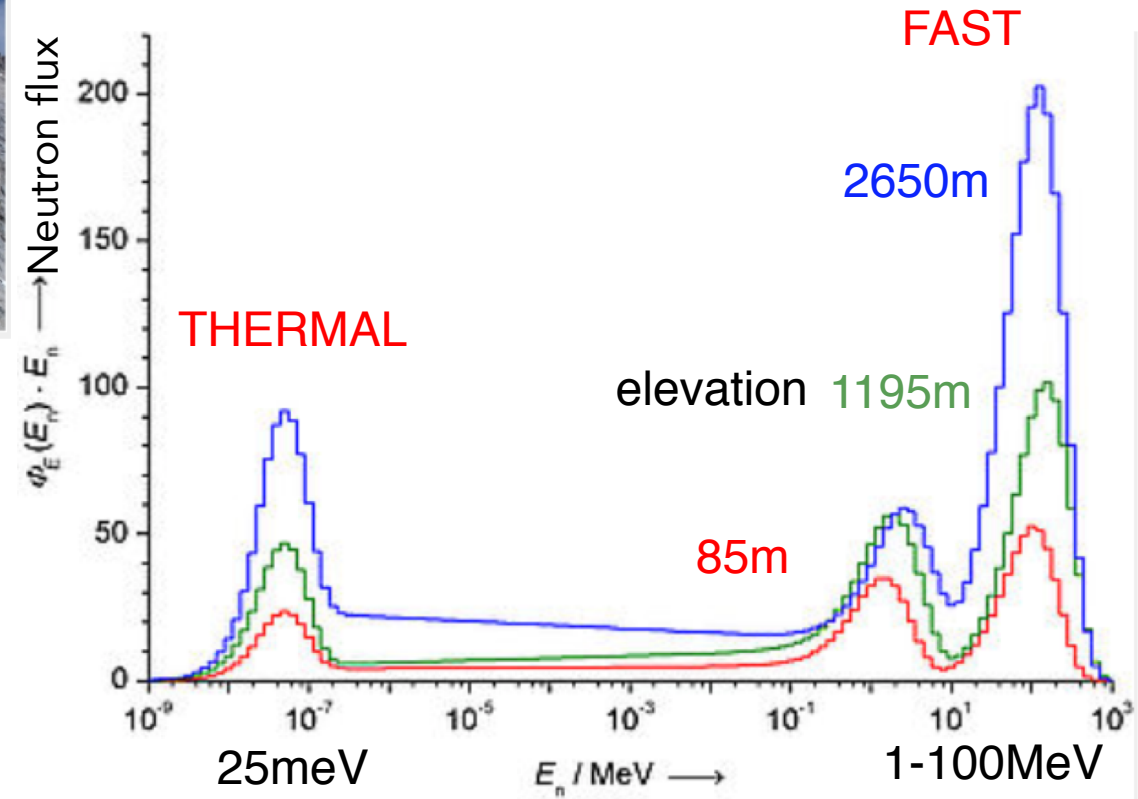
Physikalisch-Technische Bundesanstalt (PTB) - Measurements of Neutron Spectra Induced by Cosmic Radiation at Altitudes of 85m, 1195m and 2650m (2010)

BACKGROUND: cosmic neutrons

Neutrons are created by cosmic ray spallation in the high atmosphere.
Energies from 10^{-9} to 10^3 MeV



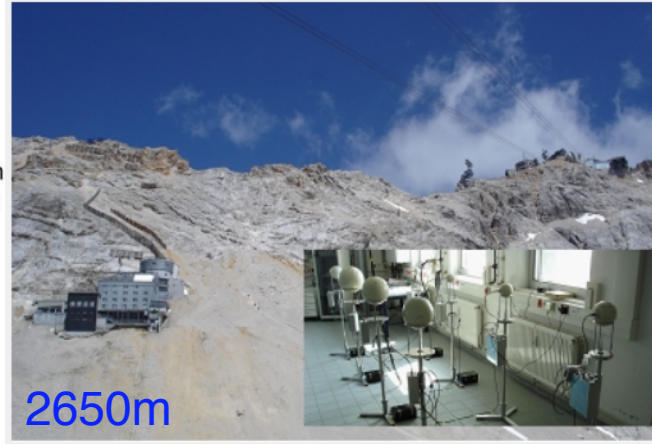
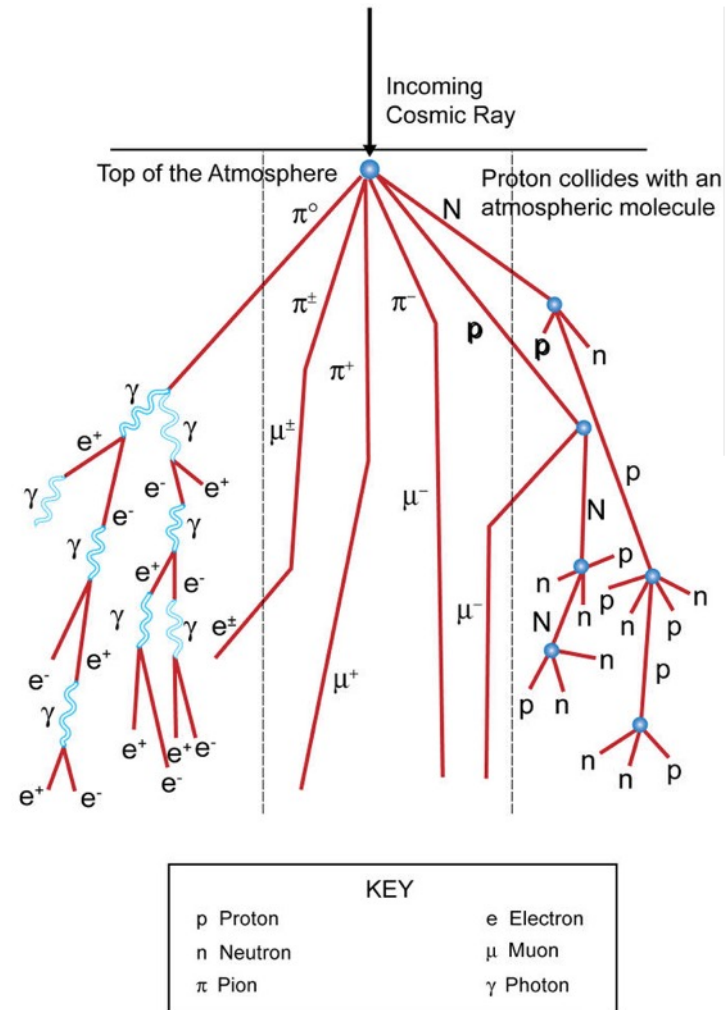
Measured with Bonner spheres in the whole energy spectrum



Physikalisch-Technische Bundesanstalt (PTB) - Measurements of Neutron Spectra Induced by Cosmic Radiation at Altitudes of 85m, 1195m and 2650m (2010)

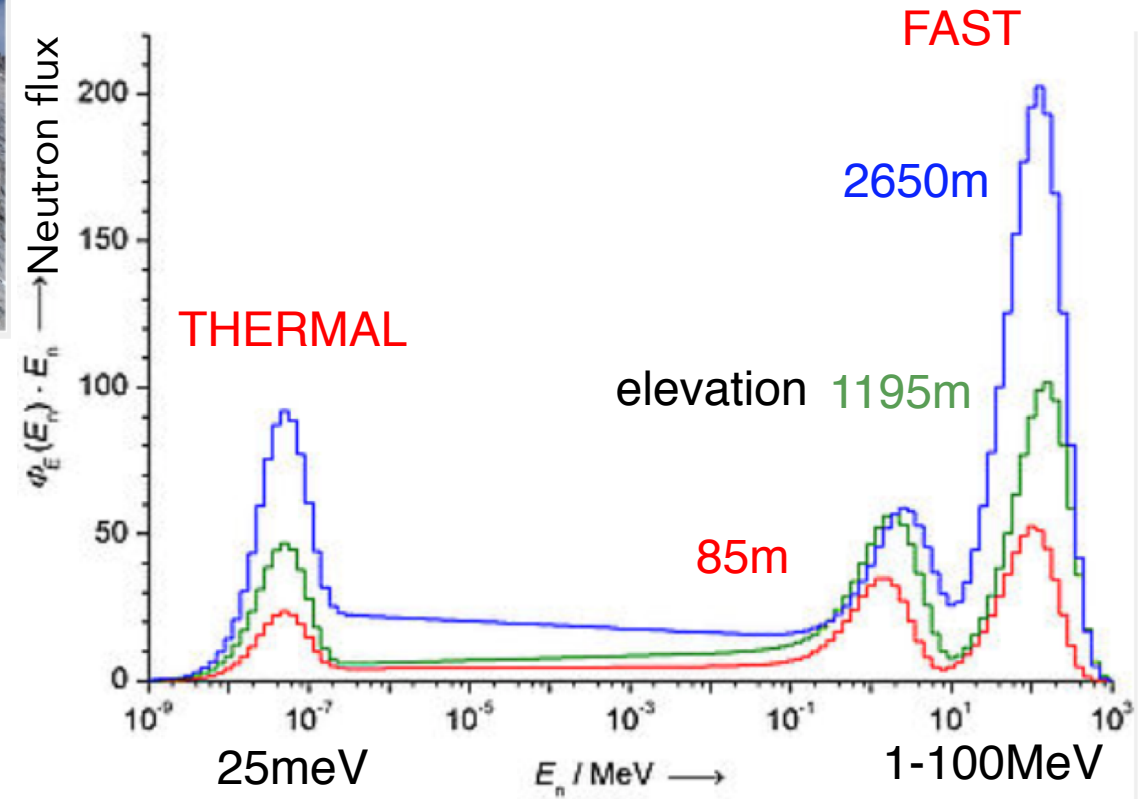
BACKGROUND: cosmic neutrons

Neutrons are created by cosmic ray spallation in the high atmosphere.
Energies from 10^{-9} to 10^3 MeV



Measured with Bonner spheres in the whole energy spectrum

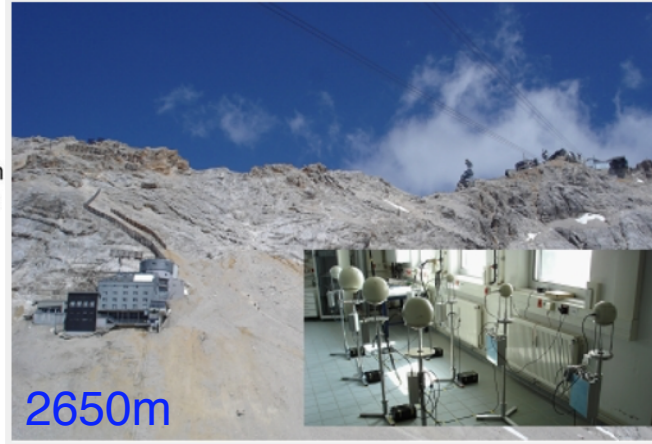
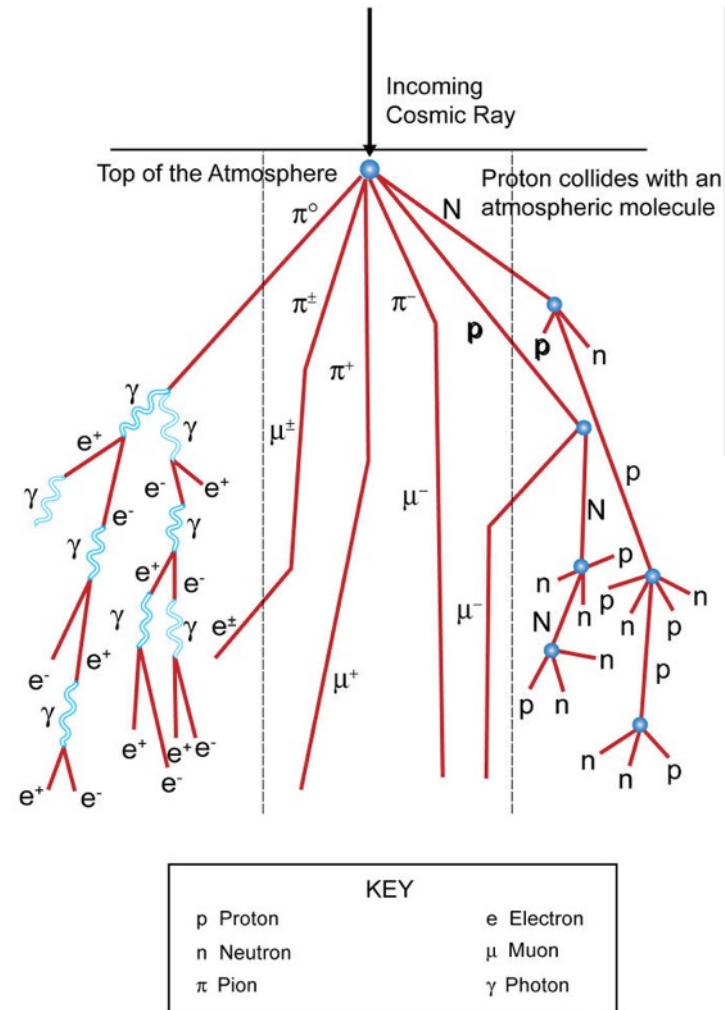
elevation 85m ~ ESS
135 Hz/m² (10^{-9} to 10^3 MeV) by PTB



Physikalisch-Technische Bundesanstalt (PTB) - Measurements of Neutron Spectra Induced by Cosmic Radiation at Altitudes of 85m, 1195m and 2650m (2010)

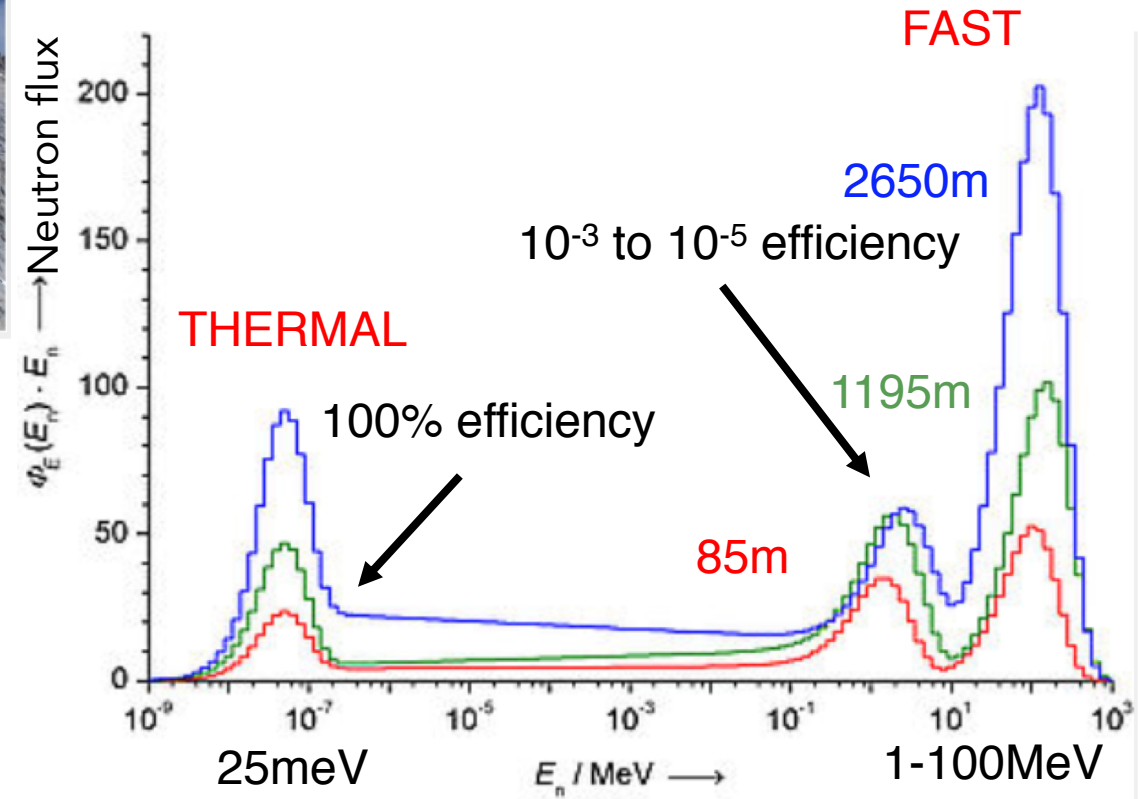
BACKGROUND: cosmic neutrons

Neutrons are created by cosmic ray spallation in the high atmosphere.
Energies from 10^{-9} to 10^3 MeV



Measured with Bonner spheres in the whole energy spectrum

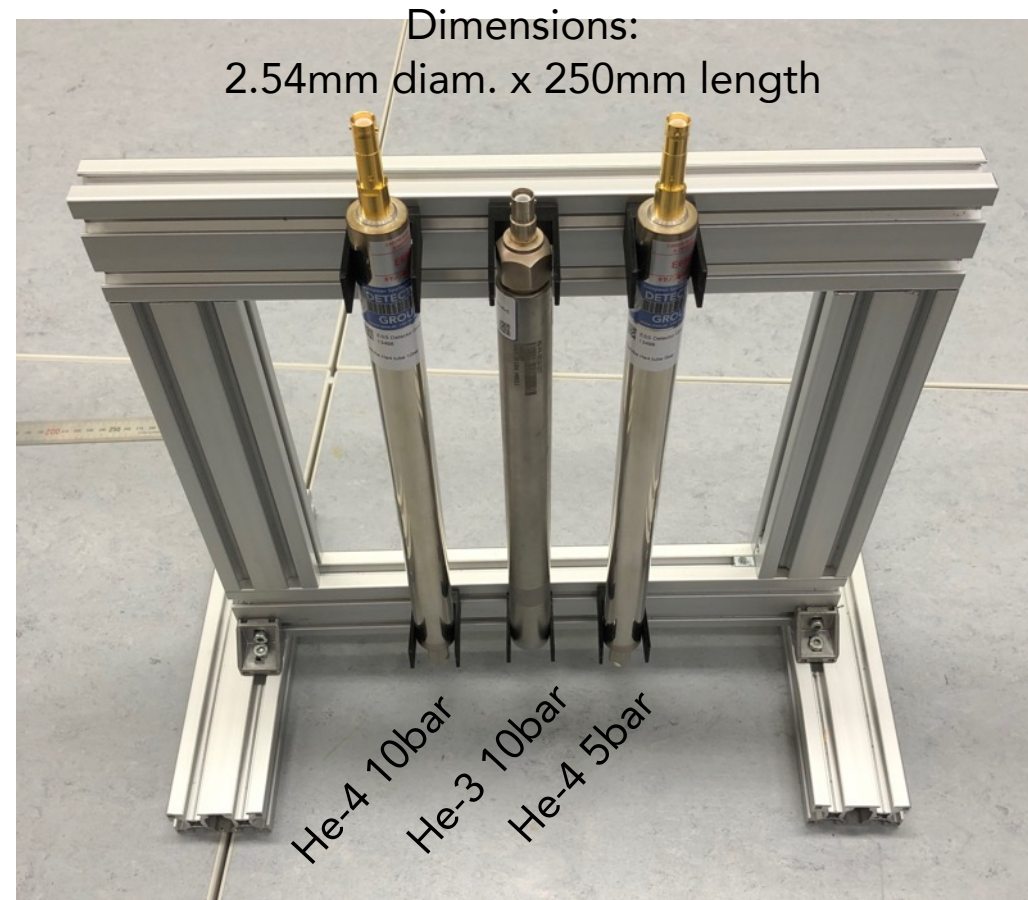
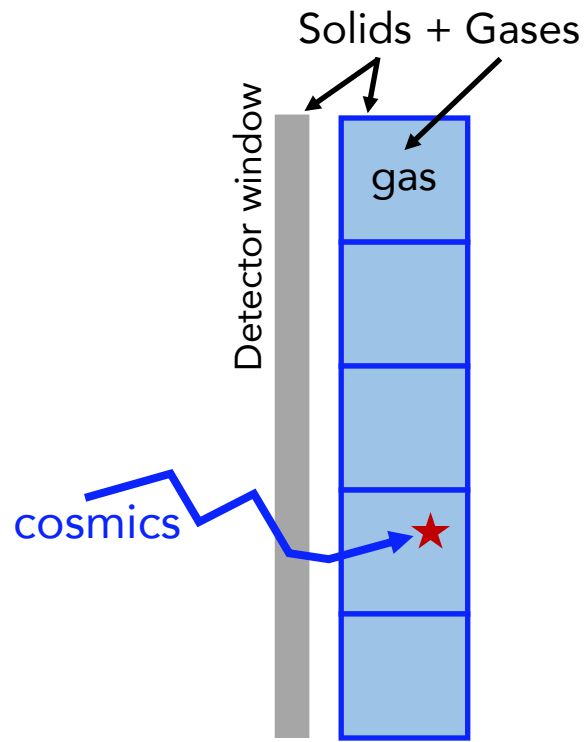
elevation 85m ~ ESS
135 Hz/m² (10^{-9} to 10^3 MeV) by PTB



Physikalisch-Technische Bundesanstalt (PTB) - Measurements of Neutron Spectra Induced by Cosmic Radiation at Altitudes of 85m, 1195m and 2650m (2010)

BACKGROUND: cosmic neutrons

Measurements done in Utgård

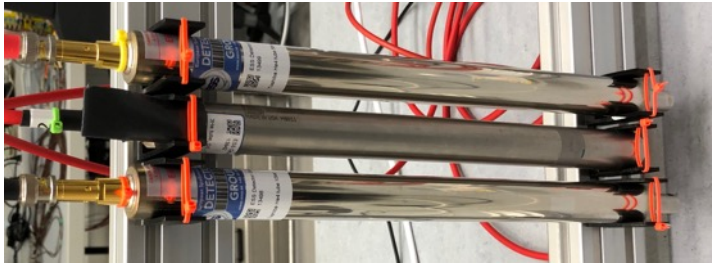


He-3 and He-4 detectors

*Thanks to Toshiba/Canon Electron Tubes & Devices Co. LTD for the He-4 tubes.

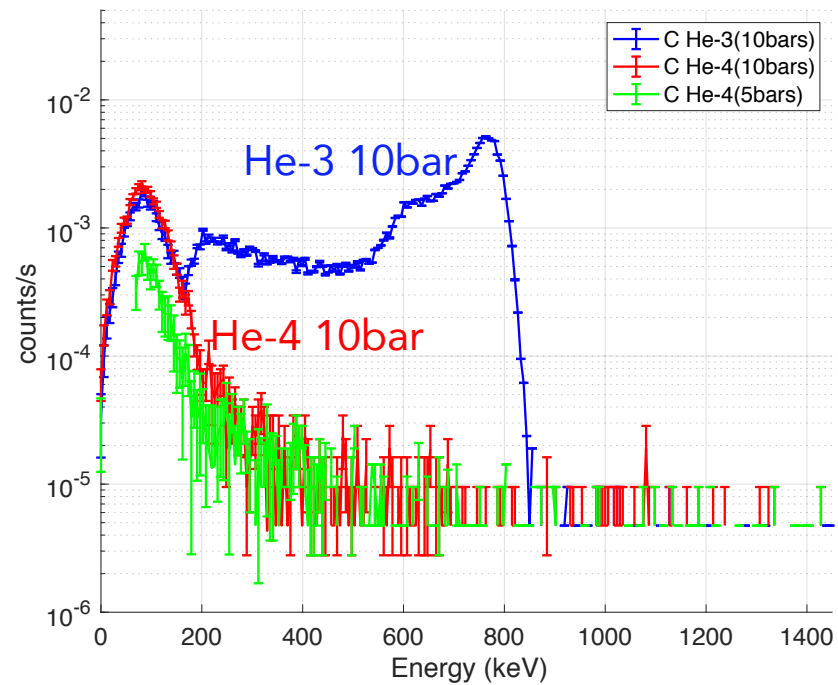
BACKGROUND: cosmic neutrons

Bare tubes



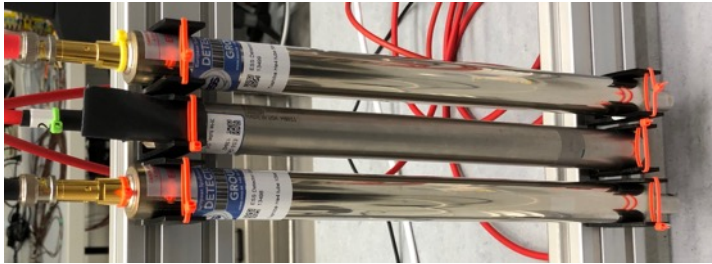
He-4 5 bar
He-3 10 bar
He-4 10 bar

Pulse-Height-Spectrum



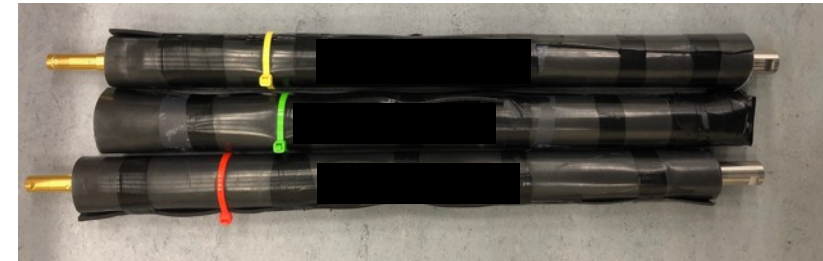
BACKGROUND: cosmic neutrons

Bare tubes

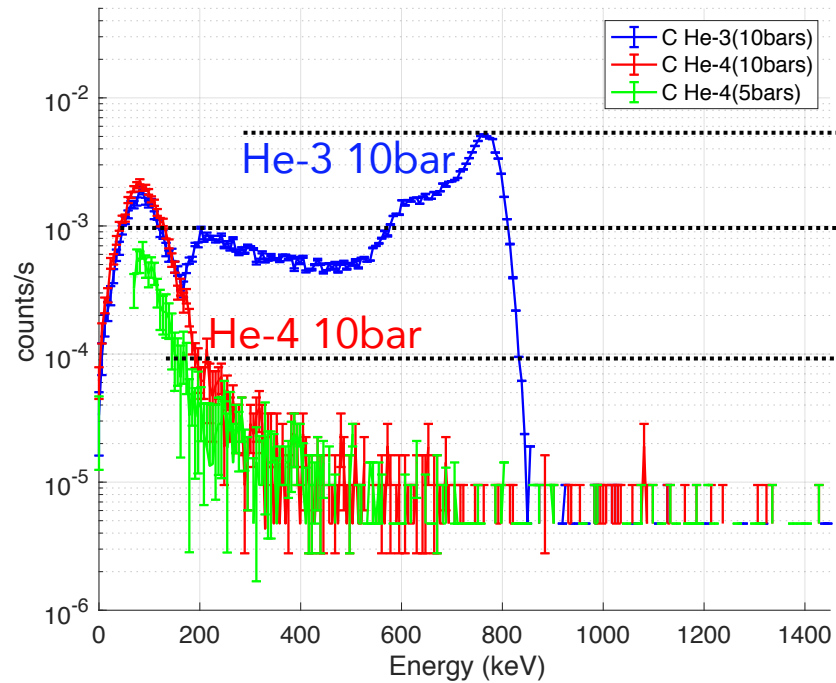


He-4 5 bar
He-3 10 bar
He-4 10 bar

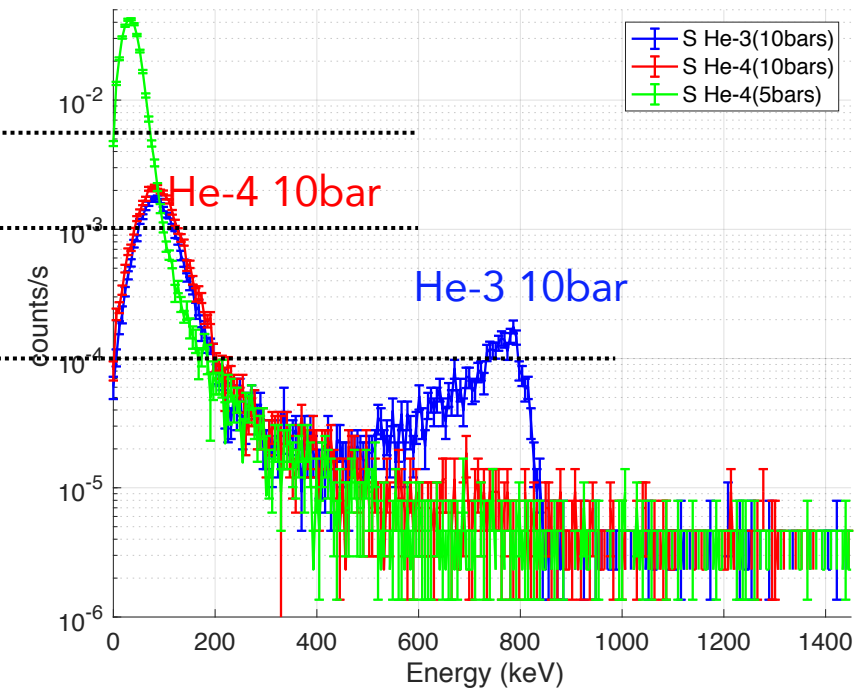
Covered with 2mm Mirrobor



Pulse-Height-Spectrum

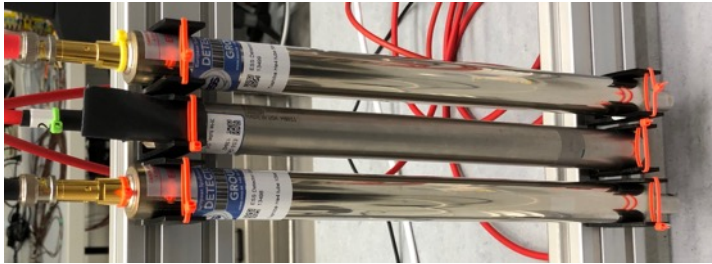


Pulse-Height-Spectrum



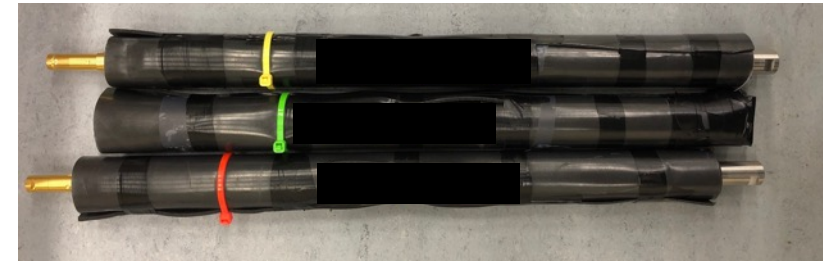
BACKGROUND: cosmic neutrons

Bare tubes

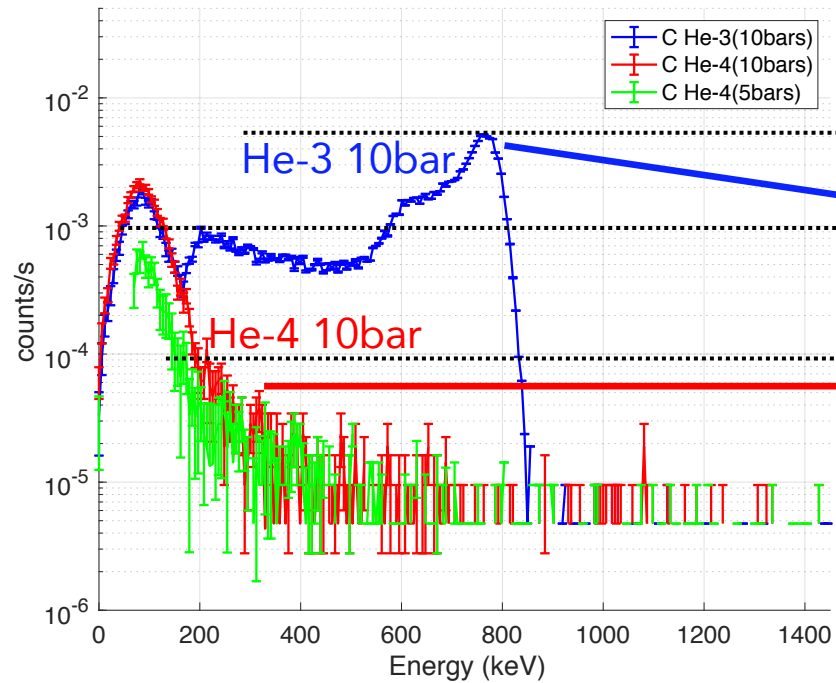


He-4 5 bar
He-3 10 bar
He-4 10 bar

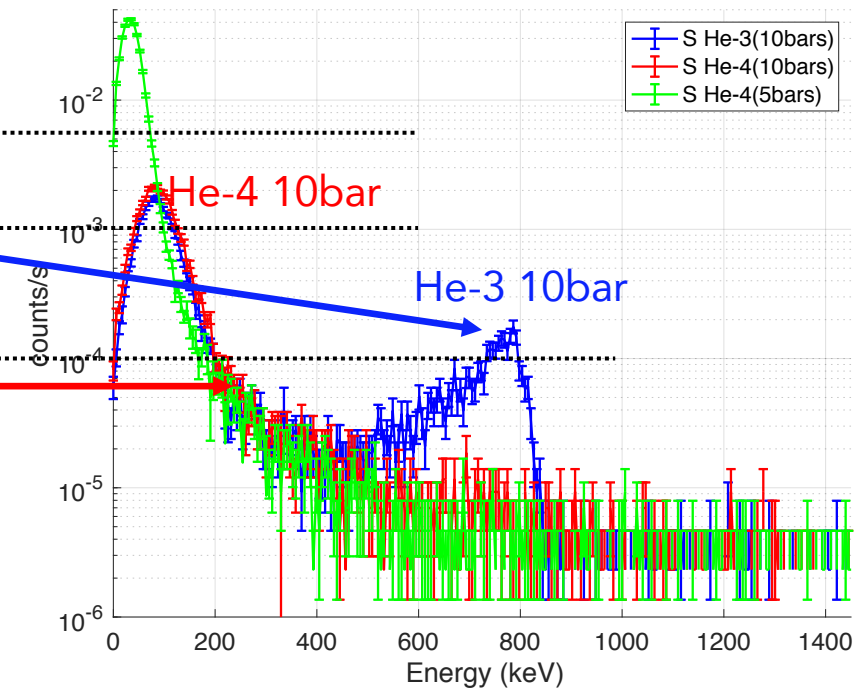
Covered with 2mm Mirrobor



Pulse-Height-Spectrum

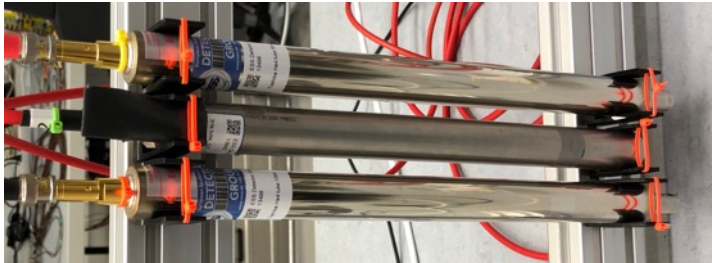


Pulse-Height-Spectrum



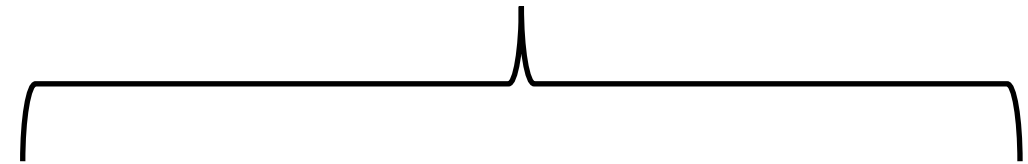
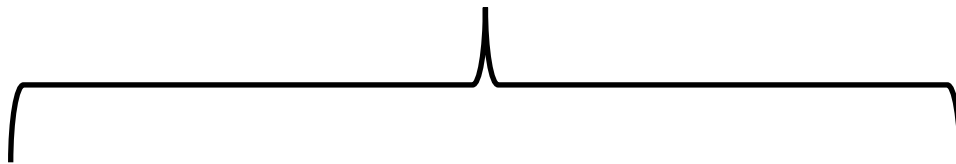
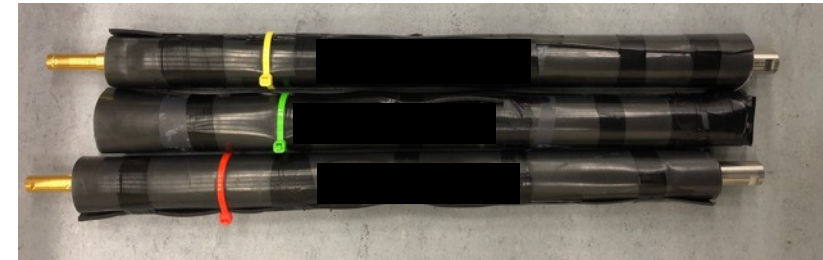
BACKGROUND: cosmic neutrons

Bare tubes



He-4 5 bar
 He-3 10 bar
 He-4 10 bar

Covered with 2mm Mirrobor



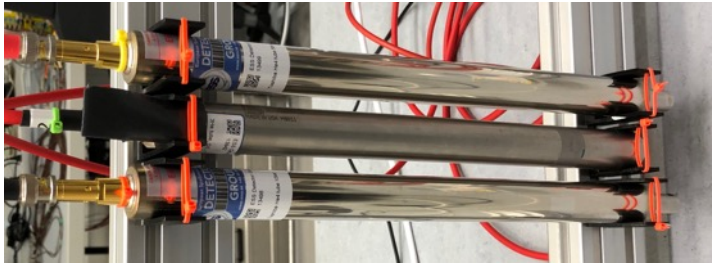
Rate Hz	rate per area Hz/m ²	rate per volume Hz/m ³	detector	Rate Hz	rate per area Hz/m ²	rate per volume Hz/m ³
0.148	23	292	He-3 10bar	0.007	1	14
0.0043	0.7	8.5	He-4 10bar	0.005	0.8	10

0.029 Hz / (bar · litre) of He3

0.0014 Hz / (bar · litre) of He3

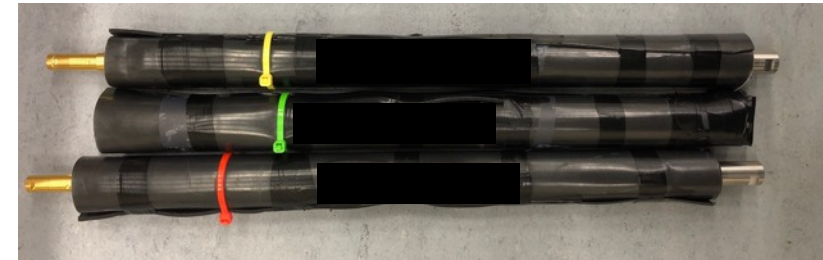
BACKGROUND: cosmic neutrons

Bare tubes



He-4 5 bar
 He-3 10 bar
 He-4 10 bar

Covered with 2mm Mirrobor



Rate Hz	rate per area Hz/m ²	rate per volume Hz/m ³	detector	Rate Hz	rate per area Hz/m ²	rate per volume Hz/m ³
0.148	23	292	He-3 10bar	0.007	1	14
0.0043	0.7	8.5	He-4 10bar	0.005	0.8	10

0.029 Hz / (bar · litre) of He3

0.0014 Hz / (bar · litre) of He3

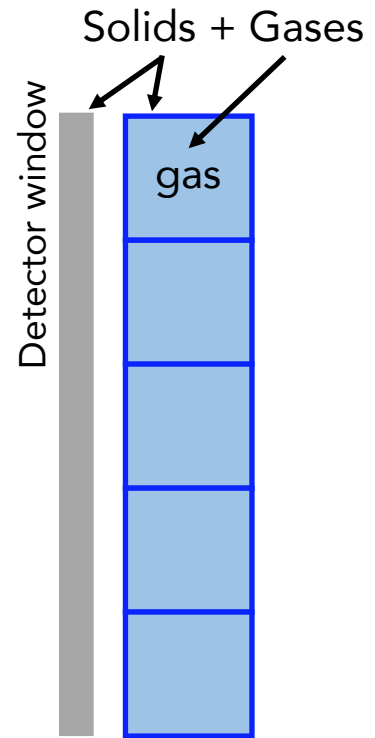
elevation 85m ~ ESS
 135 Hz/m² (10⁻⁹ to 10³ MeV) by PTB

Agreement with rates at FRMII
 K. Zeitelhack – private communication

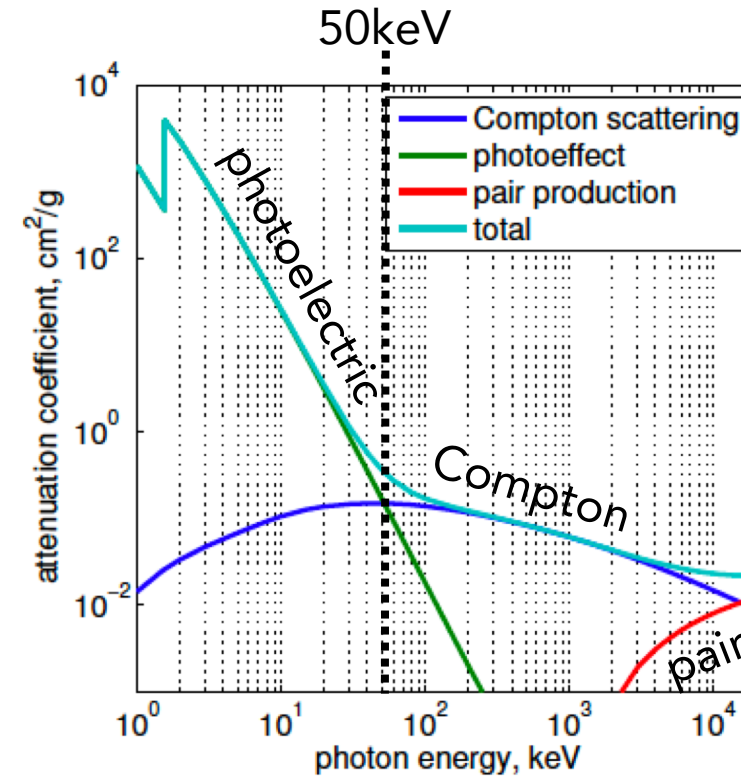
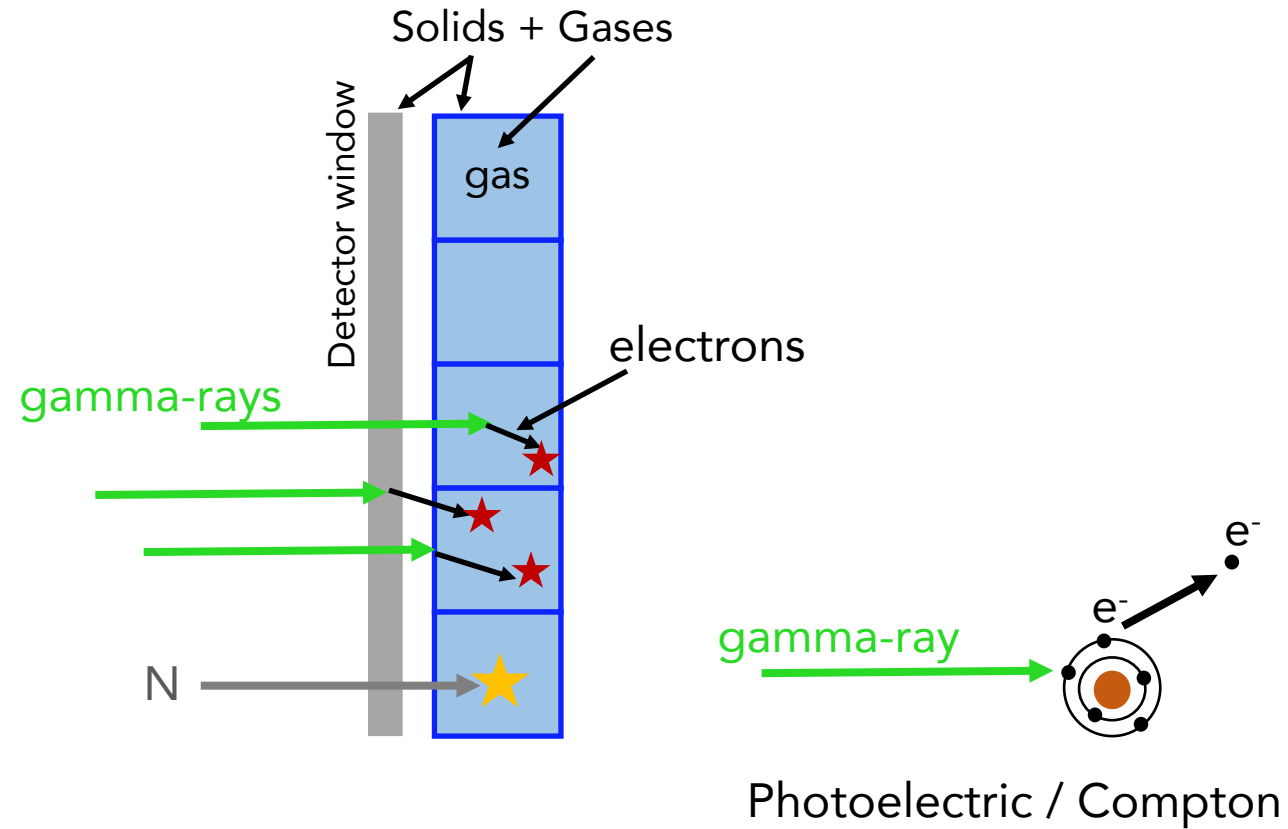
Agreement with rates measured with MG at
 Utgård (See next talk from Alex)

GAMMA-RAY BACKGROUND

BACKGROUND: **gamma-rays**

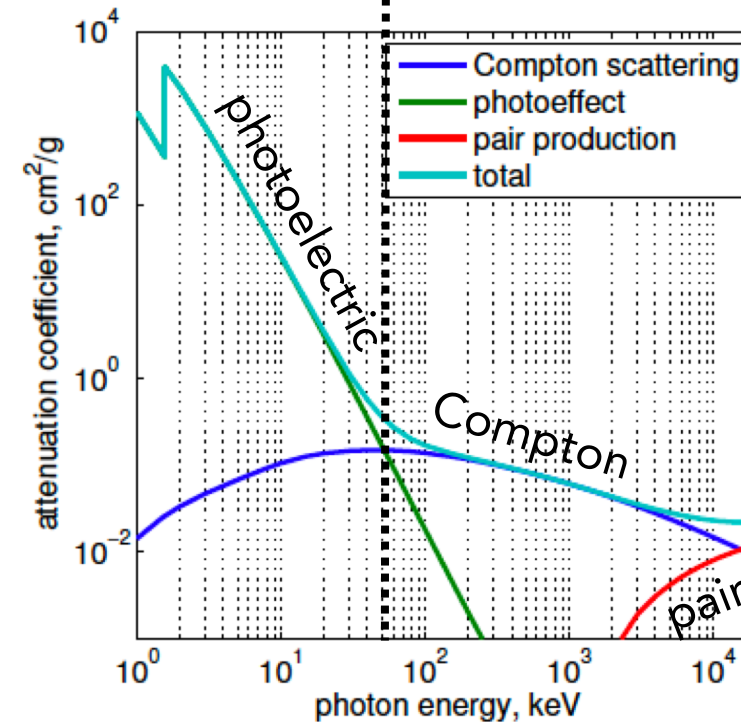
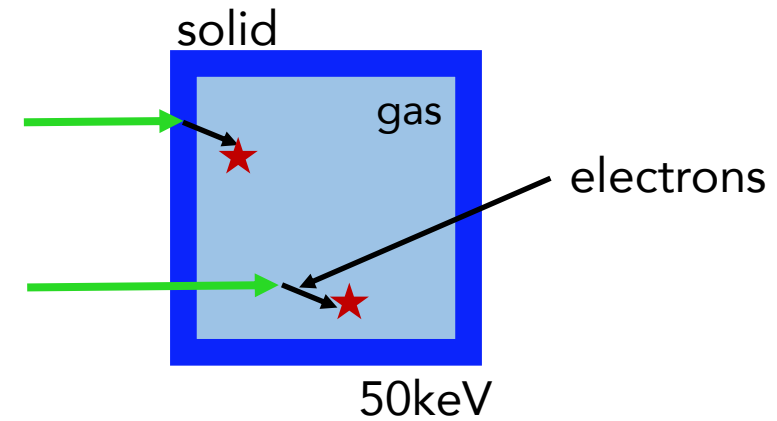
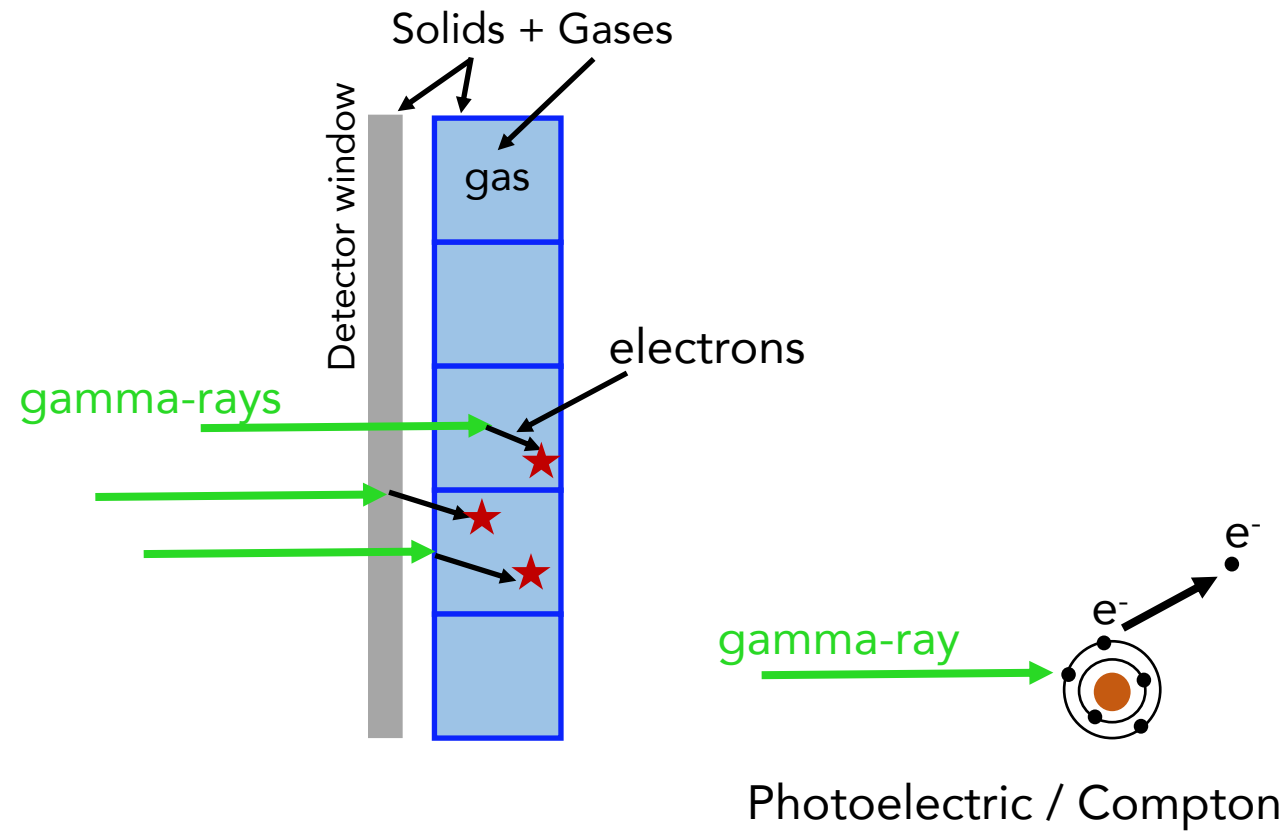


BACKGROUND: gamma-rays



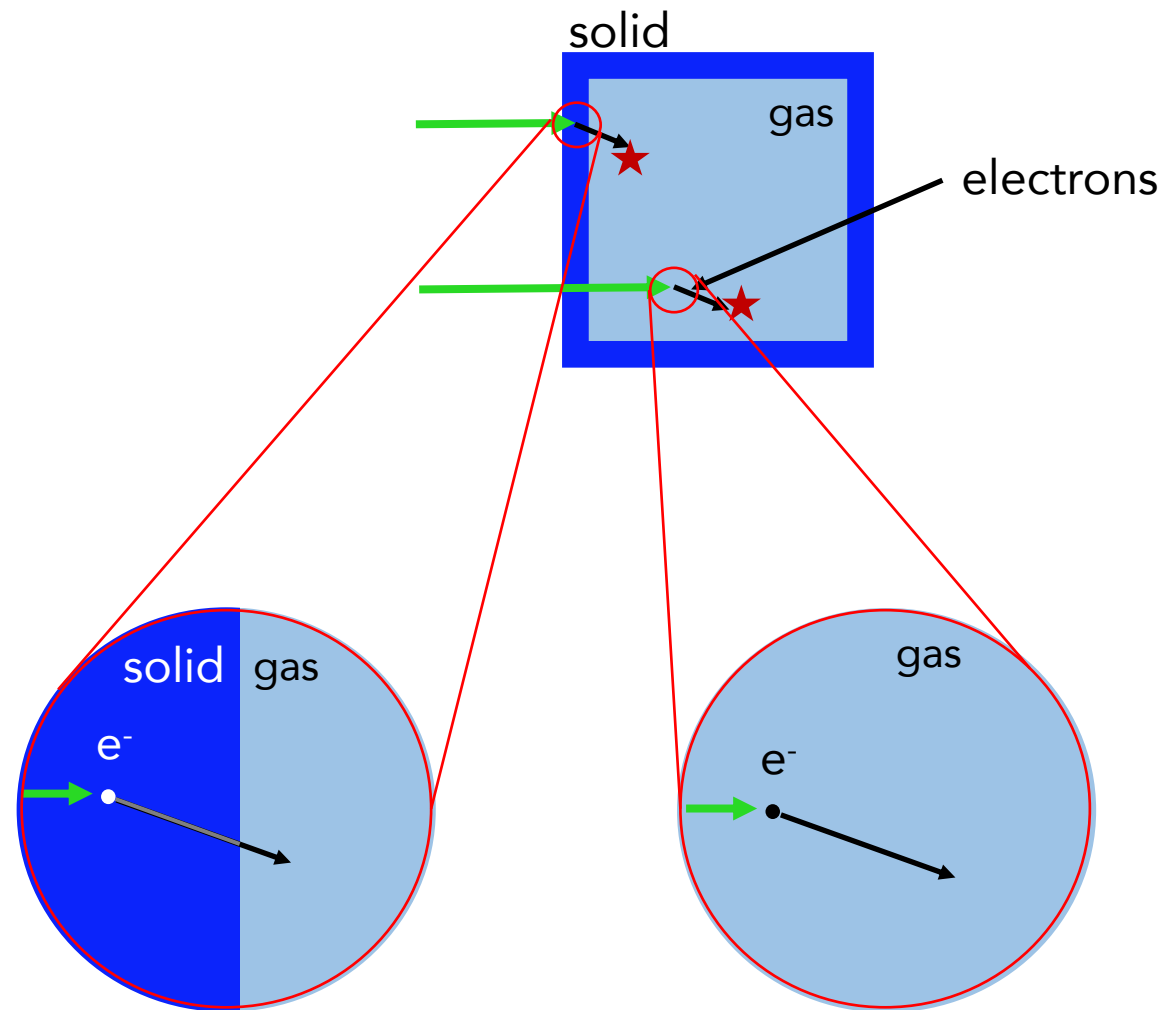
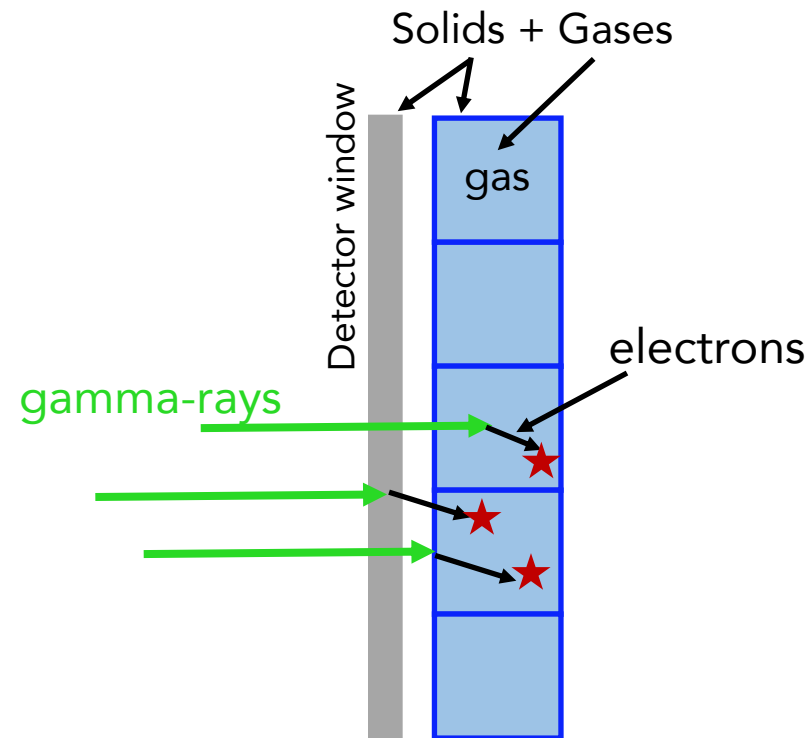
Photoelectric interaction most probable at low energy, Compton scattering is the majority of the background

BACKGROUND: gamma-rays



Photoelectric interaction most probable at low energy, Compton scattering is the majority of the background

BACKGROUND: gamma-rays

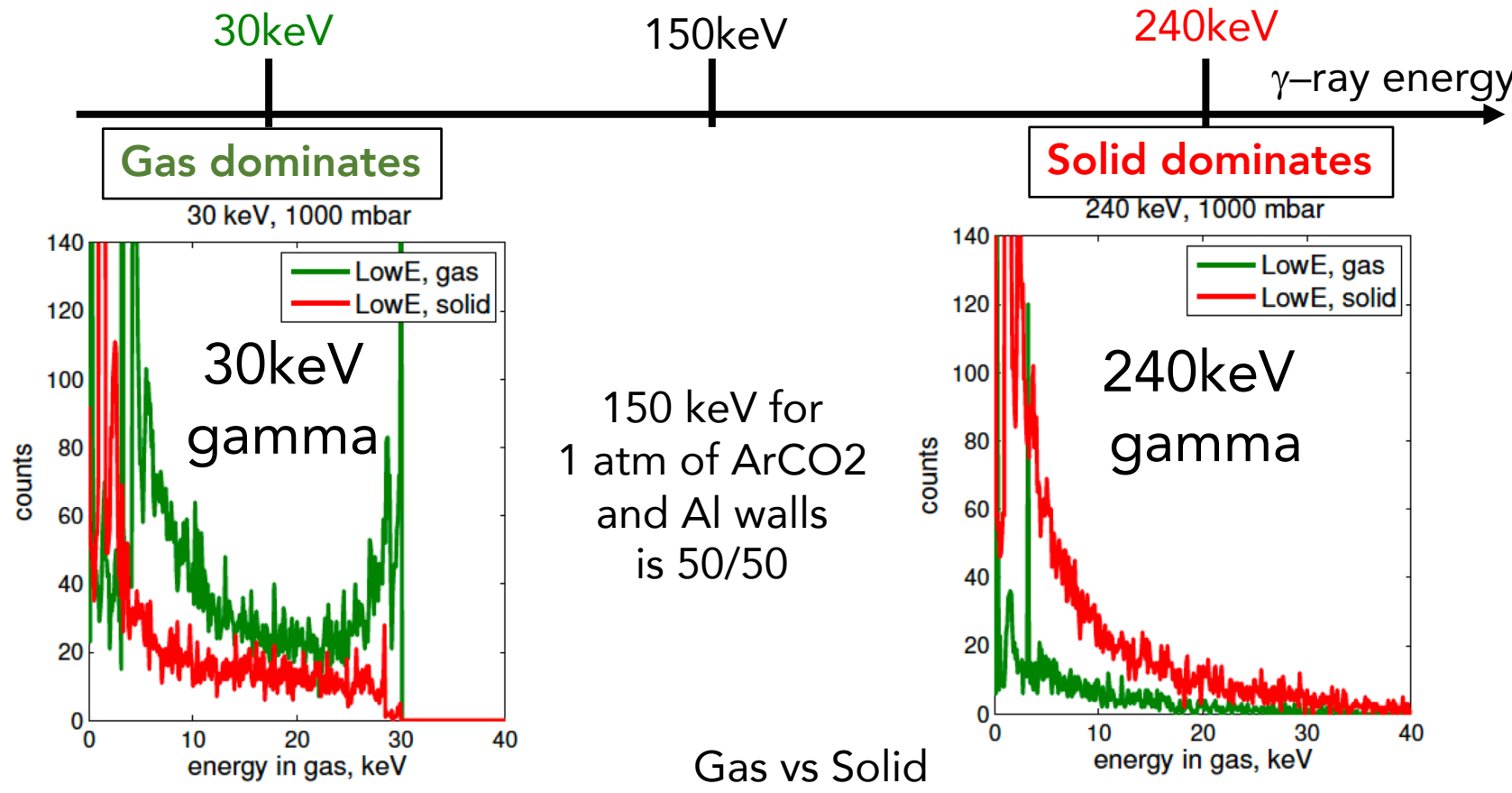
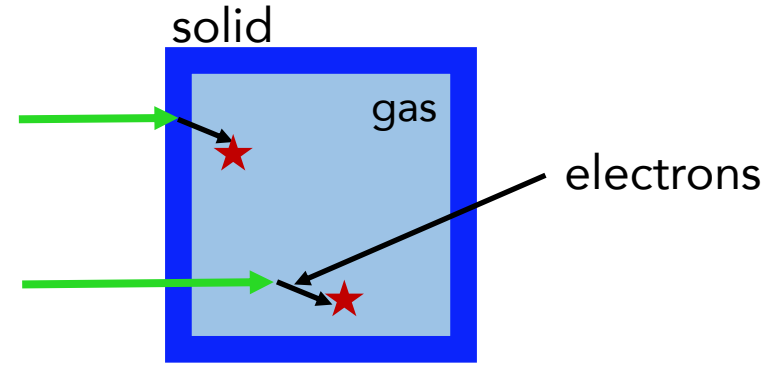
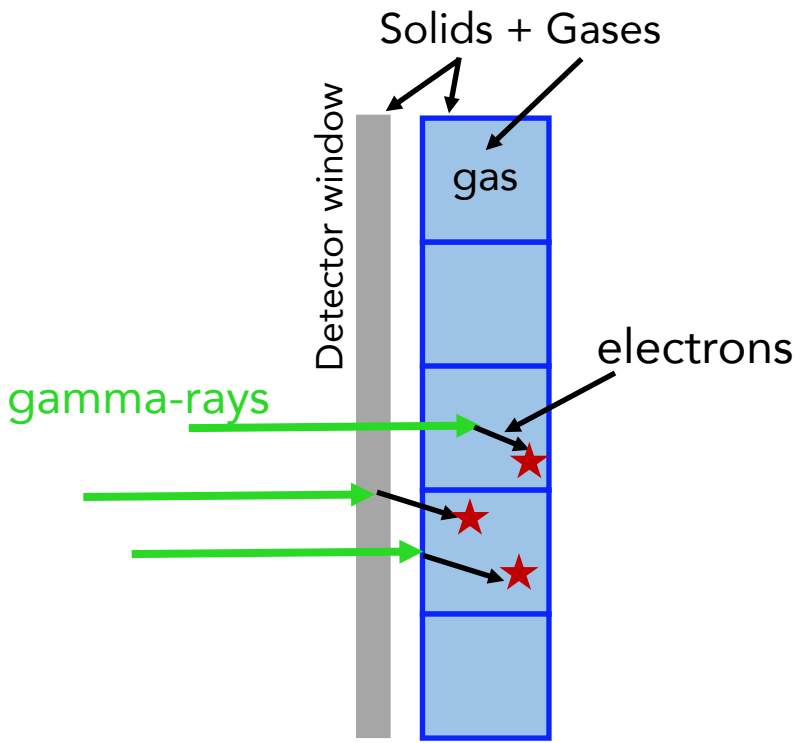


gammas in the walls can only result in a signal if the electron reaches the gas

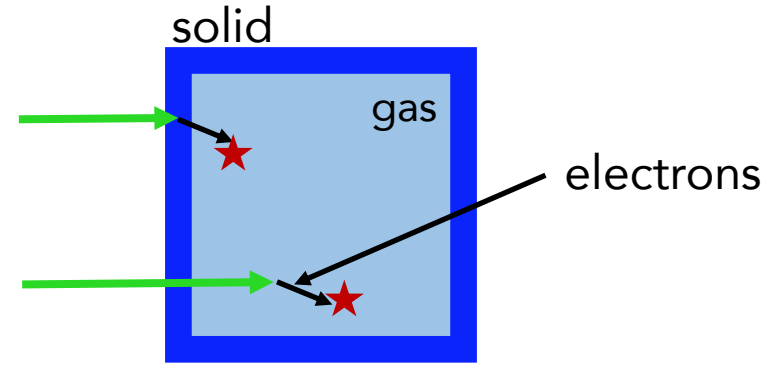
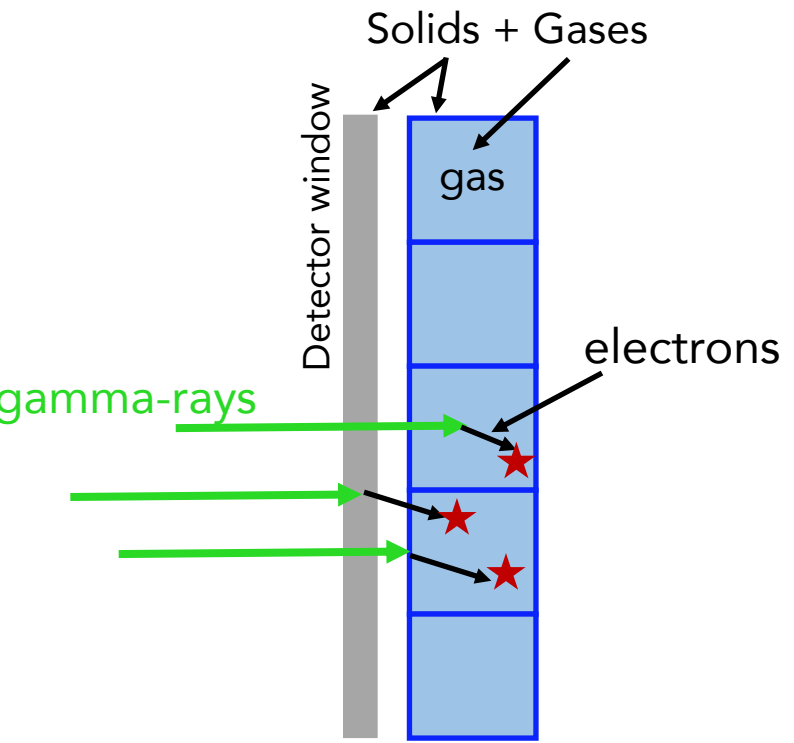
photons interacting with the gas are guaranteed to generate a signal

Gas vs Solid

BACKGROUND: gamma-rays



BACKGROUND: gamma-rays



Energy threshold generally 100-200keV

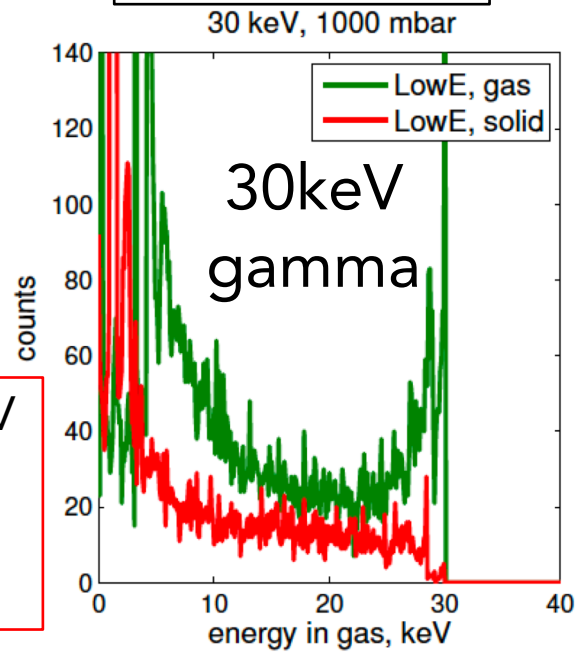
↓

Solid dominates



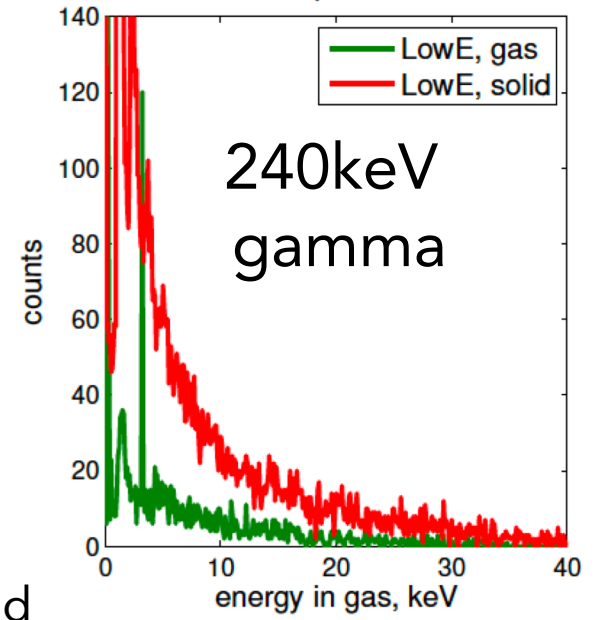
Gas dominates

Solid dominates

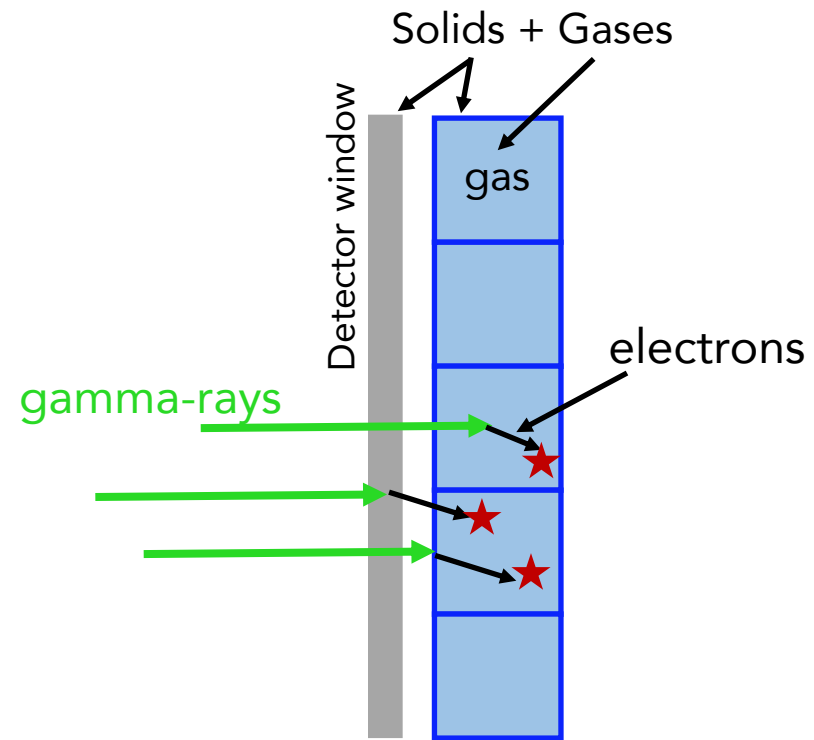


150 keV for
1 atm of ArCO₂
and Al walls
is 50/50

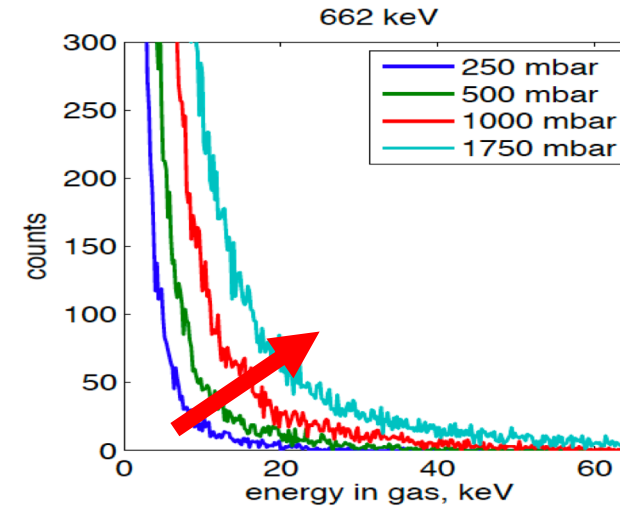
Gas vs Solid



BACKGROUND: gamma-rays

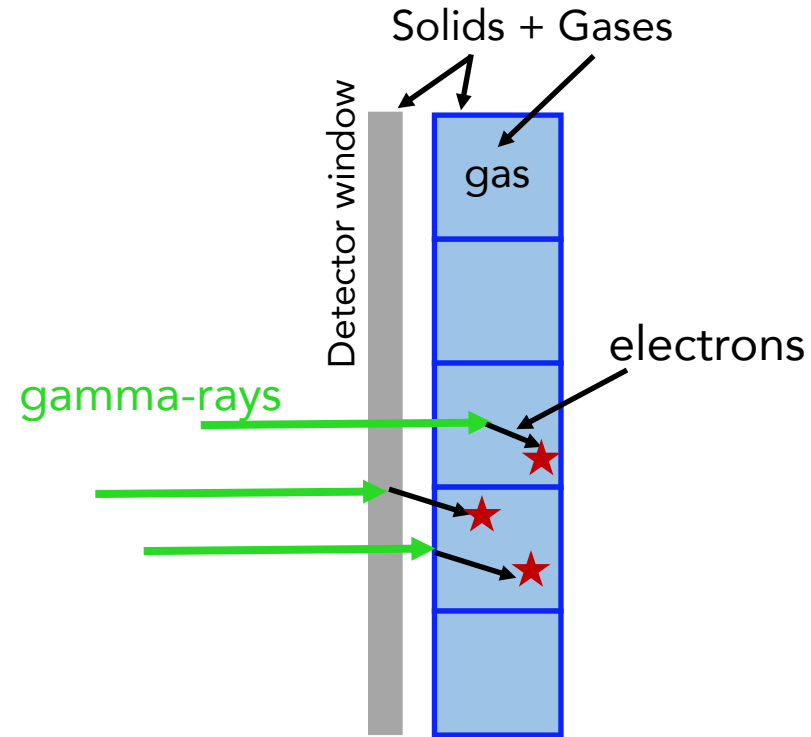


Higher pressure = higher energy deposition

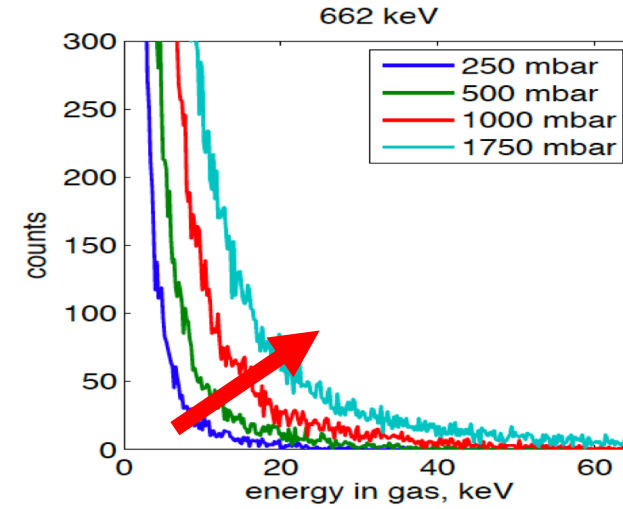


(Simulations)

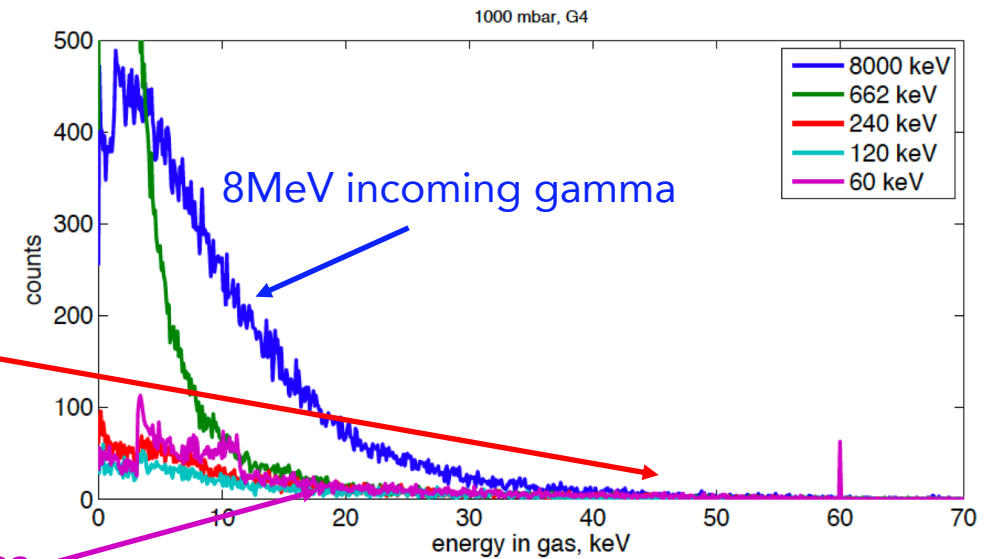
BACKGROUND: gamma-rays



Higher pressure = higher energy deposition



Higher gamma energy = higher energy deposition



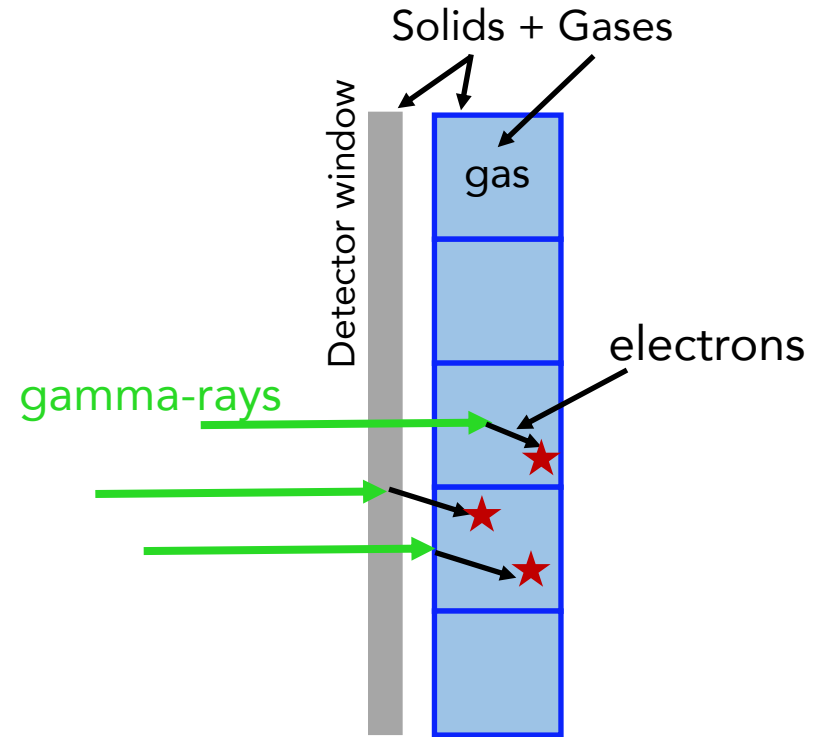
... BUT deposited energies rarely exceed a few tens of keV regardless of the primary energy

60keV incoming gamma

(Simulations)

BACKGROUND: gamma-rays

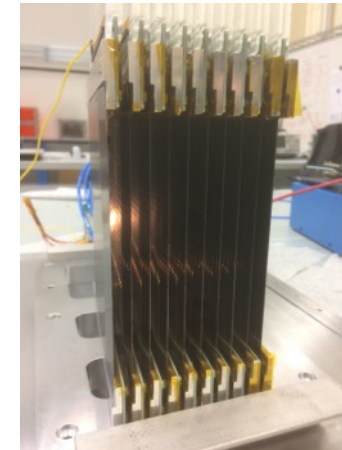
Solid dominates no matter if He-3 or B-10



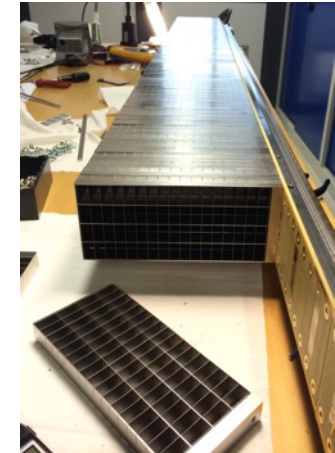
Measurements



He-3/He-4 detectors



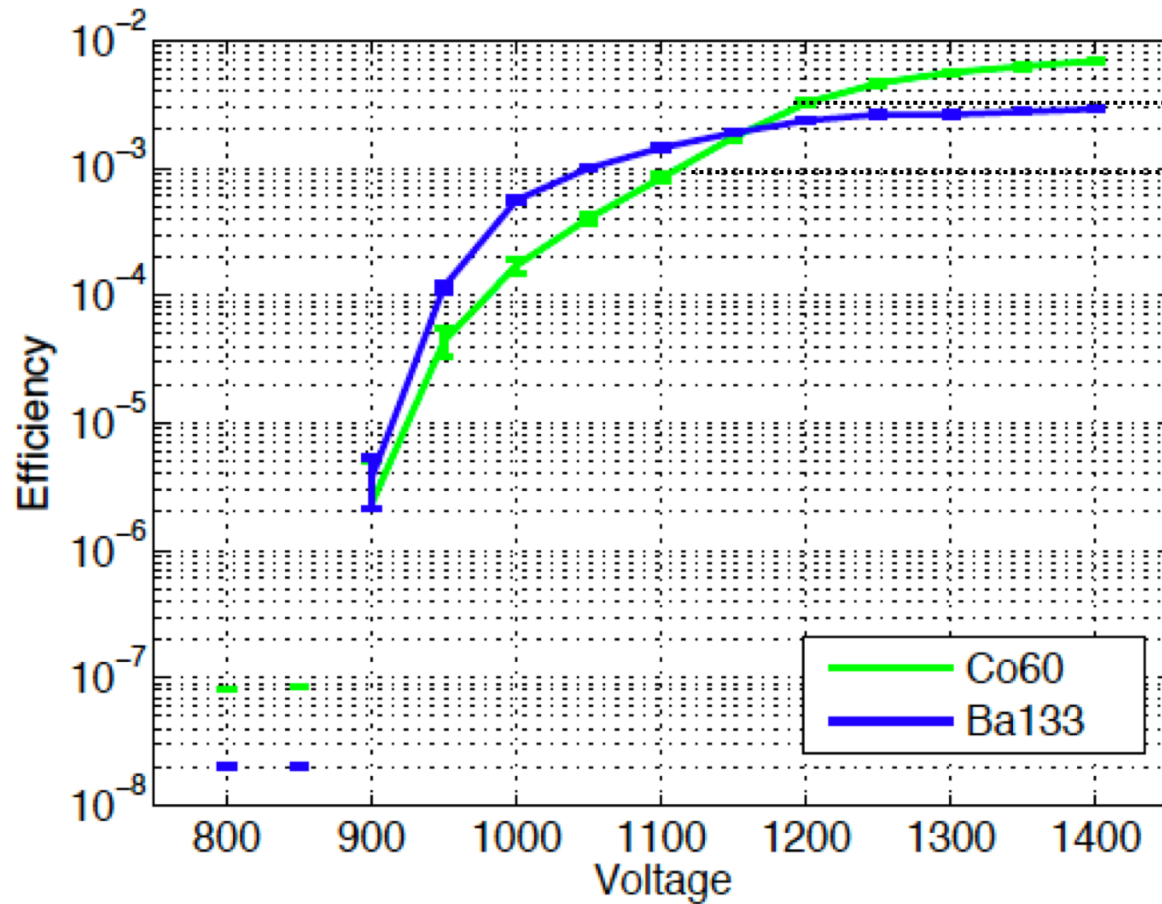
Multi-Blade B-10
(reflectometers)



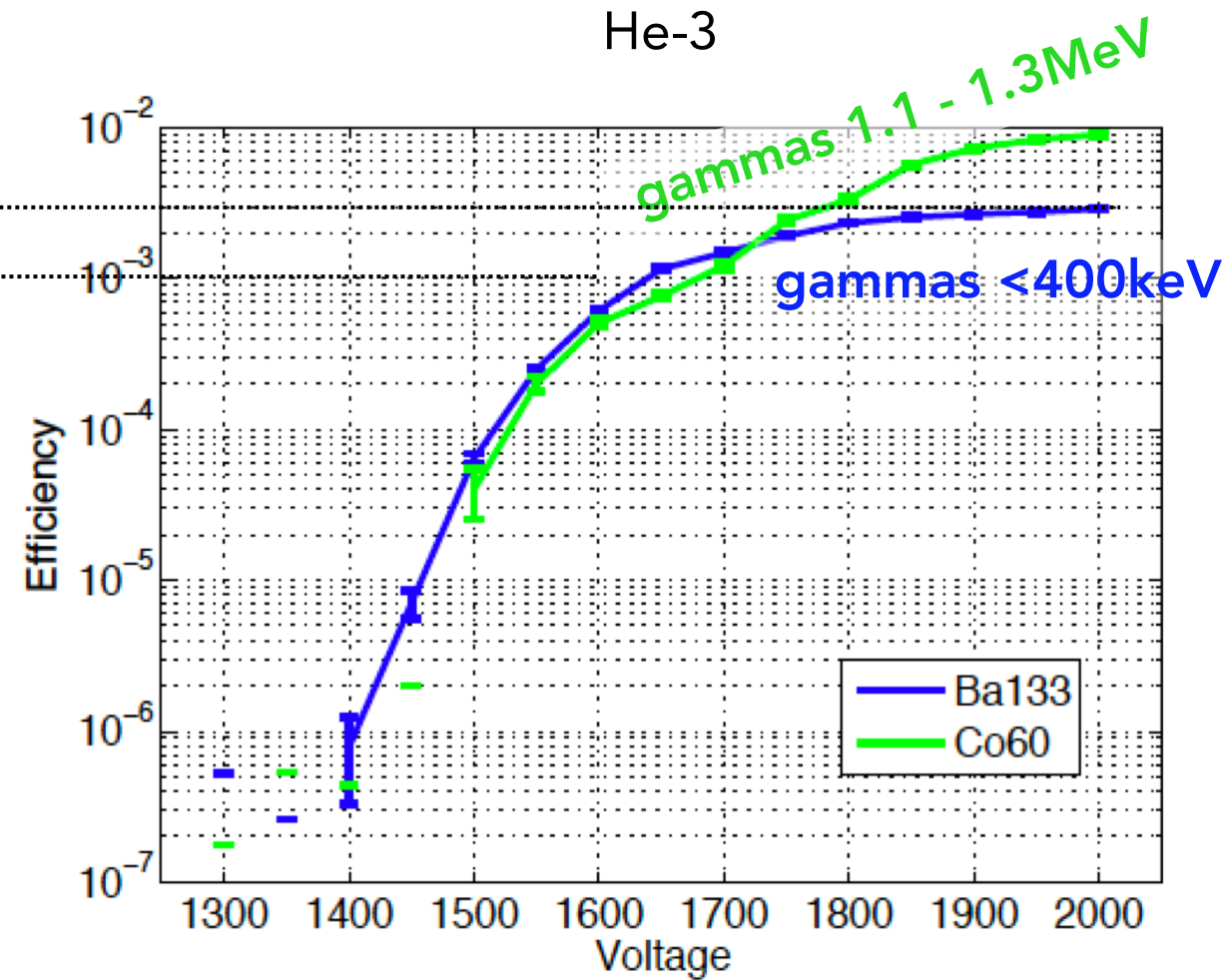
Multi-Grid B-10
(spectrometers)

Solid dominates no matter if He-3 or B-10

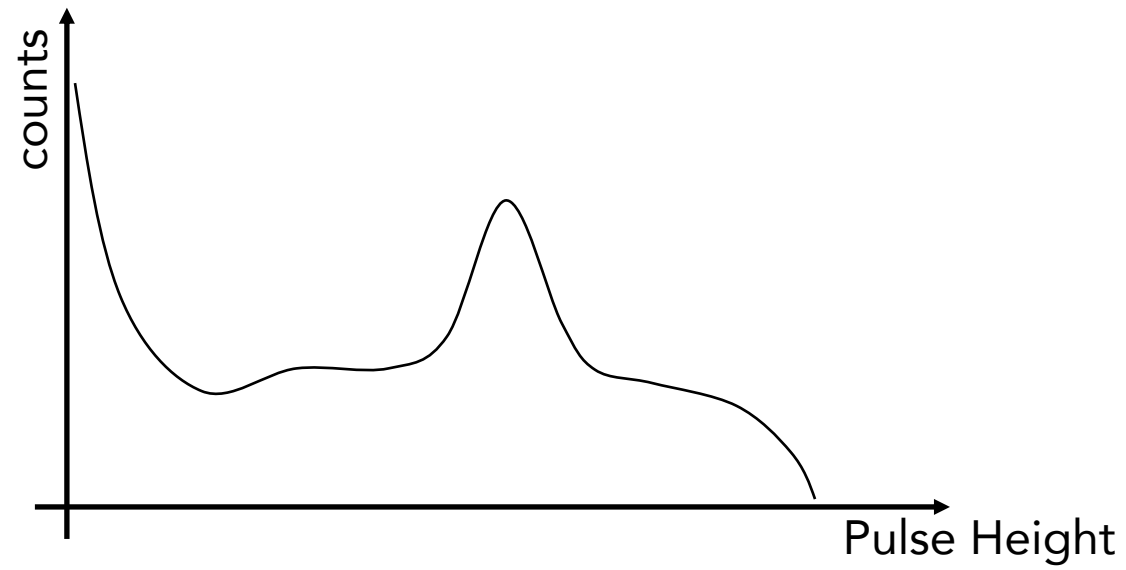
Multi-Grid B-10



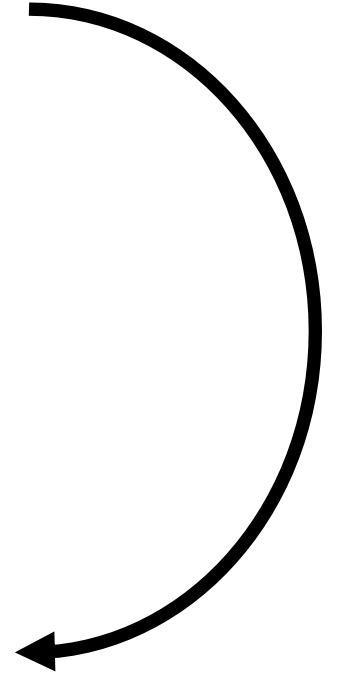
He-3



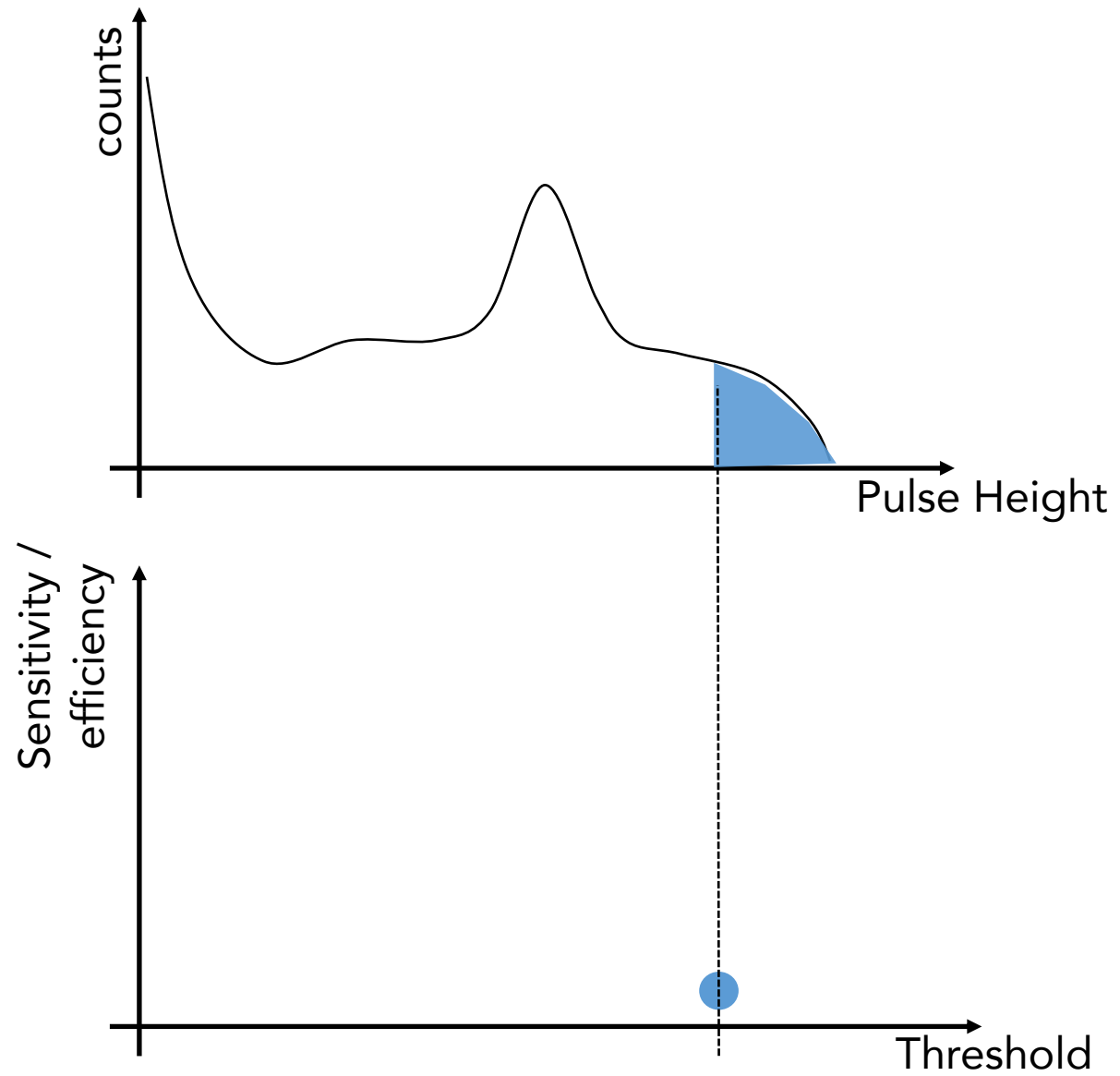
Pulse Height Spectrum - PHS



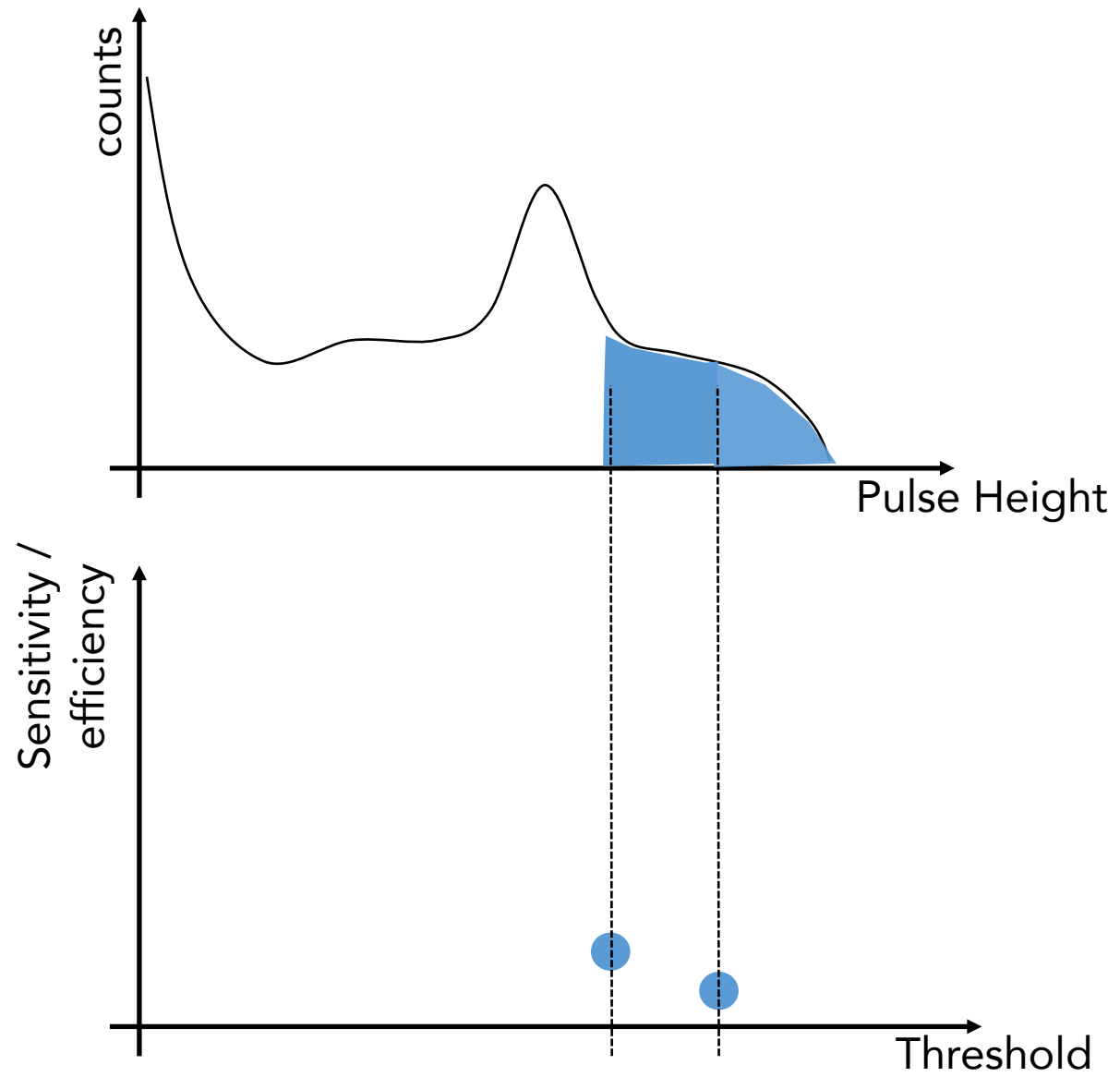
Pulse Height Spectrum - PHS

Sensitivity /
efficiency

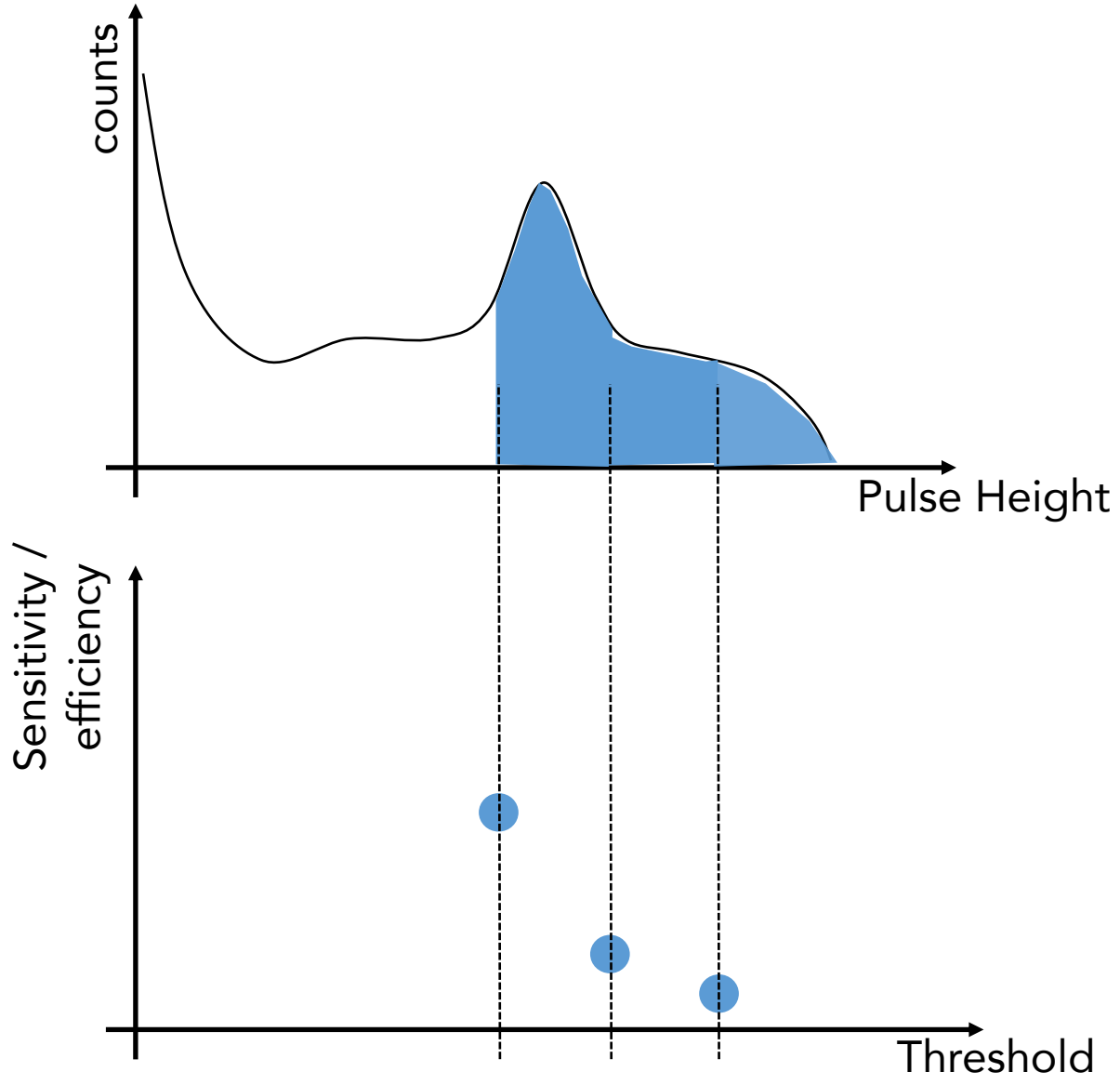
Pulse Height Spectrum - PHS



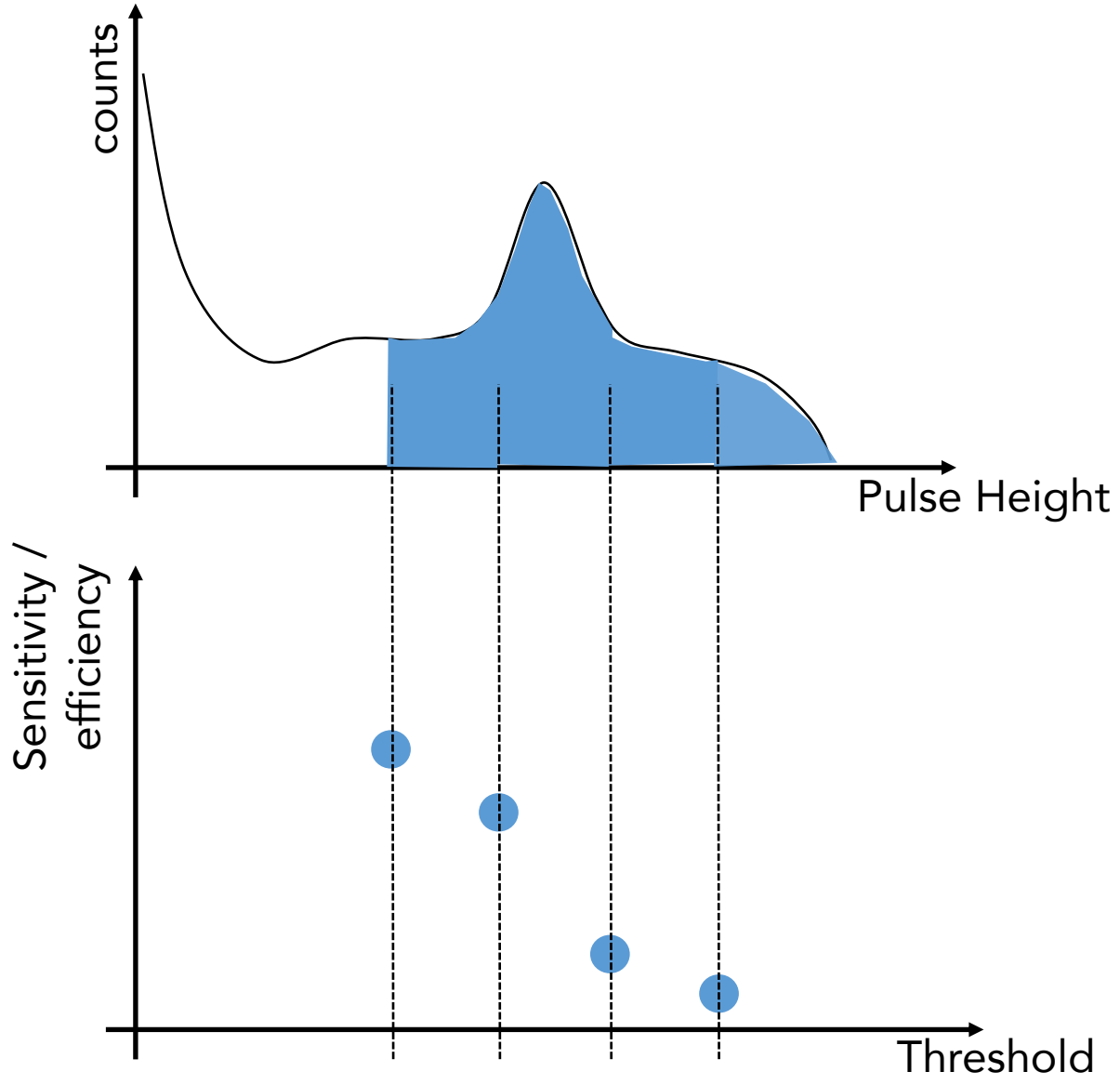
Pulse Height Spectrum - PHS



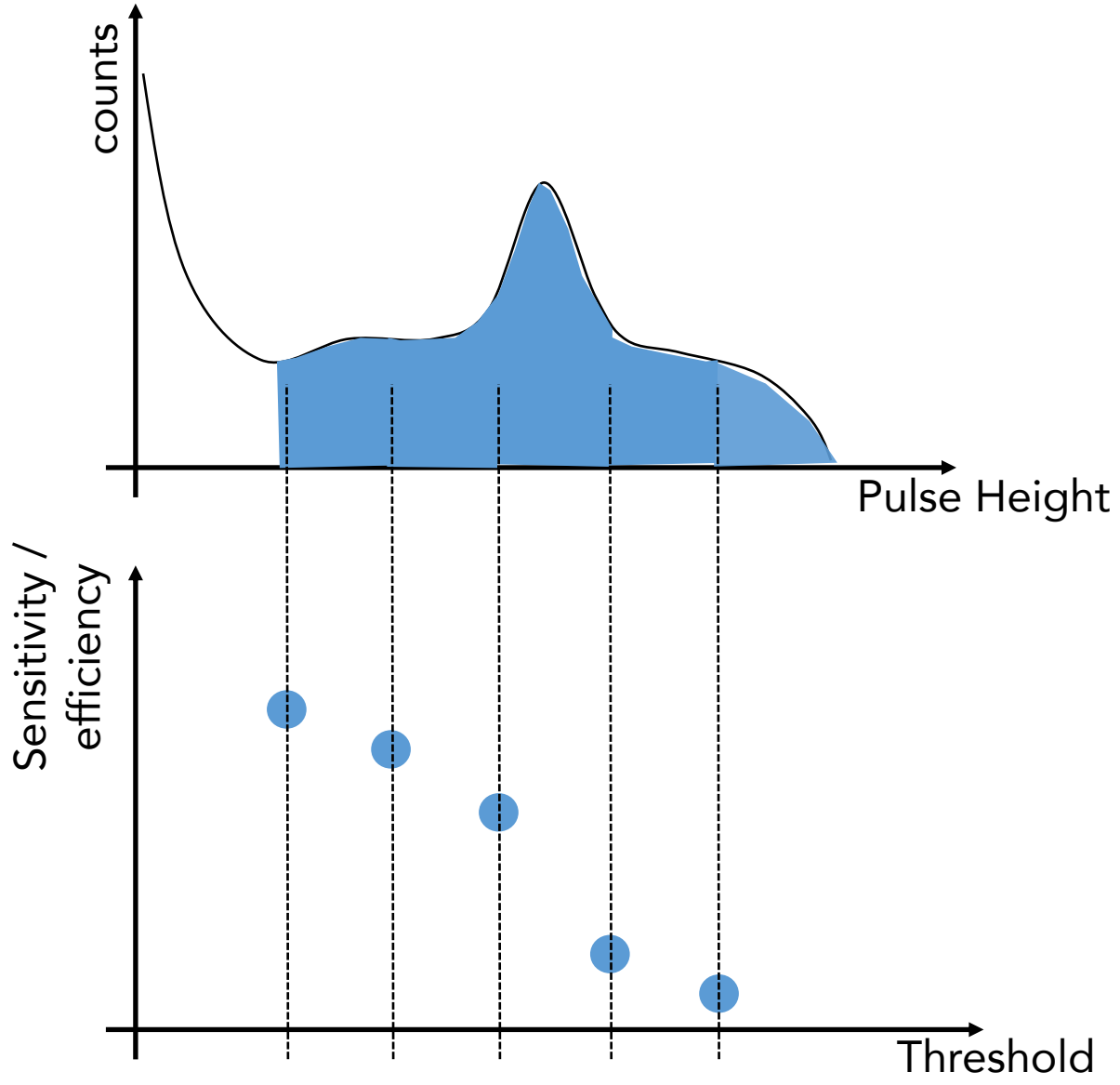
Pulse Height Spectrum - PHS



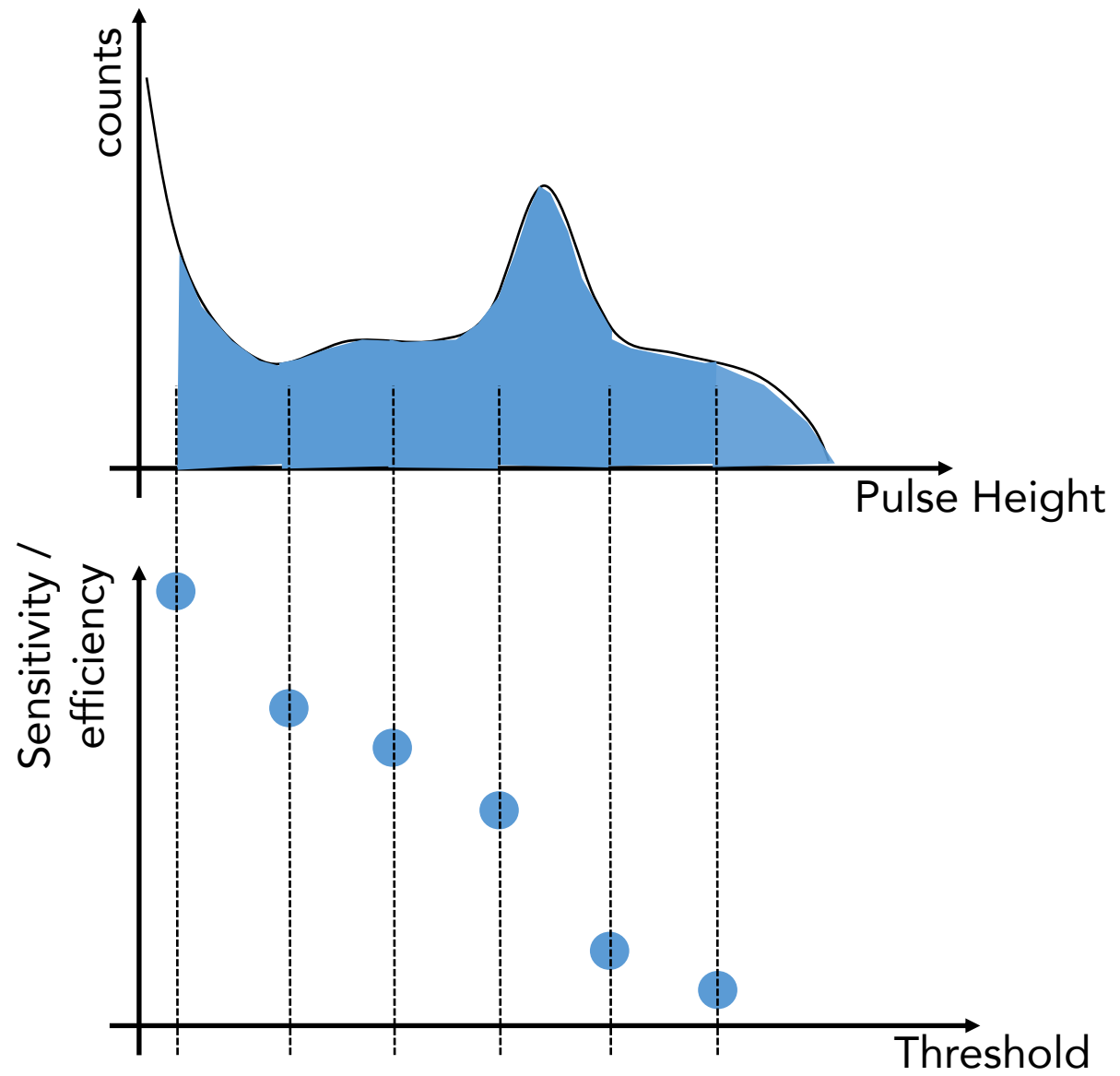
Pulse Height Spectrum - PHS



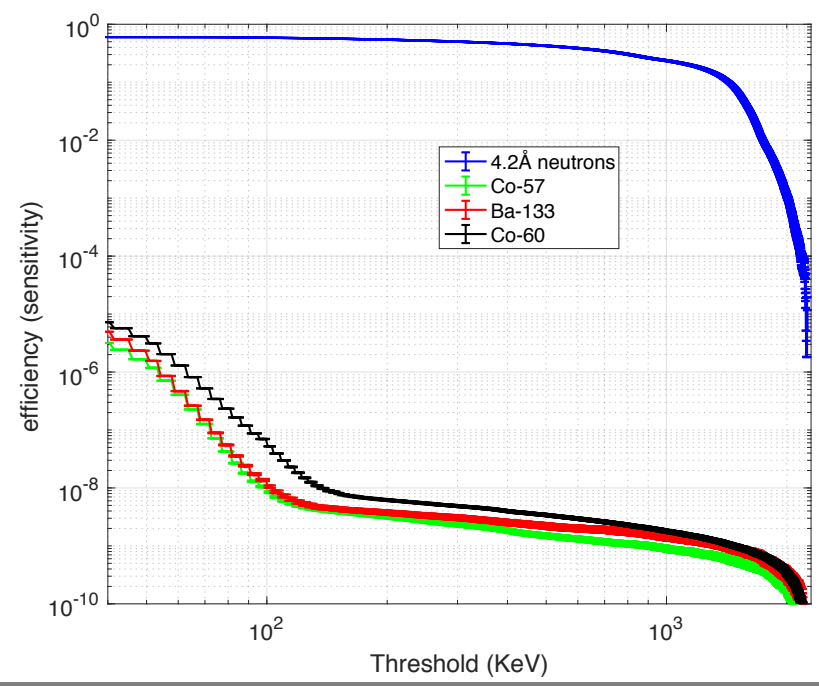
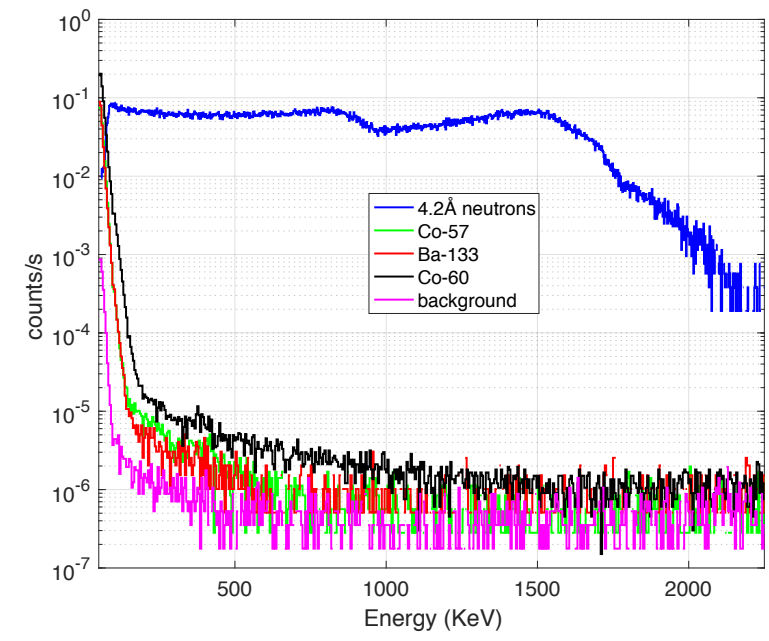
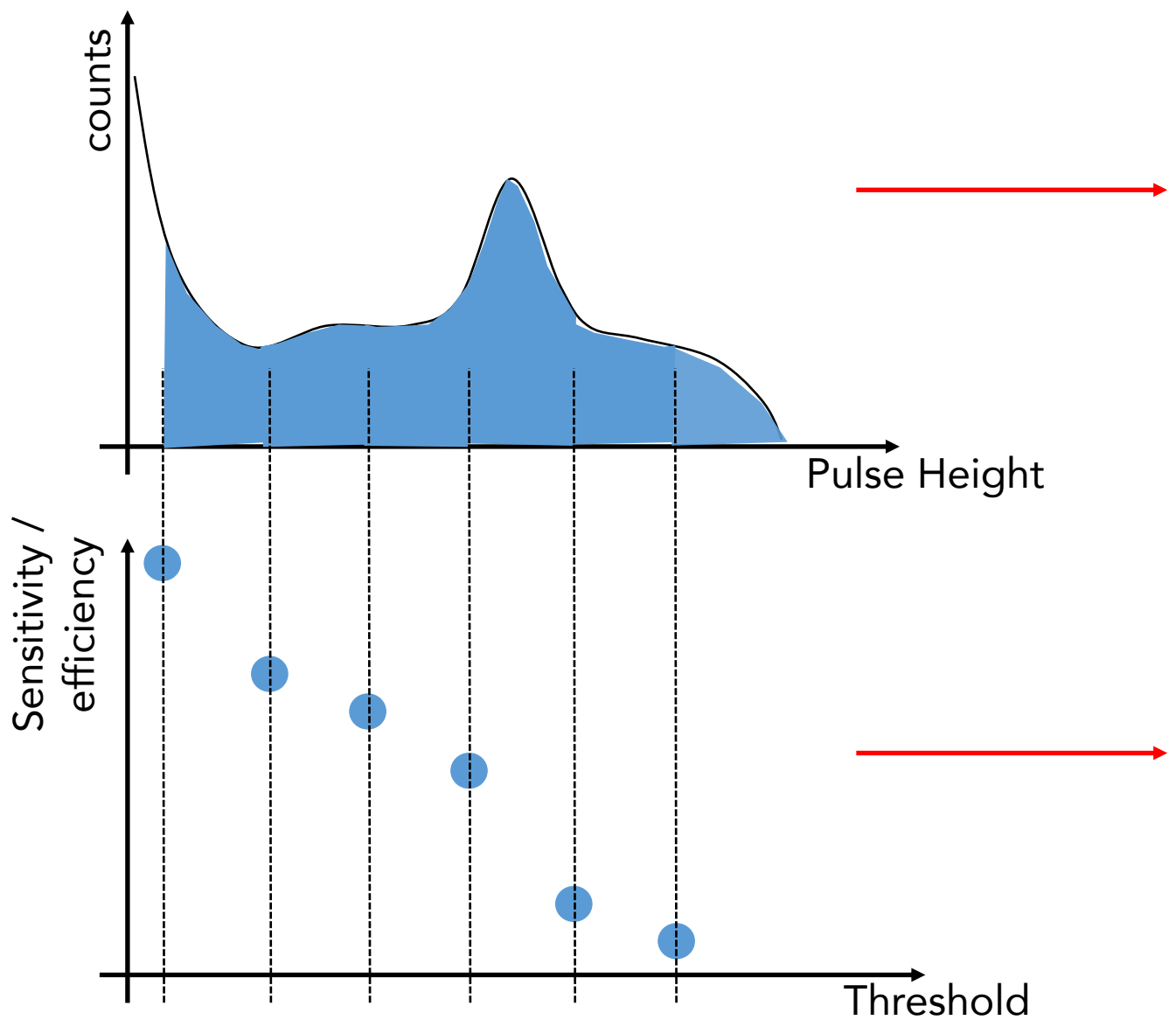
Pulse Height Spectrum - PHS



Pulse Height Spectrum - PHS

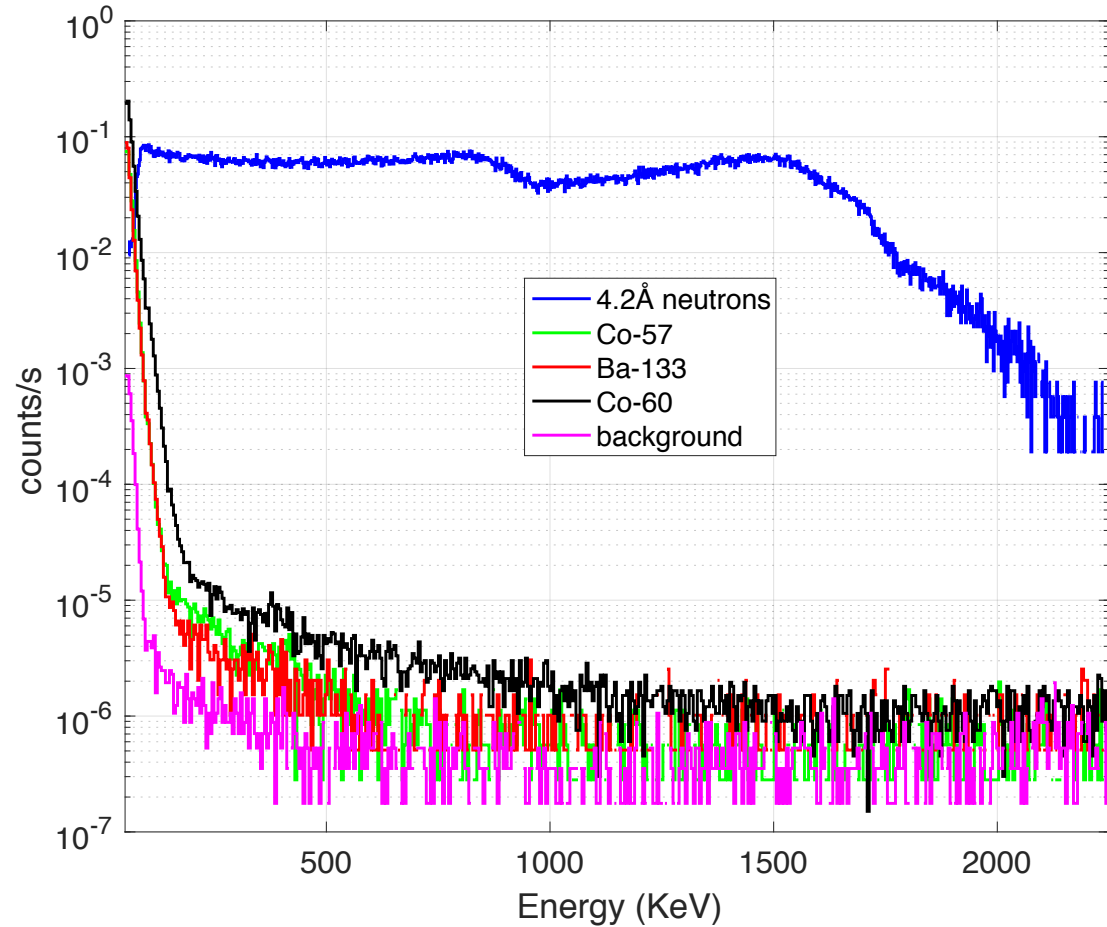


Pulse Height Spectrum - PHS

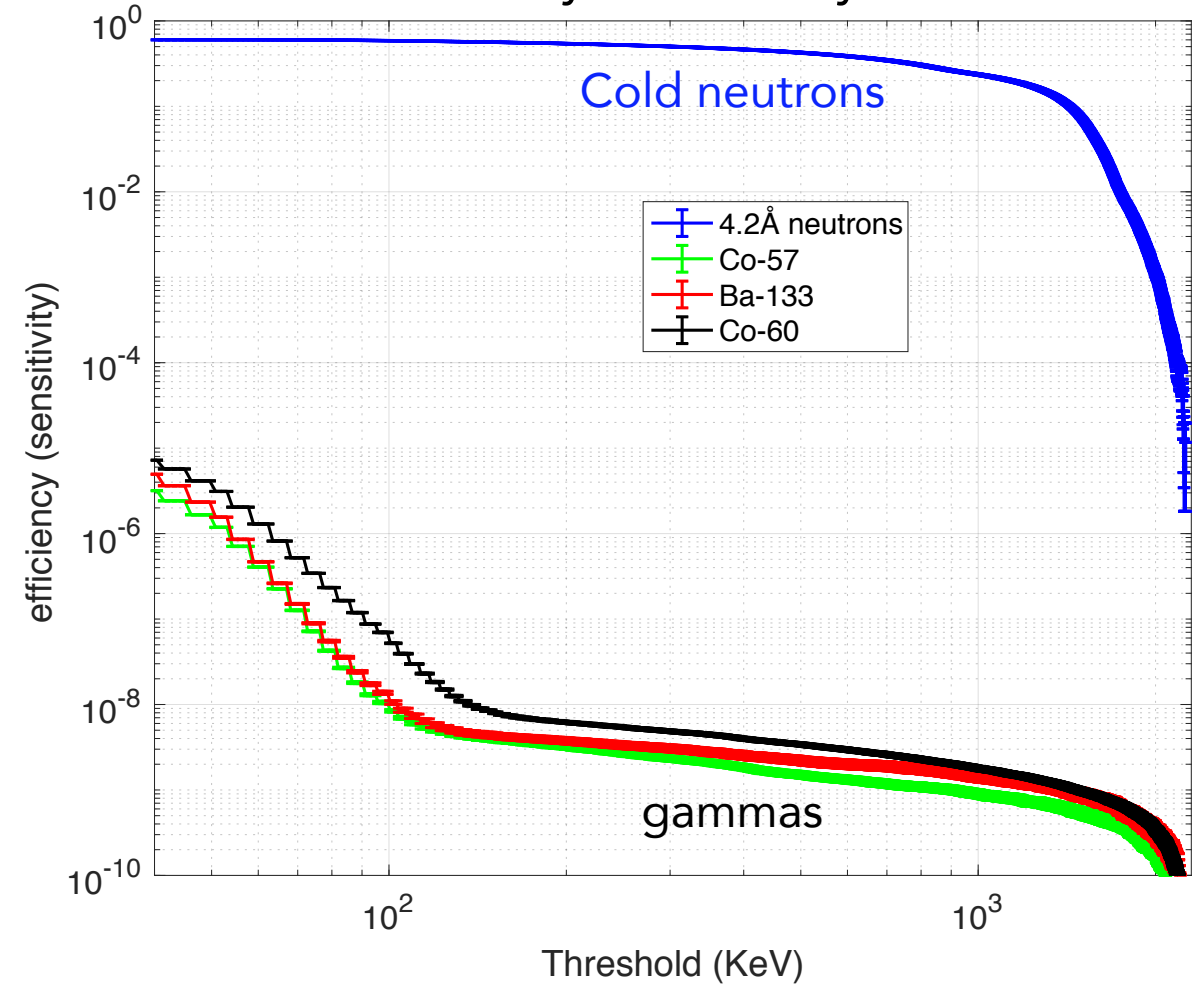


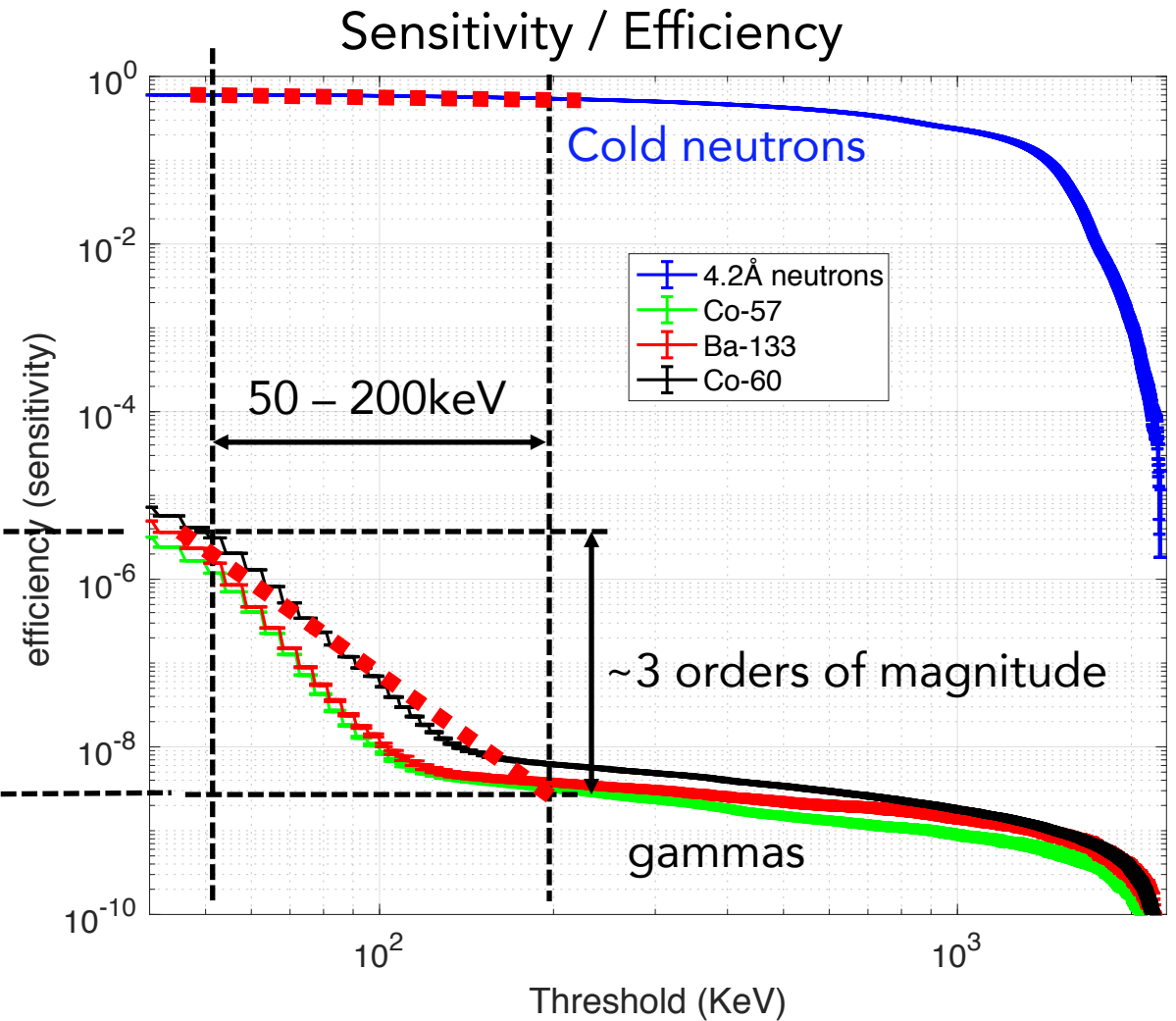
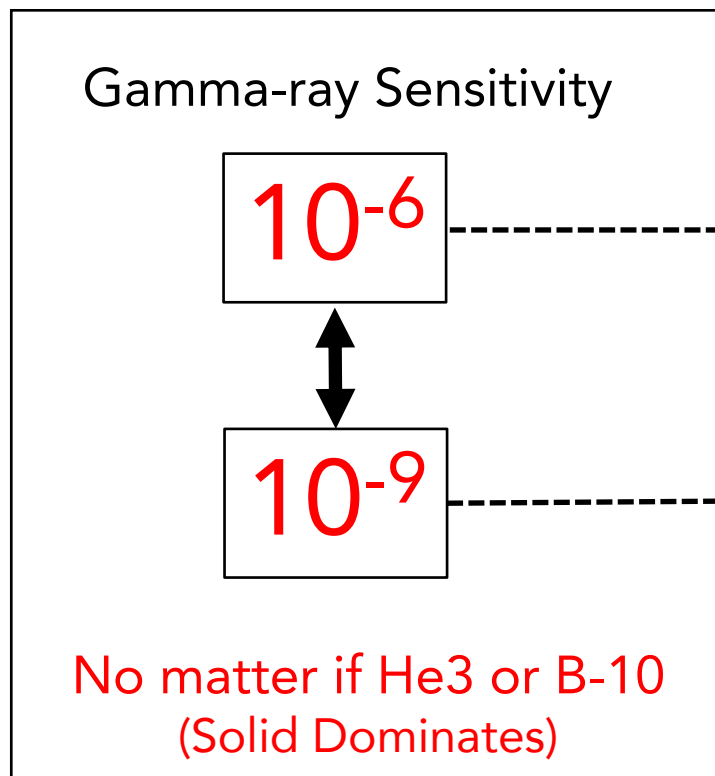
Measured with Multi-Blade B-10

PHS



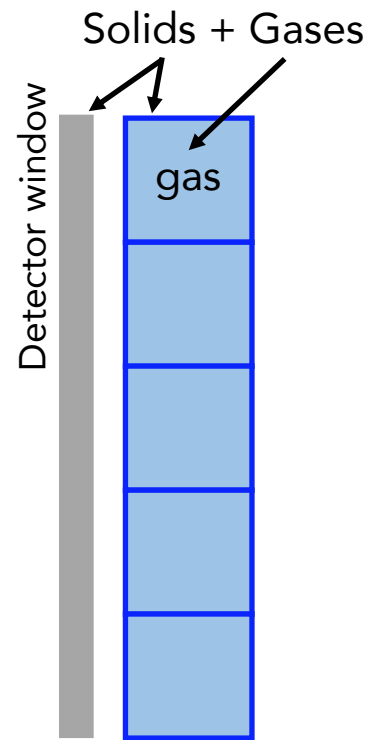
Sensitivity / Efficiency



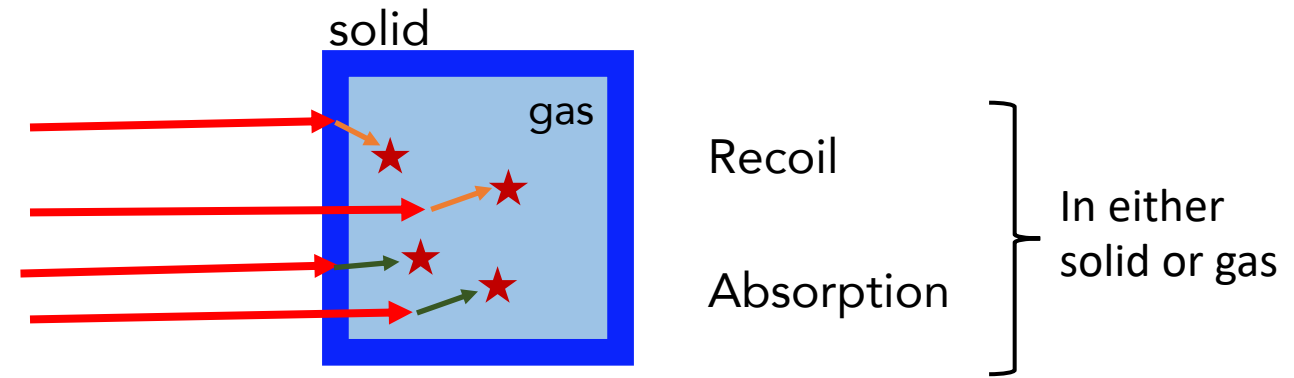
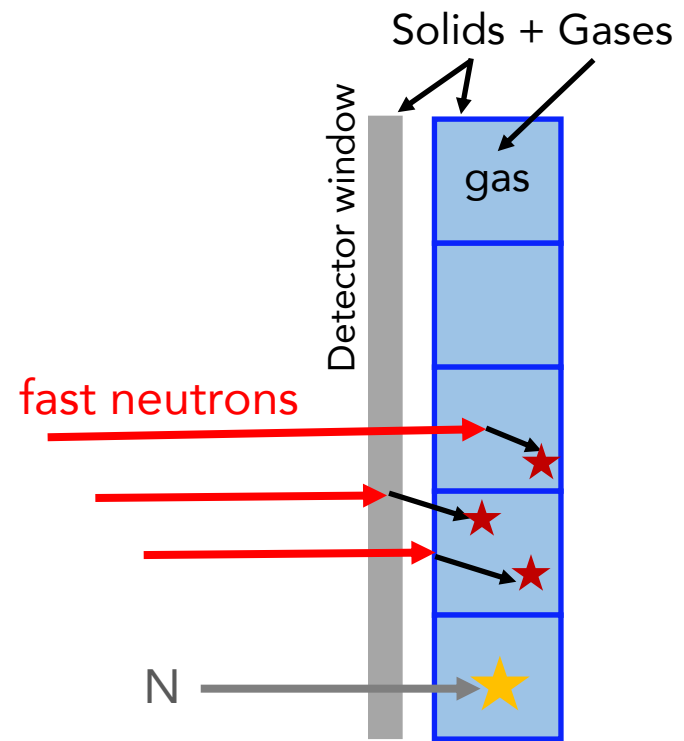


FAST NEUTRON BACKGROUND

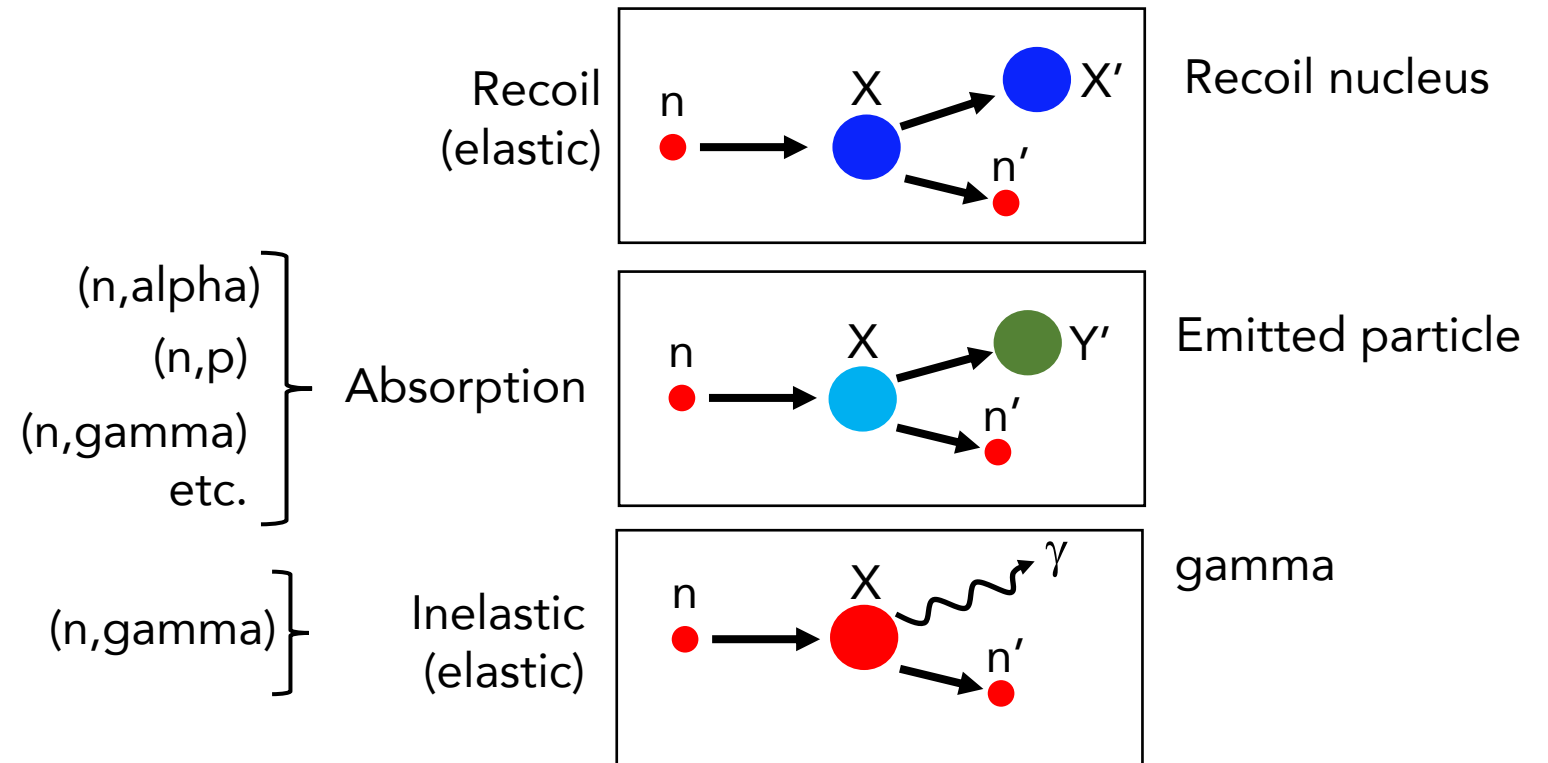
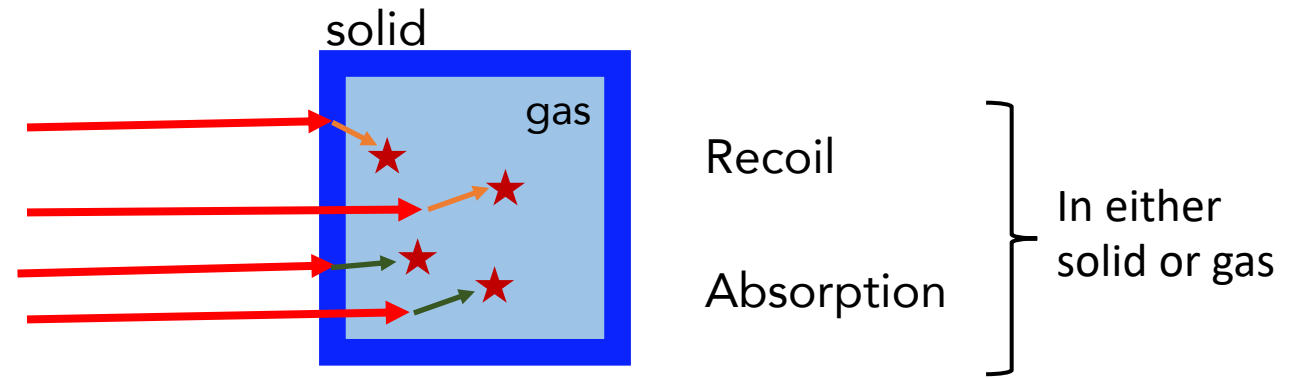
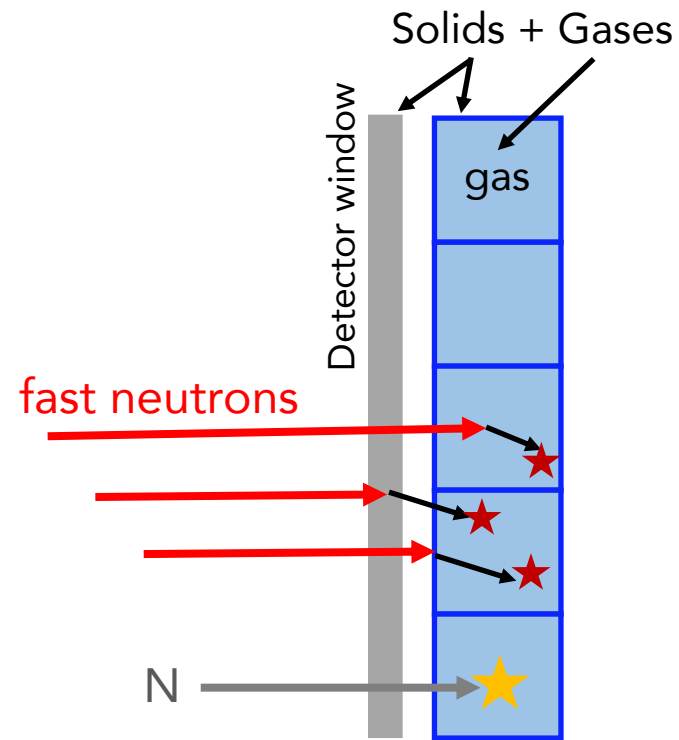
BACKGROUND: fast neutrons



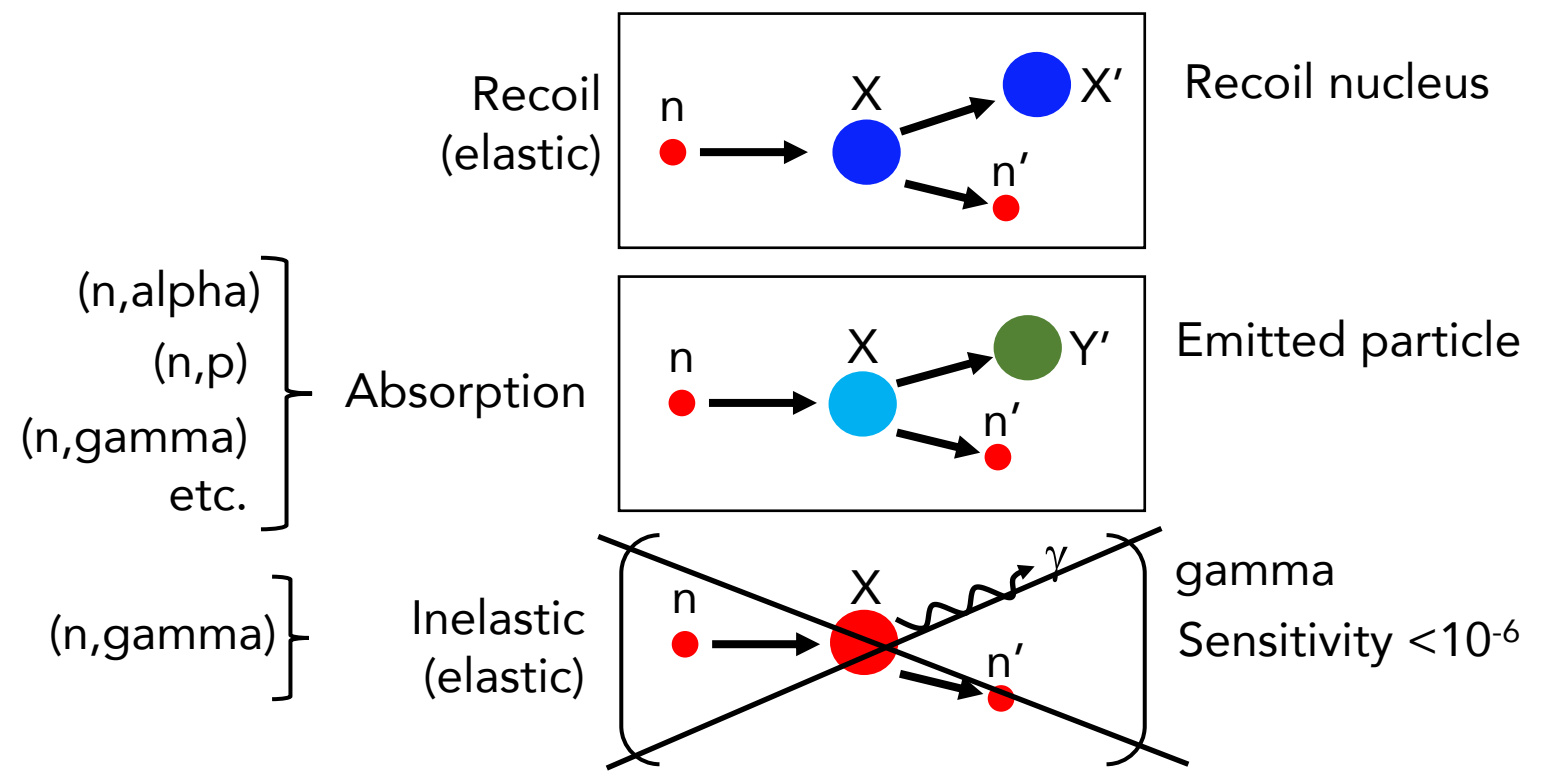
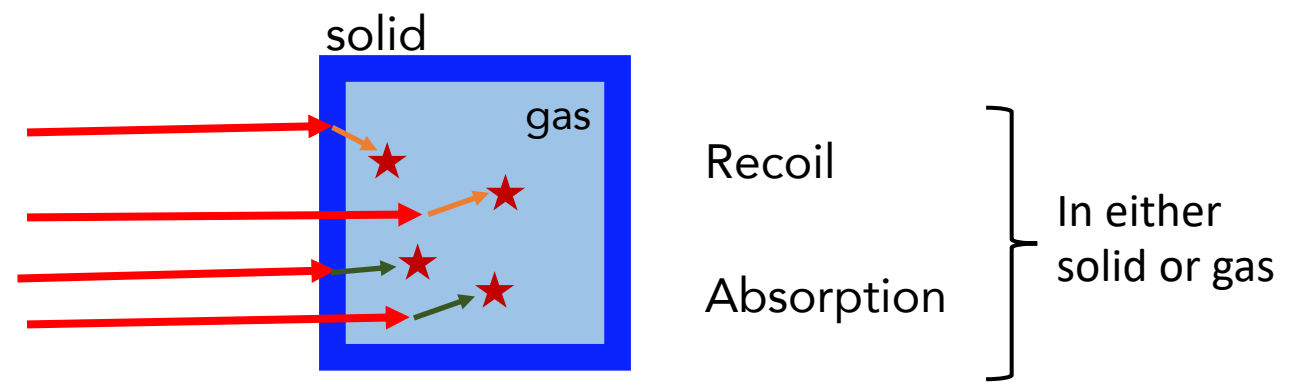
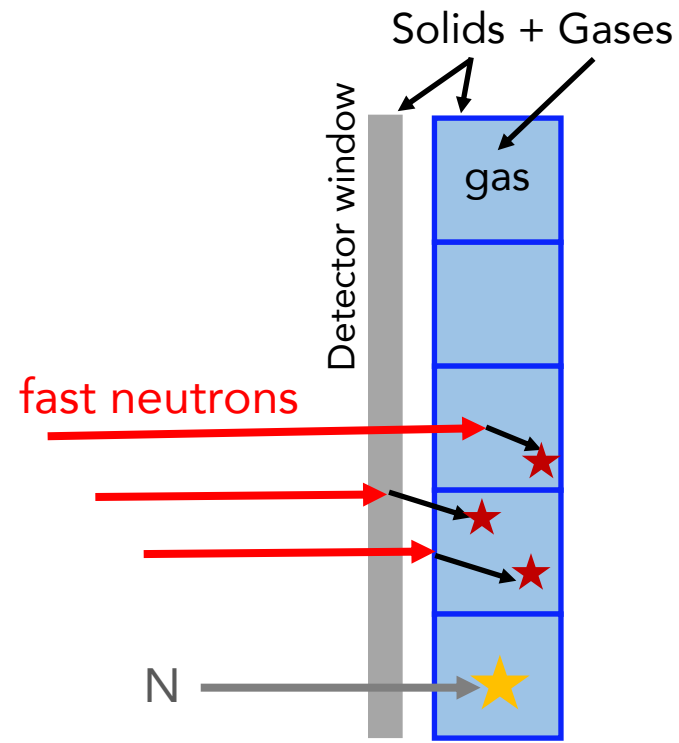
BACKGROUND: fast neutrons



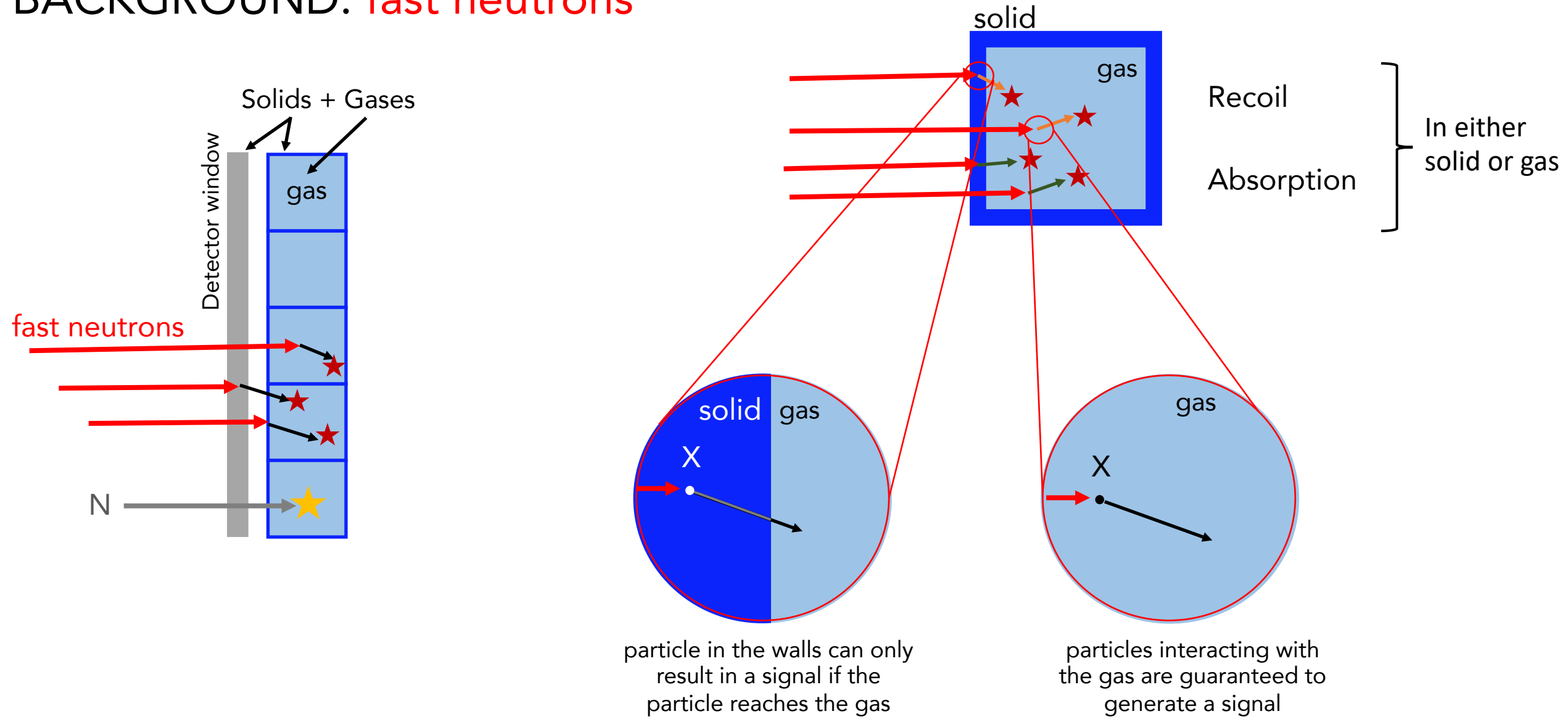
BACKGROUND: fast neutrons



BACKGROUND: fast neutrons

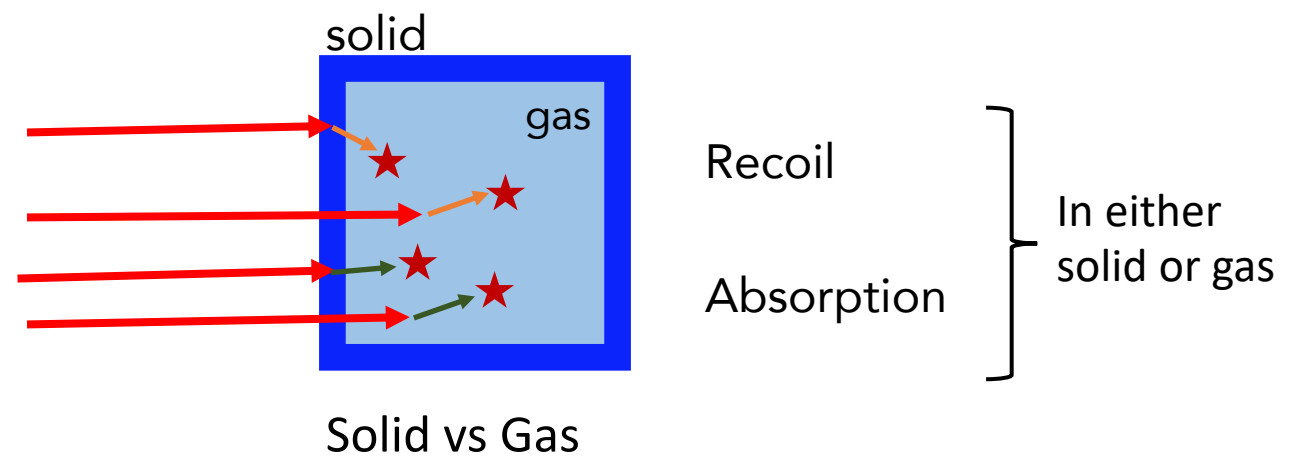
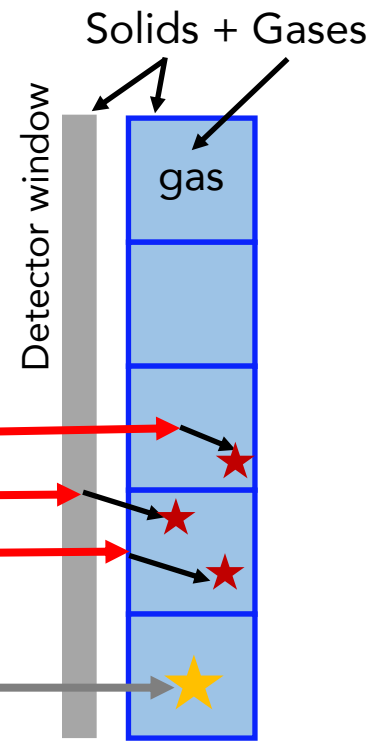


BACKGROUND: fast neutrons



Solid vs Gas

BACKGROUND: fast neutrons



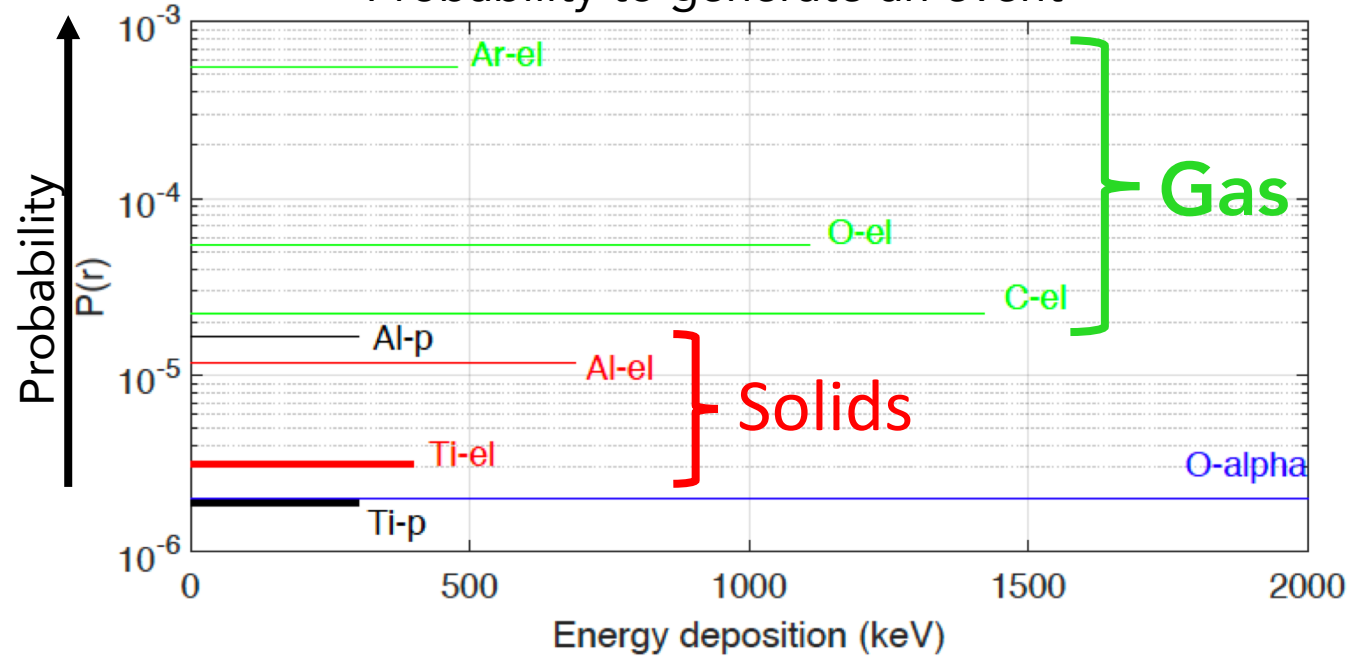
fast neutrons

N

Gas dominates
Gas = x100 Solid

True for both He-3 and B-10

Probability to generate an event



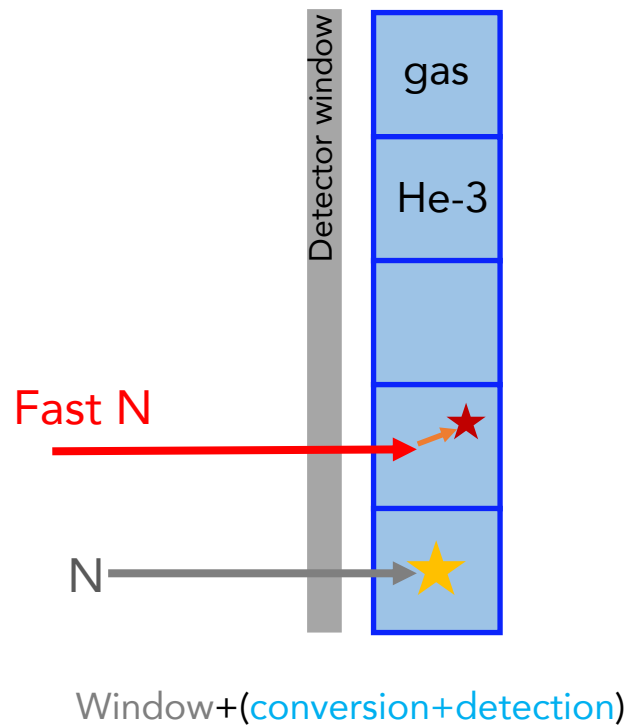
BACKGROUND: fast neutrons

Gas is main component, then He-3 differs from B-10 ...

Gas dominates

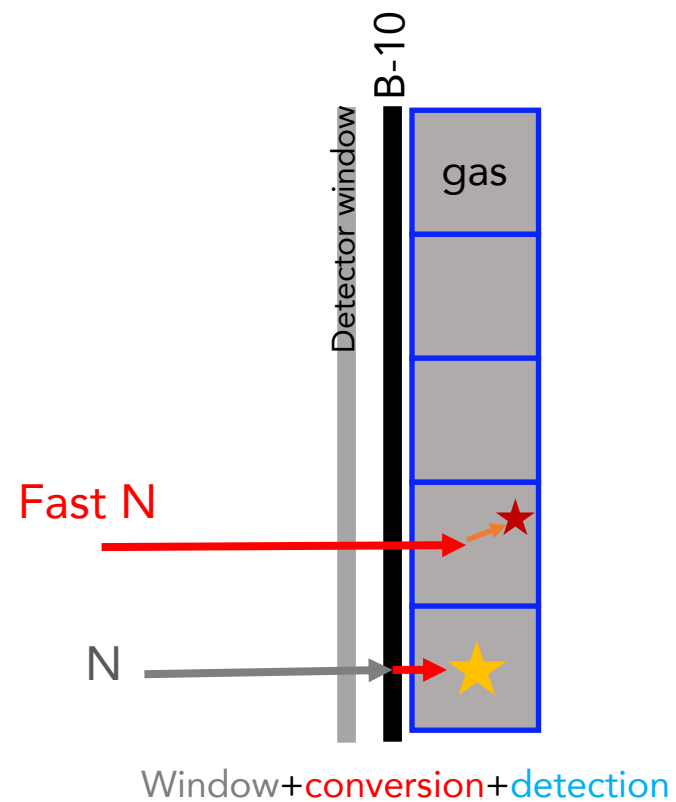
Recoil on He-3

He-3 detector



Recoil on Argon

B-10 detector



BACKGROUND: fast neutrons

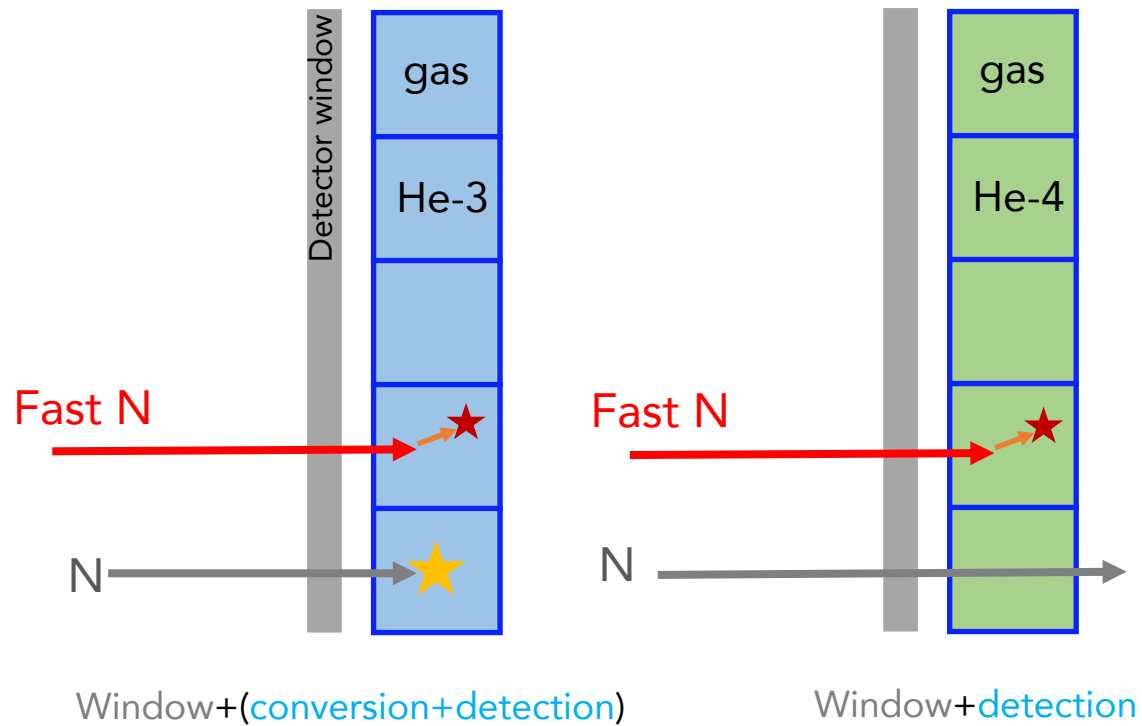
Gas is main component, then He-3 differs from B-10 ...

Gas dominates

Recoil on He-3

He-3 detector

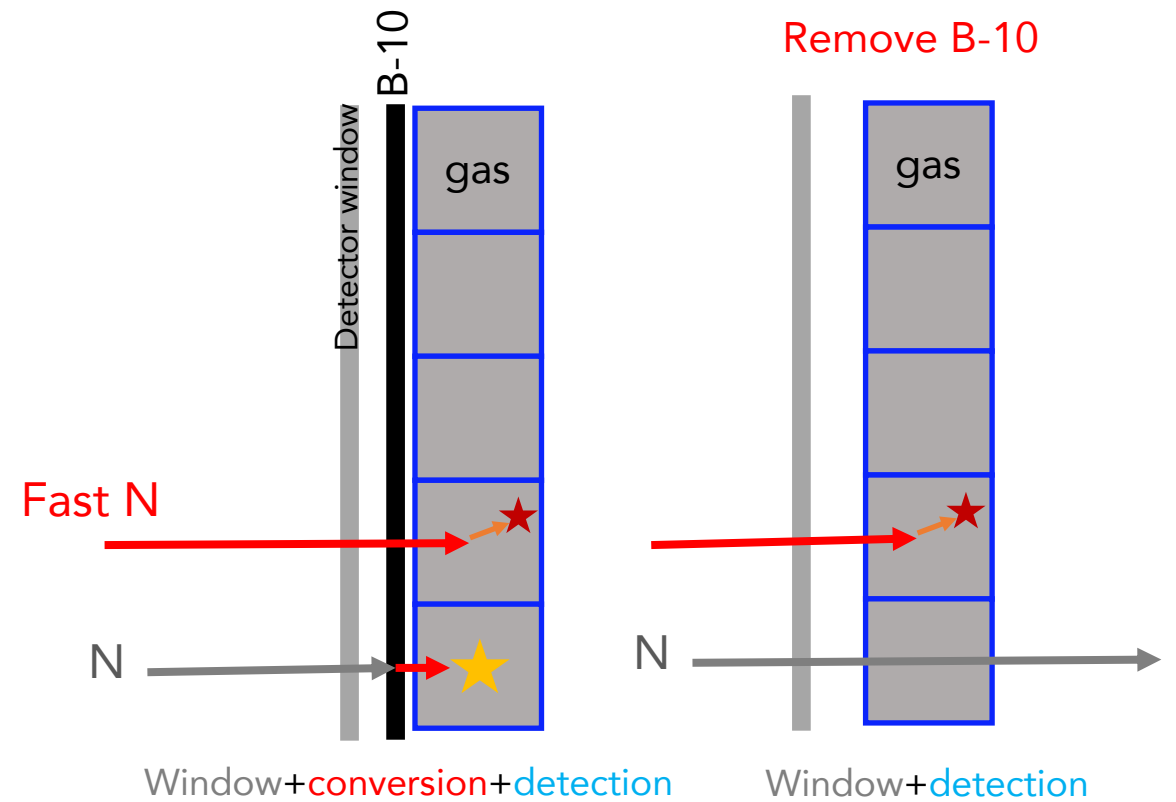
Exchange He-3 with He-4



Recoil on Argon

B-10 detector

Remove B-10



BACKGROUND: fast neutrons

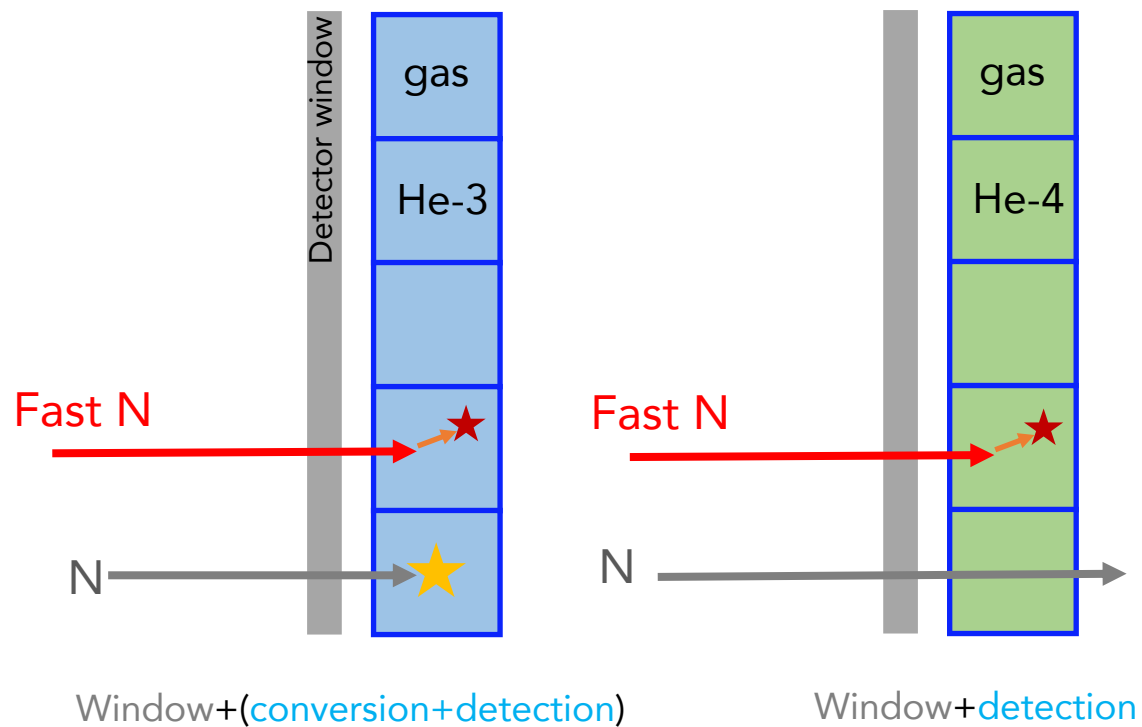
Gas is main component, then He-3 differs from B-10 ...

Gas dominates

Recoil on He-3

He-3 detector

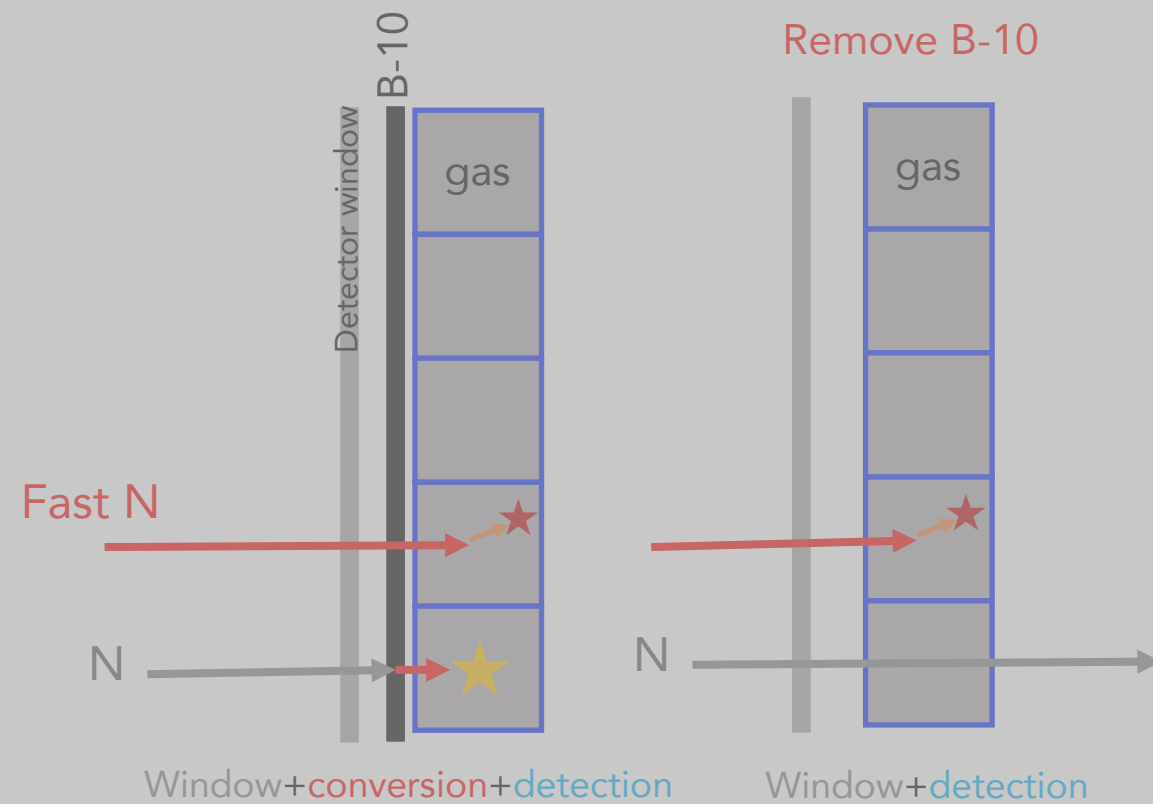
Exchange He-3 with He-4



Recoil on Argon

B-10 detector

Remove B-10



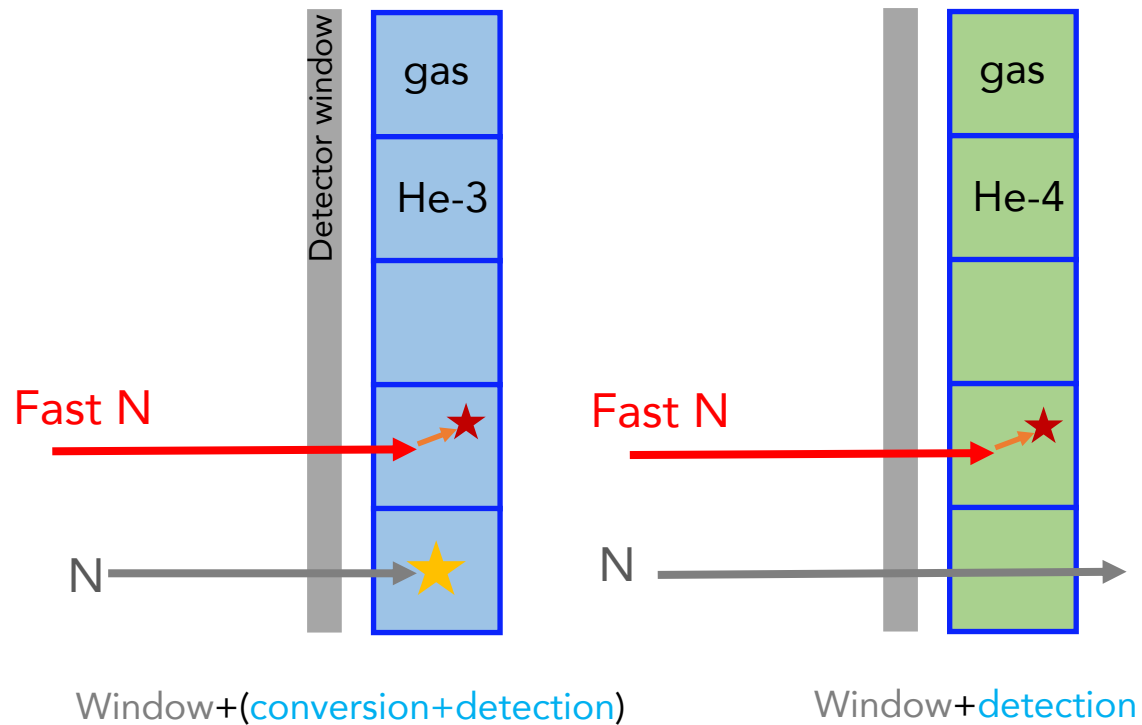
BACKGROUND: fast neutrons in He-3 detectors

*Thanks to Toshiba/Canon Electron Tubes & Devices Co. LTD for the He-4 tubes.

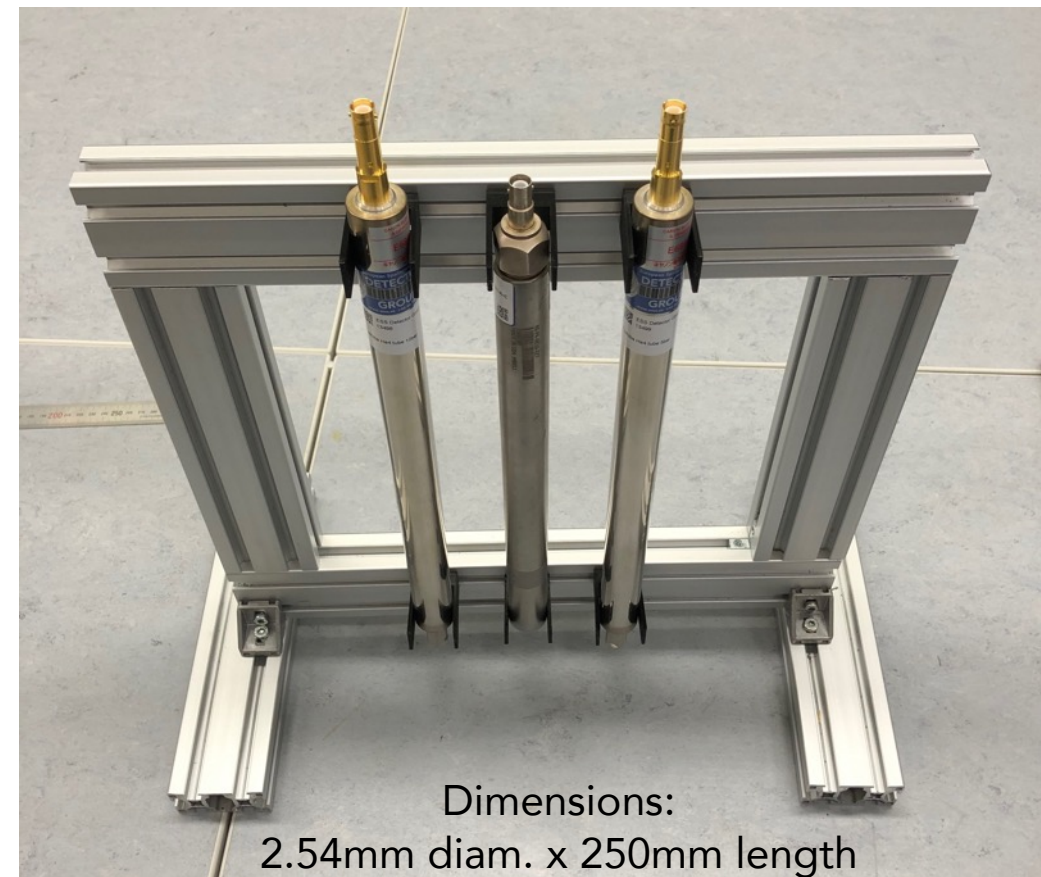
Recoil on He-3

He-3 detector

Exchange He-3 with He-4



He-4	He-3	He-4
10bar	10bar	5bar

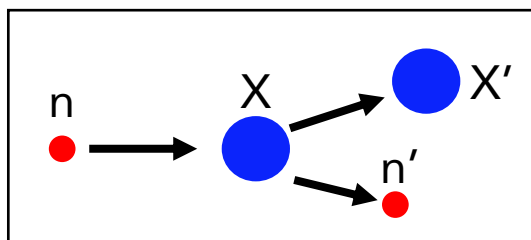


BACKGROUND: fast neutrons in He-3 detectors

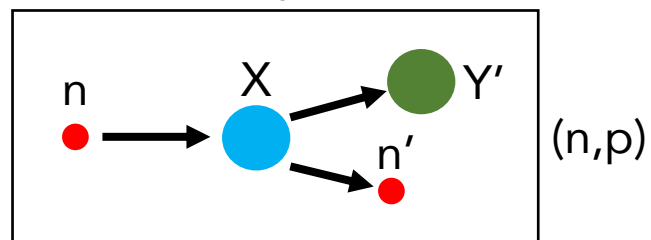
Exchanging He-3 with He-4 is a good approximation to evaluate the sensitivity of He-3 to fast n

He-3

Recoil



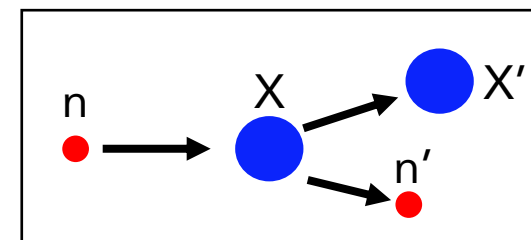
Absorption



(n,p)

He-4

Recoil



BACKGROUND: fast neutrons in He-3 detectors

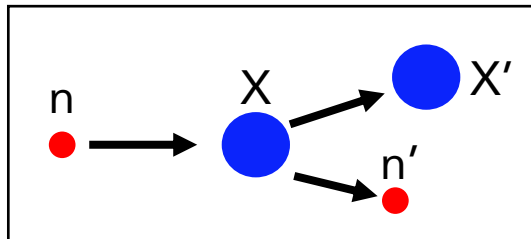
Exchanging He-3 with He-4 is a good approximation to evaluate the sensitivity of He-3 to fast n

• 1

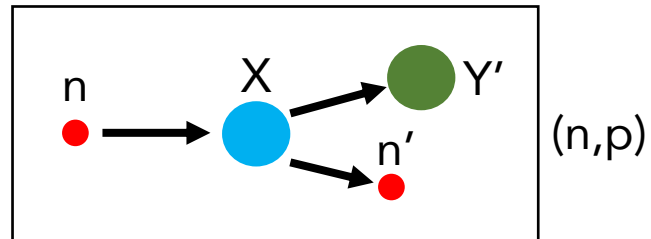
Similar total cross-section:
 $\sigma(\text{He-3}) \sim \sigma(\text{He-4})$

He-3

Recoil

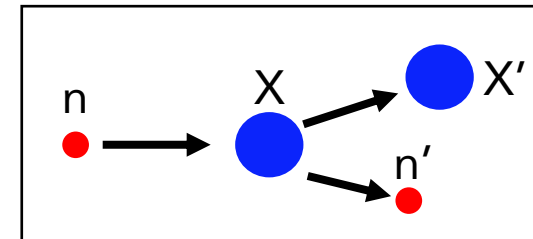


Absorption



He-4

Recoil

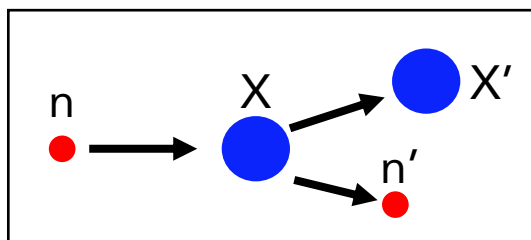


BACKGROUND: fast neutrons in He-3 detectors

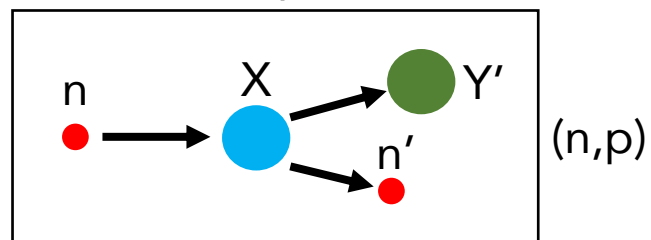
Exchanging He-3 with He-4 is a good approximation to evaluate the sensitivity of He-3 to fast n

He-3

Recoil



Absorption



• 1

Similar total cross-section:

$$\sigma(\text{He-3}) \sim \sigma(\text{He-4})$$

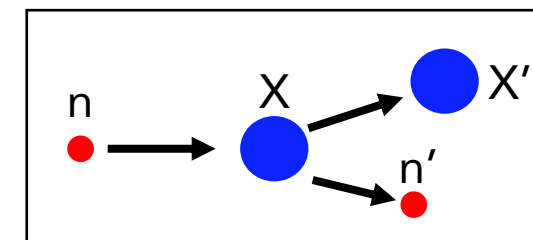
• 2

Almost identical energy deposition ion gas:

$$\text{He-3}^{++} \text{ in He-3} \sim \text{He-4}^{++} \text{ in He-4}$$

He-4

Recoil

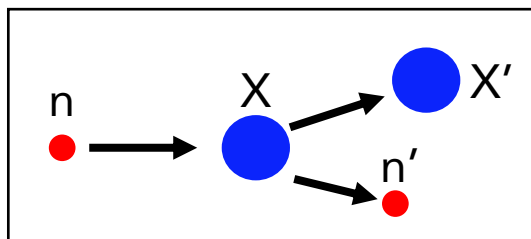


BACKGROUND: fast neutrons in He-3 detectors

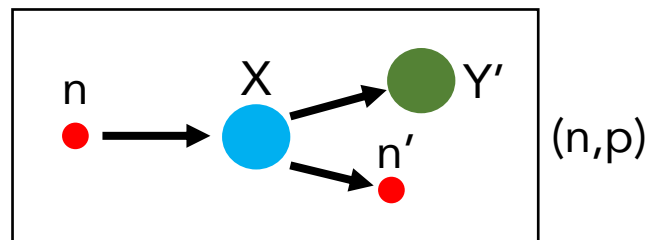
Exchanging He-3 with He-4 is a good approximation to evaluate the sensitivity of He-3 to fast n

He-3

Recoil



Absorption



• 1

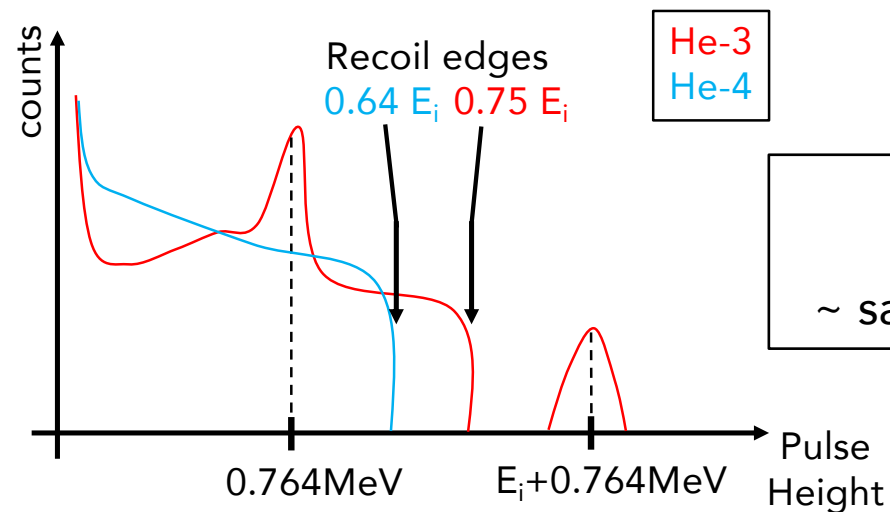
Similar total cross-section:
 $\sigma(\text{He-3}) \sim \sigma(\text{He-4})$

• 2

Almost identical energy deposition ion gas:
 $\text{He-3}^{++} \text{ in He-3} \sim \text{He-4}^{++} \text{ in He-4}$

• 3

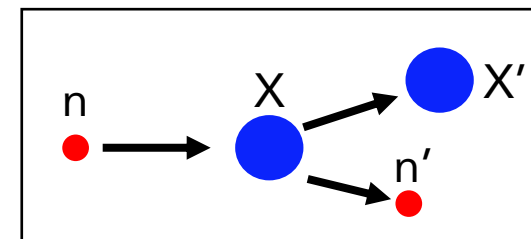
If mono-energetic beam E_i



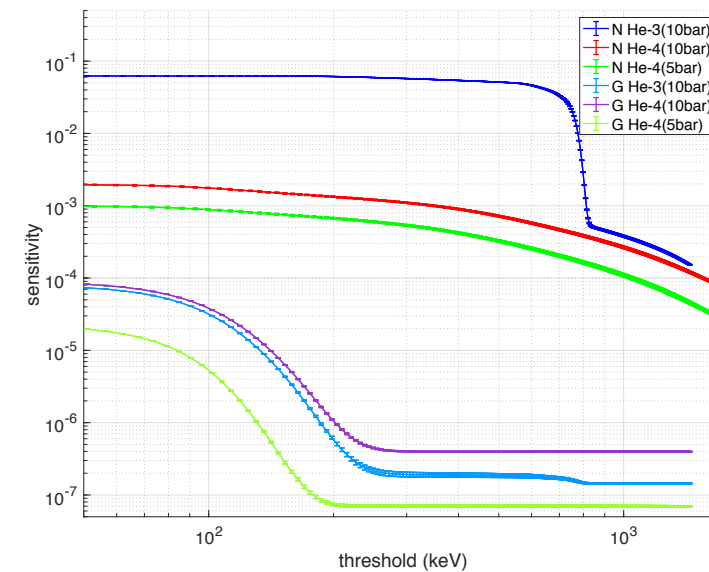
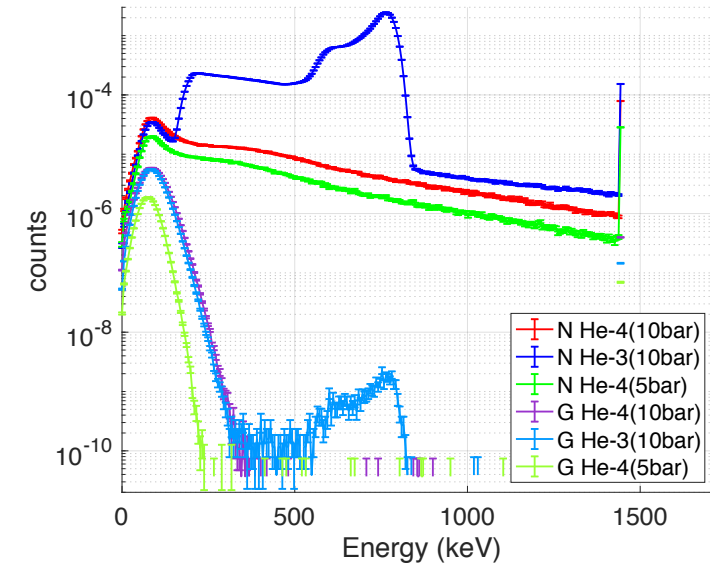
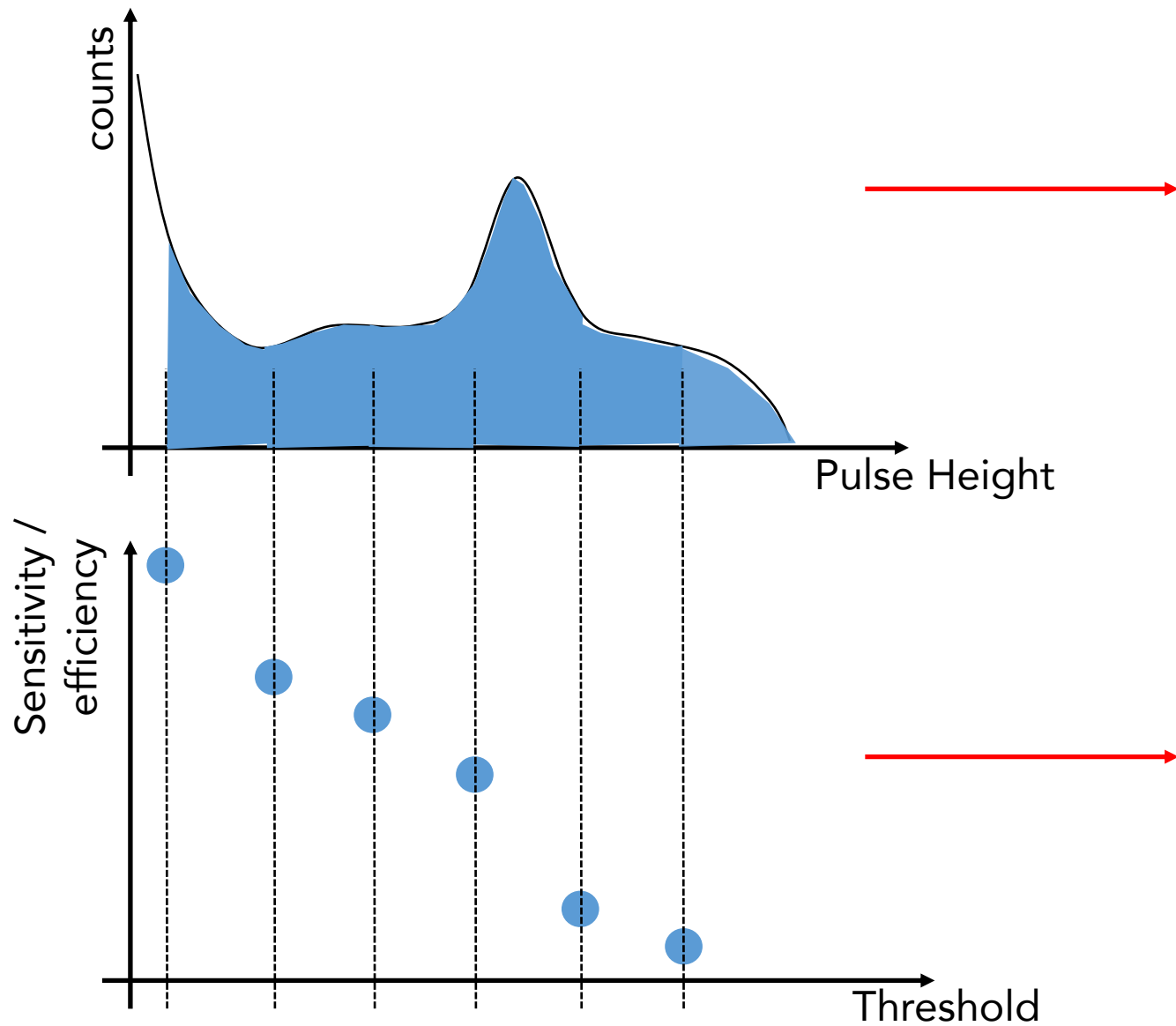
Different PHS
but
~ same number of events !

He-4

Recoil

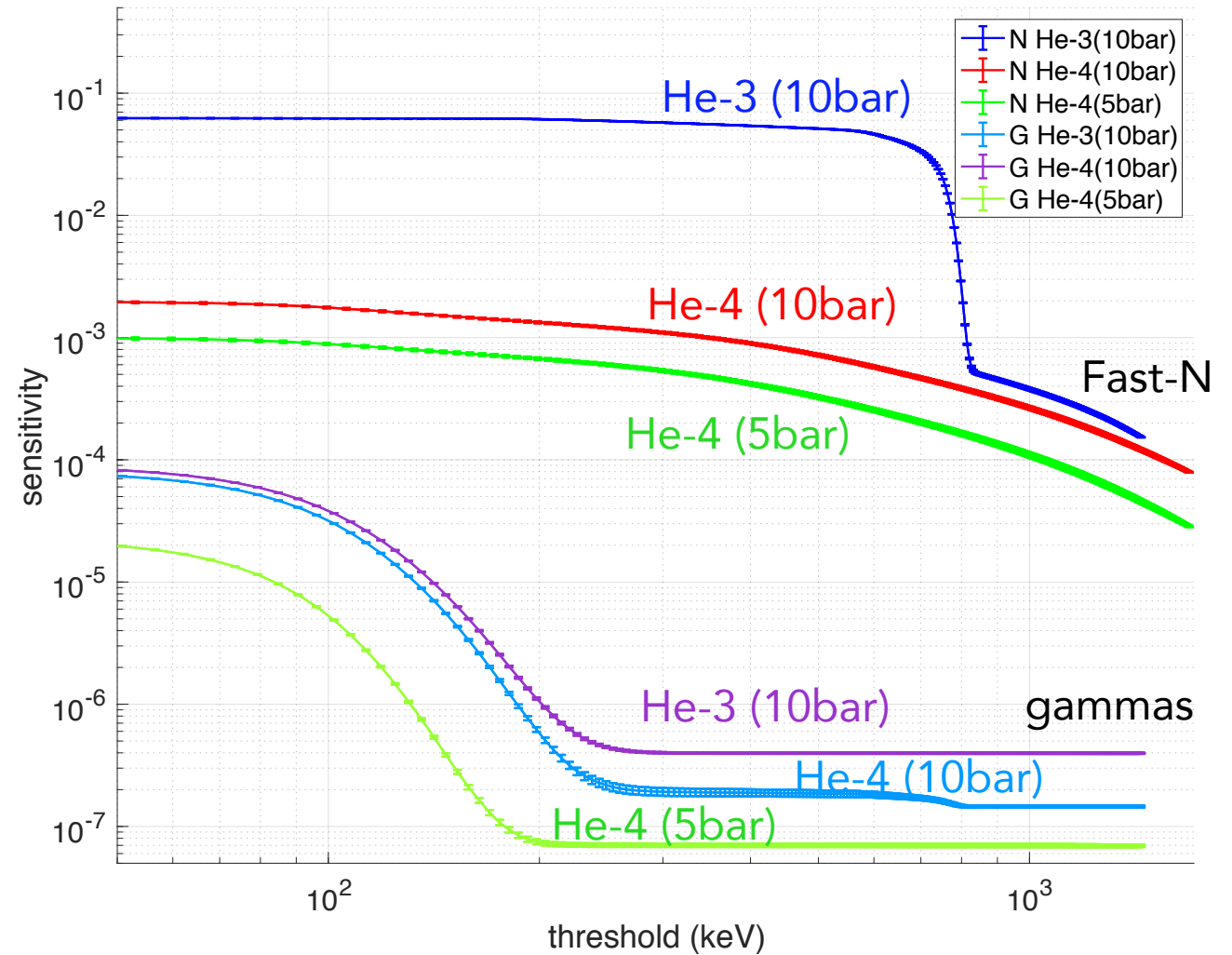
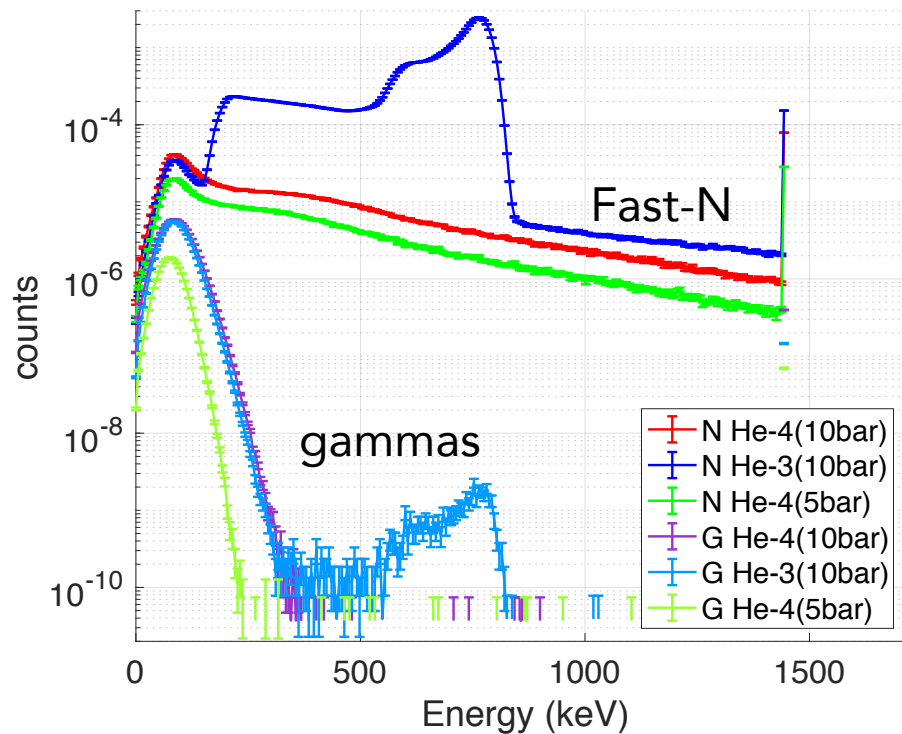


Pulse Height Spectrum - PHS



Sensitivity / Efficiency

Pulse Height Spectrum - PHS



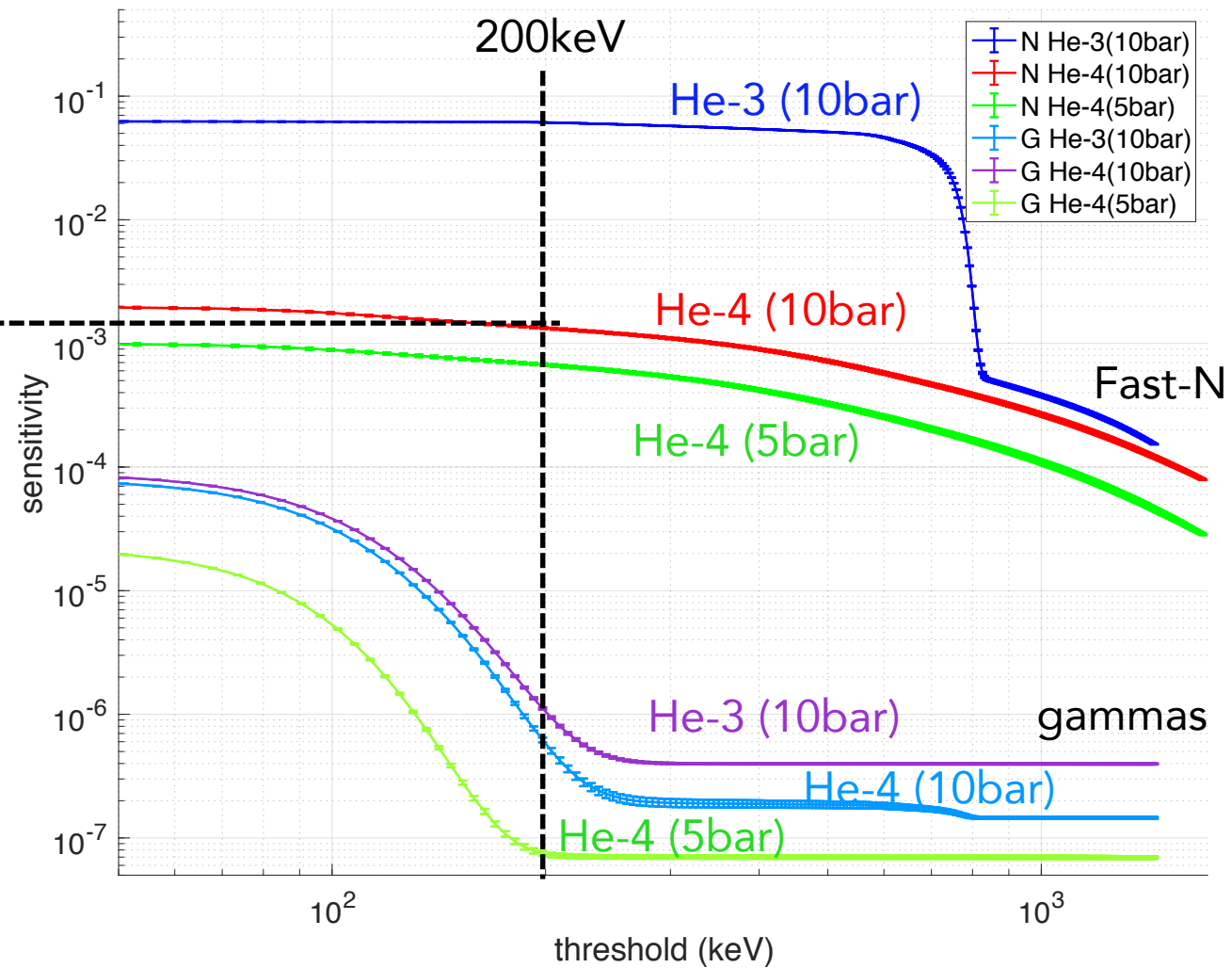
Fast Neutron Sensitivity

He-3

10^{-3}

Difference between He3 or B-10
(Gas Dominates)

Sensitivity / Efficiency



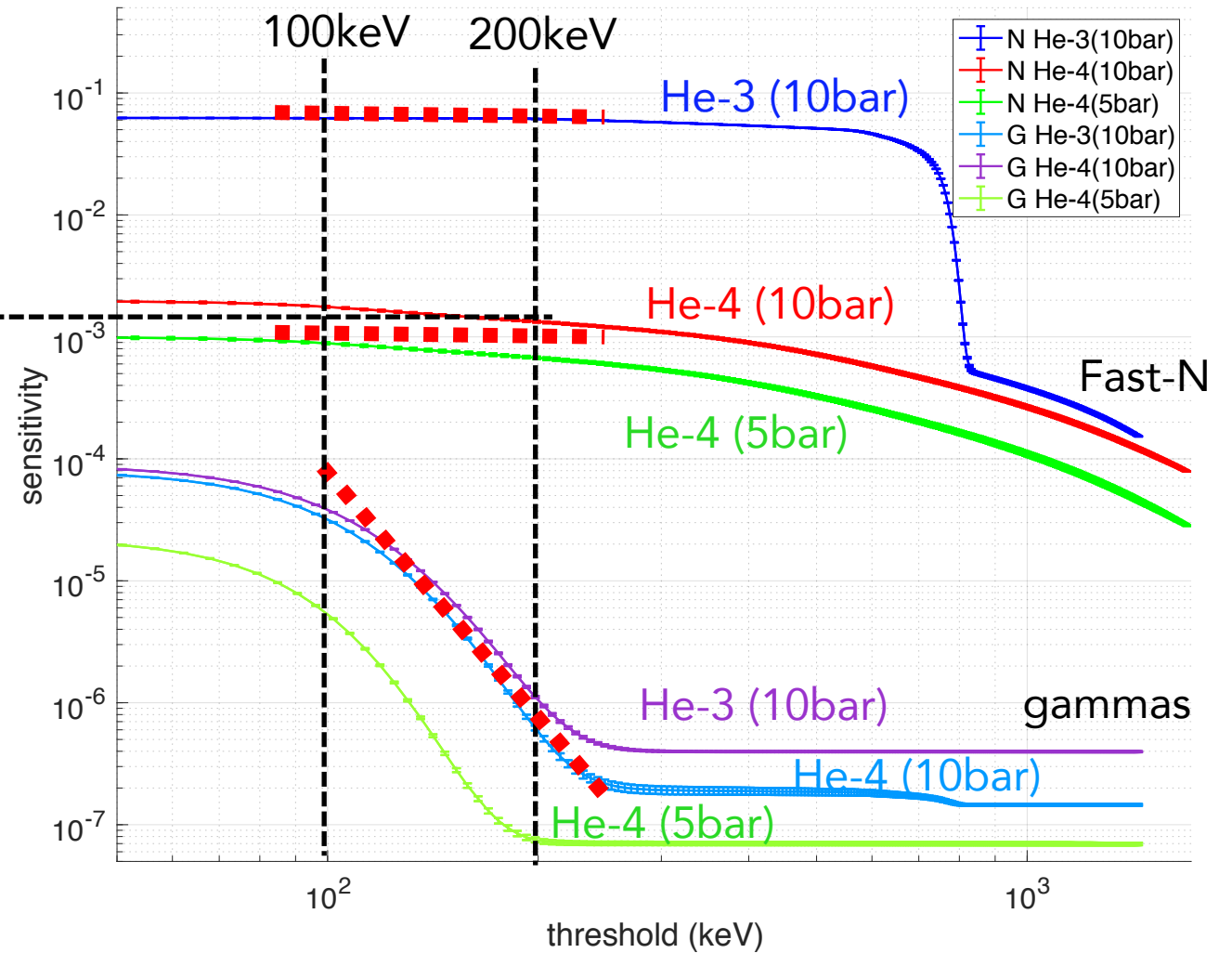
Fast Neutron Sensitivity

He-3

10^{-3}

Difference between He3 or B-10
(Gas Dominates)

Sensitivity / Efficiency



Fast Neutron Sensitivity

He-3

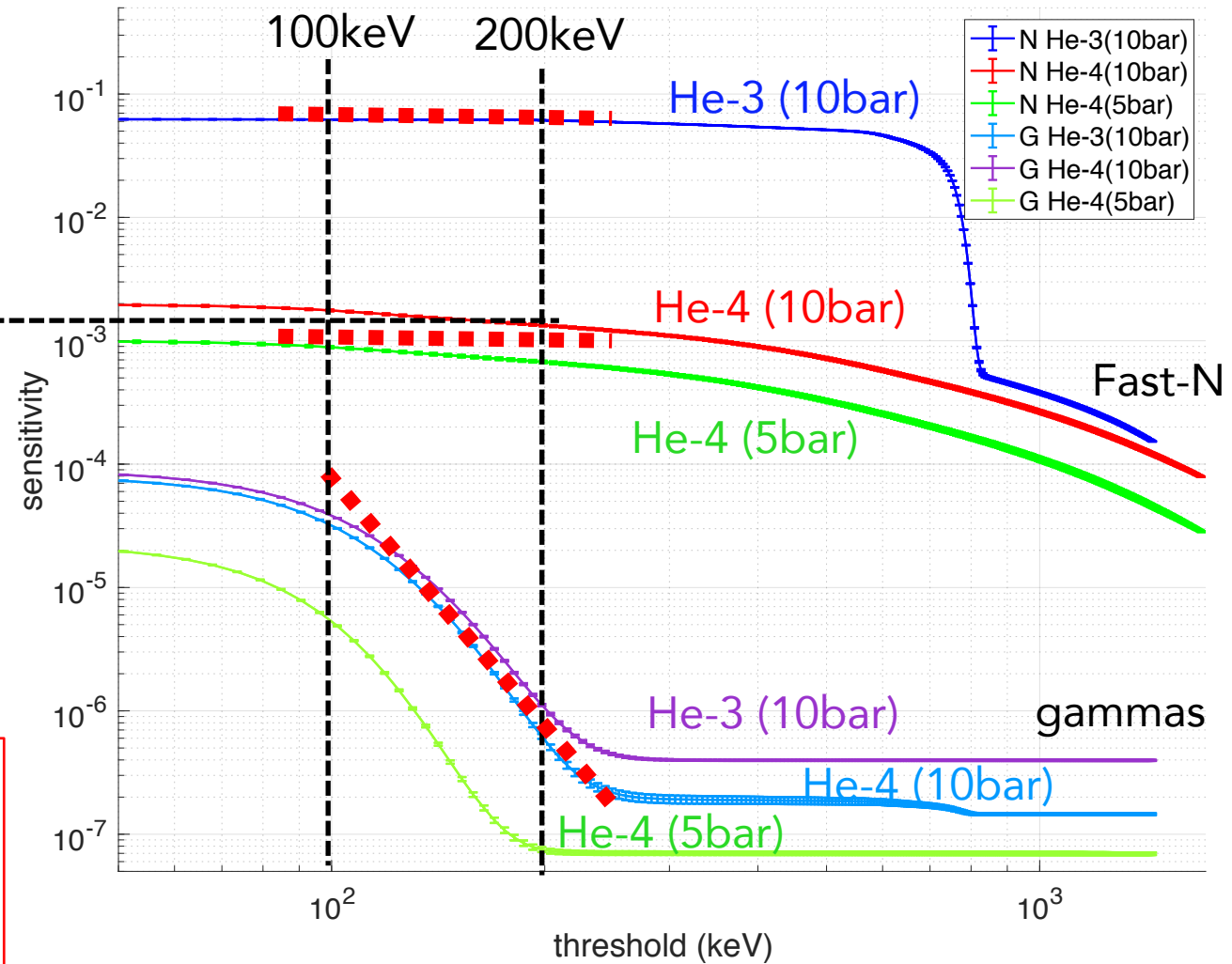
10^{-3}

Difference between He3 or B-10
(Gas Dominates)

NOTE:

high pressure He-3 detectors,
little gain in thermal n efficiency with
drastic increase of fast n sensitivity!
S/B matters!

Sensitivity / Efficiency



BACKGROUND: fast neutrons

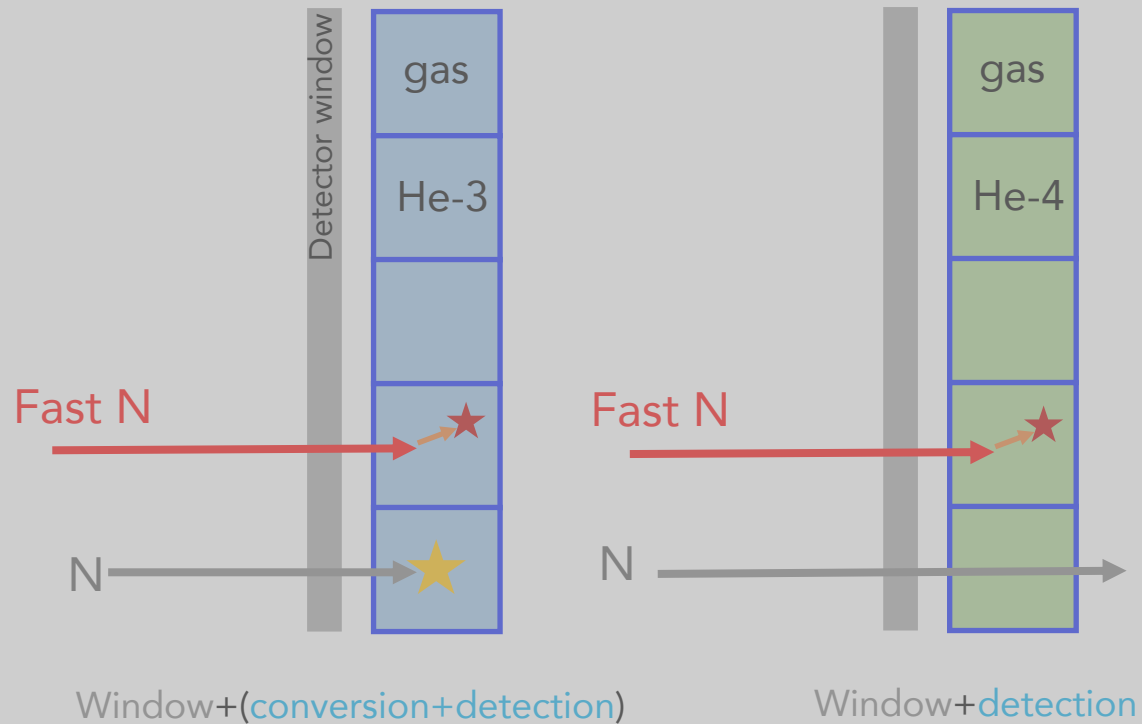
Gas is main component, then He-3 differs from B-10 ...

Gas dominates

Recoil on He-3

He-3 detector

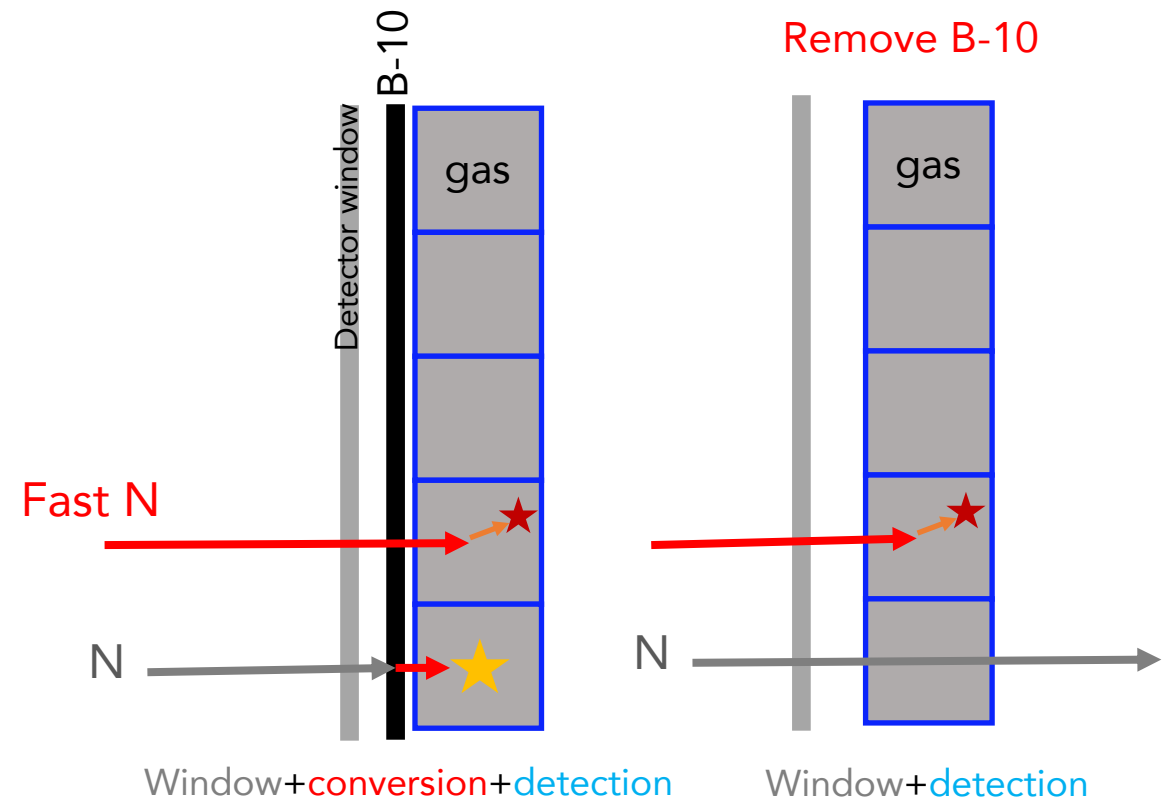
Exchange He-3 with He-4



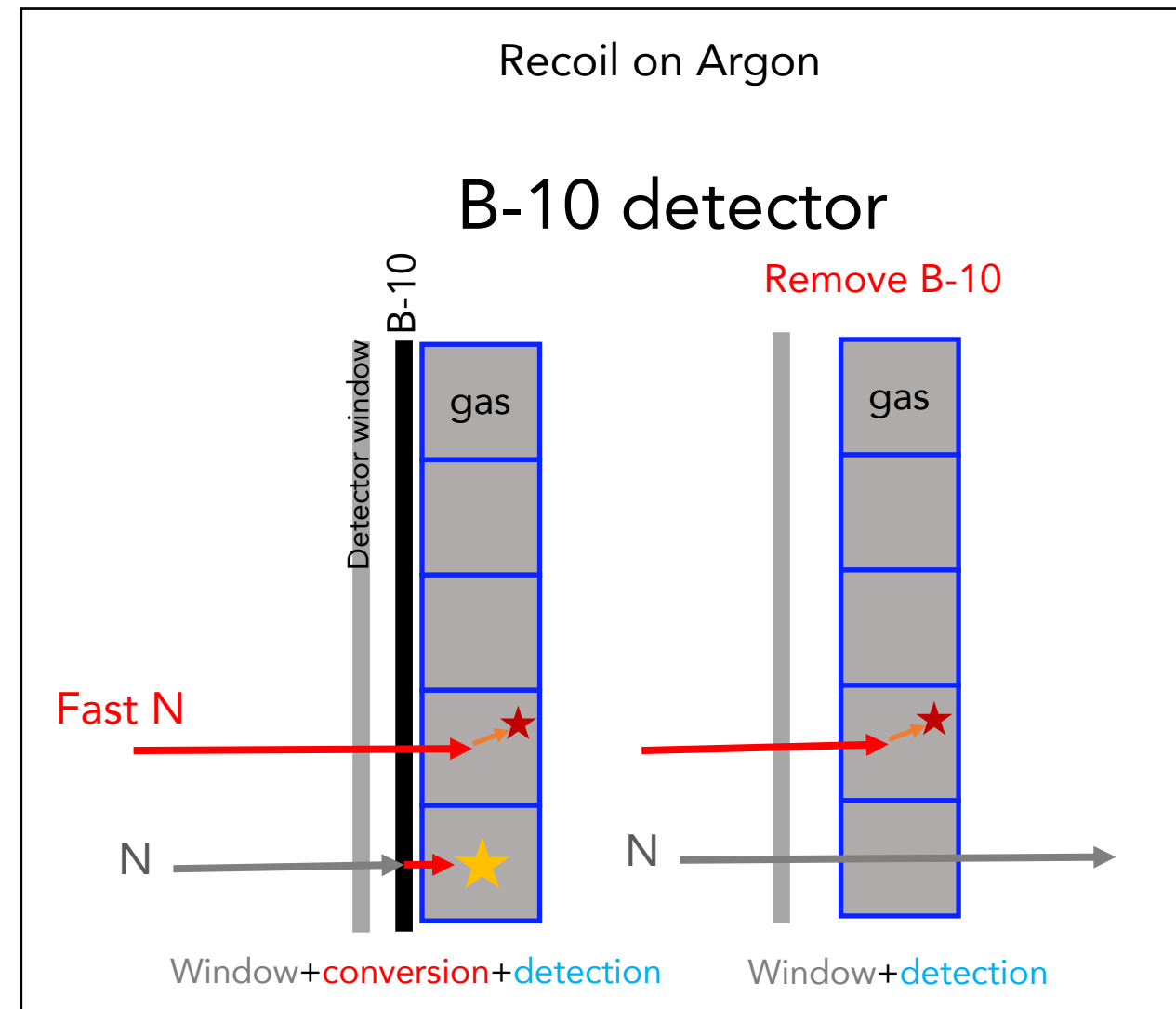
Recoil on Argon

B-10 detector

Remove B-10



BACKGROUND: fast neutrons in B-10 detectors



BACKGROUND: fast neutrons in B-10 detectors

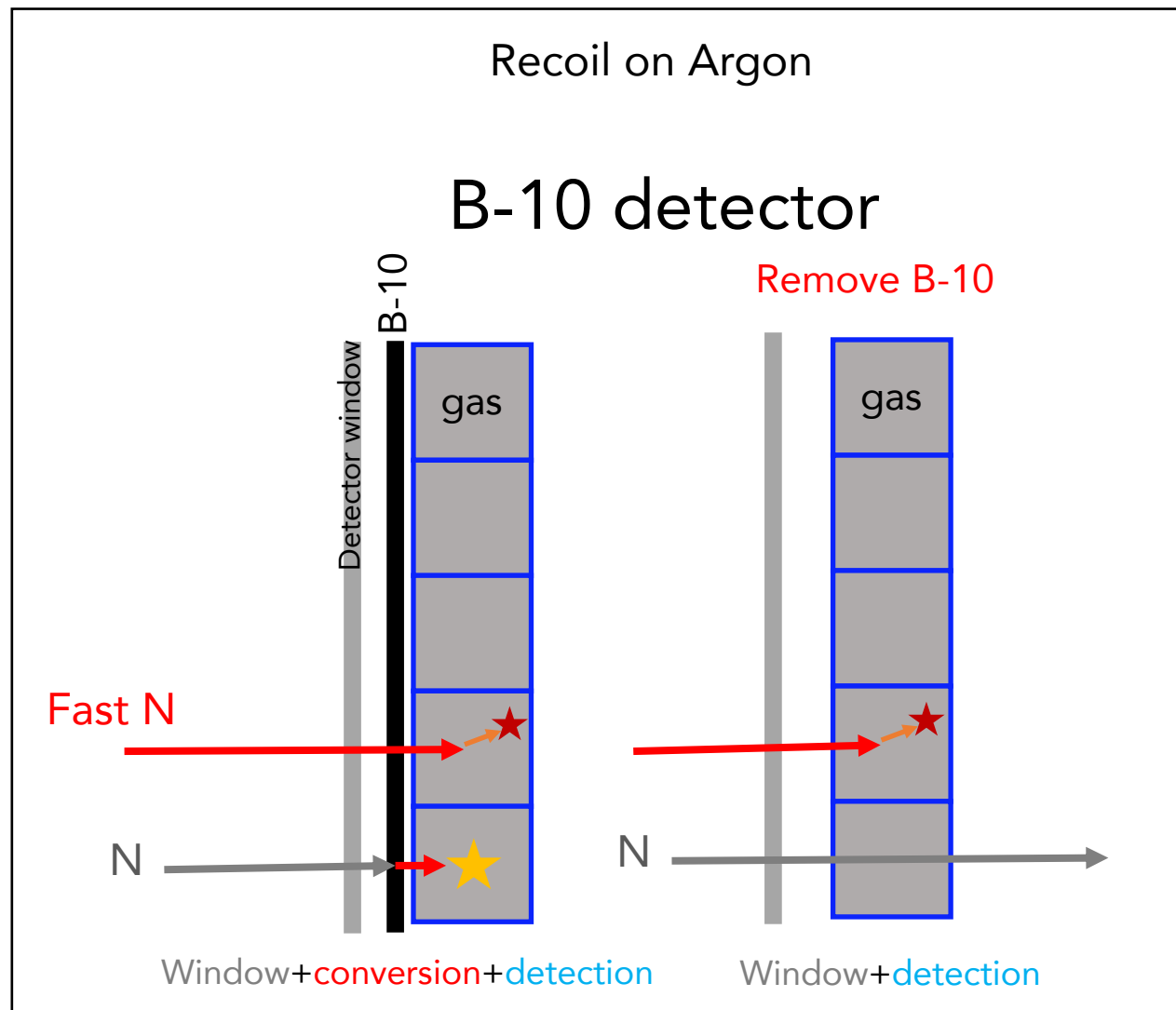
Gas dominates

Gas = x100 Solid

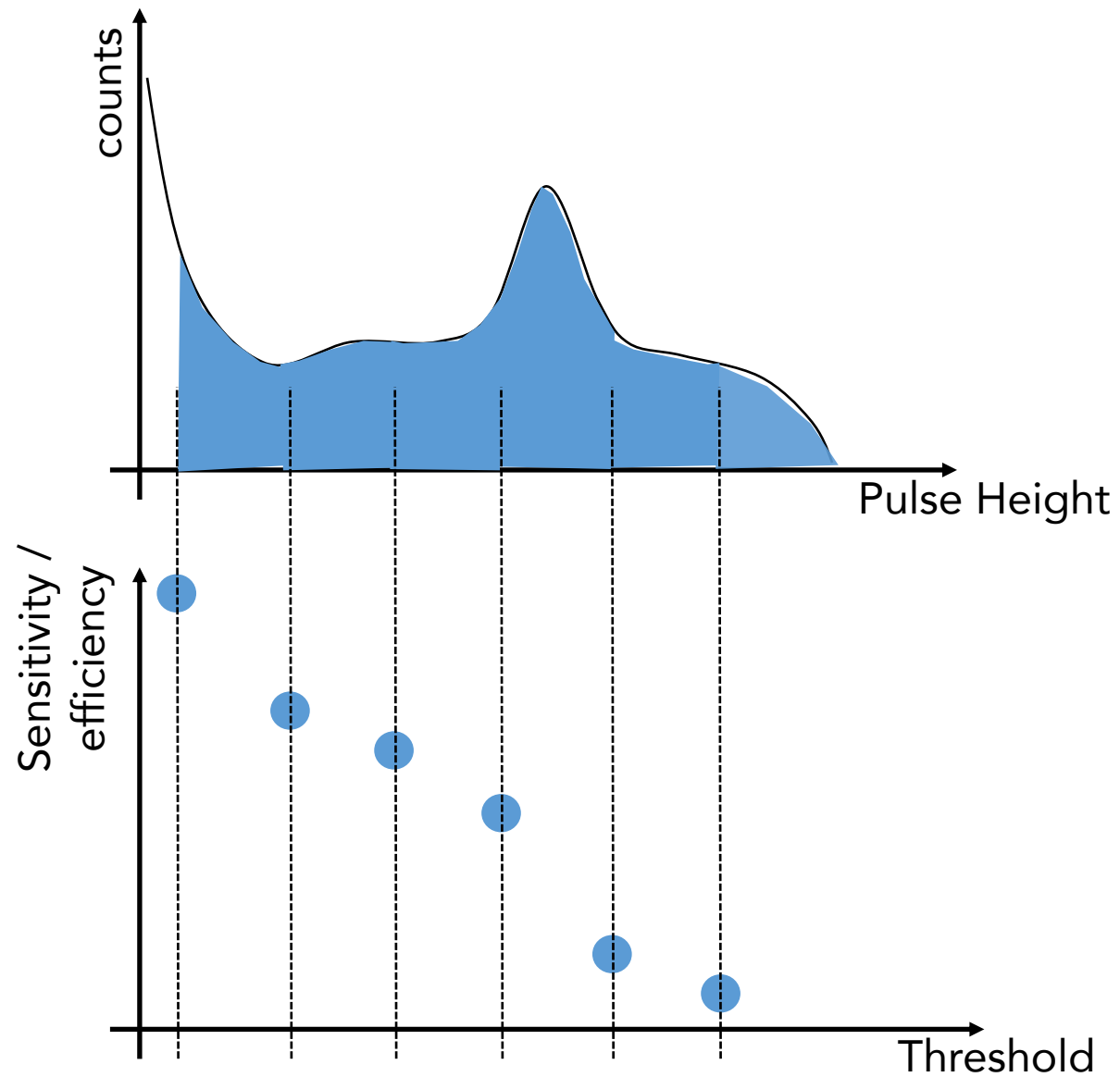
Compared Titanium, Aluminum, Copper, Kapton



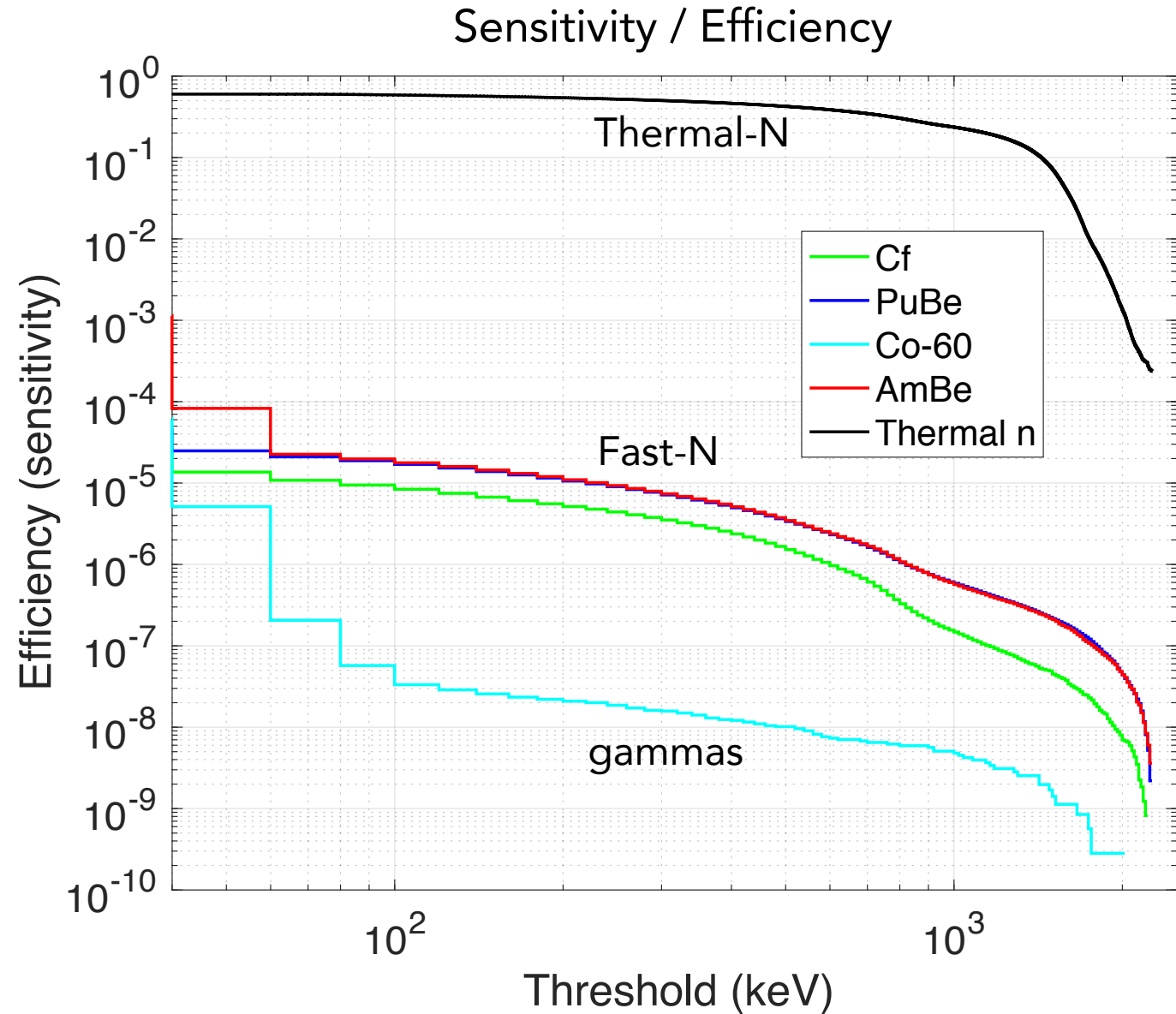
Experimentally (+simulations) proved -> no difference!



Pulse Height Spectrum - PHS



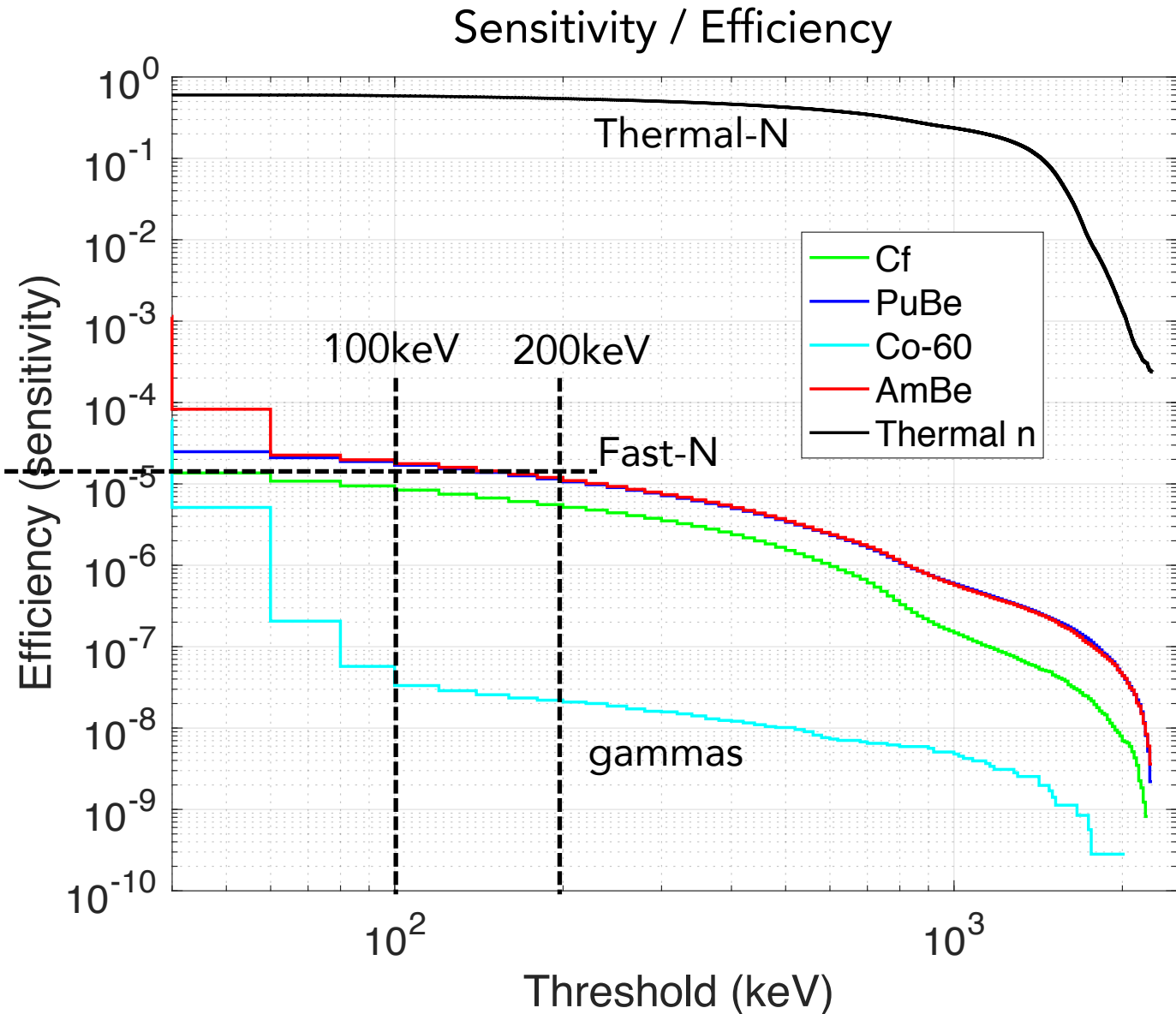
BACKGROUND: fast neutrons in B-10 detectors



BACKGROUND: fast neutrons in B-10 detectors

Fast Neutron Sensitivity
B-10
 10^{-5}

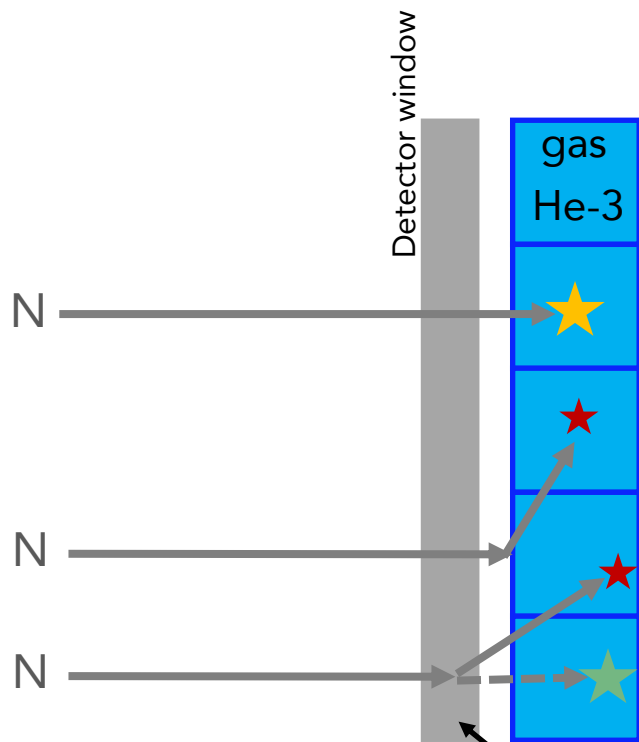
Difference between He3 or B-10
 (Gas Dominates)



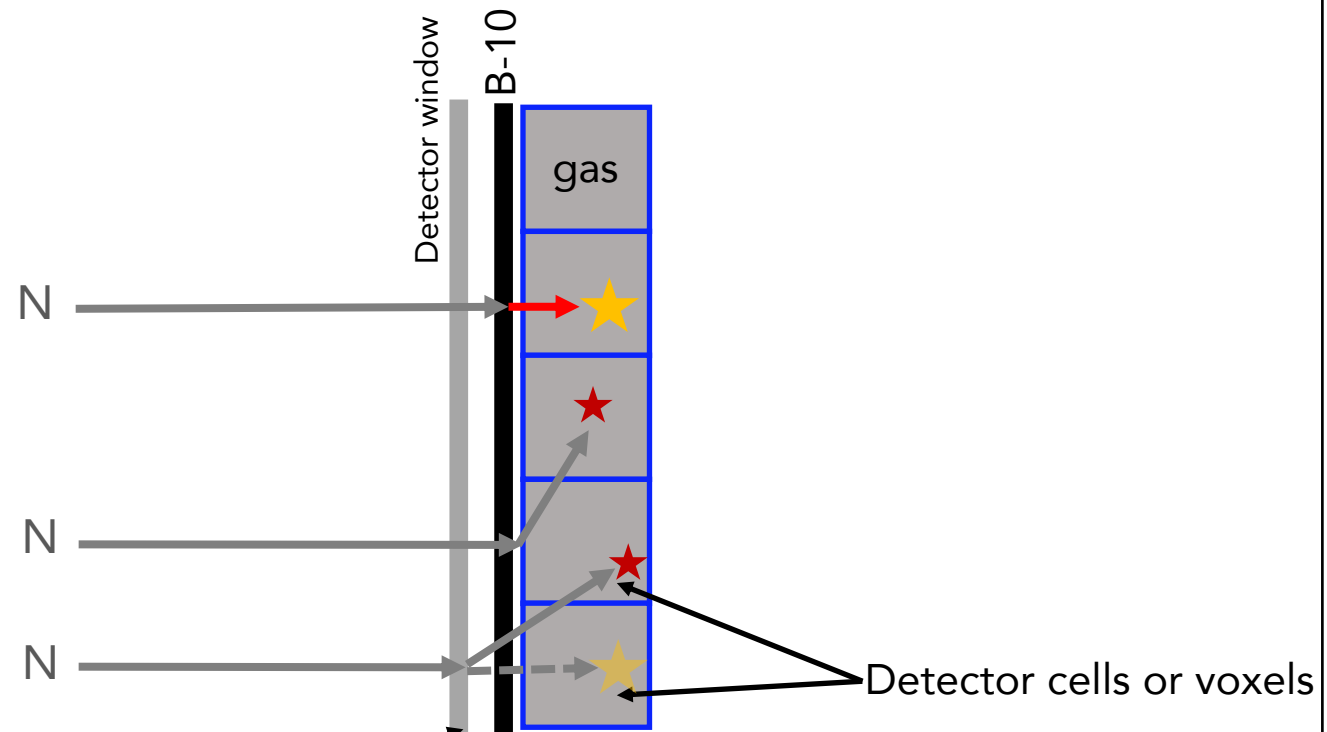
BACKGROUND DUE TO SCATTERED NEUTRONS

BACKGROUND: scattering

He-3 detector

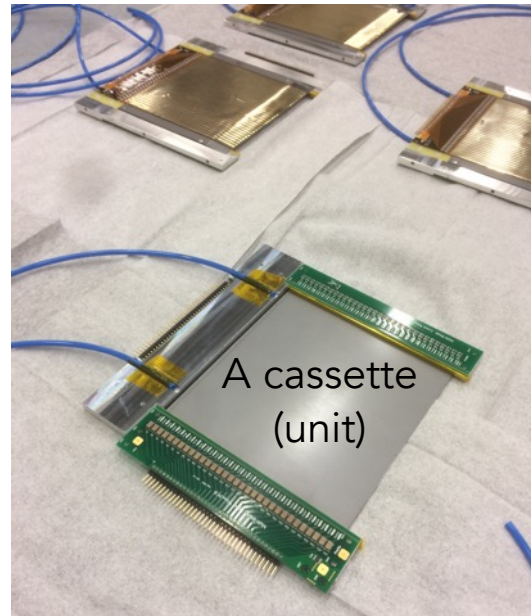
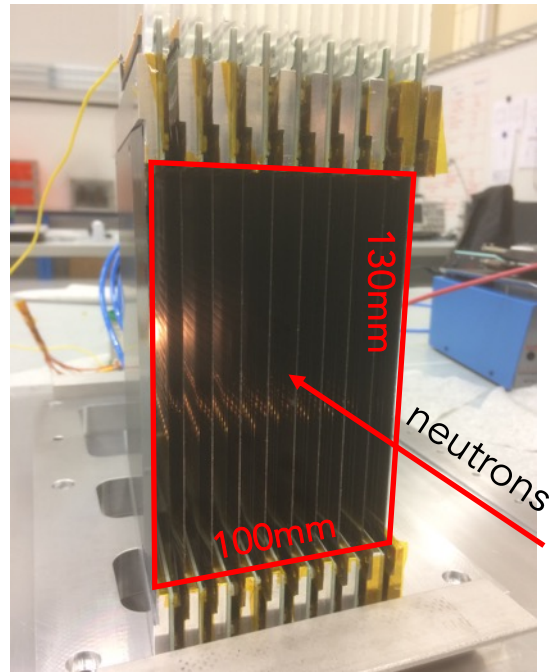


B-10 detector

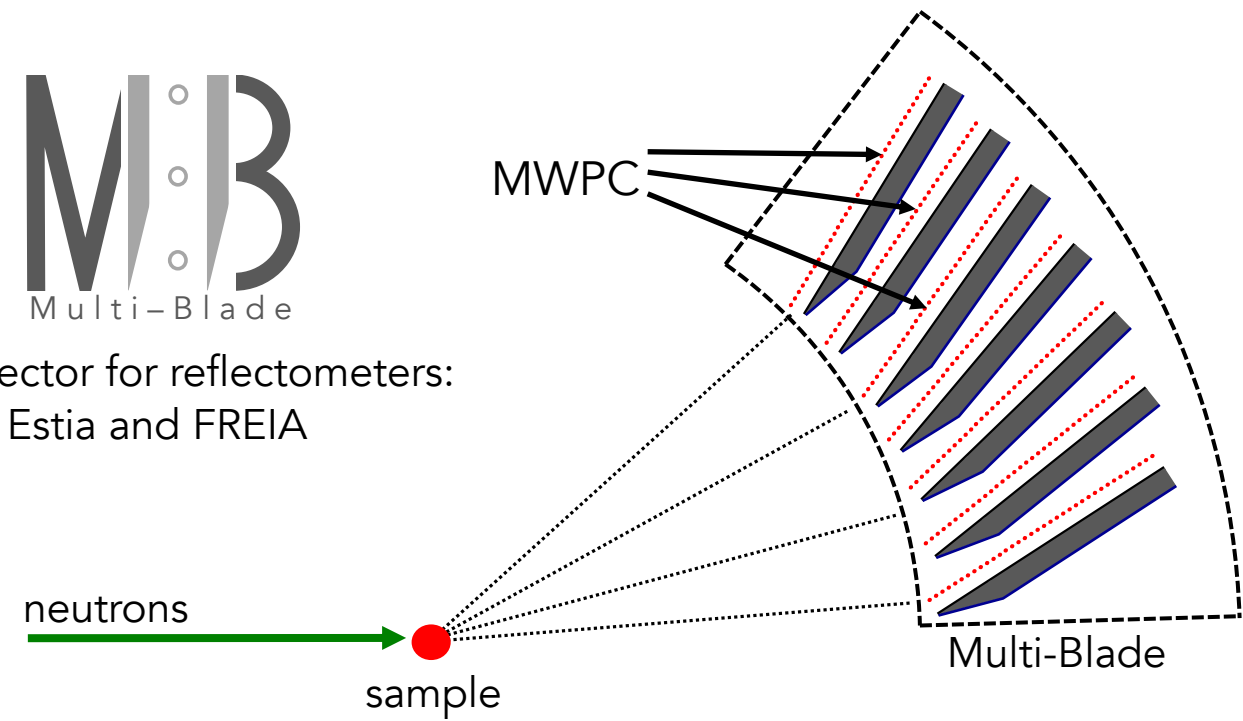


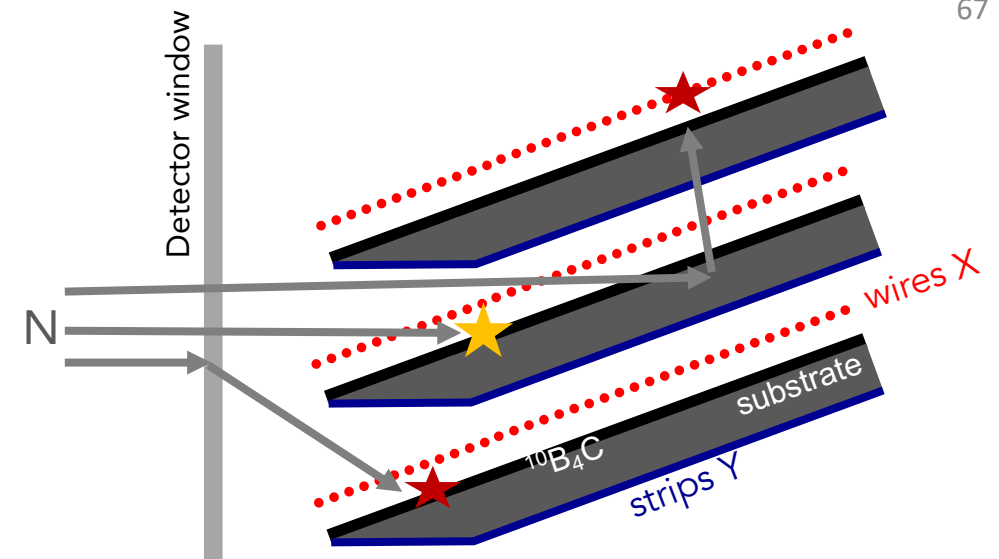
- Scattering at the detector window: B10 has generally thinner windows due to lower pressure
- Scattering at other parts of the detector

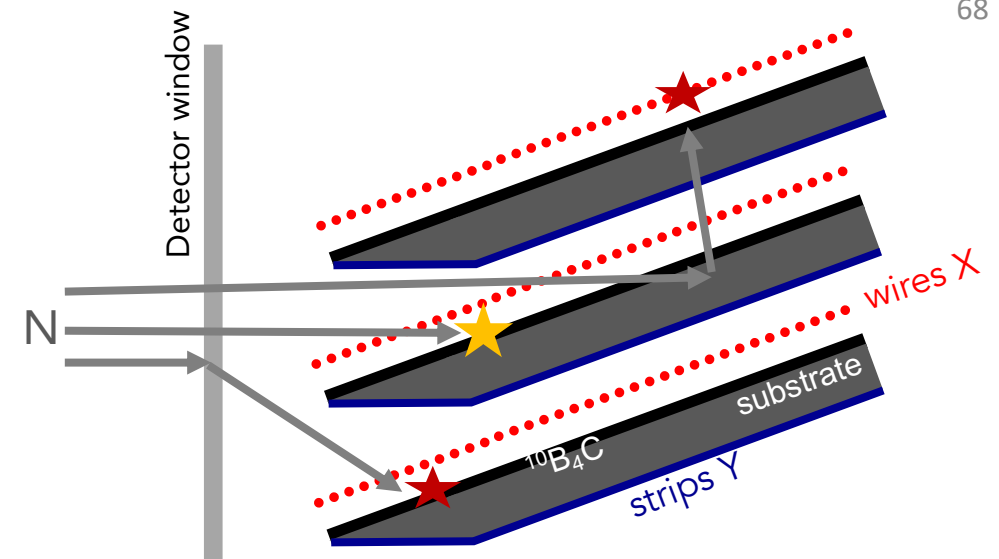
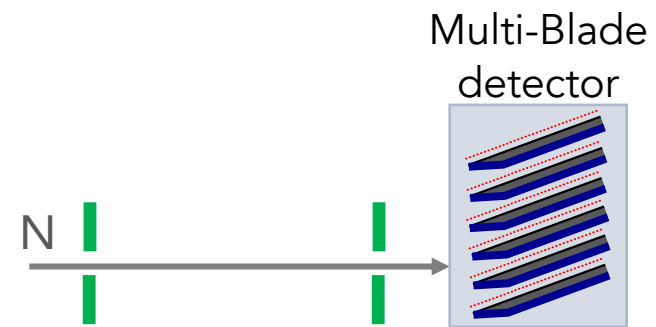
BACKGROUND: scattering

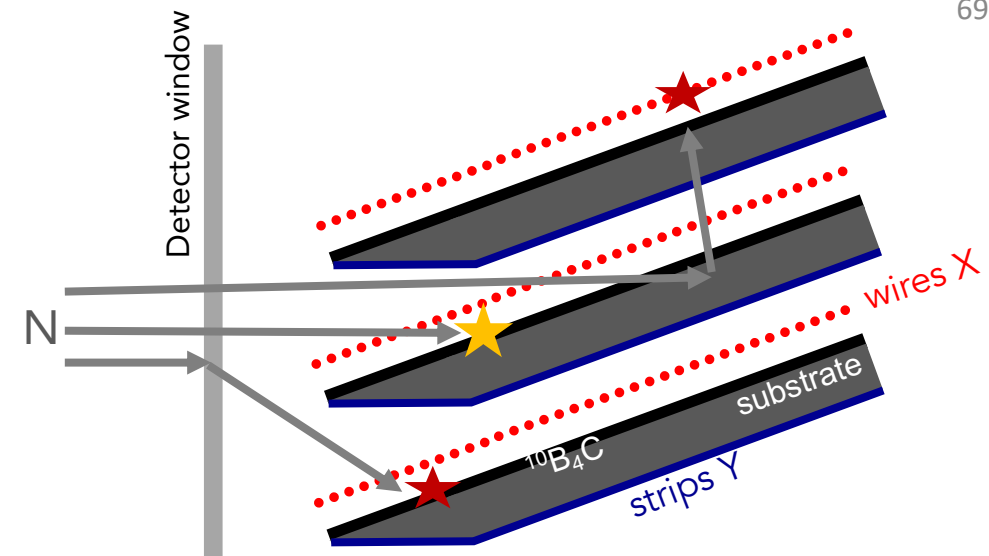
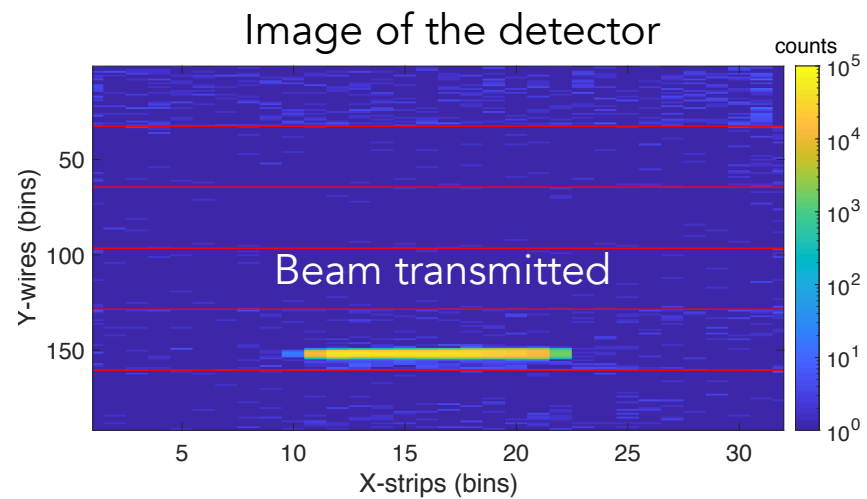
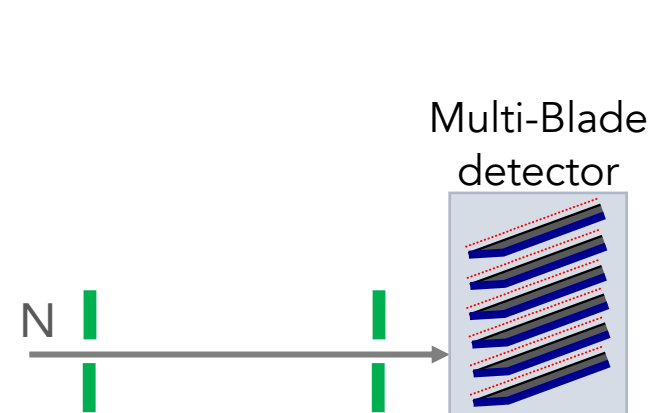


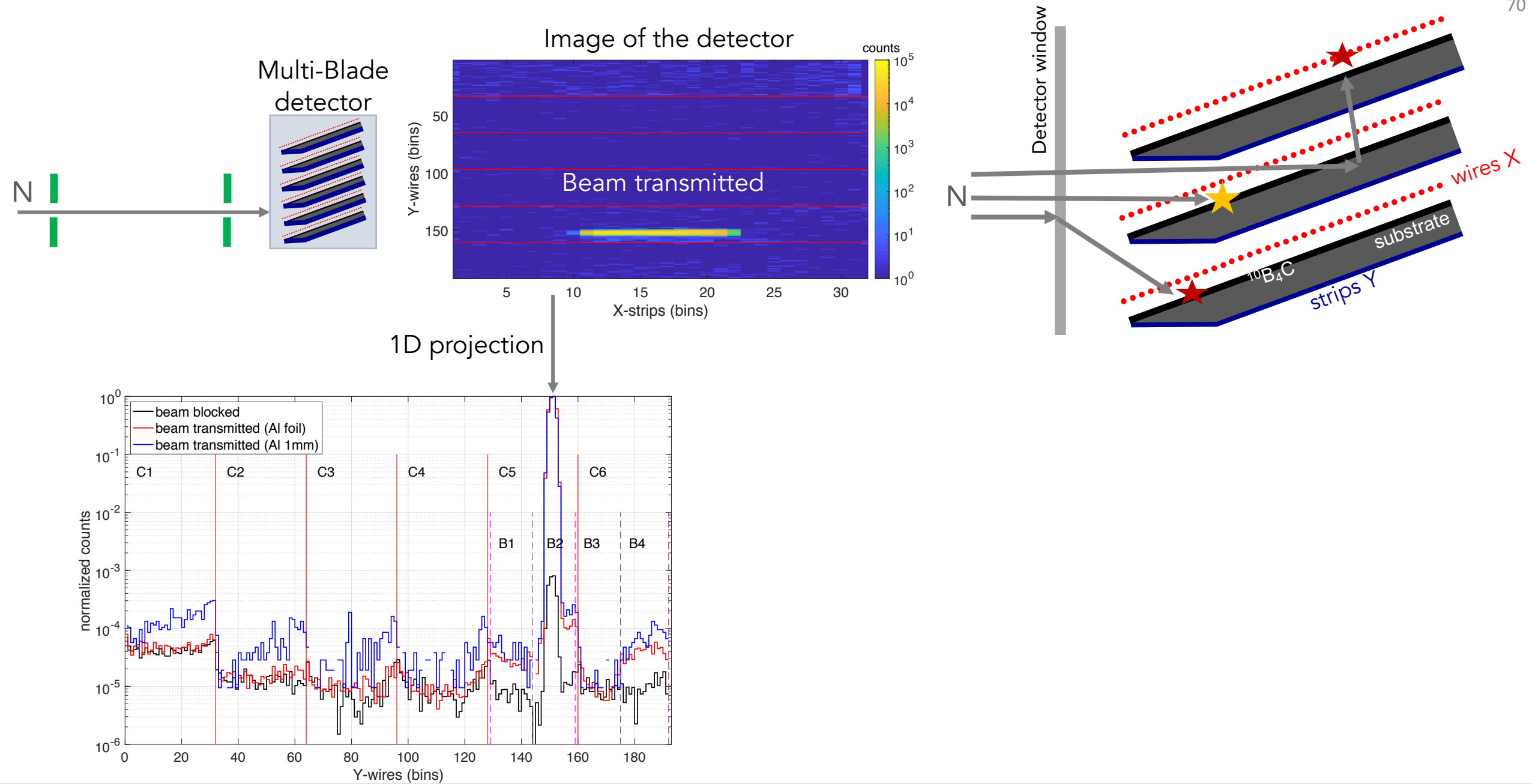

 Multi-Blade
 ^{10}B -detector for reflectometers:
 Estia and FREIA

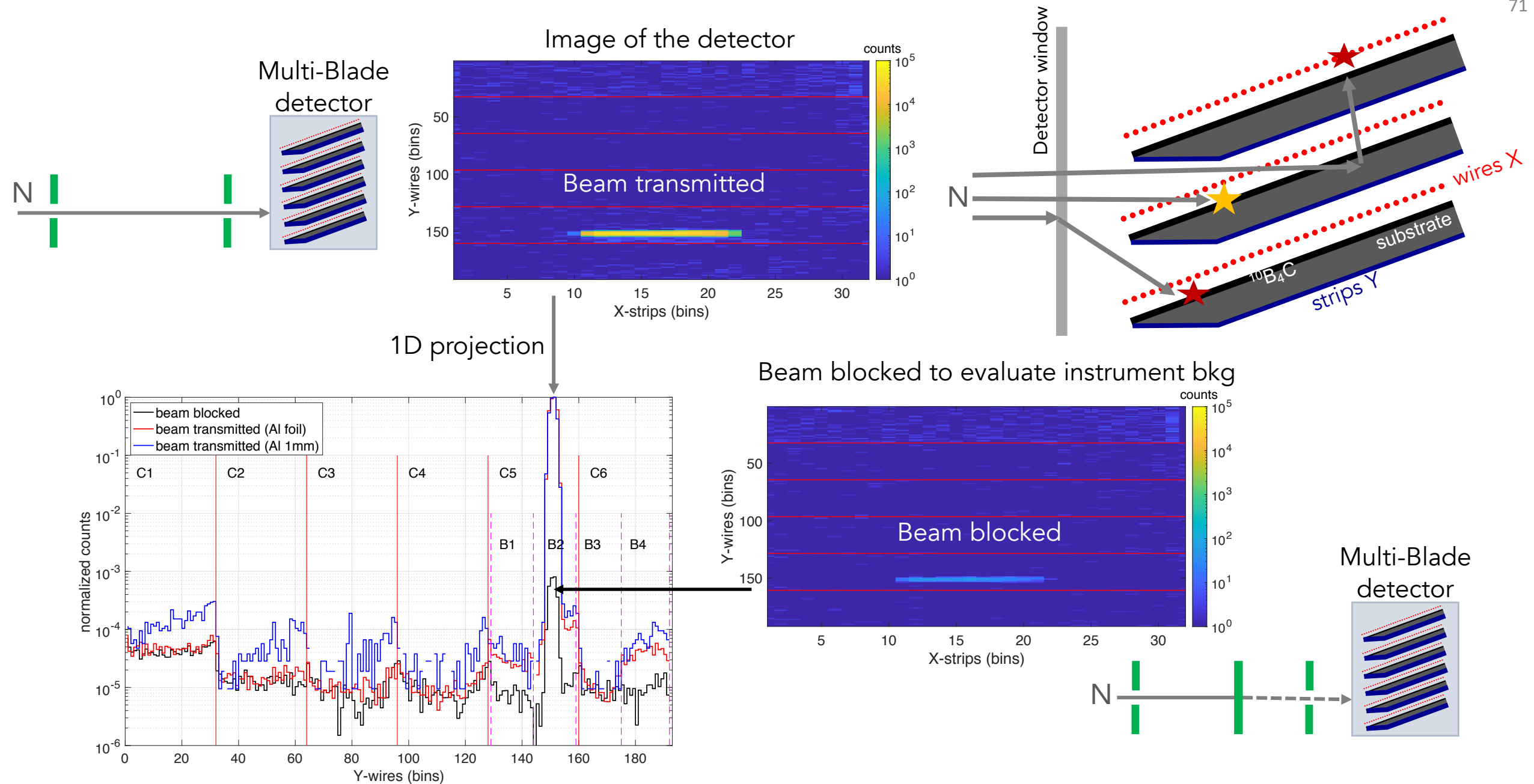


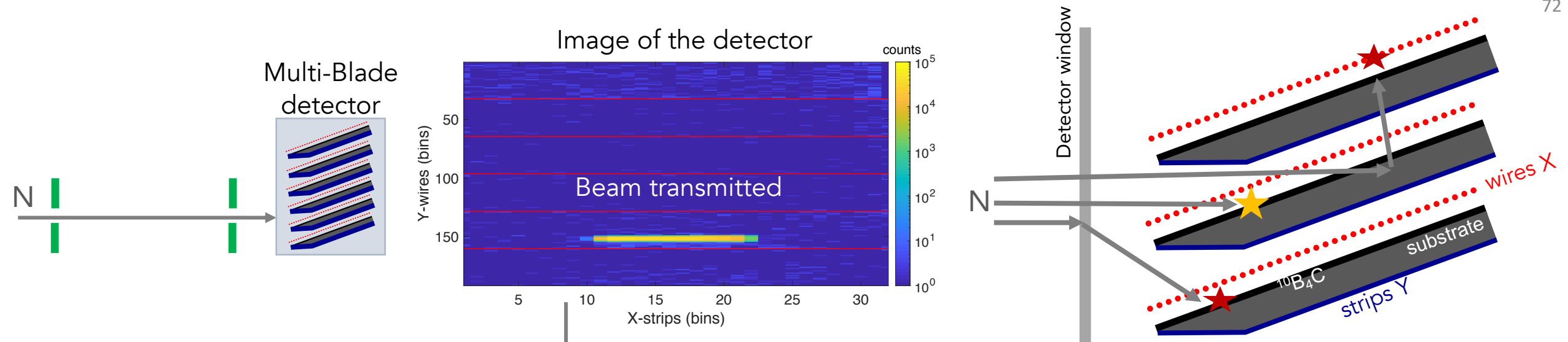




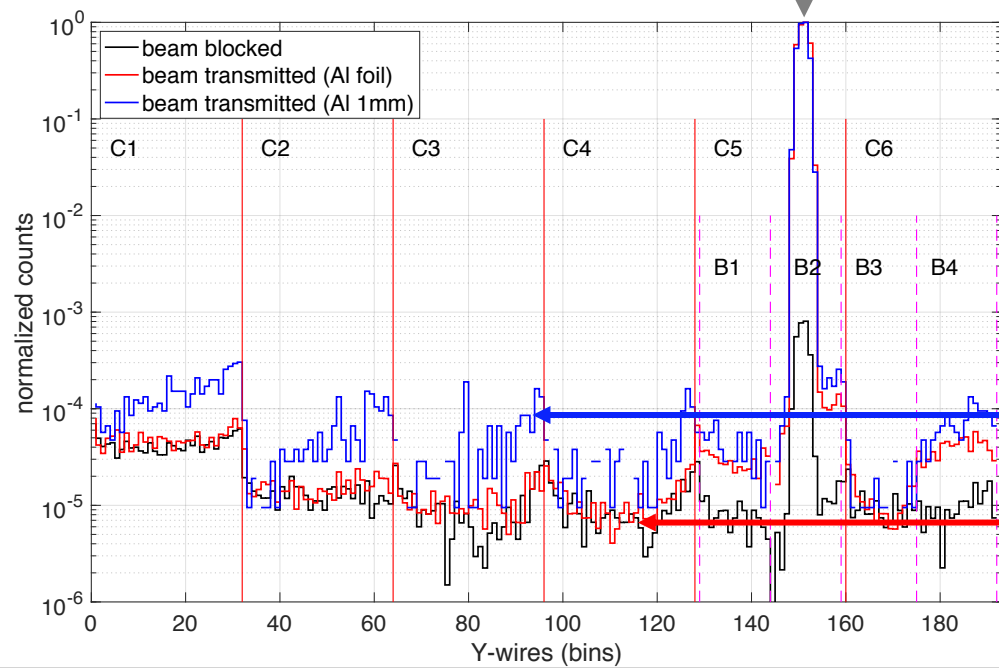






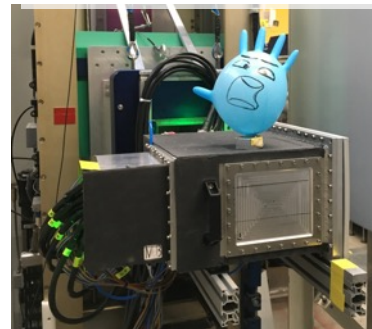


1D projection

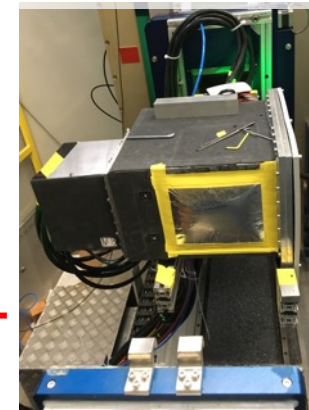


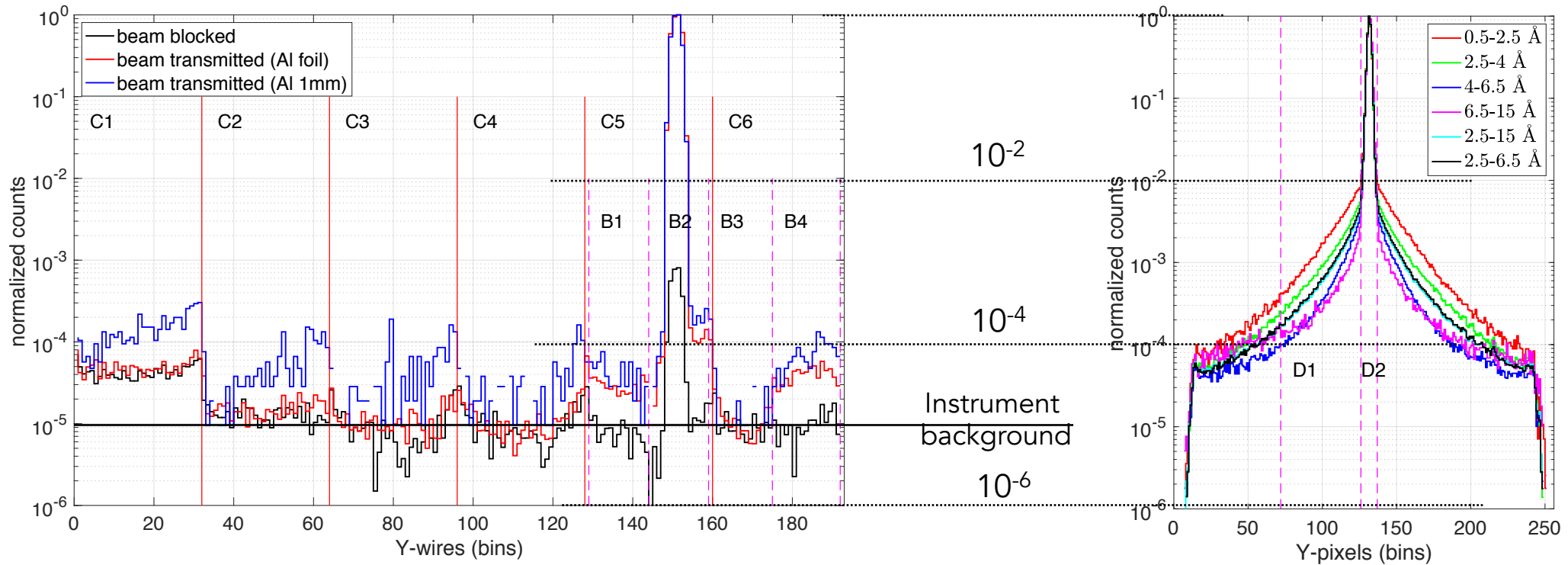
DETECTOR WINDOW

1mm Aluminium



50um Aluminium foil



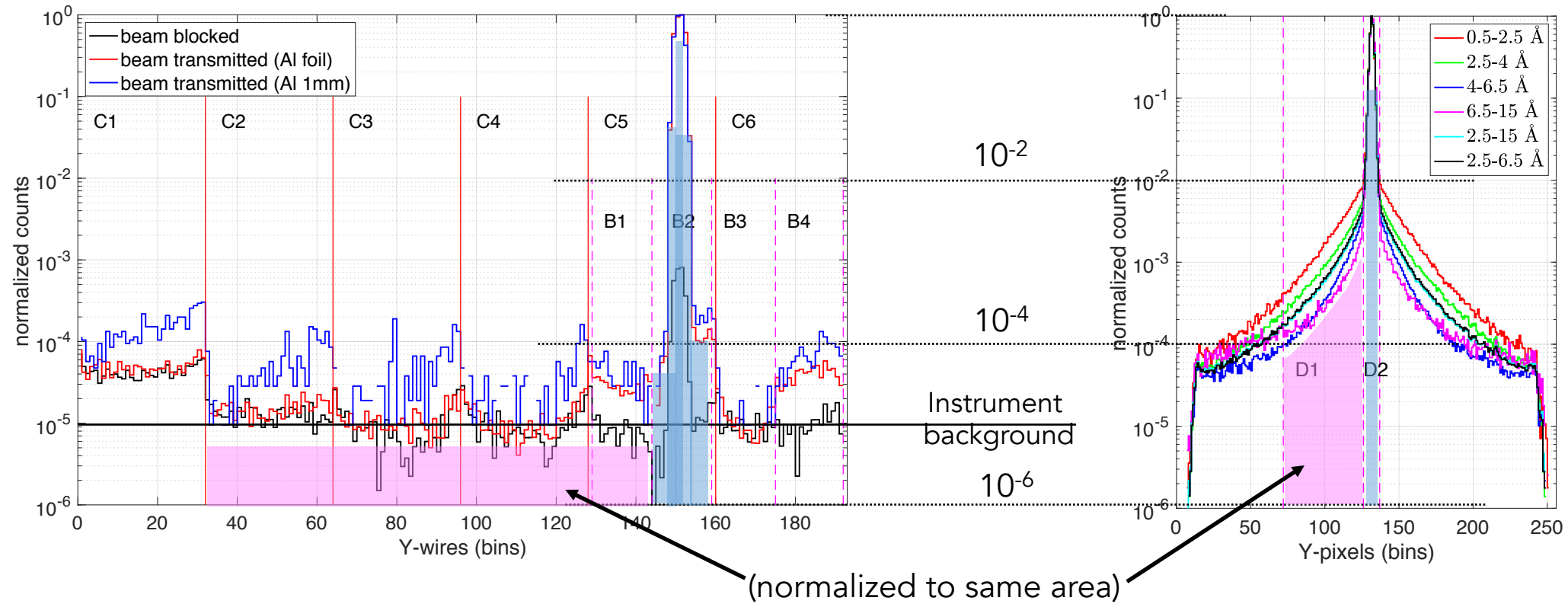


* Thanks to A. Glavic and J. Stahn

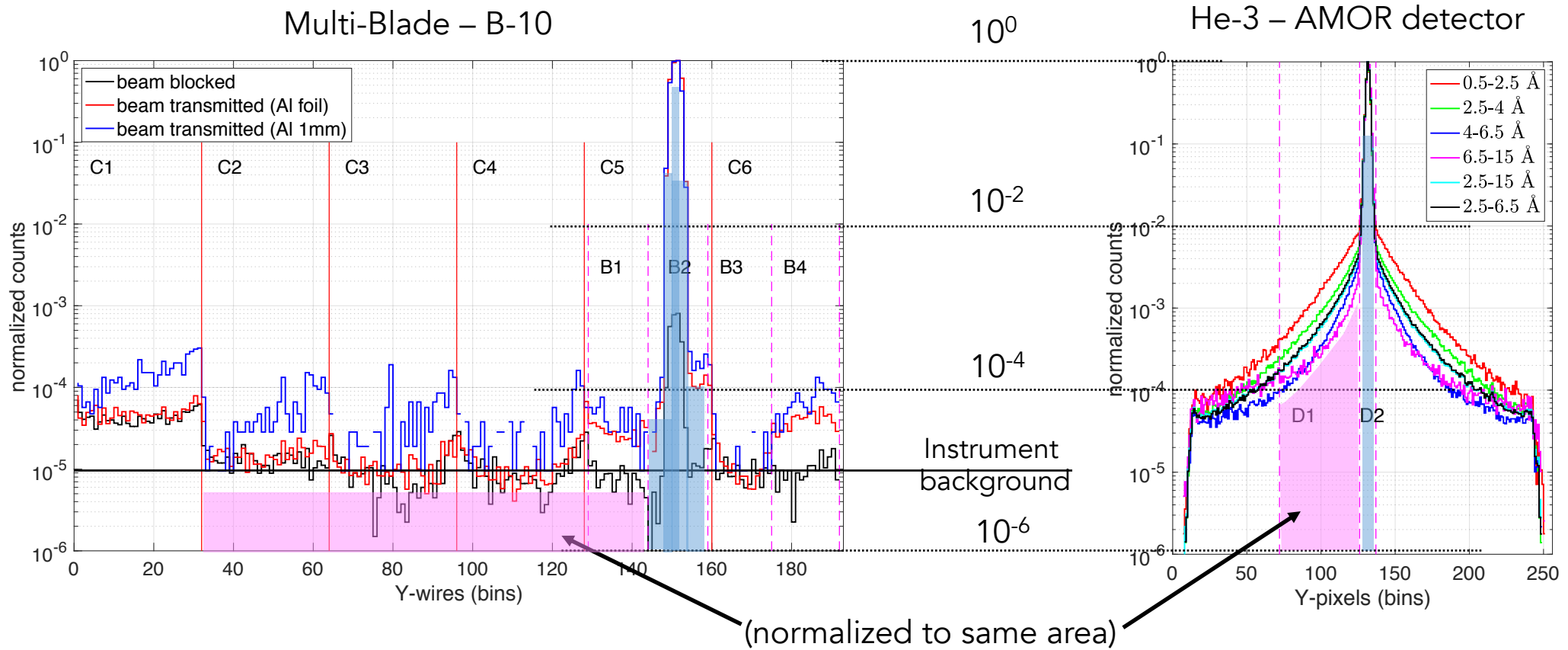
Multi-Blade – B-10

 10^0

He-3 – AMOR detector

Figure-of-Merit \propto Signal-to-Background

$$FoM = \frac{D2}{D1}$$



Detector	FoM (2.5-15Å)
He-3 (a few mm Al window)	4000
Multi-Blade B10 – Al 1mm window	73000
Multi-Blade B10 – Al foil window	211000

Figure-of-Merit \propto Signal-to-Background

$$FoM = \frac{D2}{D1}$$

x20

x50

- **1** @ESS 23 Hz/m² cosmic thermal neutrons and 1 Hz/m² with shielding

Conclusions

- **1** @ESS 23 Hz/m² cosmic thermal neutrons and 1 Hz/m² with shielding
- **2** Aluminium foil as a detector window reduces x50 the background generated by scattered neutrons

Conclusions

(*results are general and applicable for any facility)

- **1** @ESS 23 Hz/m² cosmic thermal neutrons and 1 Hz/m² with shielding
- **2** Aluminium foil as a detector window reduces x50 the background generated by scattered neutrons

- **3**

Boron-10		Helium-3
0.5 – 0.8	Thermal N Efficiency	0.6 - 1
10^{-5}	Fast neutron sensitivity (gas dominates)	10^{-3}
$10^{-6} - 10^{-9}$	Gamma-ray sensitivity (solid dominates)	$10^{-6} - 10^{-9}$

Highly affected by a small variation of the threshold

Gamma-ray sensitivity

- A. Khaplanov et al., Investigation of gamma-ray sensitivity of neutron detectors based on thin converter films, JINST 8 P10025 (2013) (arxiv: 1306:6247)
- F. Piscitelli et al., The Multi-Blade Boron-10-based Neutron Detector for high intensity Neutron Reflectometry at ESS, JINST 12 P03013 (2017) (arxiv: 1701.07623)

Fast neutron sensitivity

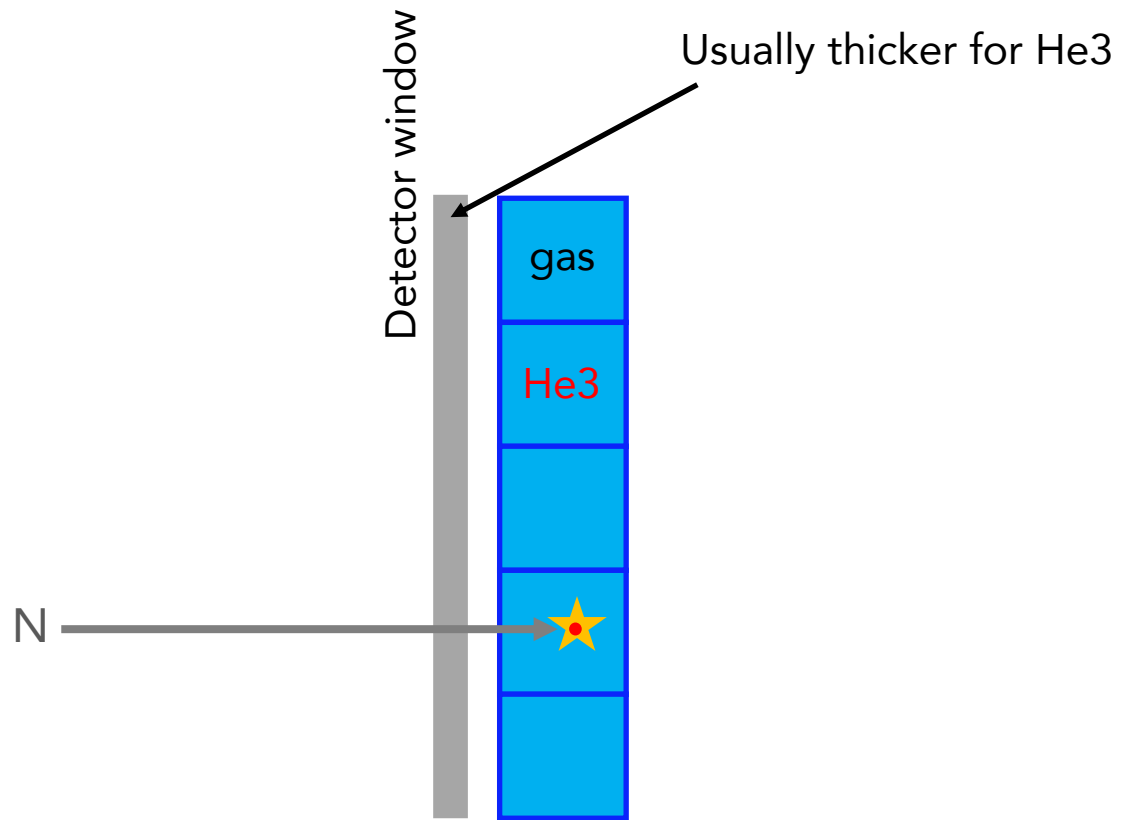
- F. Piscitelli et al., Verification of He-3 proportional counters fast neutron sensitivity through a comparison with He-4 detectors, sub. to EPJ Plus (2020) (arxiv: 2002.08153)
- G. Mauri et al., Fast neutron sensitivity for ^3He detectors and comparison with Boron-10 based neutron detectors, EPJ TI 6, no. 1, p. 3, (2019) (arxiv: 1902:09870)
- G. Mauri et al., Fast neutron sensitivity of neutron detectors based on Boron-10 converter layers, JINST 13 P03004 (2018) (arxiv: 1712.05614)

Scattered Neutron Background

- G. Mauri et al., The Multi-Blade Boron-10-based neutron detector performance using a focusing reflectometer, Accepted for publ. in JINST (2020) (arxiv: 2001:02965).
- F. Piscitelli et al., Characterization of the Multi-Blade 10B-based detector at the CRISP reflectometer at ISIS, JINST 13 P05009 (2018) (arxiv: 1803.09589)
- G. Galgoczi et al., Investigation of neutron scattering in the Multi-Blade detector with GEANT4 simulations, JINST 13 P12031 (2018) (arxiv: 1810:06241)
- E. Rossi, Master. thesis, Characterisation of the Spatial Resolution and the Gamma-ray Discrimination of Helium-3 Proportional Counters (2015) (arxiv: 1702:06501)
- J. Birch et al., Investigation of background in large-area neutron detectors due to alpha emission from impurities in aluminium, JINST 10, P10019 (2015) (arxiv: 1507:00607)
- J. Birch et al., In-beam test of the Boron-10 Multi-Grid neutron detector at the IN6 Time-of-Flight spectrometer at the ILL, J. Phys. Conf. Ser. 528:1, 012040 (2014)

BACKUP SLIDES

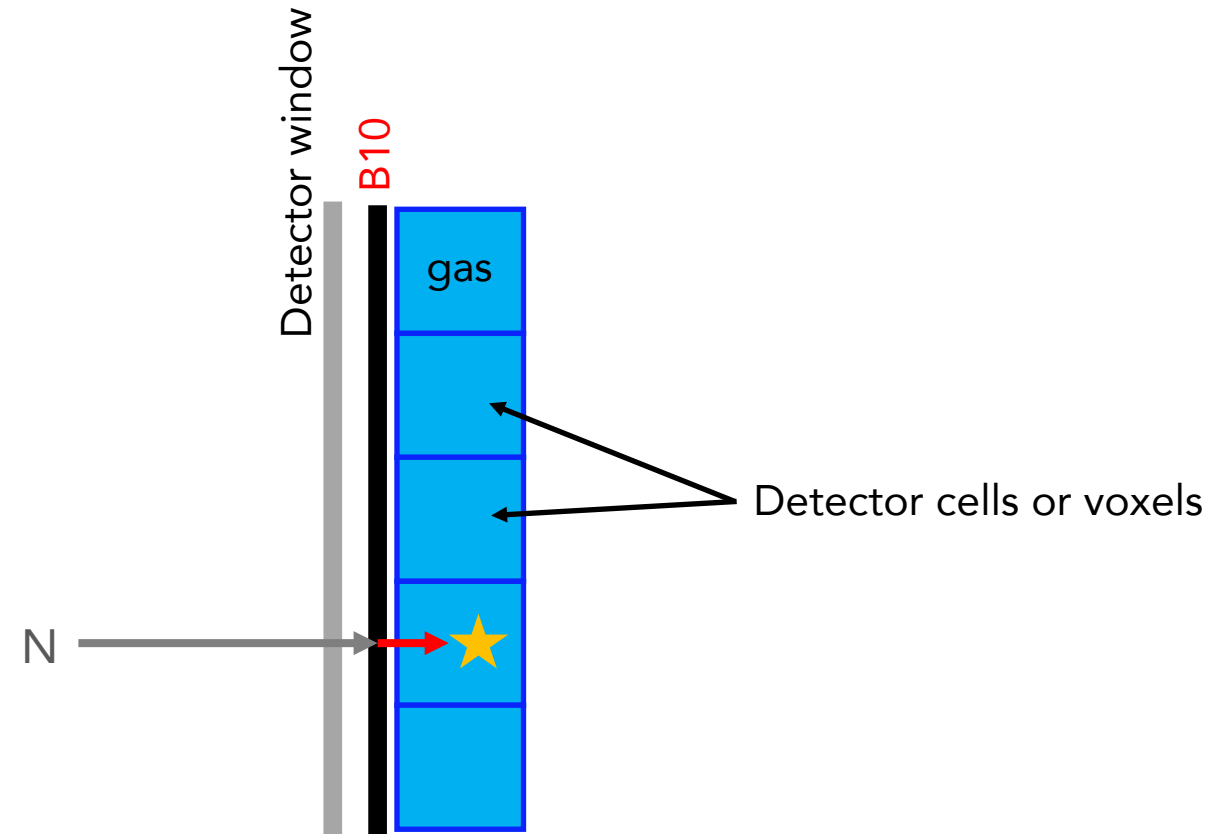
He3 detector



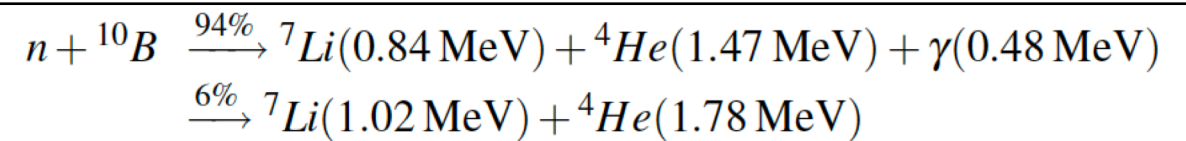
Window+(conversion+detection)



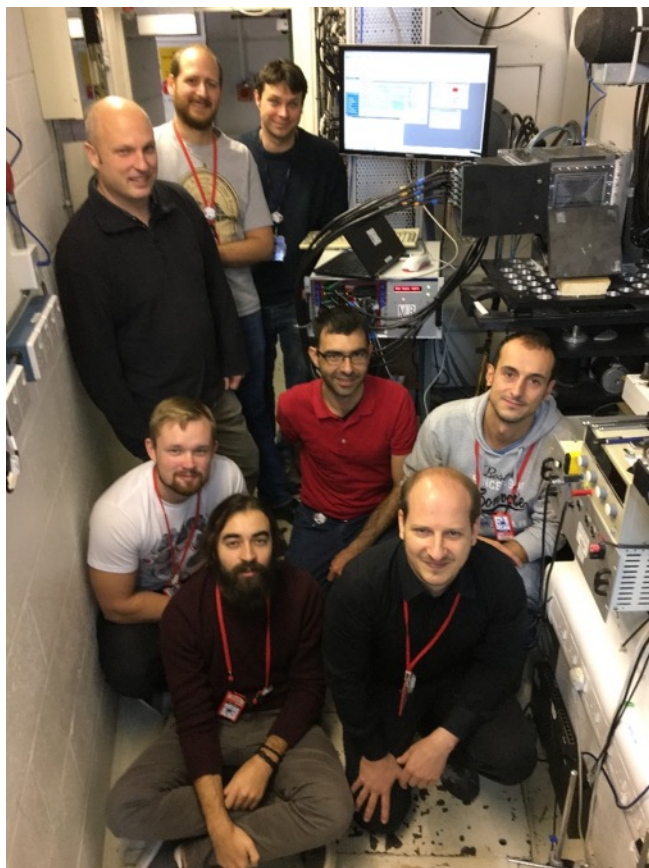
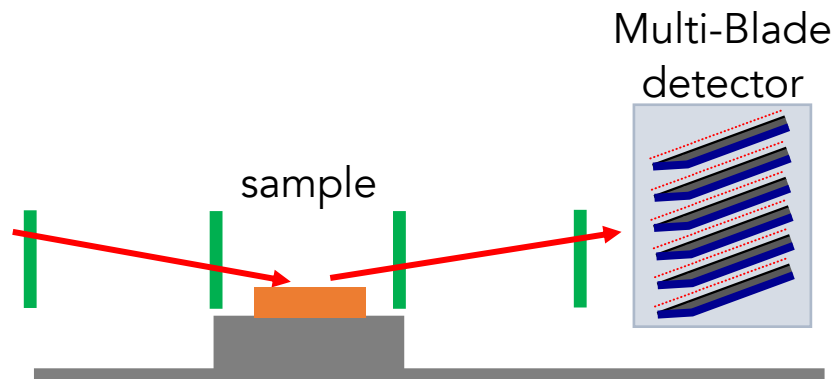
B10 detector



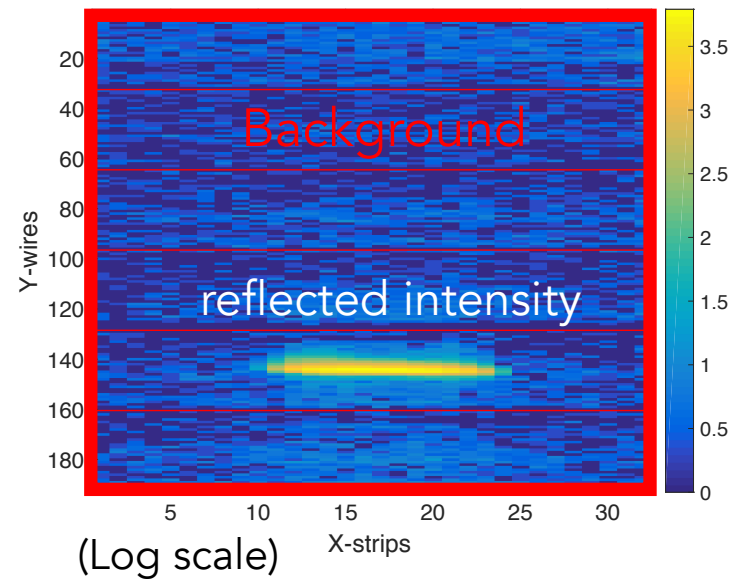
Window+conversion+detection



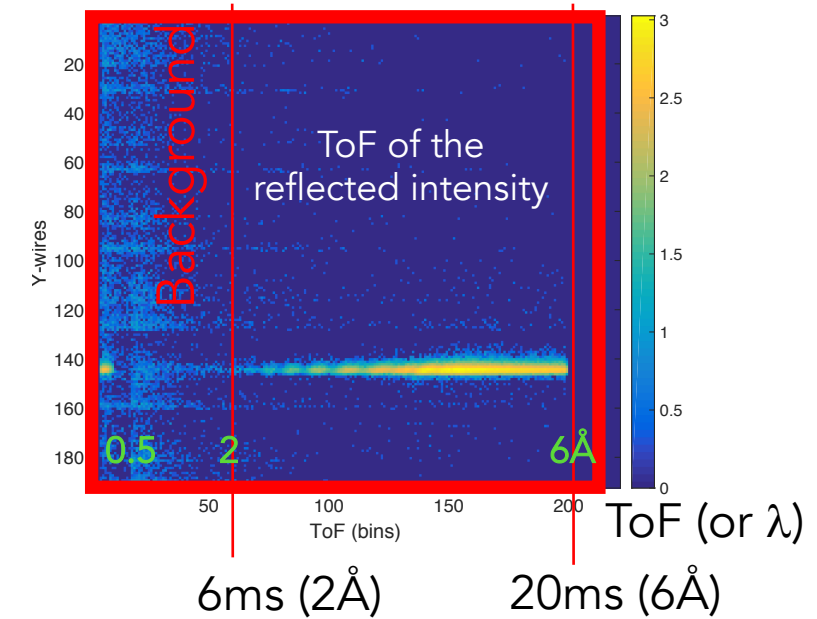
BACKGROUND DUE TO SCATTERED NEUTRONS



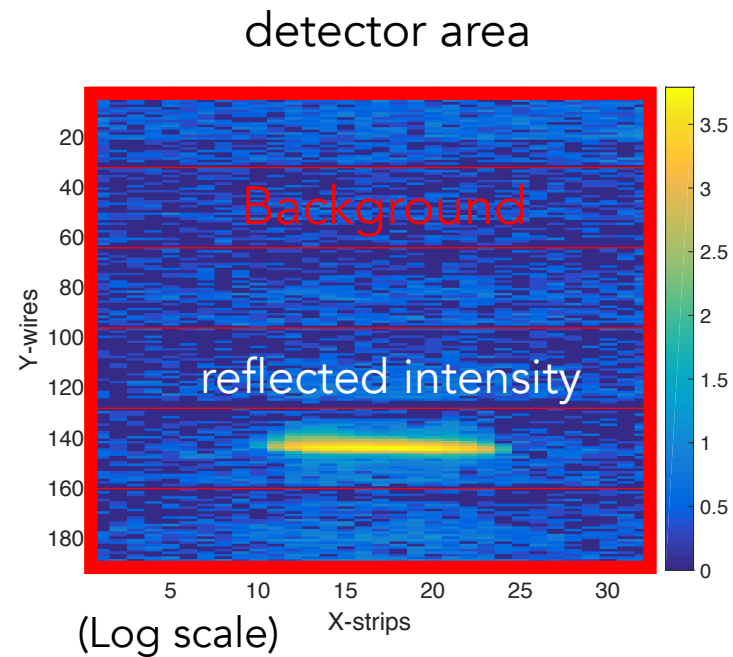
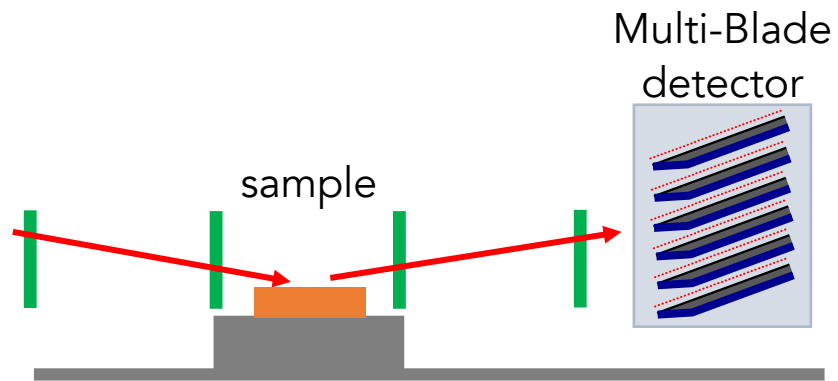
detector area



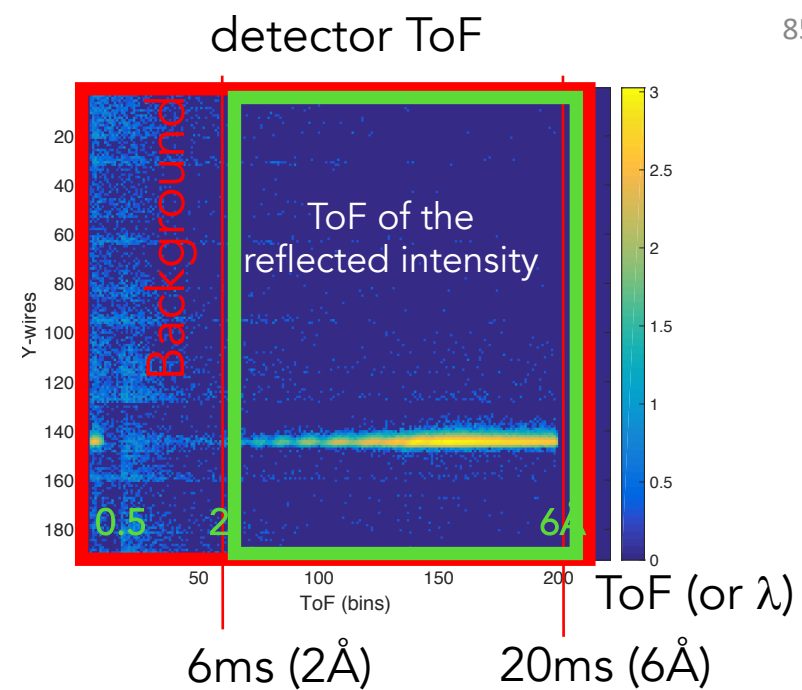
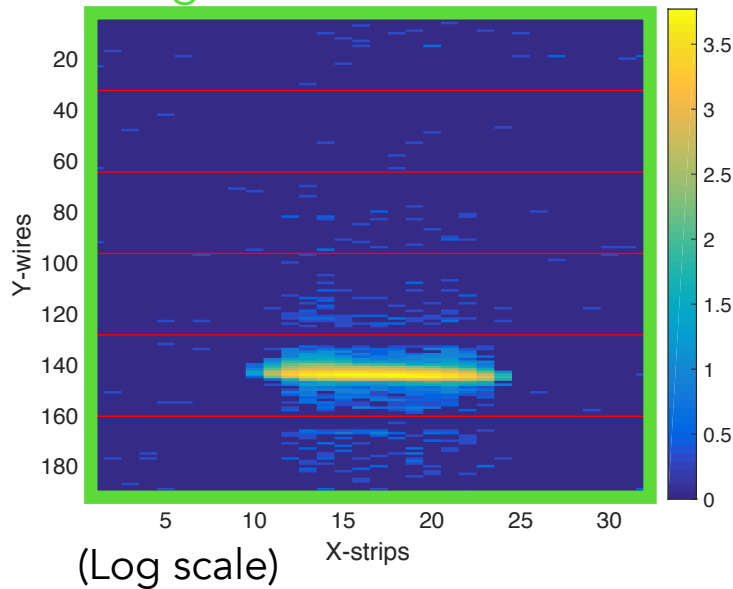
detector ToF

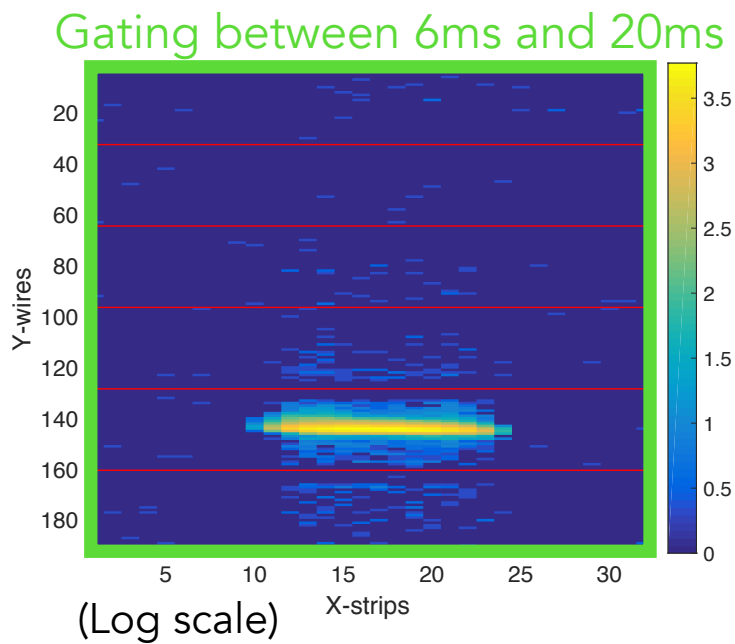
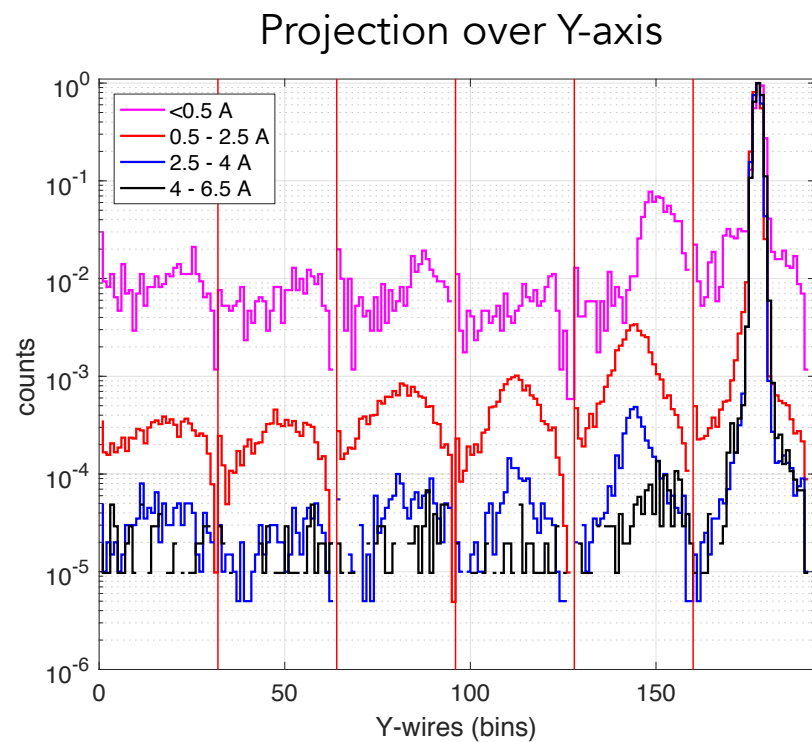
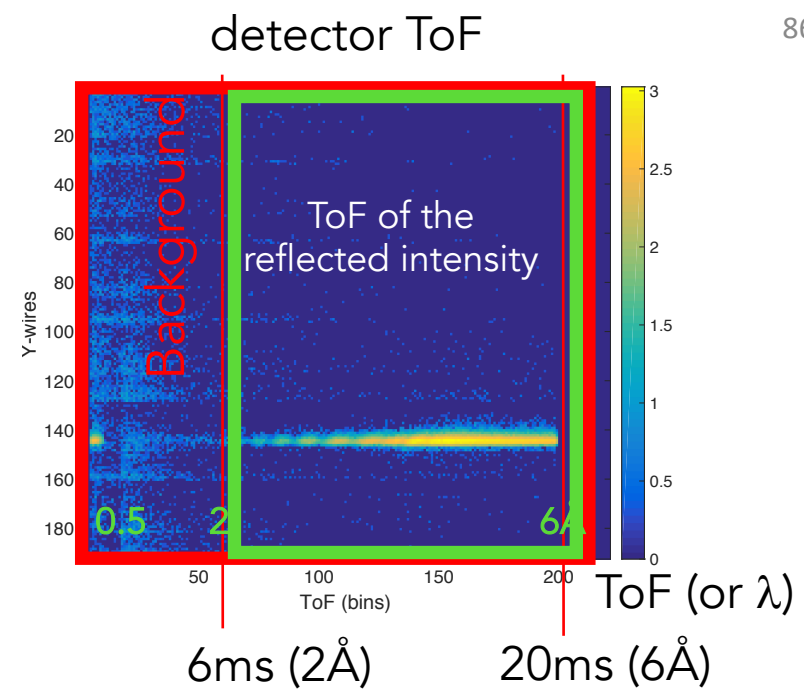
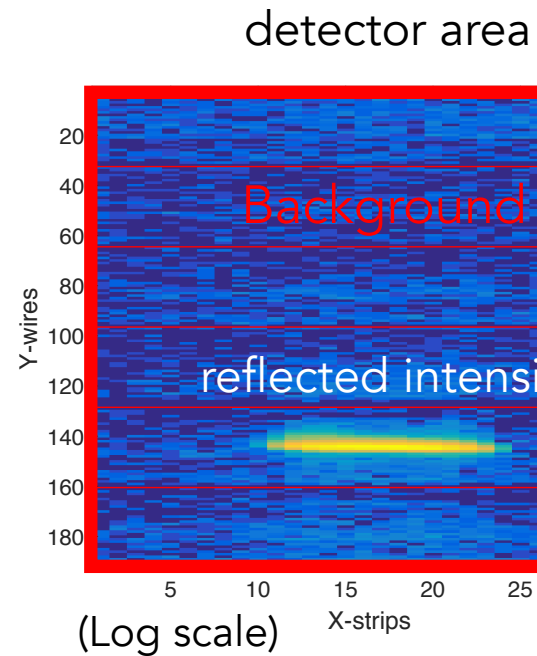
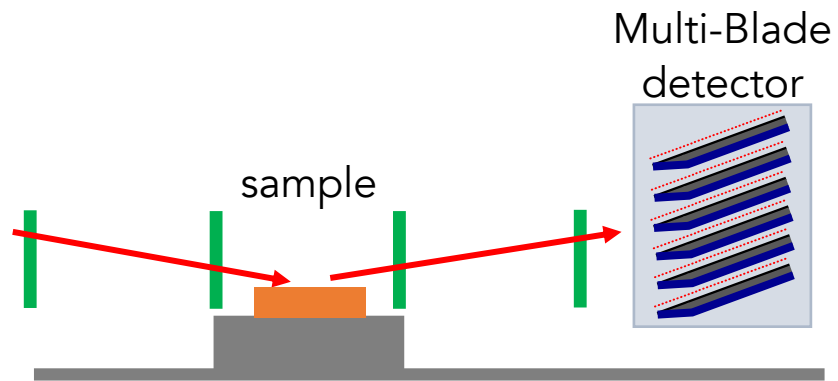


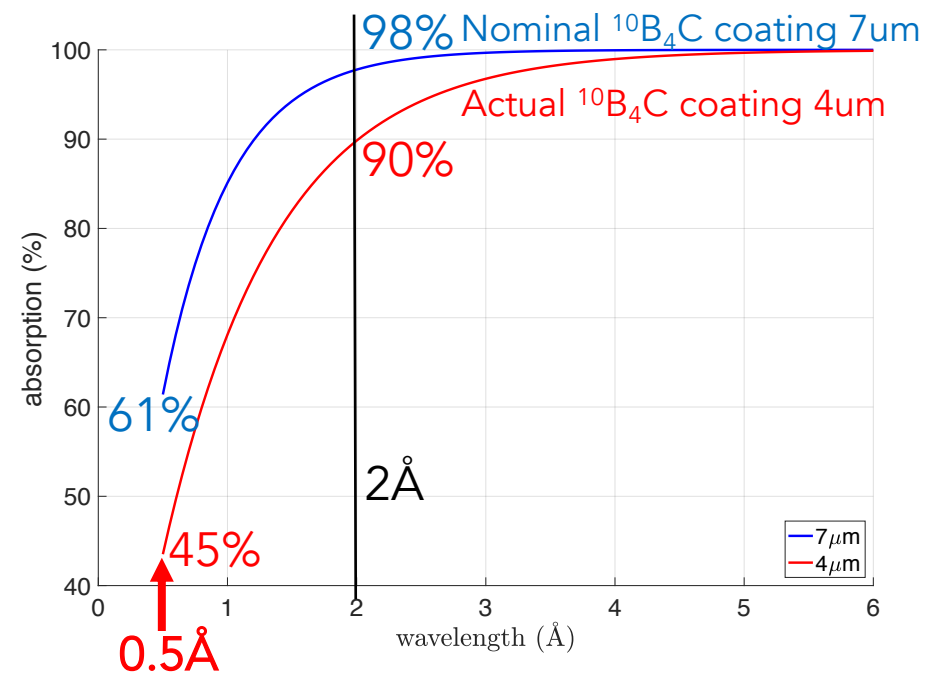
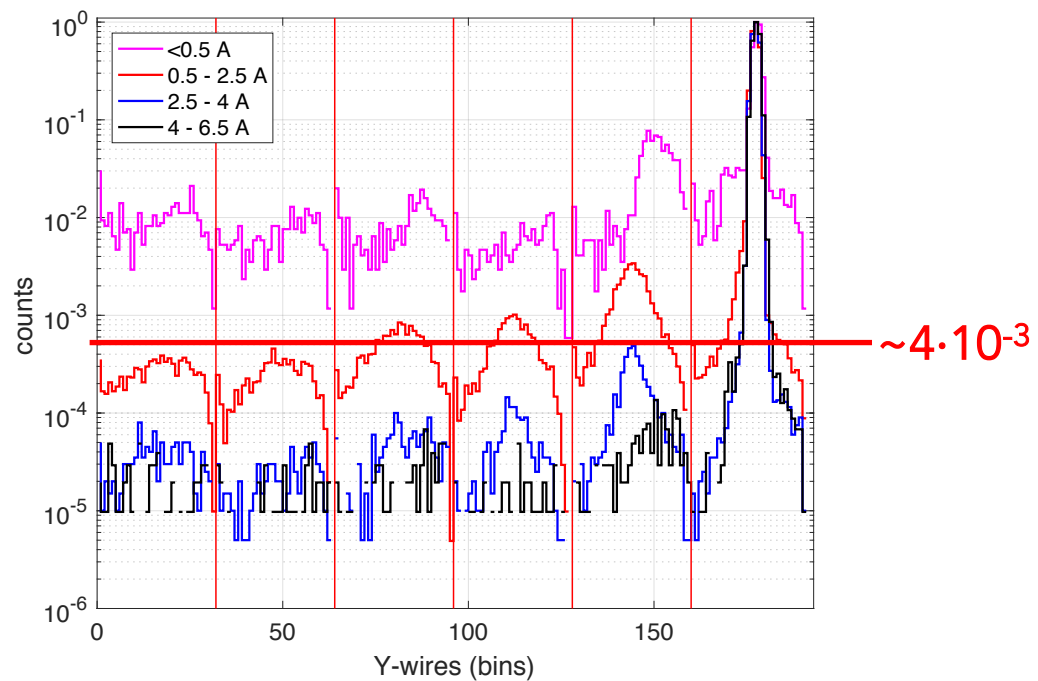
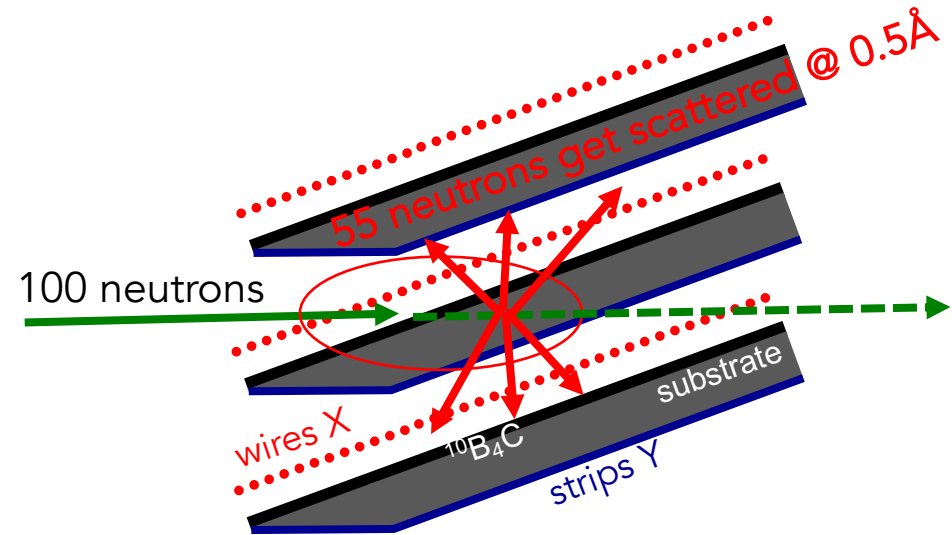
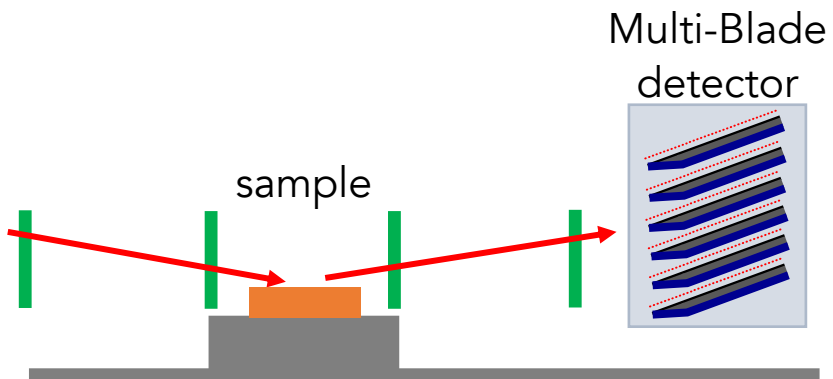
October 2017 - CRISP reflectometer @ ISIS

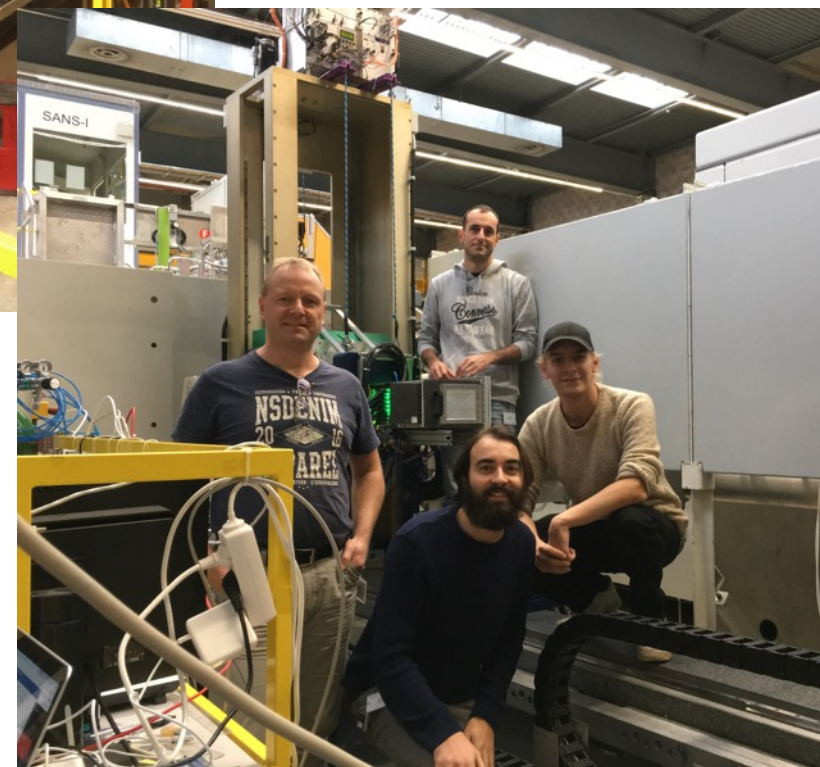
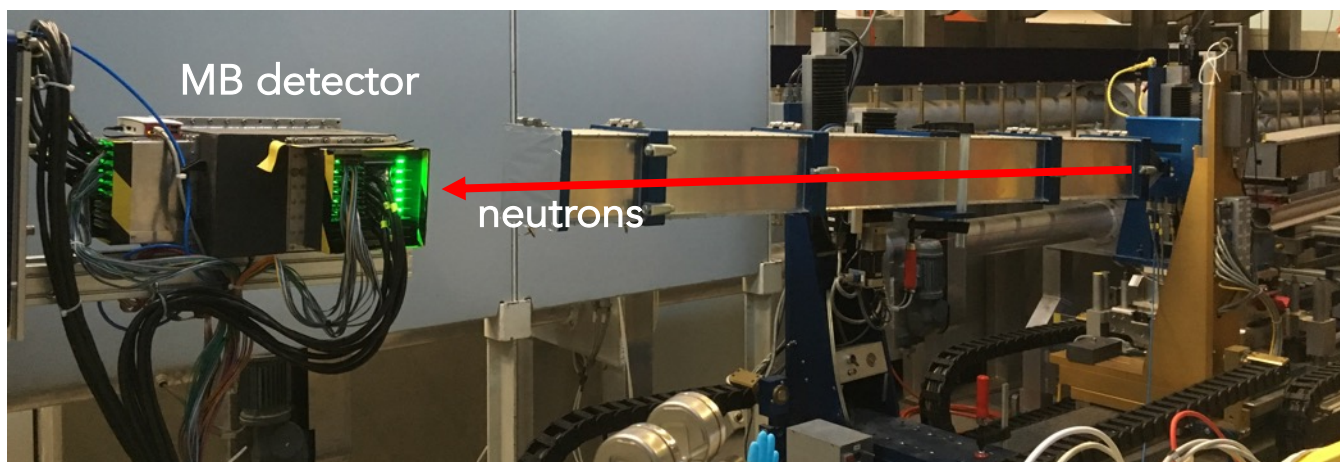
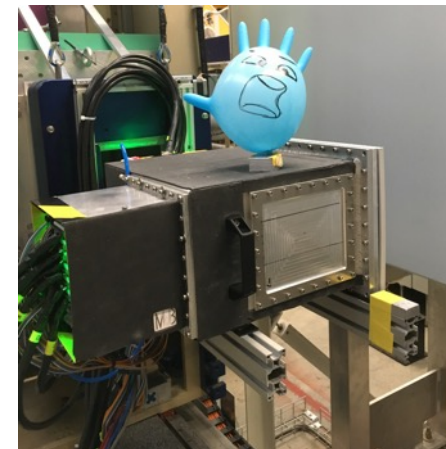
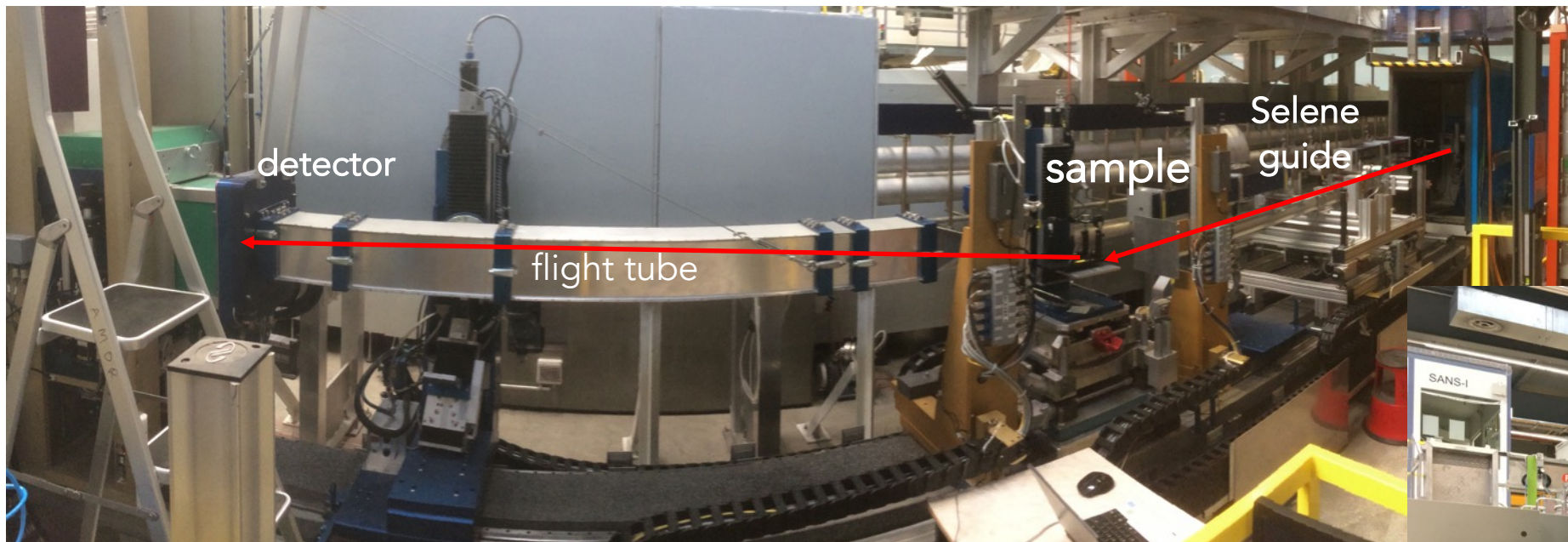


Gating between 6ms and 20ms

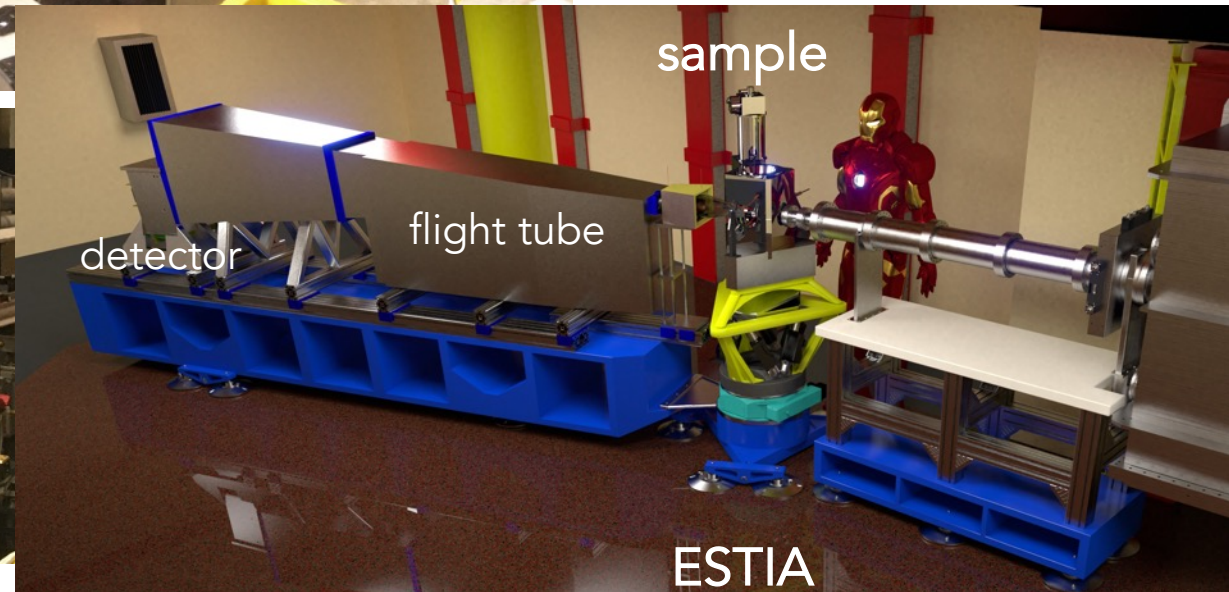
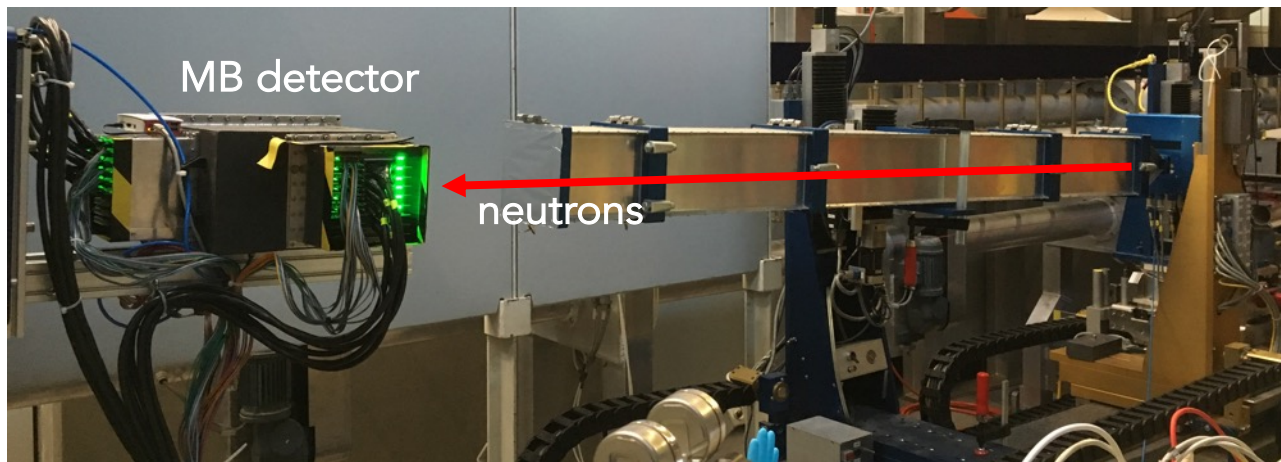
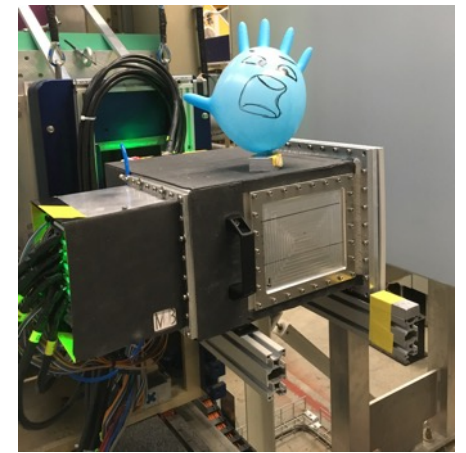
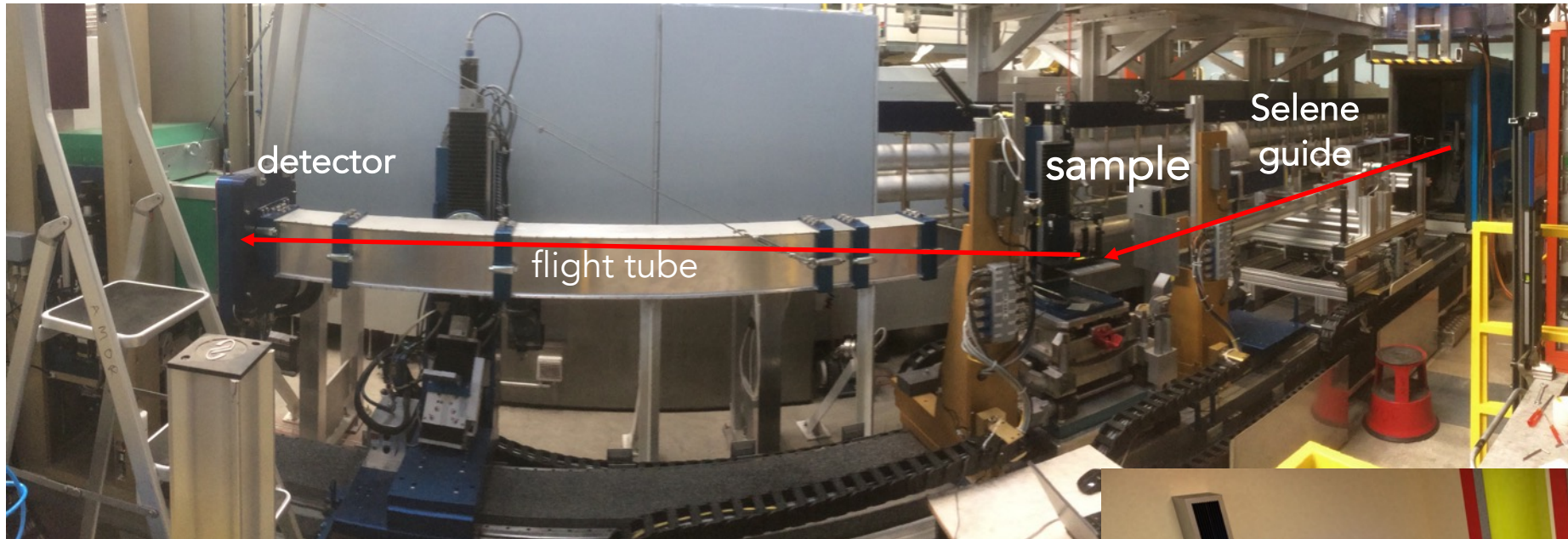


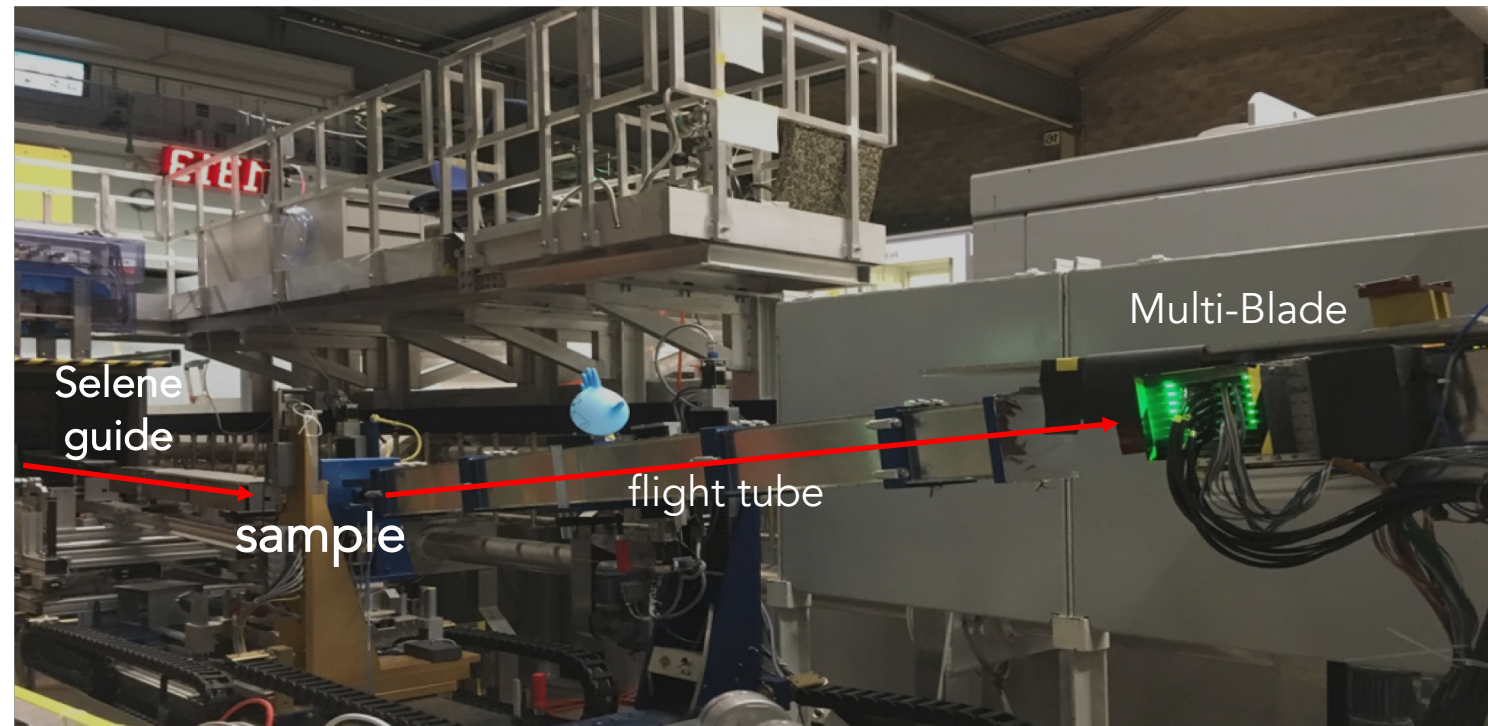
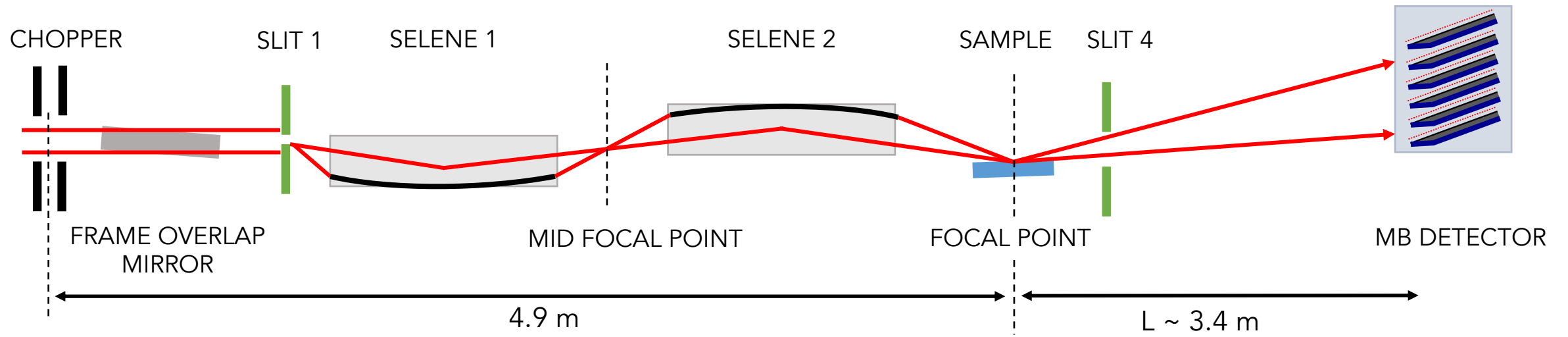


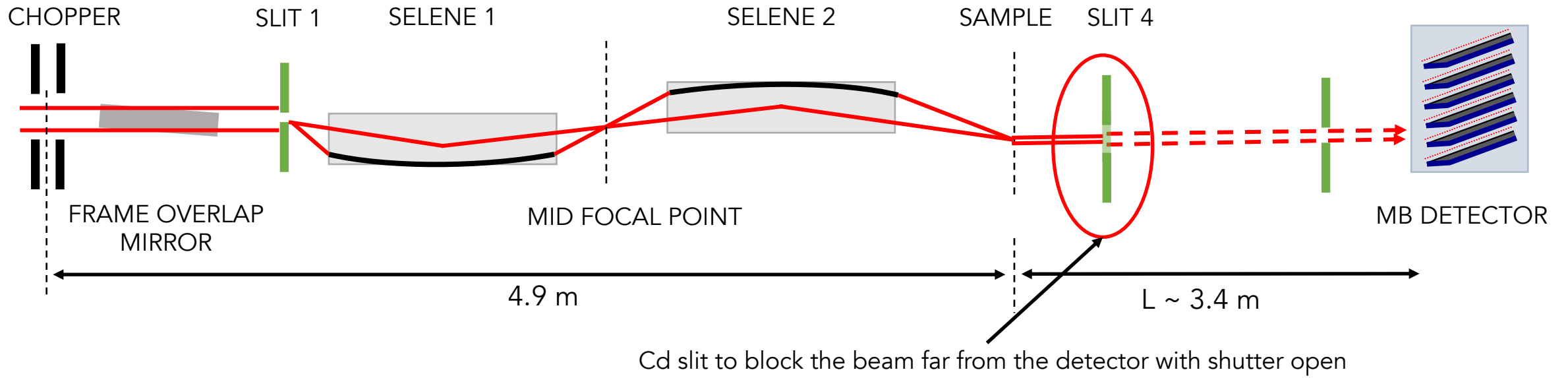




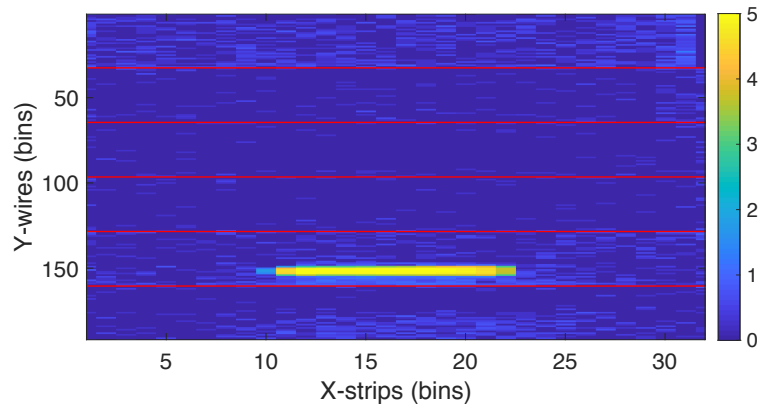
WP5 Data Acquisition software chain has been developed during BrightnESS (WP5 - i.e. DMSC/Data) and tested @ AMOR



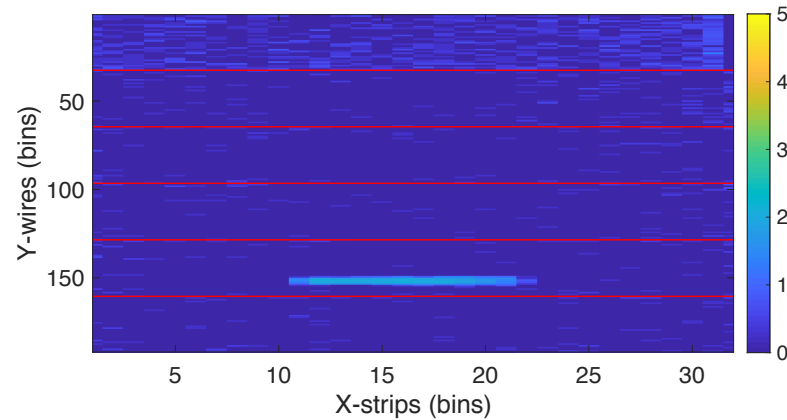


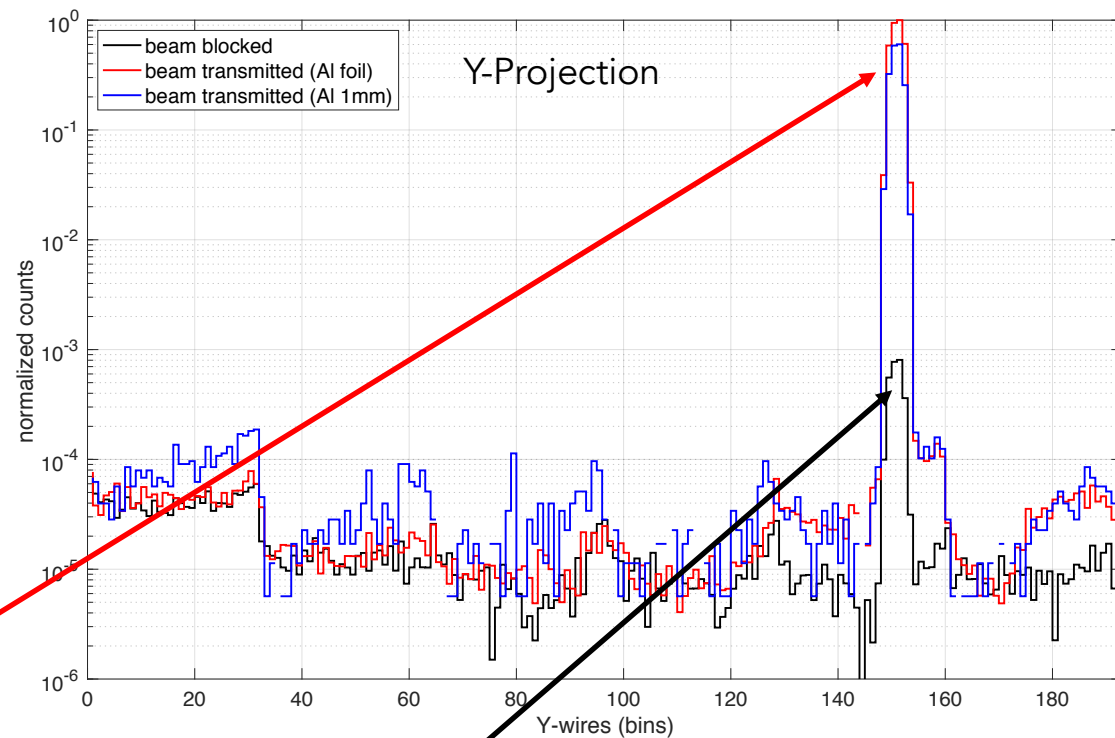


Beam transmitted

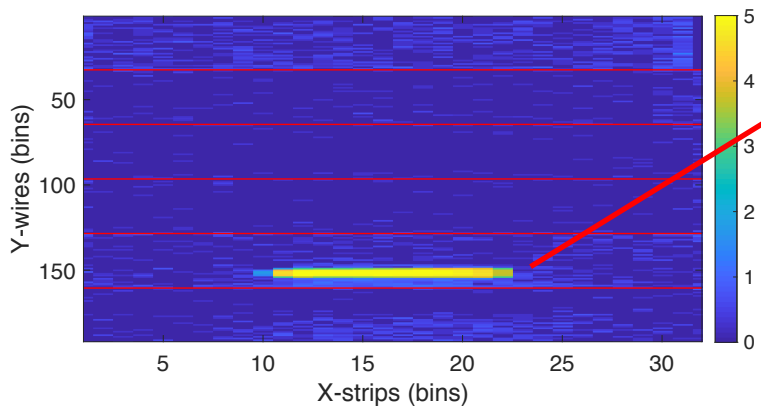


Beam blocked

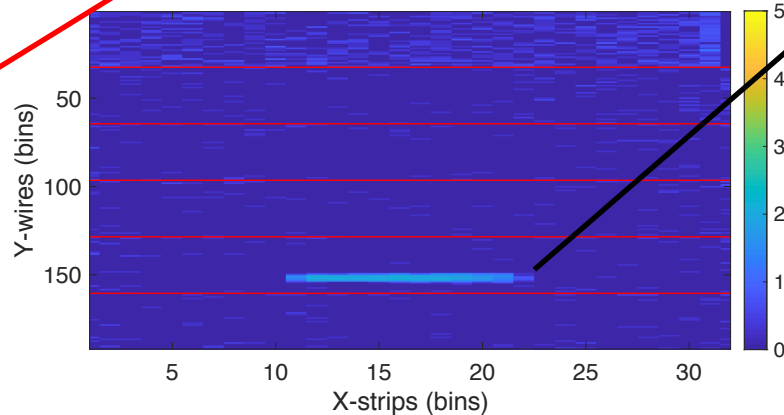




Beam transmitted

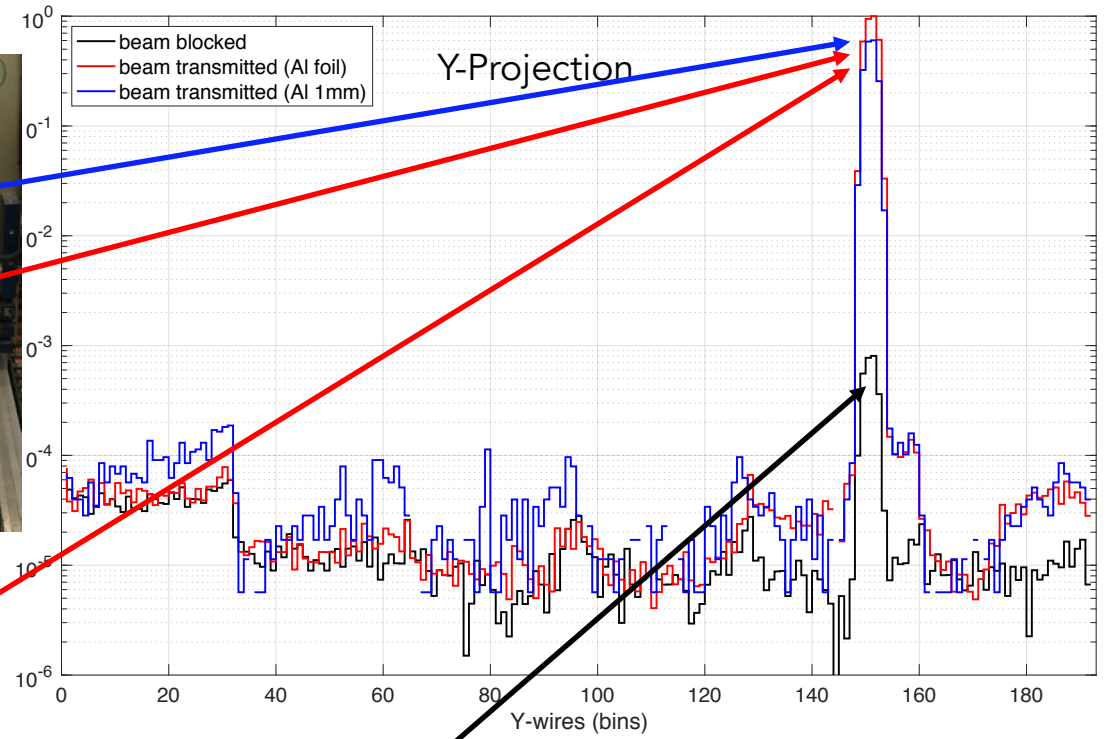
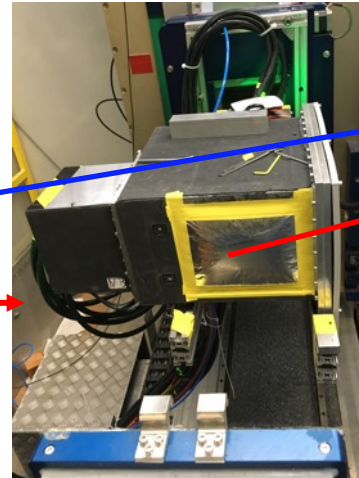
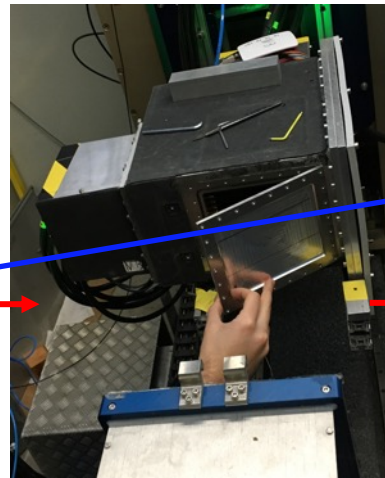
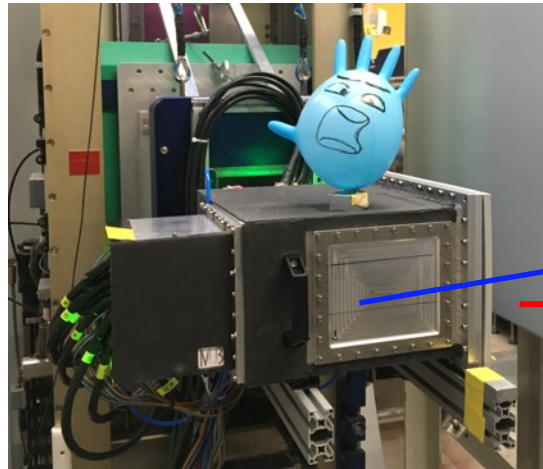


Beam blocked



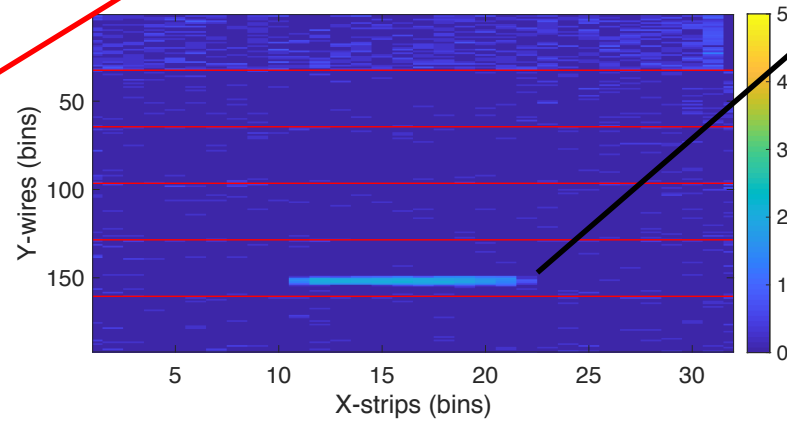
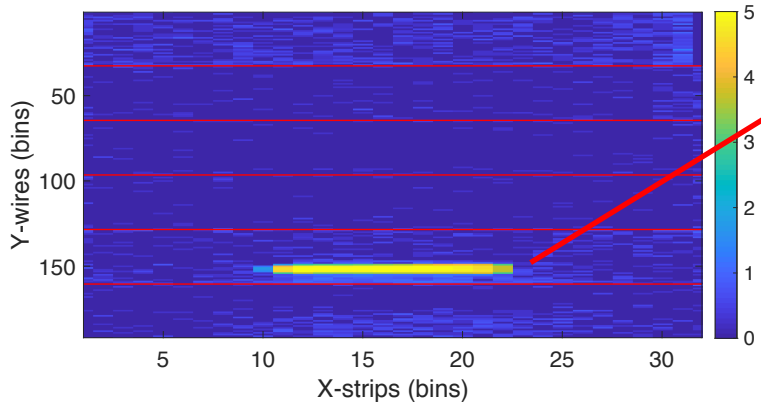
1mm Al window

50um Al foil

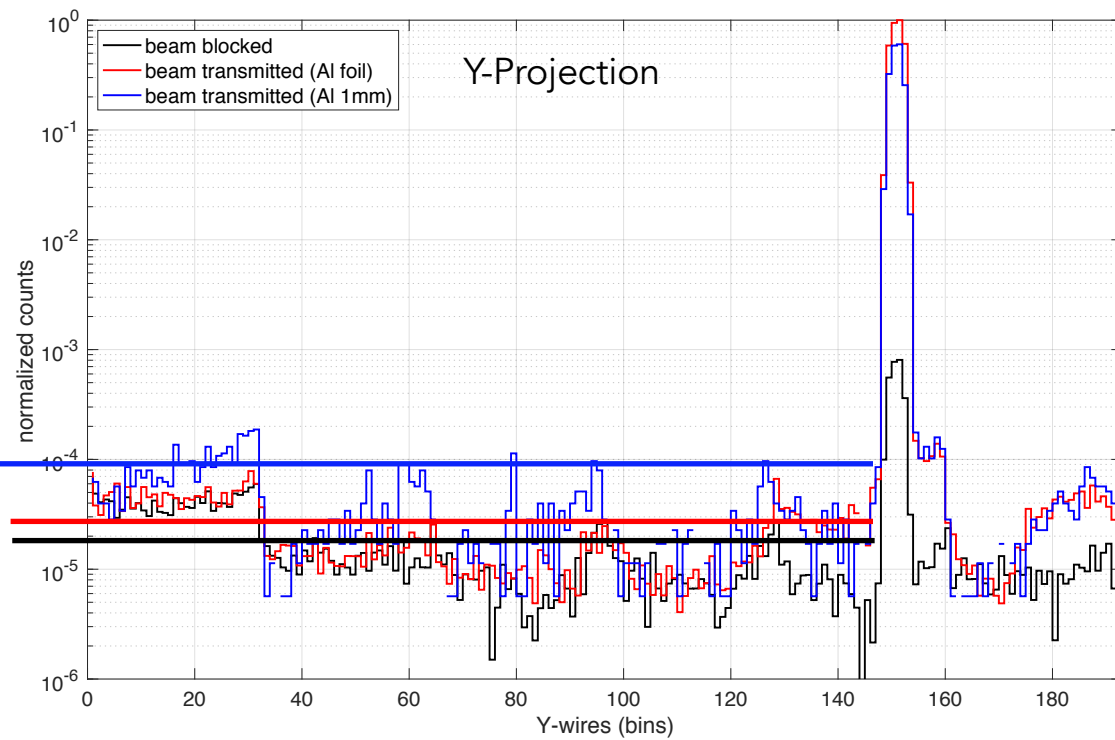


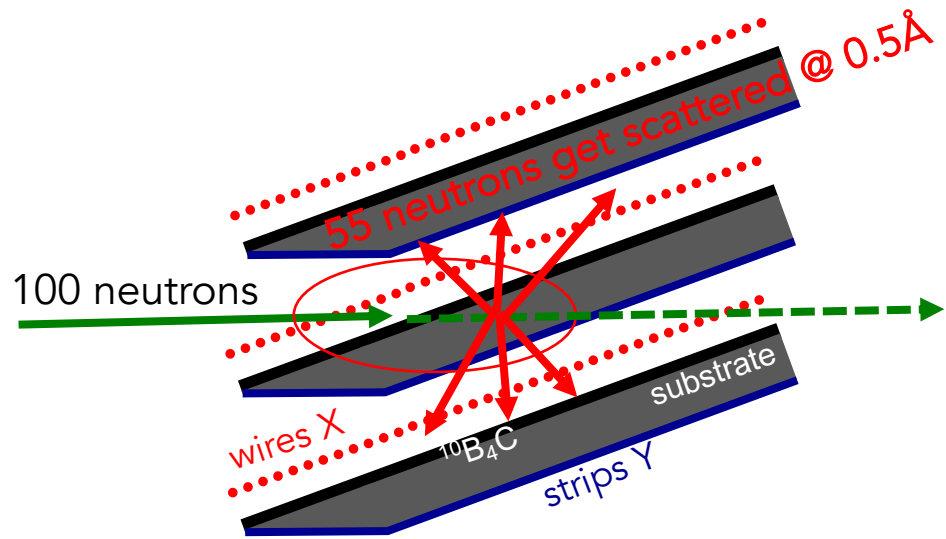
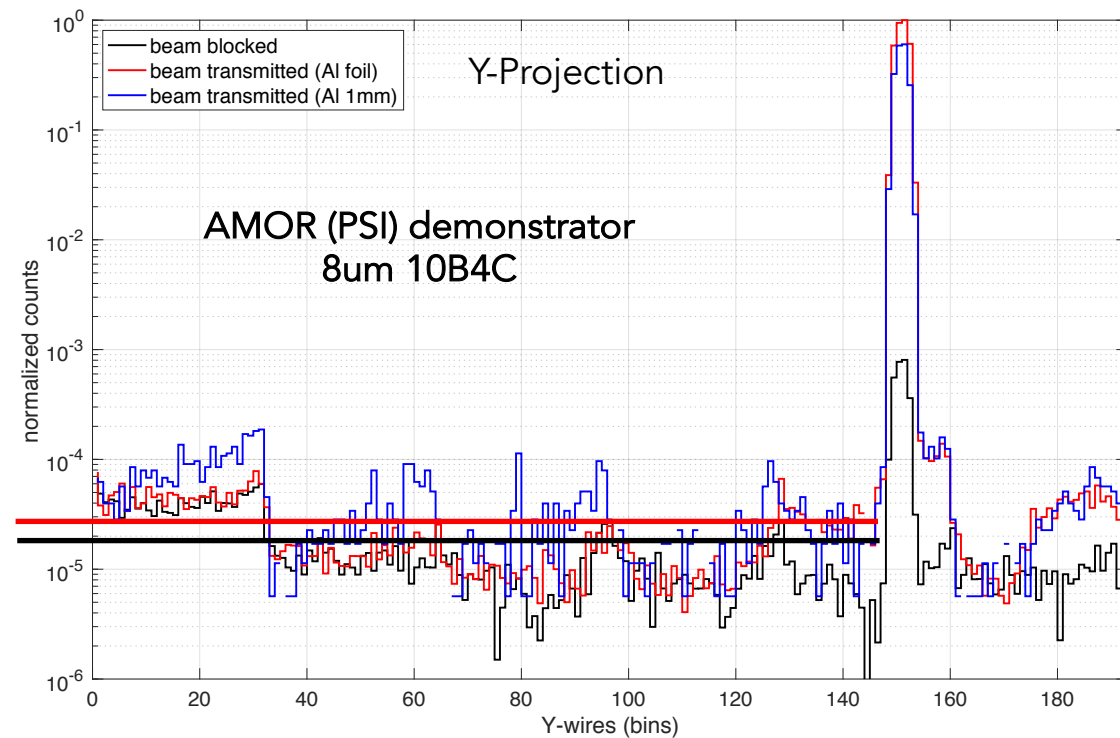
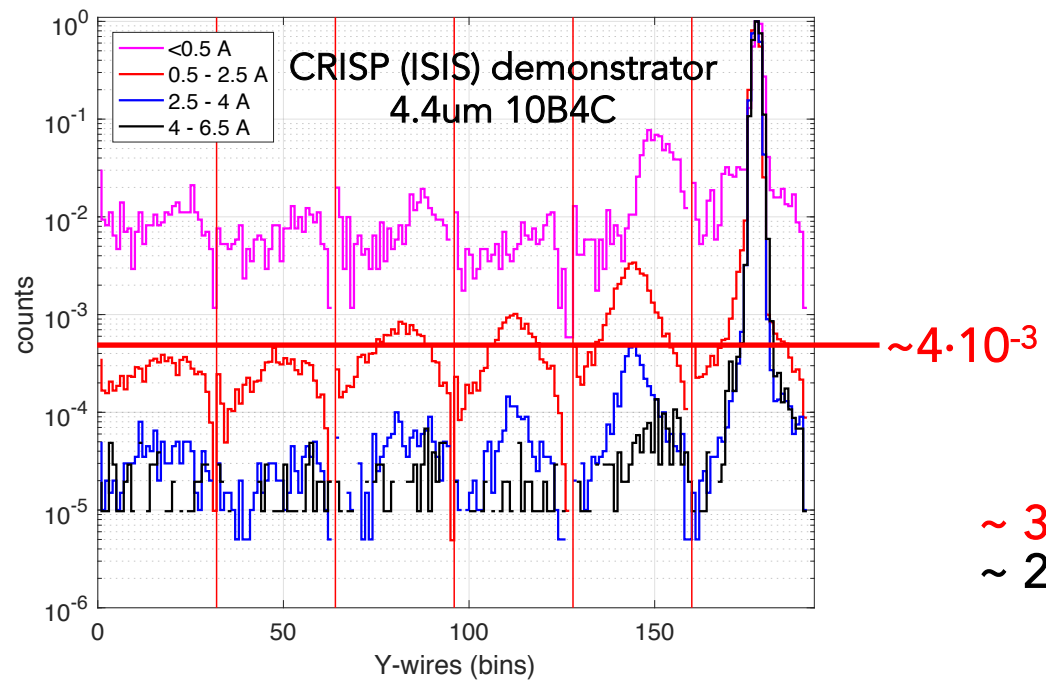
Beam transmitted

Beam blocked



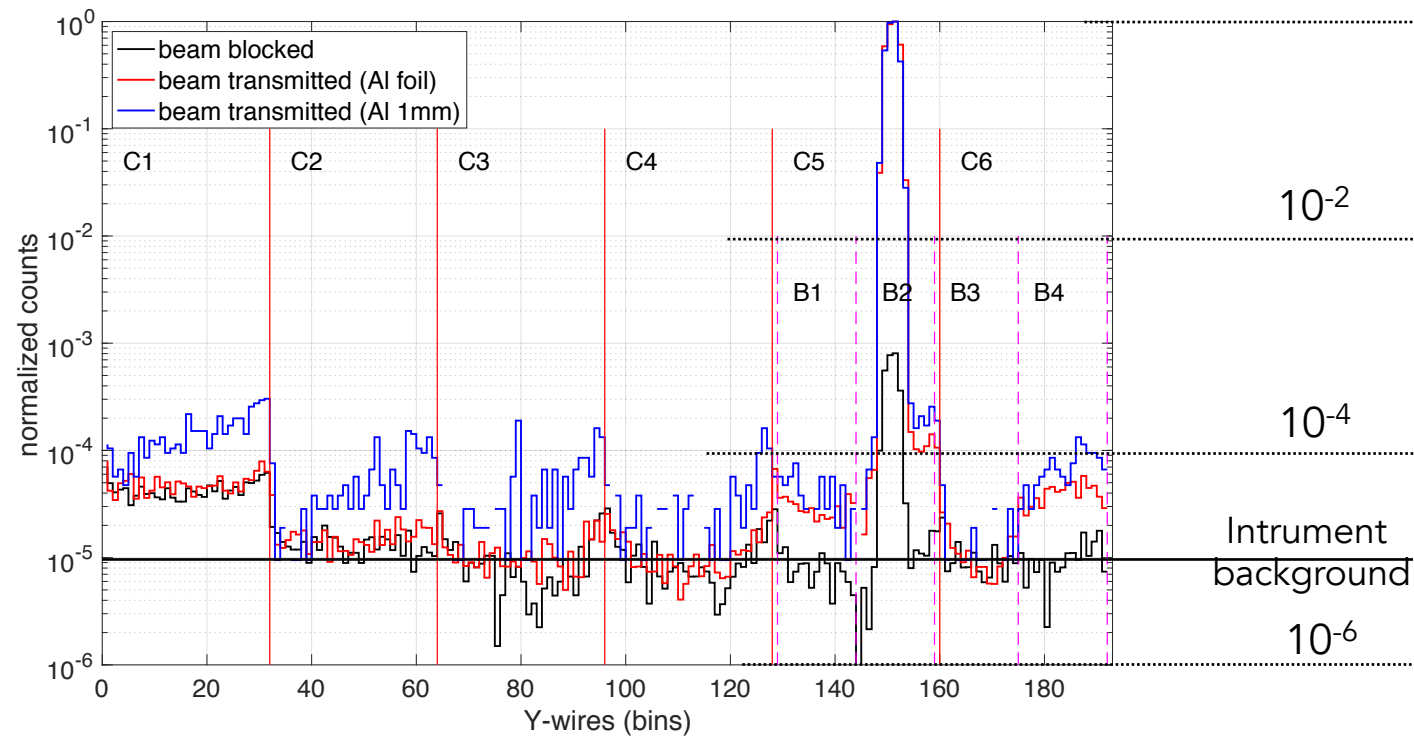
1mm window $\sim 10^{-4}$
Foil window $\sim 3 \cdot 10^{-5}$
Instrument background $\sim 2 \cdot 10^{-5}$



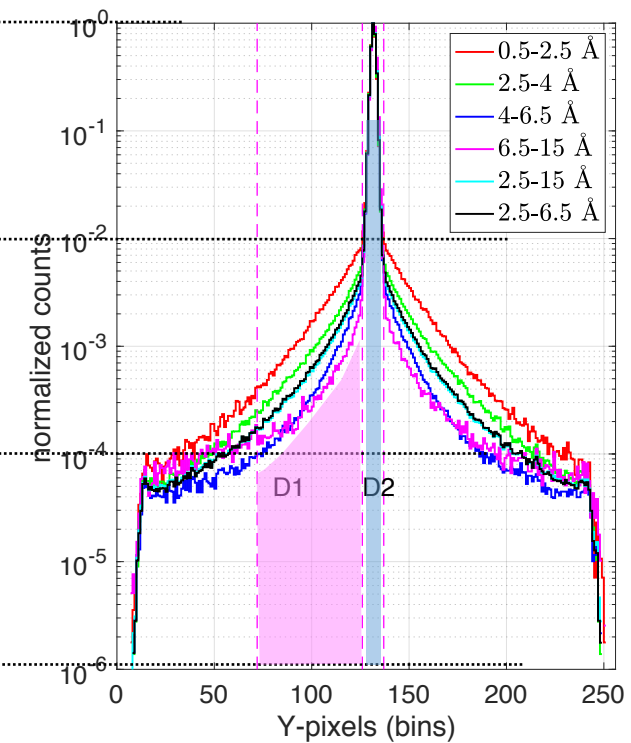


Problem Solved!

Multi-Blade - B10

10⁰

He3 – AMOR detector



$$FoM = \frac{D2}{D1}$$

(normalized to same area)

FoM ($\times 10^4$)	wavelength range (\AA)					
	0.5-2.5	2.5-4	4-6.5	2.5-6.5	6.5-15	2.5-15
MB@AMOR ($>7.5\mu\text{m}$ coating)						
1mm Al window	n/a	7.3	8.0	7.4	7.2	7.3
Al foil window	n/a	22.6	47.7	40.5	11.3	21.1
He3@AMOR	0.1	0.2	0.5	0.3	0.8	0.4

Foil ~ x3 ~ 1mm window ~ x20 ~ He3
 Foil ~ x60 ~ He3

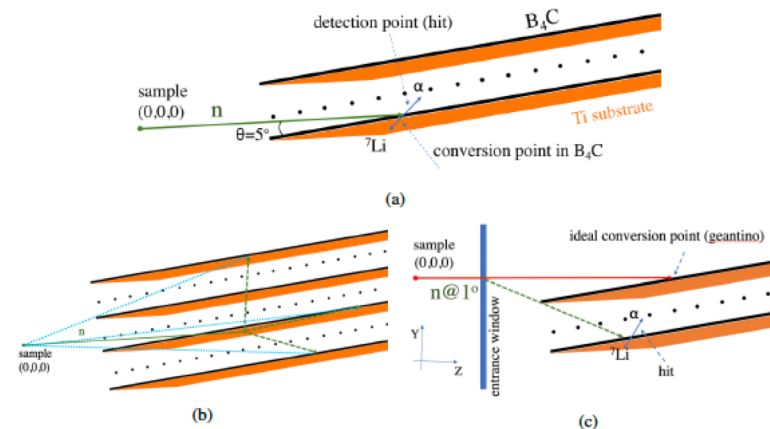


Figure 4. (a) Difference between conversion and detection point. (b) A neutron traversing the first converter layer (solid green) can scatter in the blade material and finally get converted away from the first crossing point (dashed green). This leads to the miscalculation of the distance between sample and detection point (dashed blue). (c) Similarly for a scattered neutron on the detector window. The latter is 1° inclined with respect to the vertical axis. The projection of the detection point on the converter layer is not displayed here for view simplification.

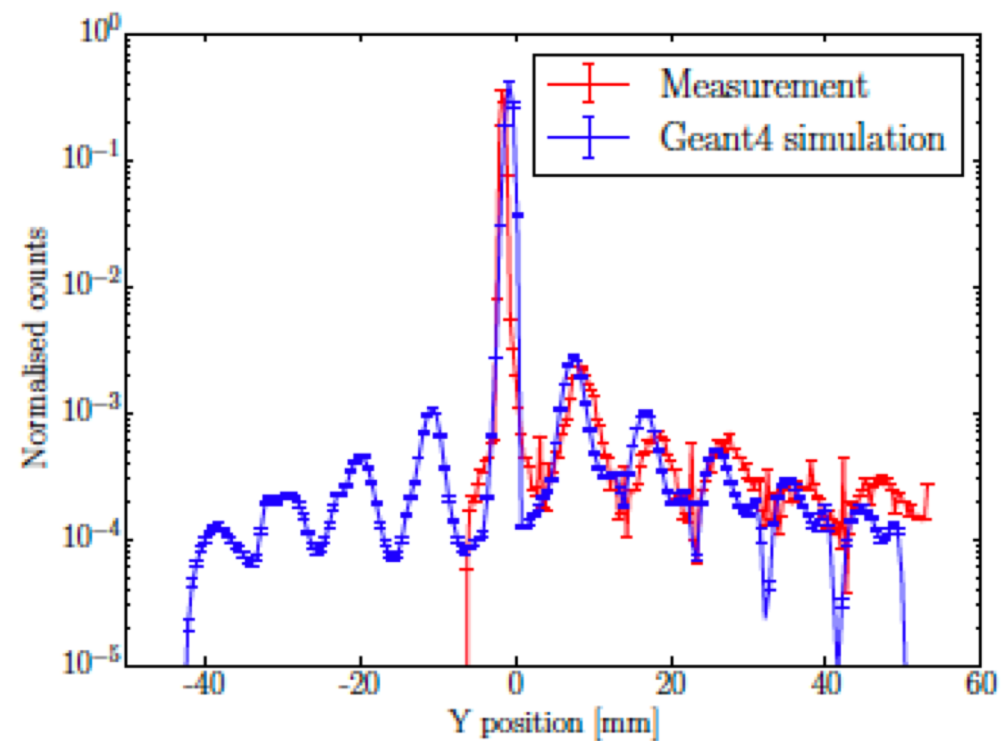


Figure 6. Comparison of Y-position of detected neutrons in measurements taken at CRISP [33] and the results of the simulation (this work). Reproduced from [33]. CC BY 4.0.

FAST NEUTRON BACKGROUND

THEORY

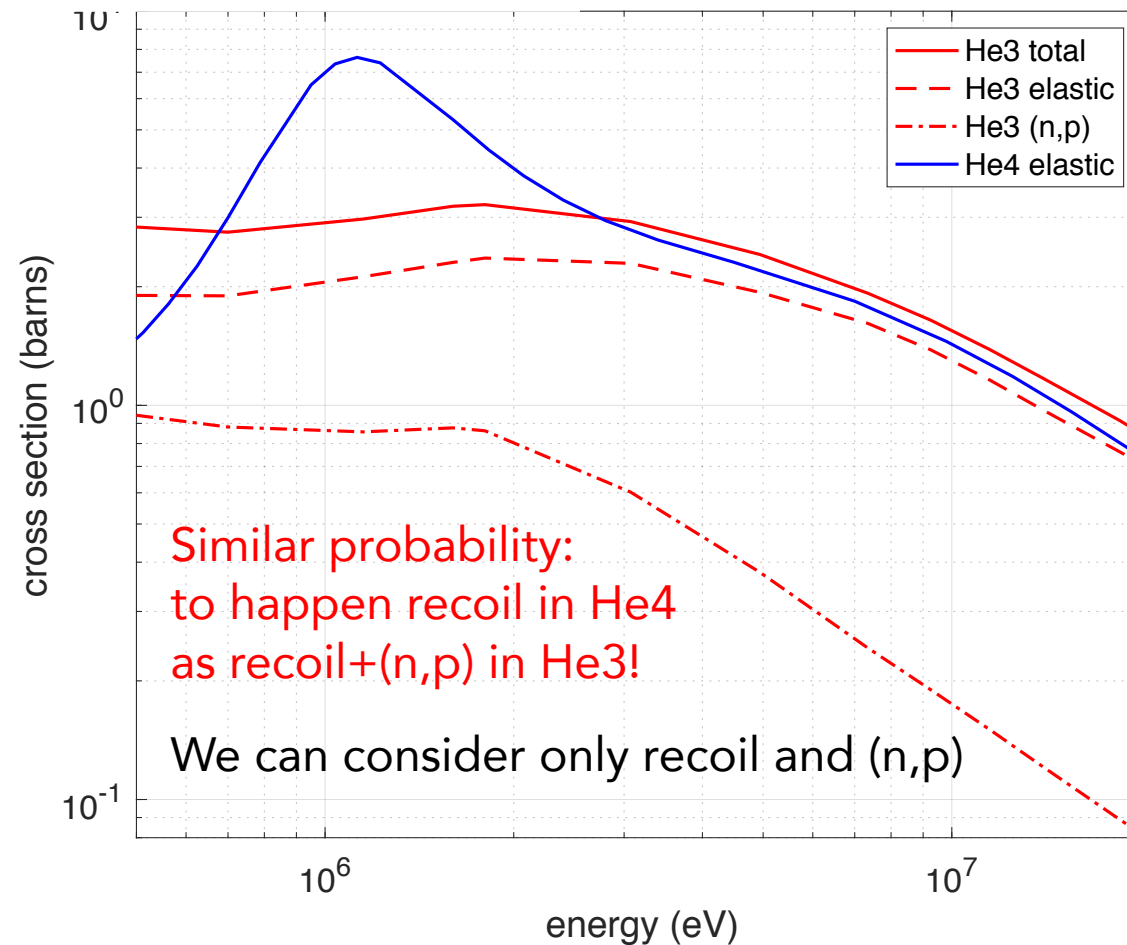
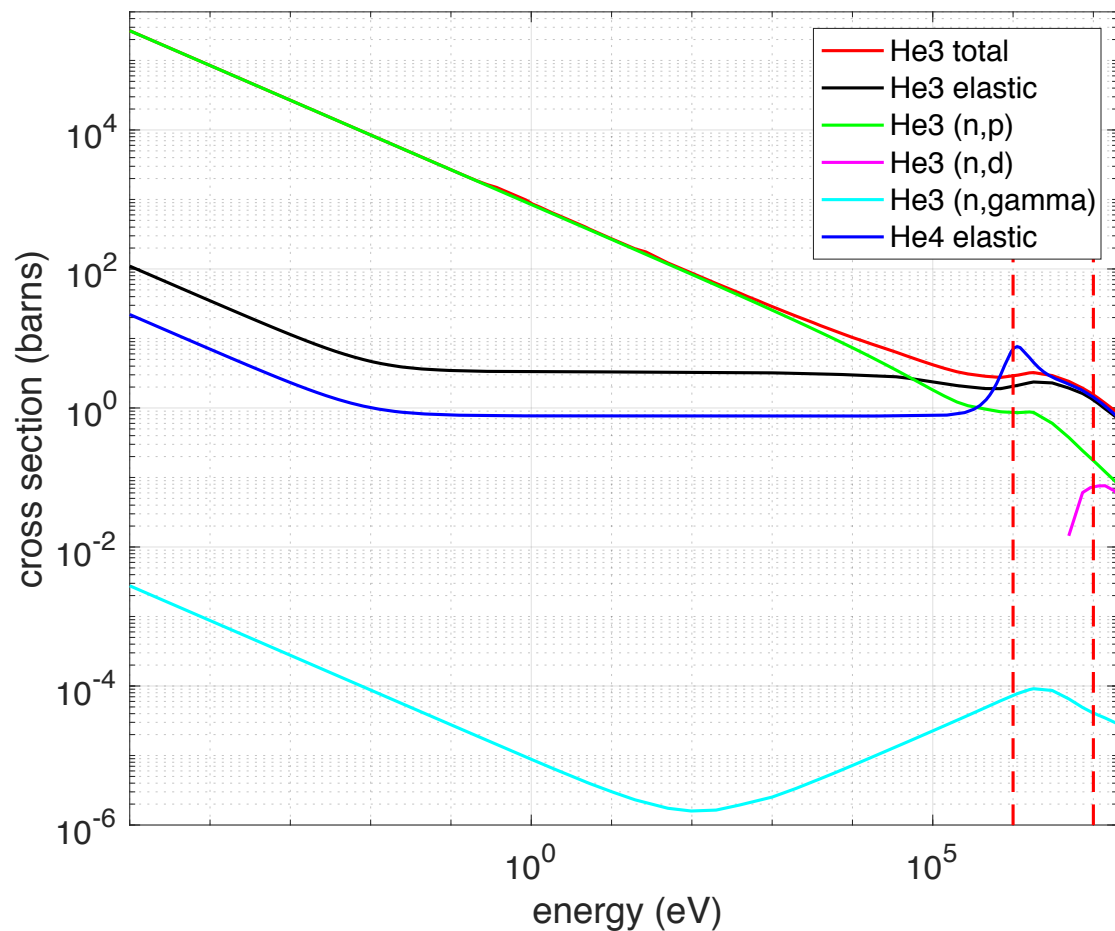
Cross-sections

- ^3He

- ^3He recoil (elastic scattering): $^3\text{He} + n \rightarrow ^3\text{He}' + n' \quad (E_{Rmax} = 0.75 E_i)$
- (n, p) : $^3\text{He} + n \rightarrow ^1\text{H} + ^3\text{H} + 0.764 \text{ MeV}$
- (n, d) : $^3\text{He} + n \rightarrow ^2\text{H} + ^2\text{H} - 3.27 \text{ MeV}$
- (n, γ) γ -rays

- ^4He

- ^4He recoil (elastic scattering): $^4\text{He} + n \rightarrow ^4\text{He}' + n' \quad (E_{Rmax} = 0.64 E_i)$

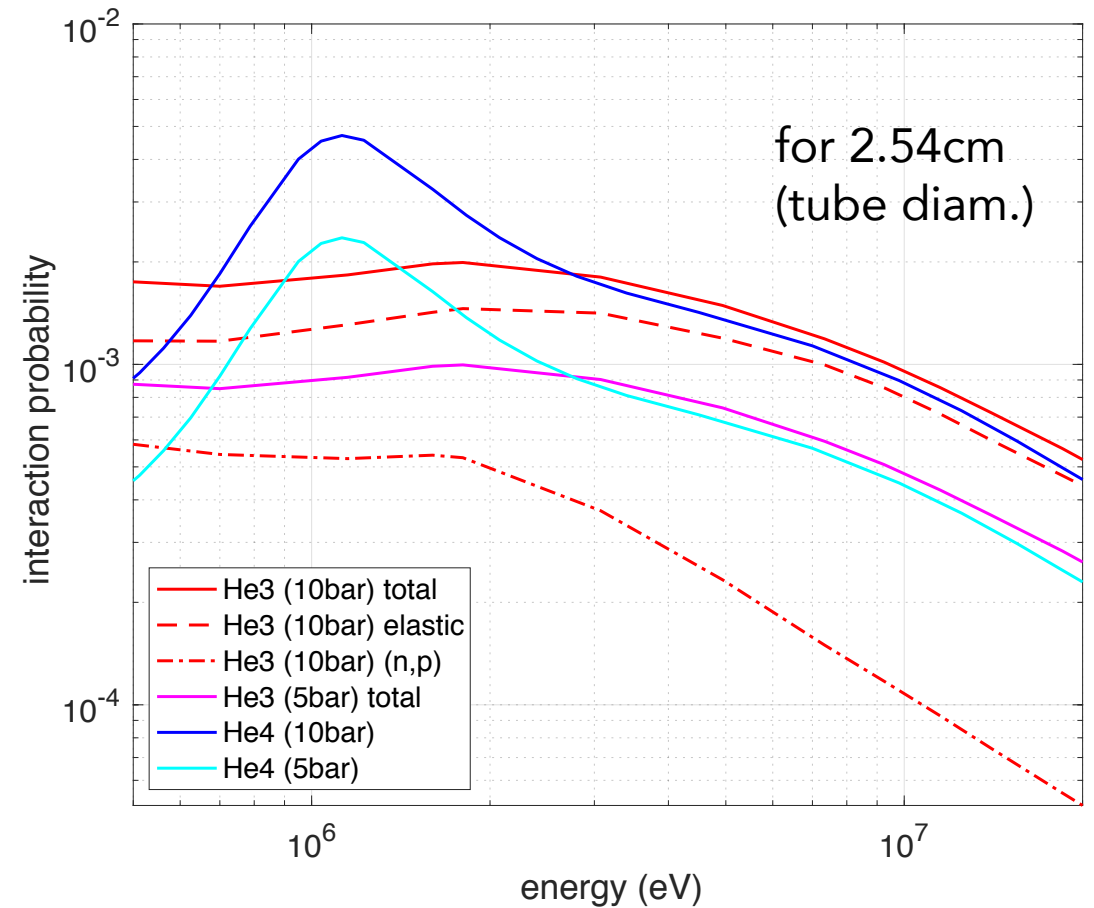
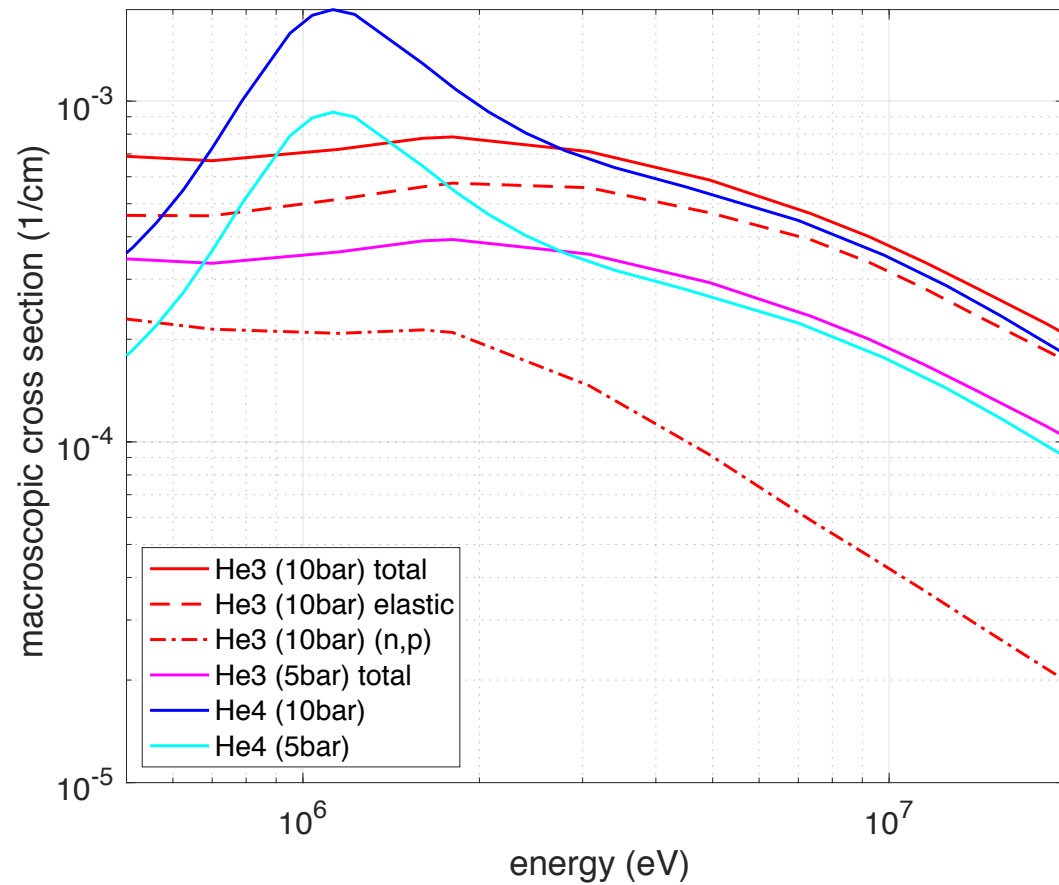


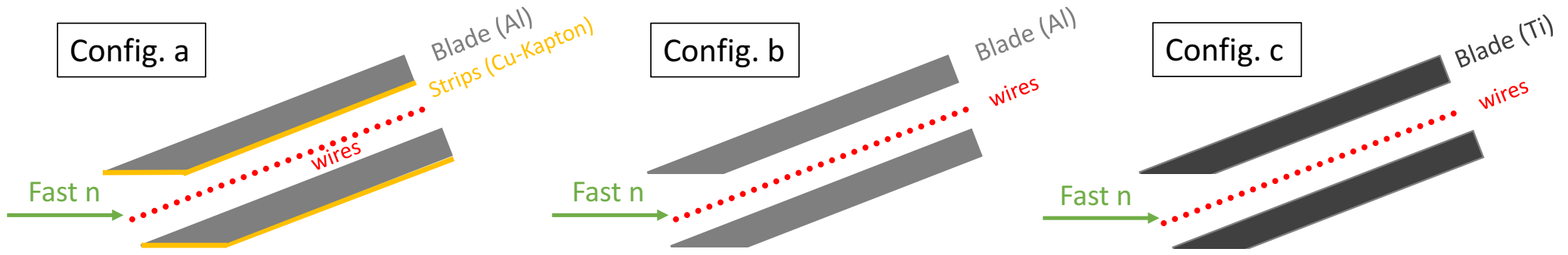
THEORY

Same number density for He3 and He4 $\rightarrow n = 2.43 \times 10^{19} \text{ 1/cm}^3 @ 1 \text{ bar}$

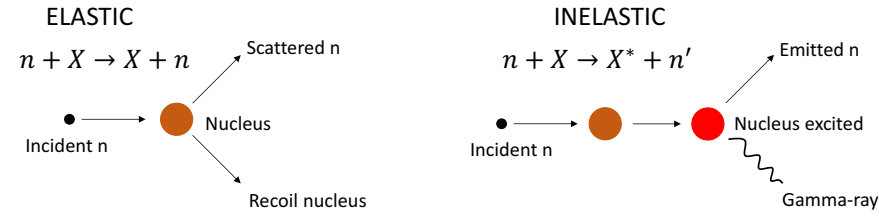
Macro cross-sect
And Probability

Mass density: He4 0.00016 g/cm³ and He3 0.00012 g/cm³ @ 1bar

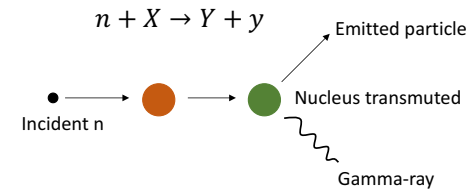


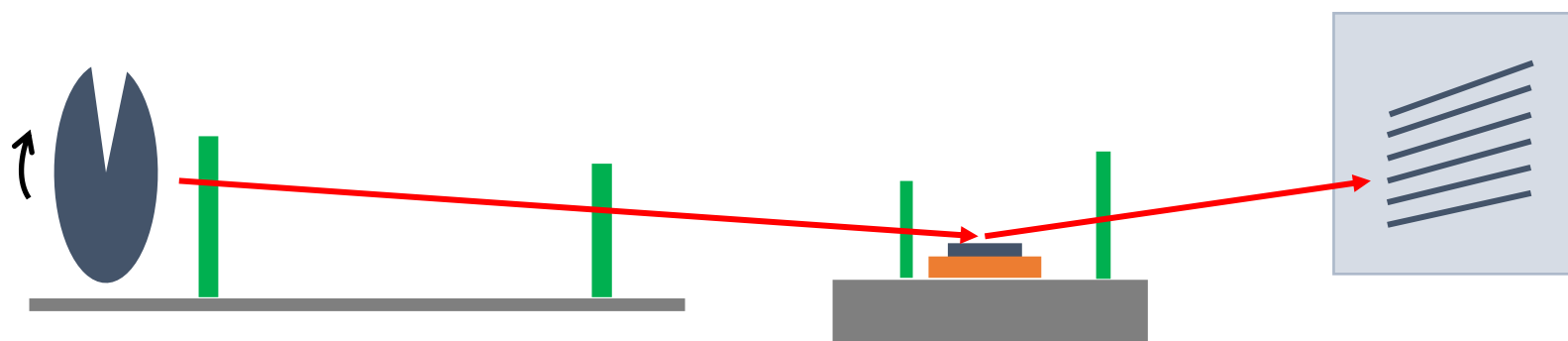


SCATTERING INTERACTION

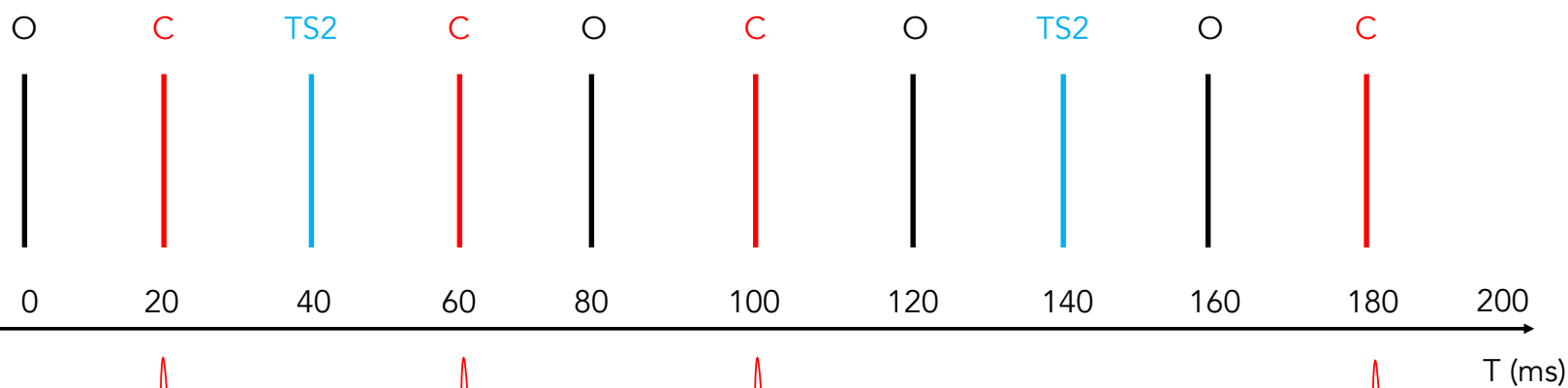


ABSORPTION INTERACTION



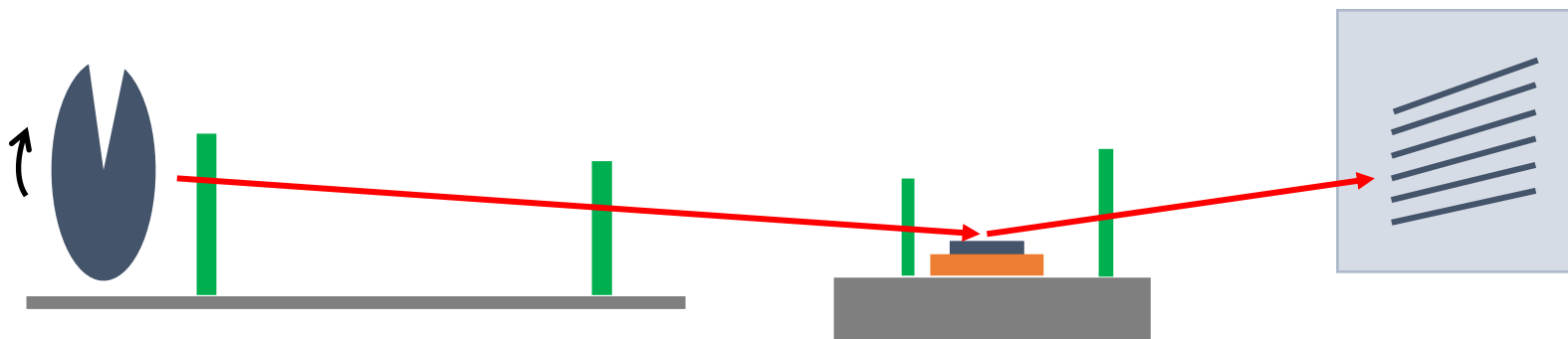


25Hz



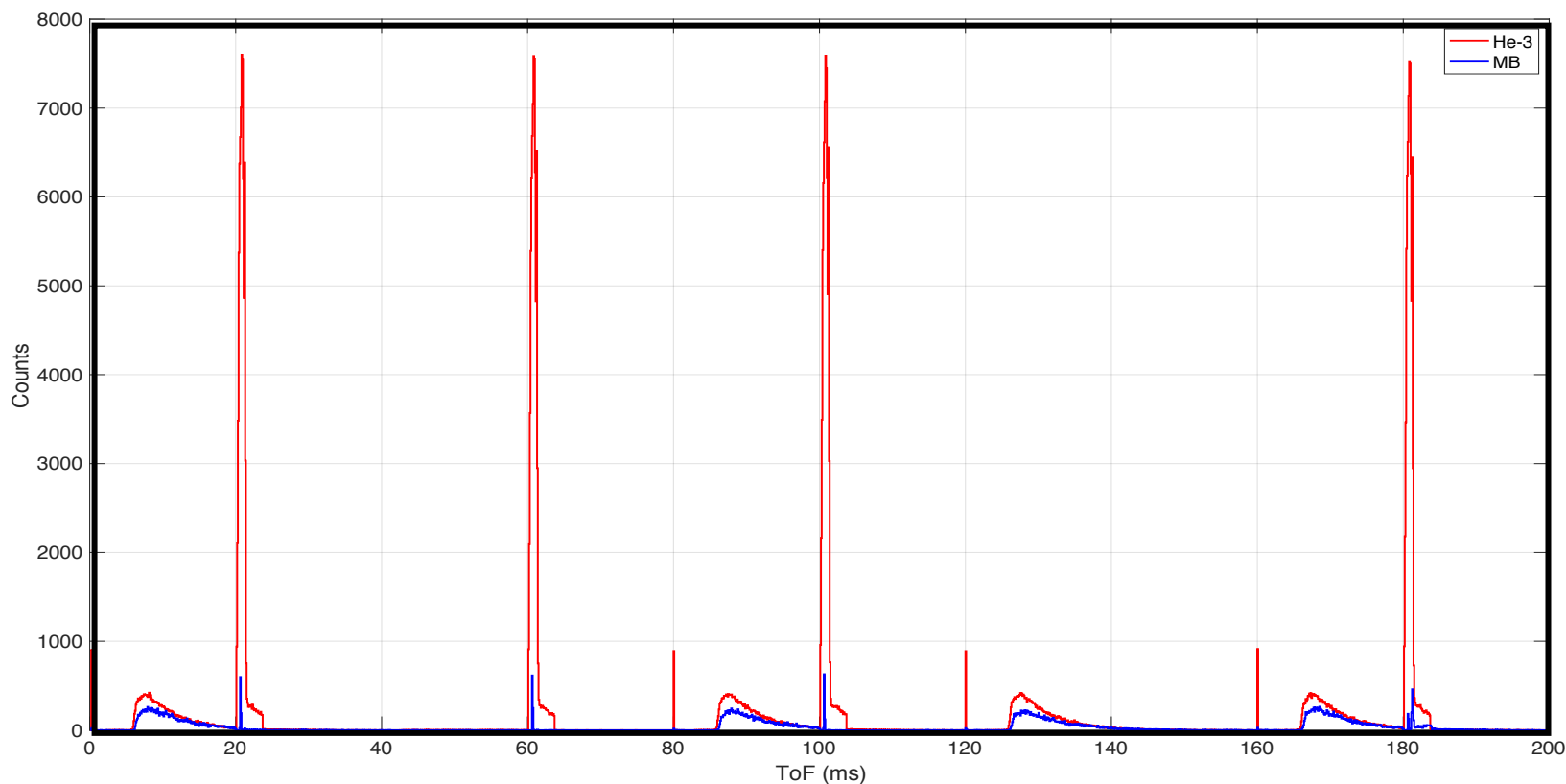
Spectrum sketch scenarios:

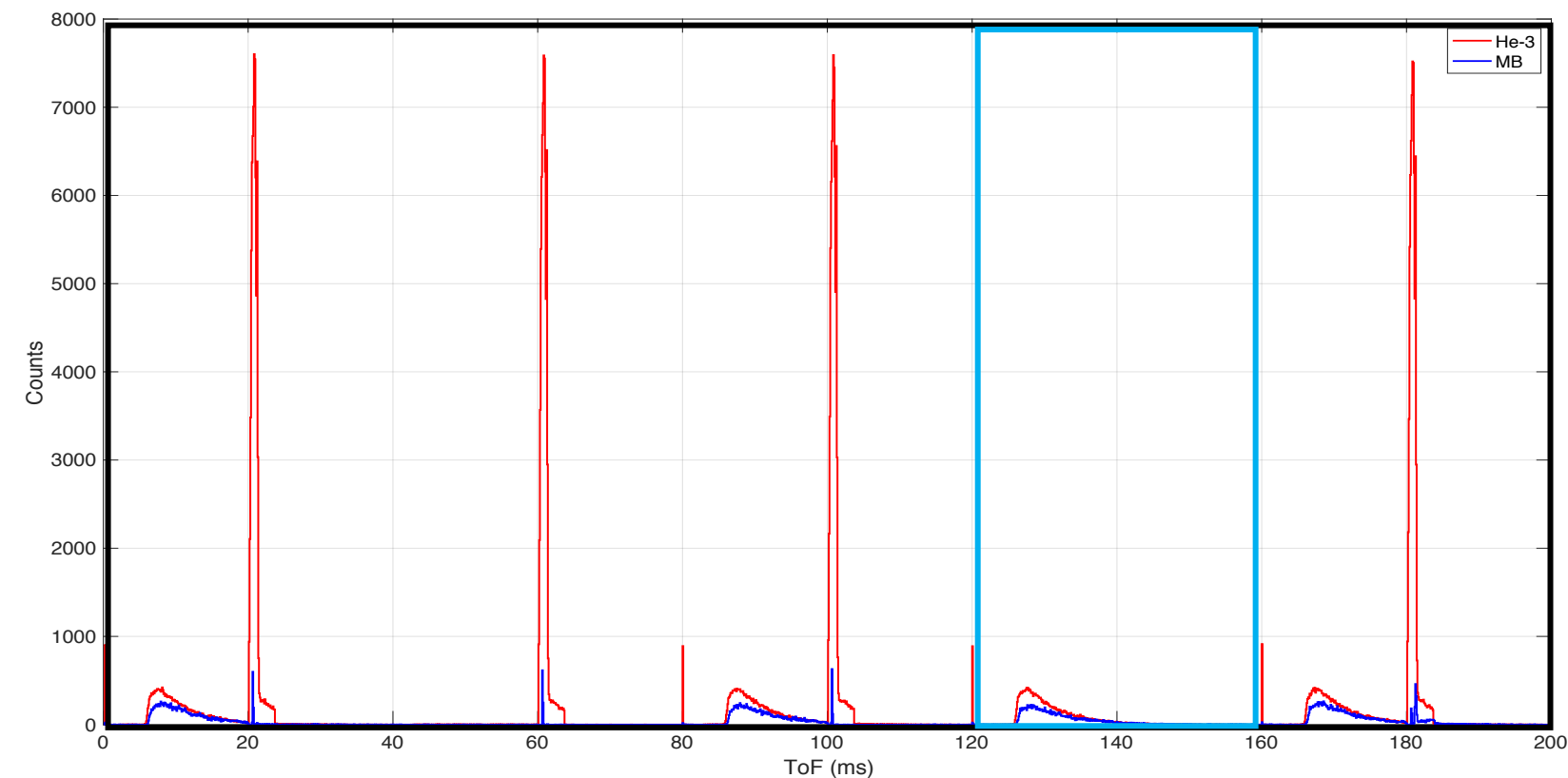
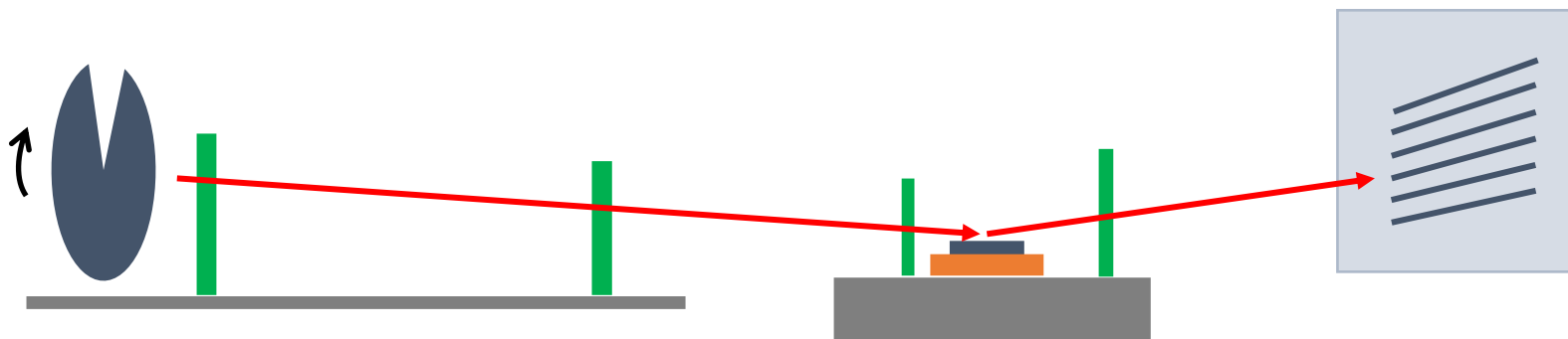
1. O: chopper in phase with the proton pulse. Beam passes through the chopper
2. C: chopper not in phase with the proton pulse. Neutron beam hits the chopper
3. TS2: 1 every 5 pulse is sent to Target Station 2



Spectrum region of interest:

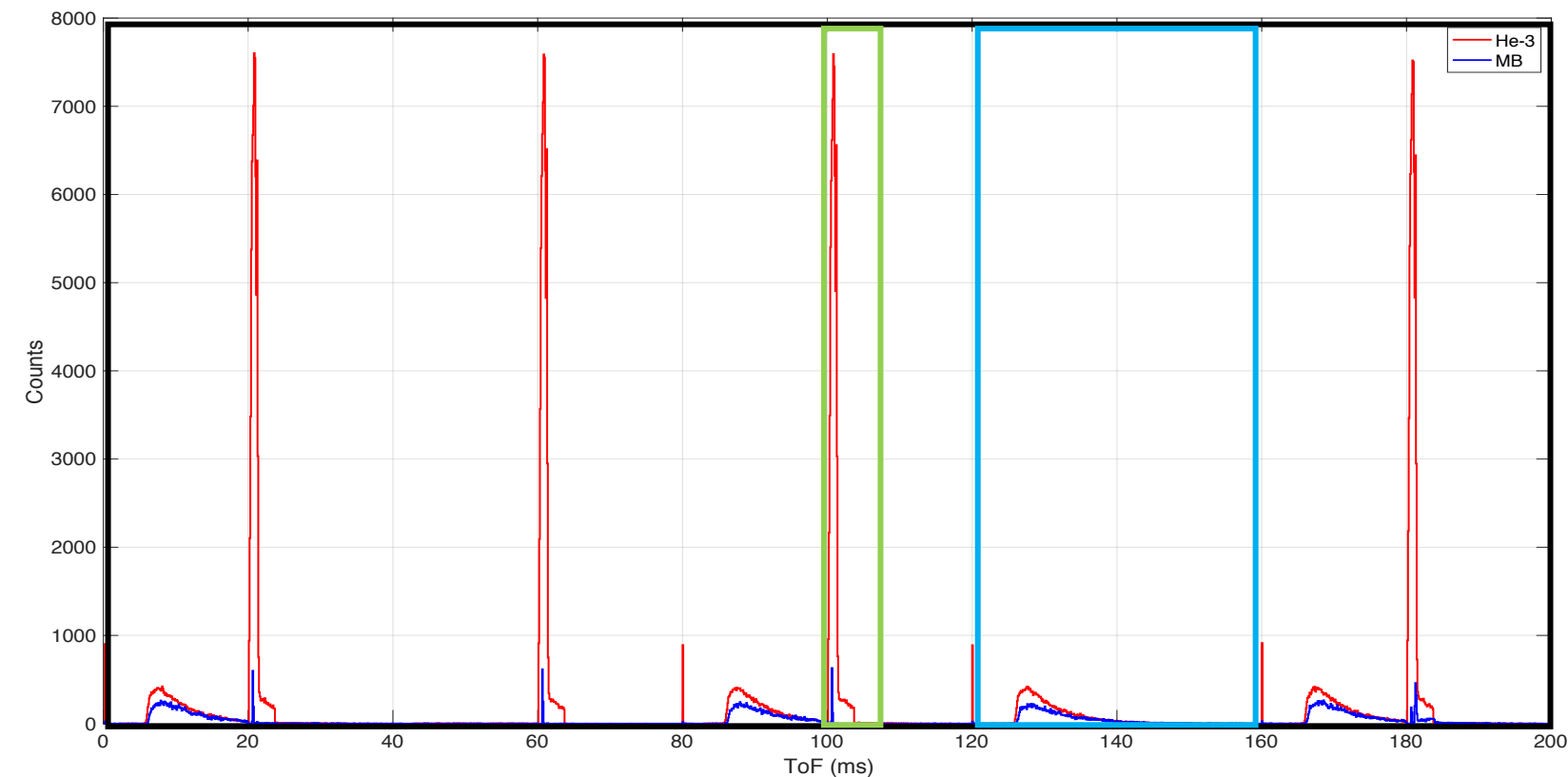
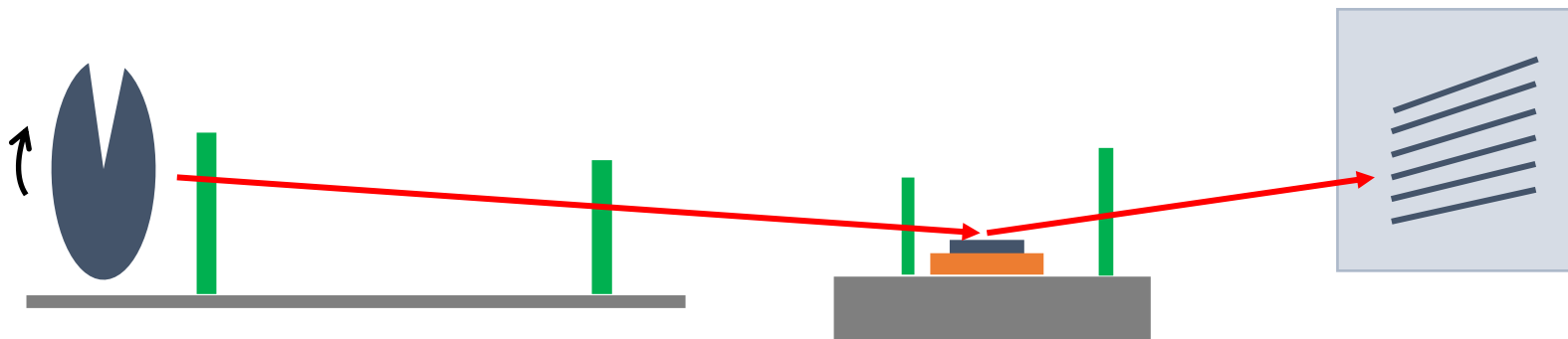
1. (T): Φ total flux integrated in the full spectrum $t = 0-200$ ms
2. (O): Φ_{tn} thermal flux integrated in $t = 120-160$ ms
3. (C): Φ_{p} background flux integrated in $t = 100-105$ ms





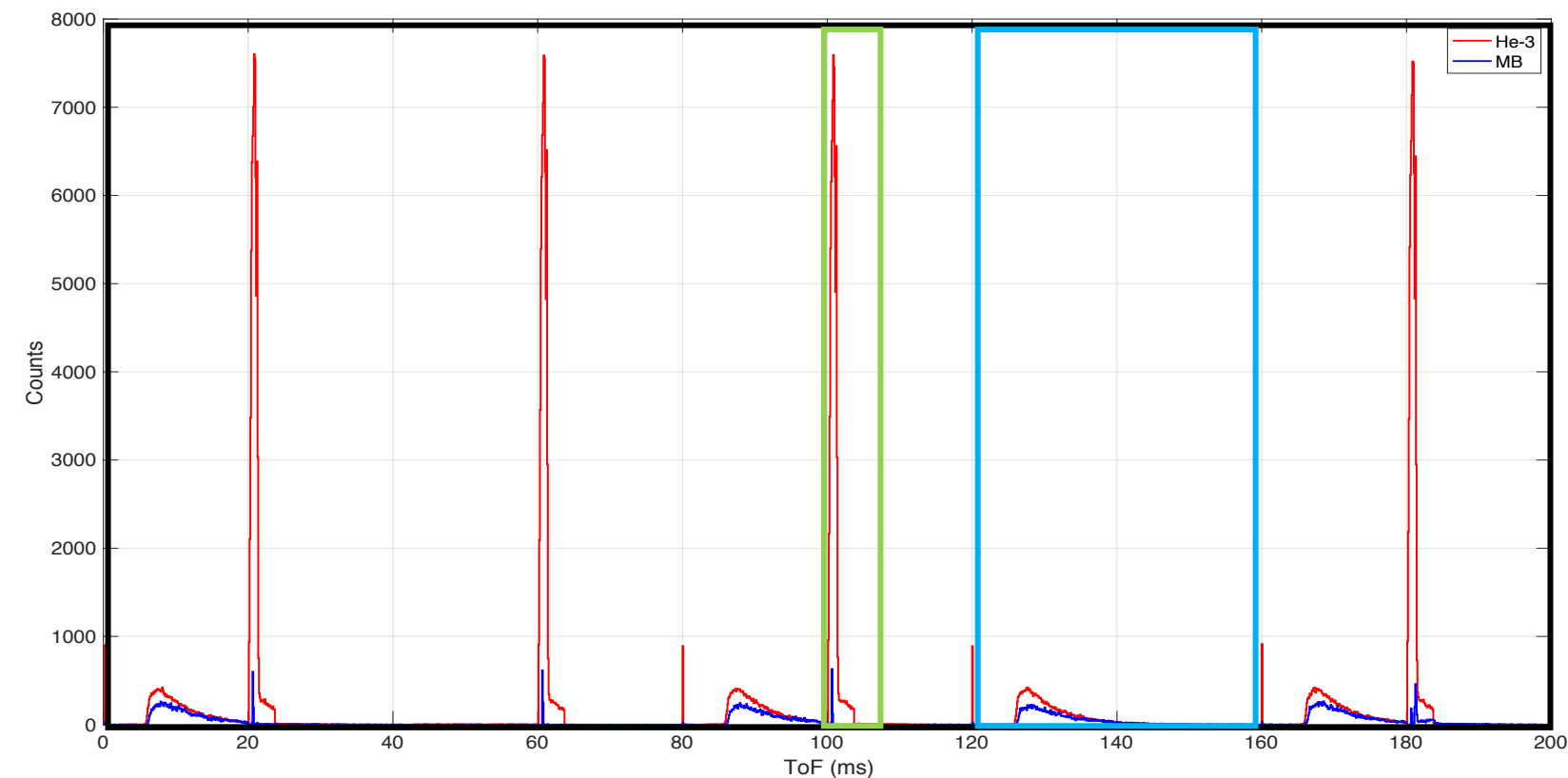
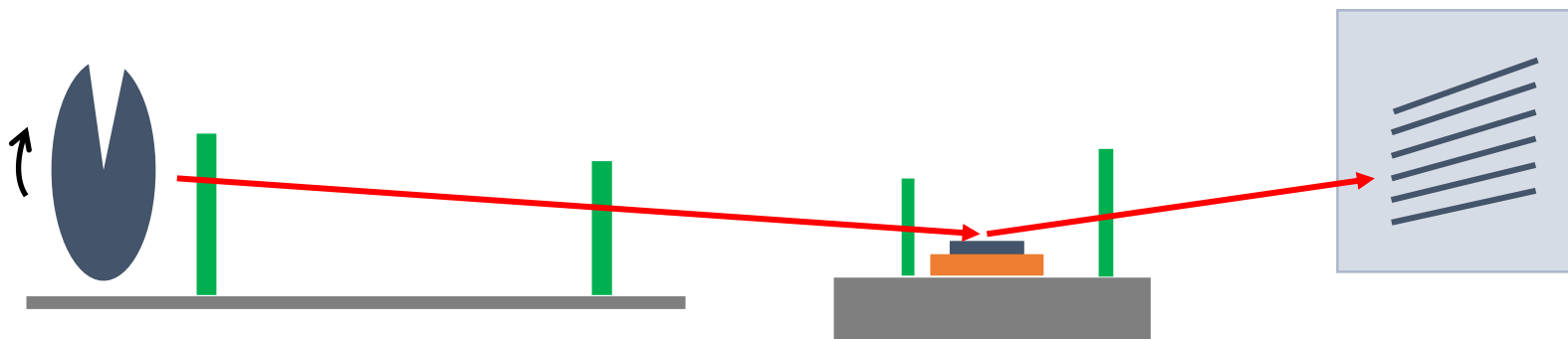
Spectrum region of interest:

1. (T): Φ total flux integrated in the full spectrum $t = 0-200$ ms
2. (O): Φ_{tn} thermal flux integrated in $t = 120-160$ ms
3. (C): Φ_p background flux integrated in $t = 100-105$ ms



Spectrum region of interest:

1. (T): Φ total flux integrated in the full spectrum $t = 0-200$ ms
2. (O): Φ_{tn} thermal flux integrated in $t = 120-160$ ms
3. (C): Φ_p background flux integrated in $t = 100-105$ ms



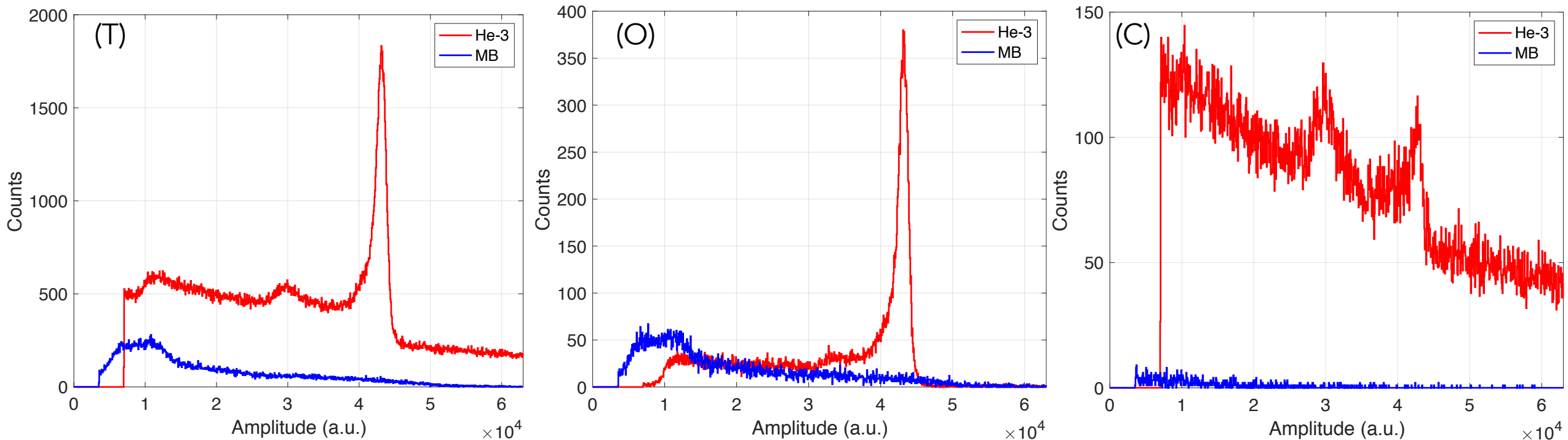
Indirect calculation

$$\Phi_i^{MB} : \epsilon_i^{MB} = \Phi_i^{He} : \epsilon_i^{He}$$

$$\epsilon_i^{He} = \frac{\Phi_i^{He} * \epsilon_i^{MB}}{\Phi_i^{MB}}$$

Calculated fast neutron flux

$$\Phi_{fn} = \frac{\Phi - (4 * \Phi_{tn})}{4} \sim \Phi_p$$



	Φ	Φ_{tn}	Φ_p	Φ_{fn}
MB	$6.6 \cdot 10^4 \pm 250$	$1.56 \cdot 10^4 \pm 120$	870 ± 30	900 ± 140
^3He	$3.95 \cdot 10^5 \pm 600$	$2.45 \cdot 10^4 \pm 150$	$7.38 \cdot 10^4 \pm 300$	$7.4 \cdot 10^4 \pm 200$

$$\epsilon_{tn}^{He} = 0.94 \pm 0.09$$

$$\epsilon_{fn}^{He} = 1.2 \cdot 10^{-3} \pm 6 \cdot 10^{-4}$$

GAMMA-RAY BACKGROUND

Source	x- or γ -ray, keV	intensity, %
^{133}Ba	31	96.1
	35	17.3
	81	32.9
	276	7.2
	303	18.3
	356	62.0
	384	9.8
^{60}Co	1173	99.8
	1332	100
^{137}Cs	32	5.6
	662	85.1

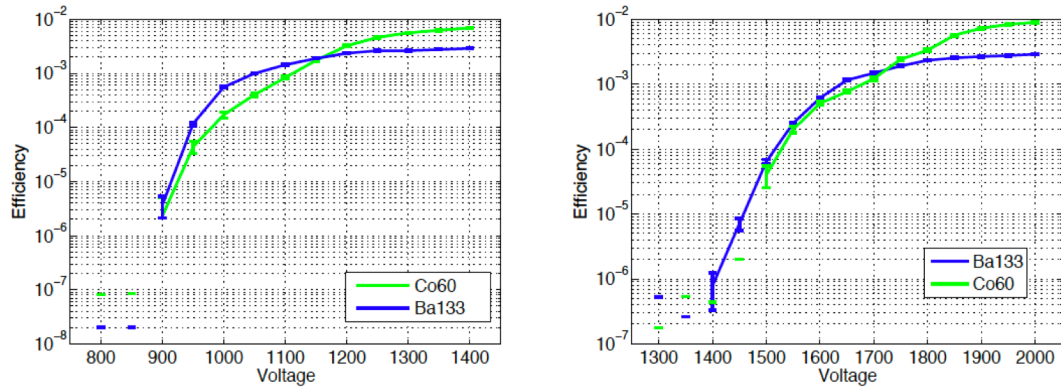


Figure 7: Plateau measurements with the Multi-Grid ^{10}B detector (left) and a Multi-Tube ^3He detector (right) with ^{133}Ba (200 kBq) and ^{60}Co (25 kBq) γ -ray sources. Detection efficiency per tube is shown. The nominal operating voltages are 850V and 1350V respectively. The disconnected points at the left ends of the curves are upper limits, where no statistically-significant counts could be detected over background.

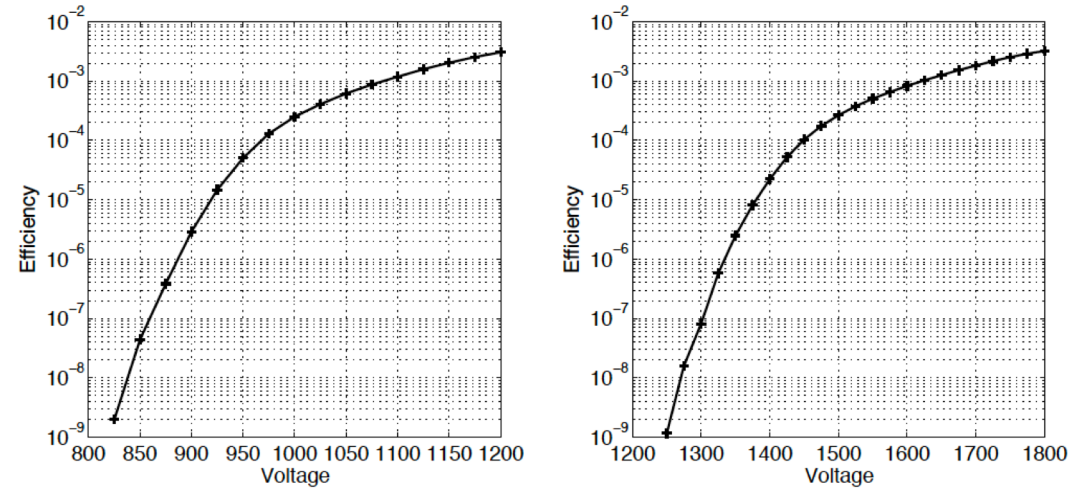


Figure 8: Plateau measurement with the Multi-Grid ^{10}B detector (left) and a Multi-Tube ^3He detector (right) with a 164 MBq ^{137}Cs source.

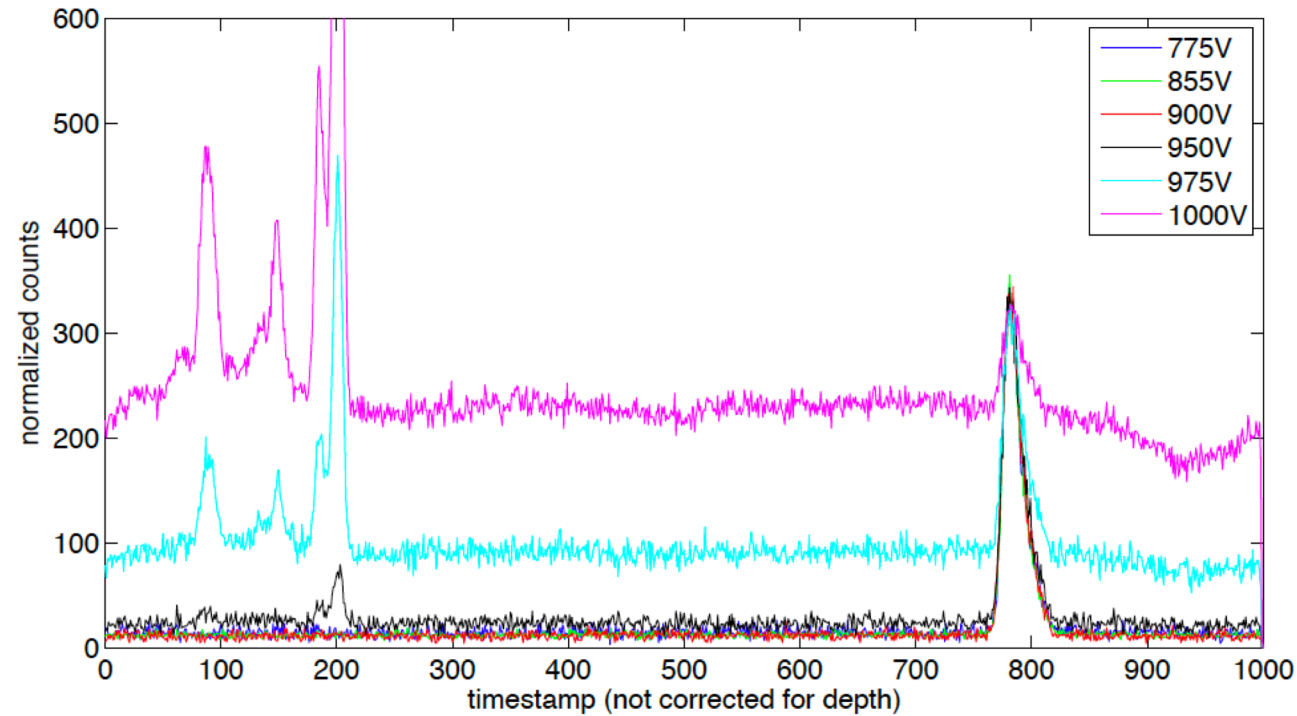


Figure 10: Time spectrum of the ^{10}B prototype for a range of bias voltages. No evidence of the γ -peaks is visible until the voltage reaches 950V. The peak at the channel numbers 770-810 is the elastic neutron peak. Note that no timing correction for the depth of the detector was performed here, since it cannot be done in a consistent way for both γ and n at the same time – and here we are interested in γ – therefore the neutron peak appears wider than it normally would.

Investigation of gamma-ray sensitivity of neutron detectors based on thin converter films

A. Khaplanov^{a,b,1}, F. Piscitelli^{b,c}, J.-C. Buffet^b, J.-F. Clergeau^b, J. Correa^b,
P. van Esch^b, M. Ferraton^b, B. Guerard^b and R. Hall-Wilton^c

^aEuropean Spallation Source,
P.O. Box 176, SE-22100 Lund, Sweden

^bInstitute Laue Langevin,
Rue Jules Horowitz, FR-38042 Grenoble, France

^cUniversity of Perugia,
Piazza Università 1, 06123 Perugia, Italy

E-mail: Anton.Khaplanov@ess.se

ABSTRACT: Currently, many detector technologies for thermal neutron detection are in development in order to lower the demand for the rare ^3He gas. Gas detectors with solid thin film neutron converters readout by gas proportional counter method have been proposed as an appropriate choice for applications where large area coverage is necessary. In this paper, we investigate the probability for γ -rays to generate a false count in a neutron measurement. Simulated results are compared to measurement with ^{10}B thin film prototypes and a ^3He detector. It is demonstrated that equal γ -ray rejection to that of ^3He tubes is achieved with the new technology. The arguments and results presented here are also applicable to gas detectors with converters other than solid ^{10}B layers, such as ^6Li layers and $^{10}\text{BF}_3$ gas.

KEYWORDS: Gaseous detectors; Neutron detectors (cold, thermal, fast neutrons); Detector modelling and simulations I (interaction of radiation with matter, interaction of photons with matter, interaction of hadrons with matter, etc)

ARXIV EPRINT: [1306.6247](https://arxiv.org/abs/1306.6247)

¹Corresponding author.

© 2013 IOP Publishing Ltd and Sissa Medialab srl

doi: [10.1088/1748-0221/8/10/P10025](https://doi.org/10.1088/1748-0221/8/10/P10025)

2013 JINST 8 P10025

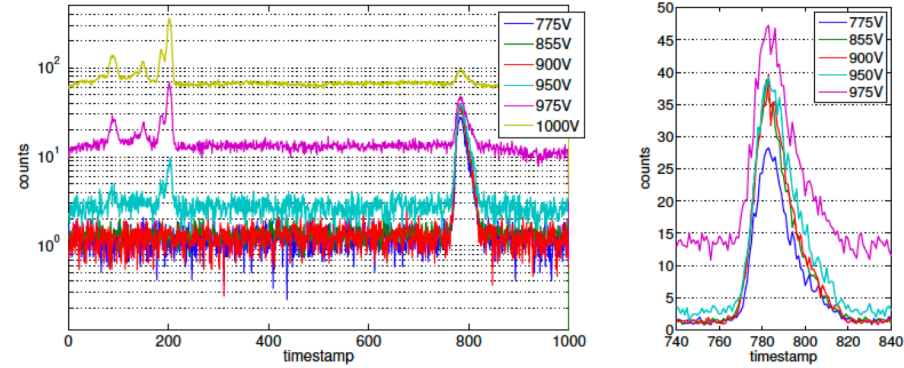


Figure 10. *Left:* time spectrum of the ^{10}B prototype for a range of bias voltages. No evidence of the γ -peaks is visible until the voltage reaches 950 V. The peak at the channel numbers 770–810 is the elastic neutron peak. *Right:* detail of the neutron peak in linear scale. Note the reduced height of the peak at the lowest voltage is due to reduced neutron efficiency. No timing correction for the depth of the detector was performed here, since it cannot be done in a consistent way for both γ and n at the same time — and here we are interested in γ — therefore the neutron peak appears wider than it normally would.

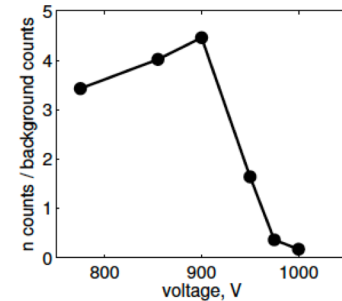


Figure 11. Evolution of the ratio of the neutron signal in the elastic peak to background corresponding to the largest γ peak.

