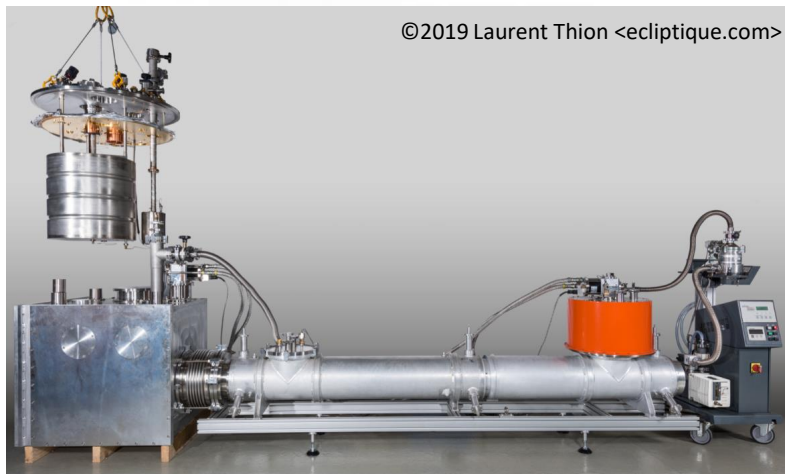
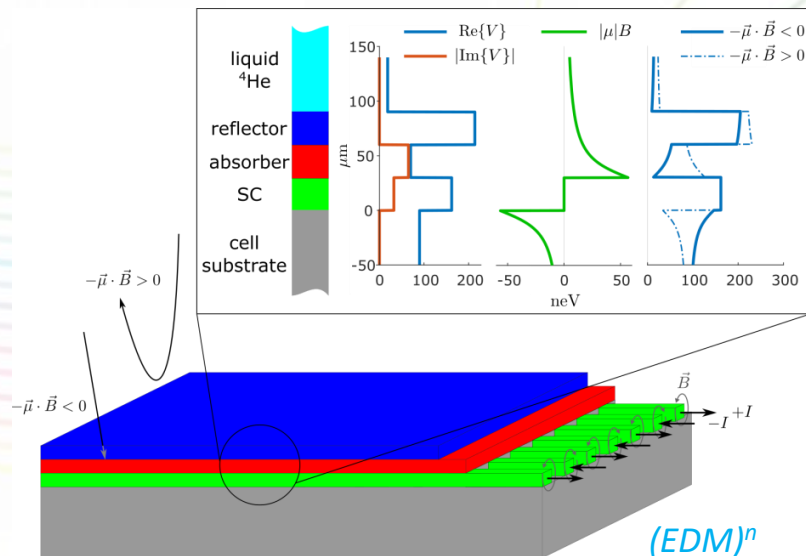


# New platforms for ultracold neutron production: SuperSUN and beyond



SuperSUN



Skyler Degenkolb, Institut Laue-Langevin

# Neutron delivery at the ILL

## In SuperSUN's converter vessel:

- $R \sim 15 \text{ UCN}/(\text{cm}^3 \text{ s})$  expected

## End of guide H523:

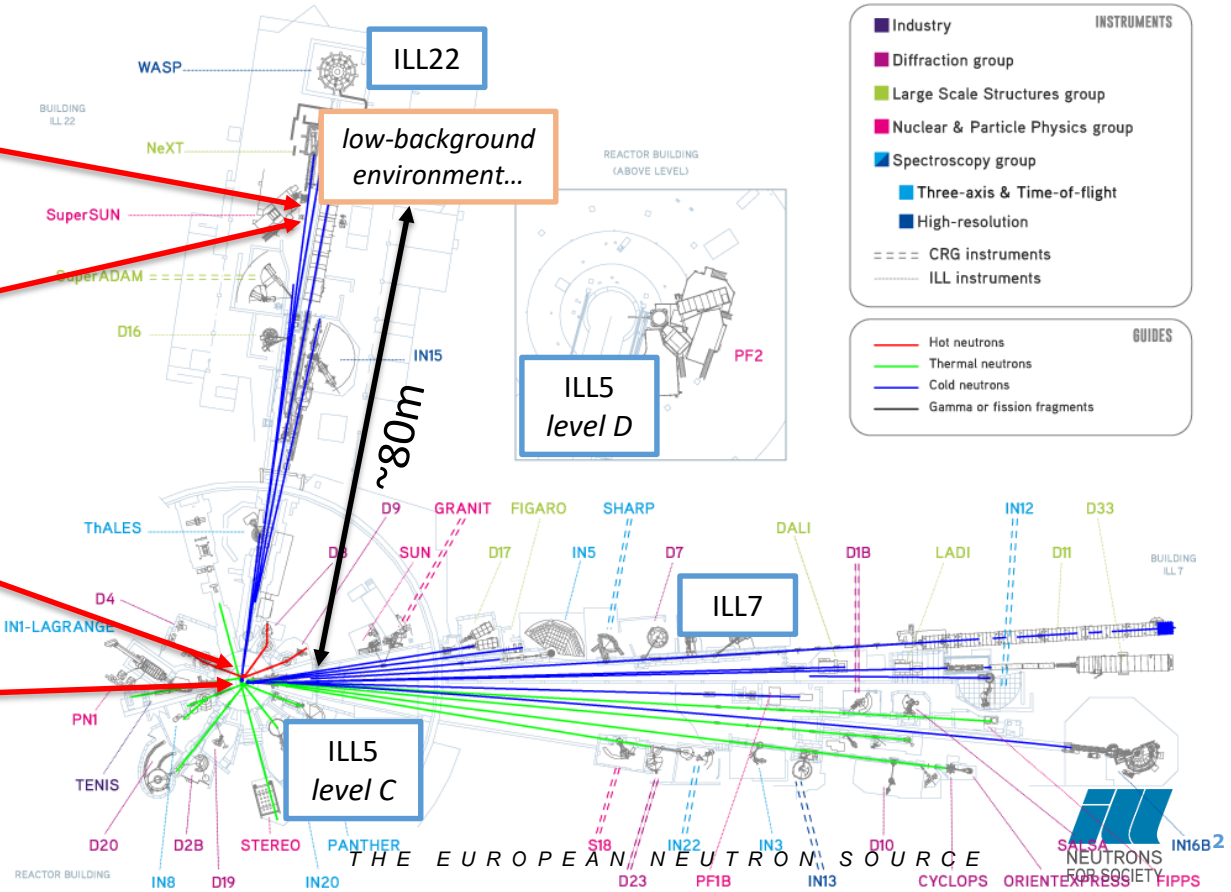
- $\Phi \sim 2 \times 10^{10} \text{ n}/(\text{cm}^2 \text{ s})$

## Horizontal Cold Source:

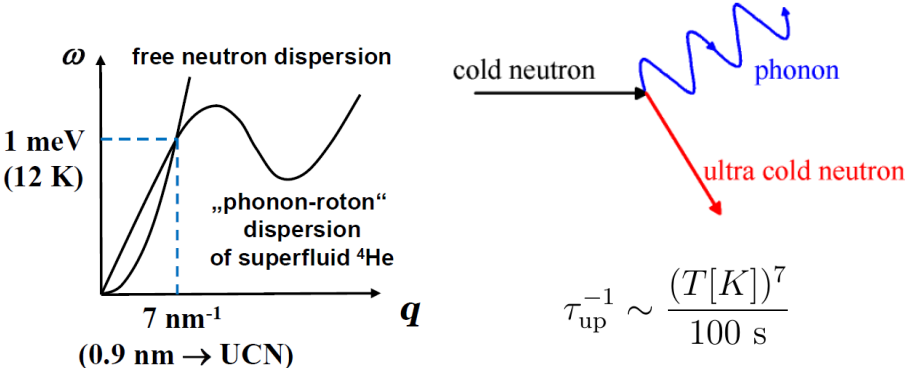
- $\Phi \sim 10^{14} \text{ n}/(\text{cm}^2 \text{ s})$

## In pile:

- $\Phi \sim 1.5 \times 10^{15} \text{ n}/(\text{cm}^2 \text{ s})$



# “Ultracold” and “Superthermal”



$$\tau_{\text{up}}^{-1} \sim \frac{(T[\text{K}])^7}{100 \text{ s}}$$

Velocity	“Temperature”	Energy
$10^0 - 10^1 \text{ m/s}$	Ultracold	5 neV – 500 neV
$10^1 - 10^2 \text{ m/s}$	Very cold	$0.5 \mu\text{eV} - 50 \mu\text{eV}$
$10^2 - 10^3 \text{ m/s}$	Cold	$50 \mu\text{eV} - 5 \text{ meV}$
$2.2 \times 10^3 \text{ m/s}$	Thermal	25 meV
$2 \times 10^3 - 2 \times 10^4 \text{ m/s}$	Hot	20 meV – 2 eV

## Production rate density, and storage losses

### “Conversion” in Superfluid $^4\text{He}$

- phase space compression
- requires dissipative physics
- flux vs. density (beam vs. storage)
- important for next generation

$$5\text{-}15 \text{ /}(\text{cm}^3 \text{ s}) \leftarrow R \sim \left( \frac{5 \times 10^{-8}}{\text{cm}^3 \text{ s}} \right) \times \left. \frac{d\Phi}{d\lambda} \right|_{8.9\text{\AA}} \times \left( \frac{V_{\text{trap}}}{233 \text{ neV}} \right)^{3/2}$$

$$\text{limited by decay, wall interactions} \leftarrow \frac{1}{\tau} = \frac{1}{\tau_{\beta}} + \frac{1}{\tau_{\text{up}}} + \frac{1}{\tau_{\text{capture}}} + \frac{1}{\tau_{\text{wall}}} + \dots$$

# UCN Storage (and nEDM)

## Statistics

- Flux vs. *density*
  - want to count many UCN, after storage
  - transport losses and dilution
- Storage time (including  $T_1/T_2$ )
- Total measurement time/repetitions
  - duty factor vs. accumulation time
  - long-term stability becomes important
- Polarization (incl. analyzing power)
- Electric field



$$N_{\text{cell}} \sim \rho_{\text{cell}} V_{\text{cell}} \sim \frac{\rho_{\text{source}} V_{\text{cell}}}{1 + \frac{V_{\text{cell}} + V_{\text{guide}}}{V_{\text{source}}}}$$

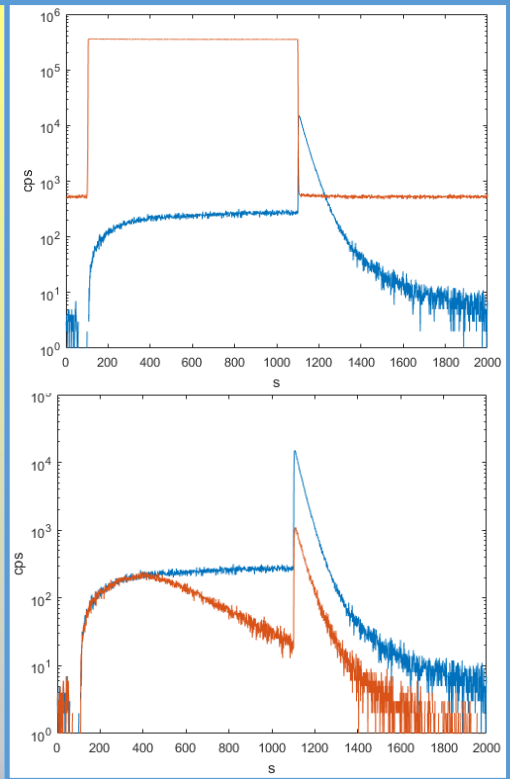
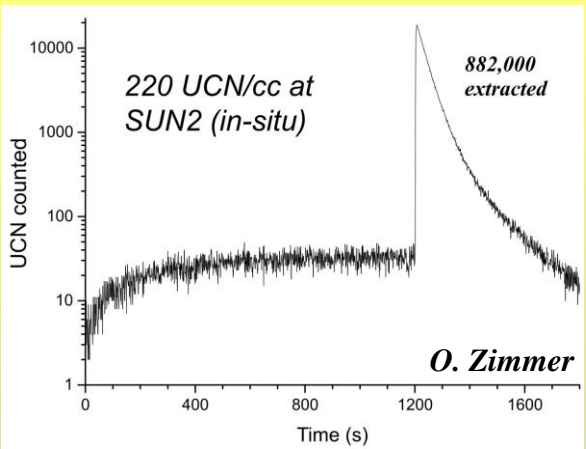
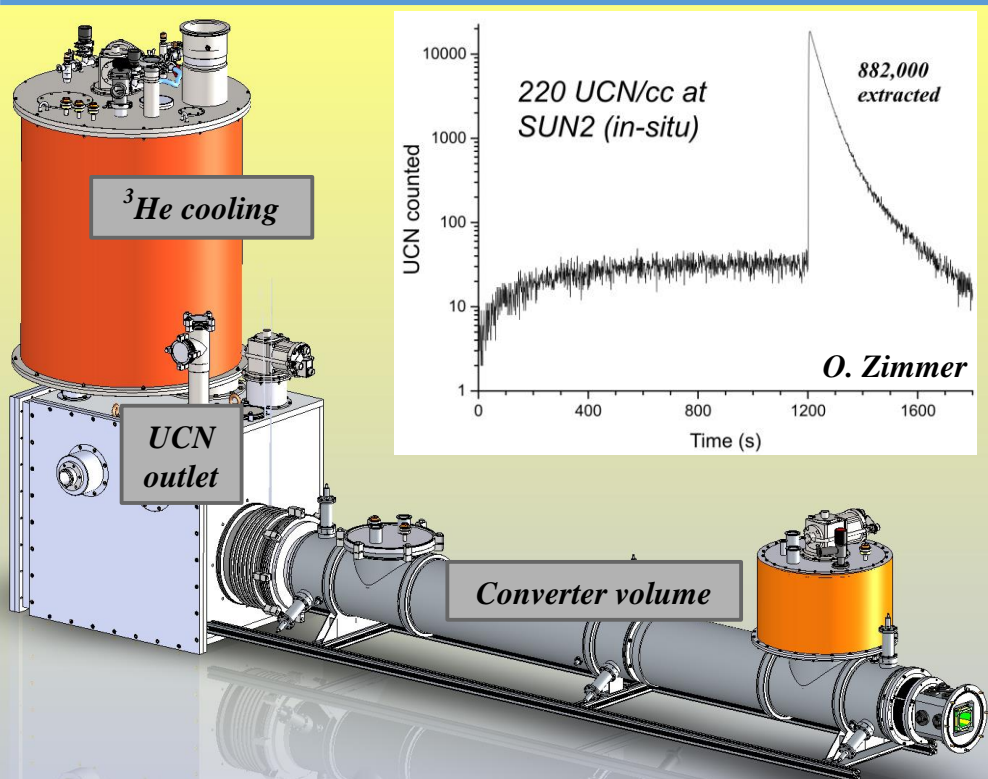


$$\frac{1}{\tau} = \frac{1}{\tau_{\beta}} + \frac{1}{\tau_{\text{up}}} + \frac{1}{\tau_{\text{capture}}} + \frac{1}{\tau_{\text{wall}}} + \dots$$

## Systematics (not exhaustive)

- Cell size and quality
- Field stability, monitor quality
- Magnetic screening
- Environment/backgrounds

# Superthermal (He-II) UCN Sources



**High-density sources**

- (high-flux: more common)
- phase space density
- beam current

**Two basic issues:**

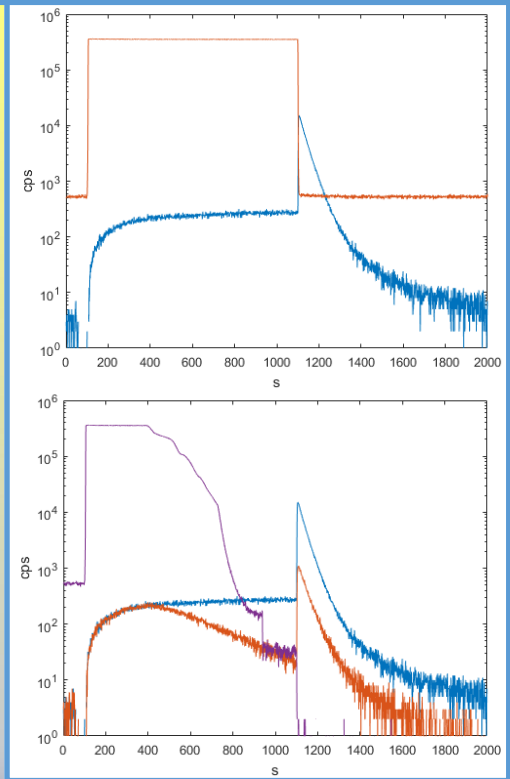
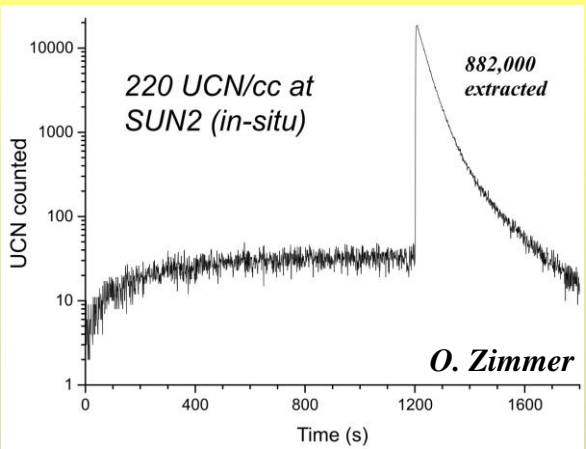
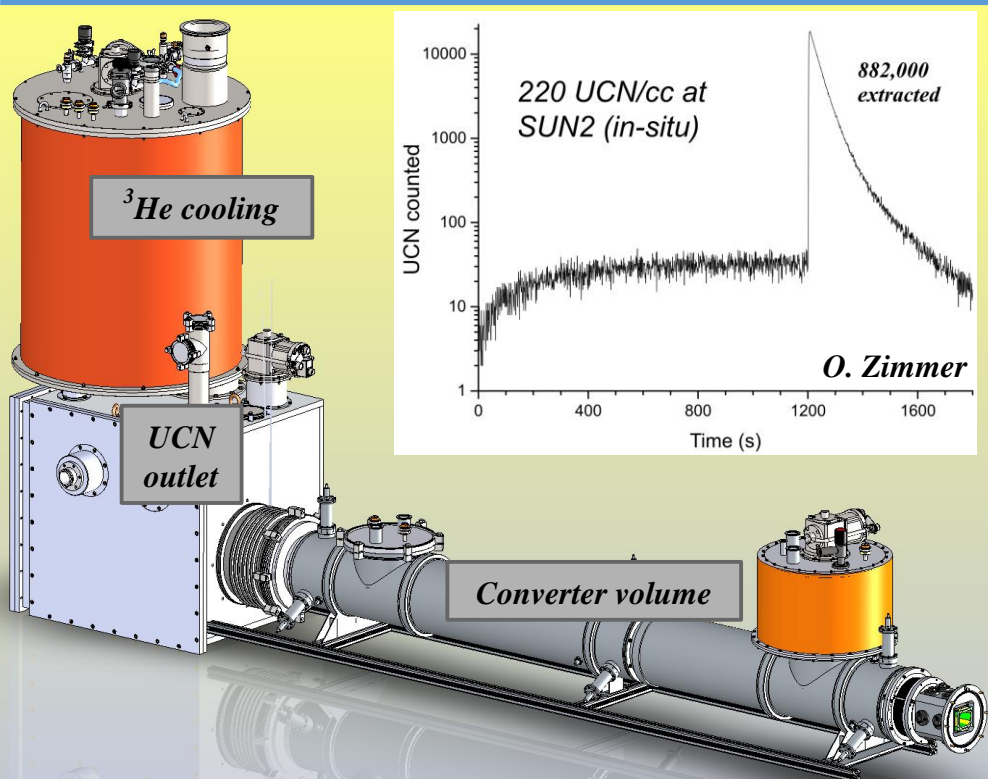
$$\rho_{UCN} = P\tau$$

- production rate
- UCN loss

**We might improve:**

- cold beam flux (usage?)
- storage/transport losses

# Superthermal (He-II) UCN Sources

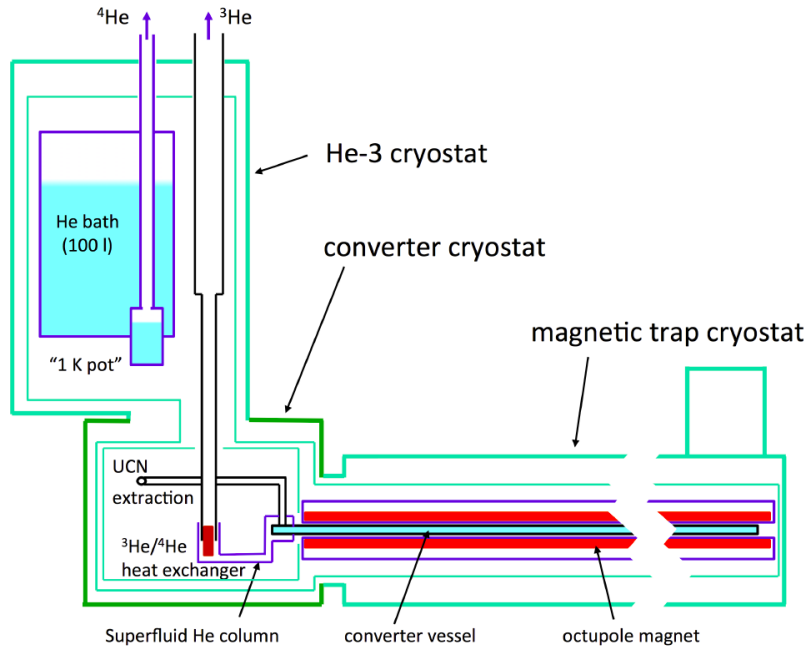


**Characteristic output:**

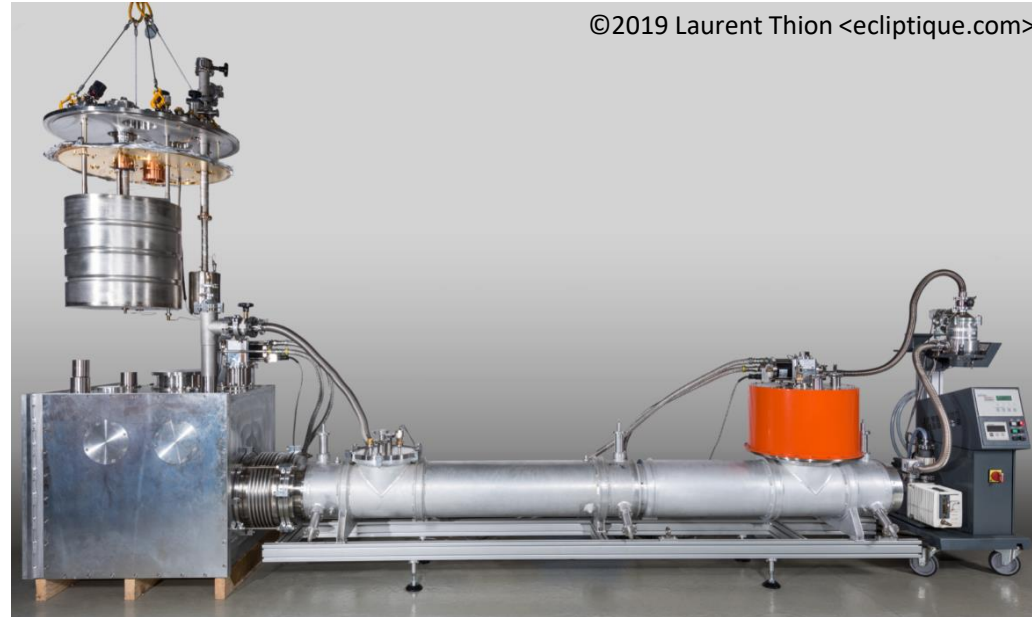
- $\lambda \sim 900 \text{ \AA}$
- $(v \sim 4 \text{ m/s})$
- $\rho \sim 2 \text{ cm}^{-3}$
- $(\sim 1 \times 10^{-10} \rho_{rest-gas})$
- $\Phi \sim 500 \text{ n/s/cm}^2$
- $(\sim 3 \times 10^{-13} \Phi_{pool})$
- $\rho_{phase-space} < 10^{-13}$
- $\sim (900 \text{ \AA})^3 (220 \text{ cm}^{-3})$

# SuperSUN (in two phases)

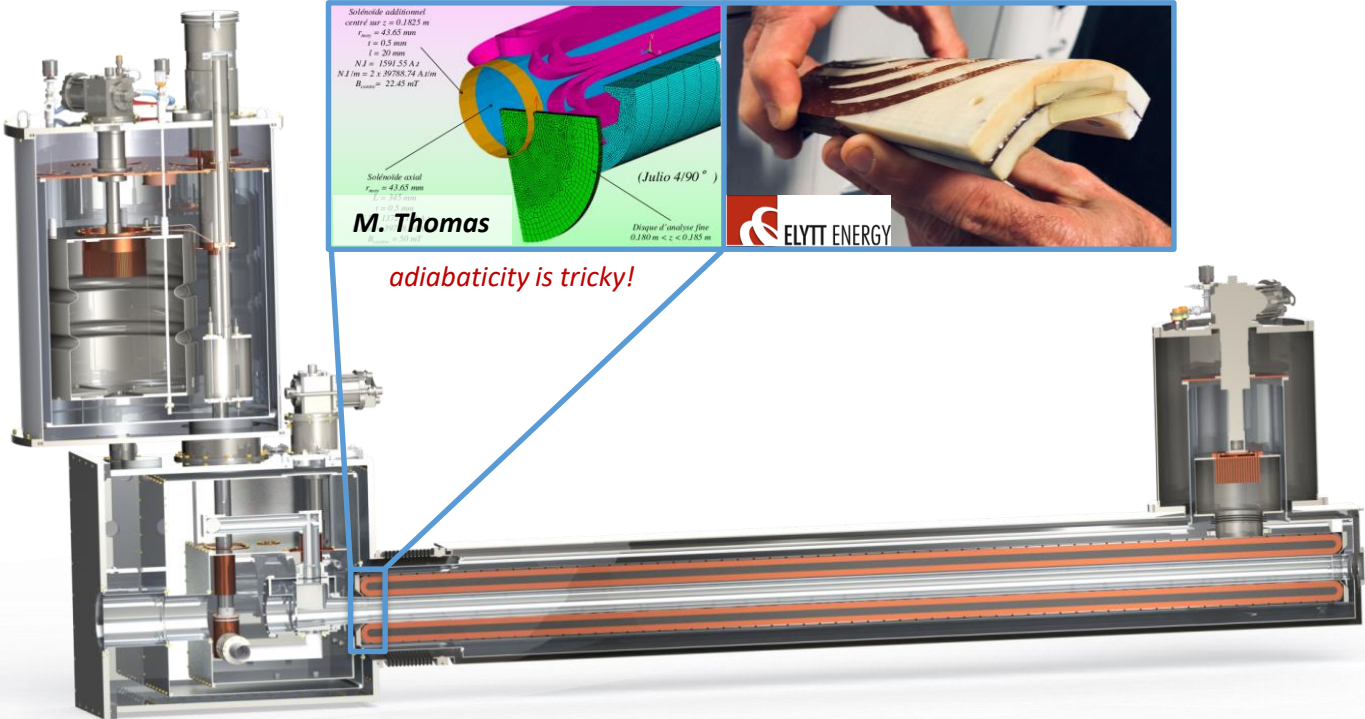
## Cryogenic concept:



## Present status:



# Magnet and Accumulation Vessel



Solénoïde additionnel  
centre sur  $z = 0,1025 \text{ m}$   
 $r_{\text{ext}} = 43,65 \text{ mm}$   
 $r = 0,5 \text{ mm}$   
 $l = 20 \text{ mm}$   
 $N/l = 1991,55 \text{ A/l}$   
 $N/l \text{ (m)} = 2 \times 39788,74 \text{ A/l(m)}$   
 $R_{\text{max}} = 22,45 \text{ mT}$

Solénoïde axial  
 $r_{\text{ext}} = 43,65 \text{ mm}$   
 $r = 0,5 \text{ mm}$

(Julio 4/90 °)

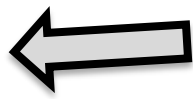
Disque d'analyse fine  
 $0,100 \text{ m} < z < 0,150 \text{ m}$

**M. Thomas**

*adiabaticity is tricky!*



*cold neutrons in*





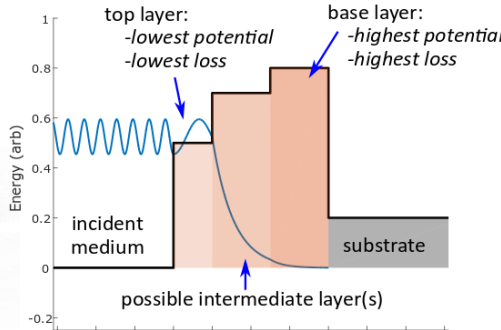
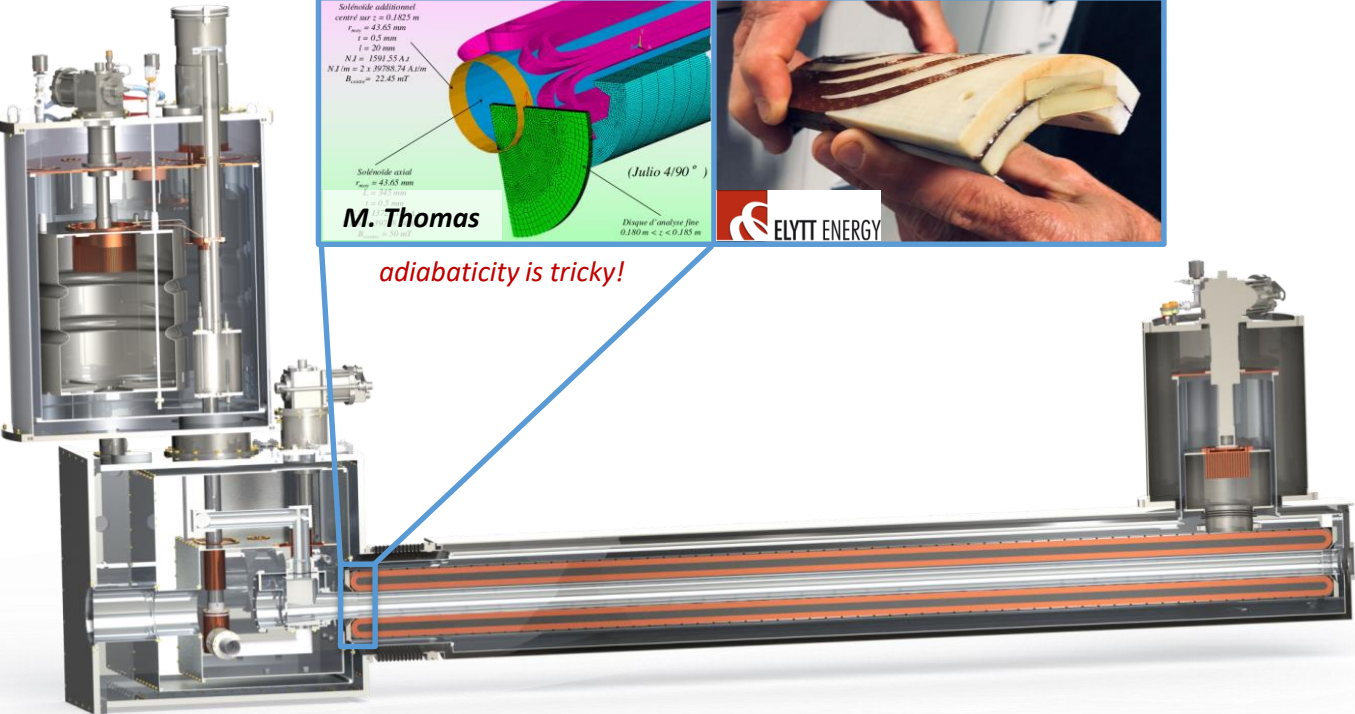
# Magnet and Accumulation Vessel



*adiabaticity is tricky!*



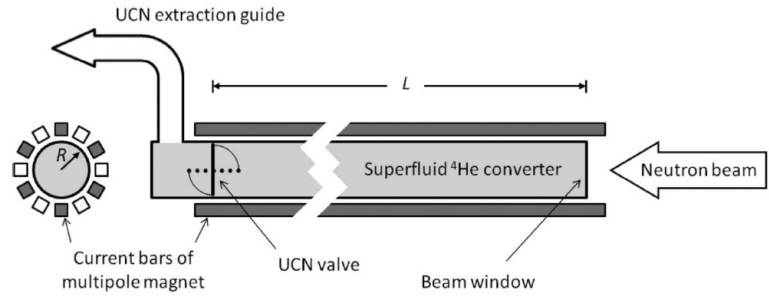
J. Hingerl, MSc.



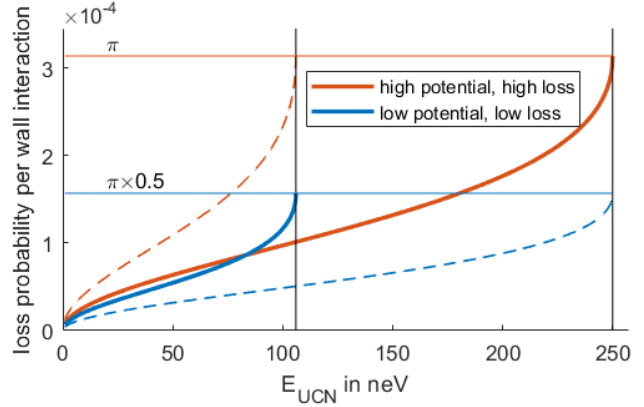
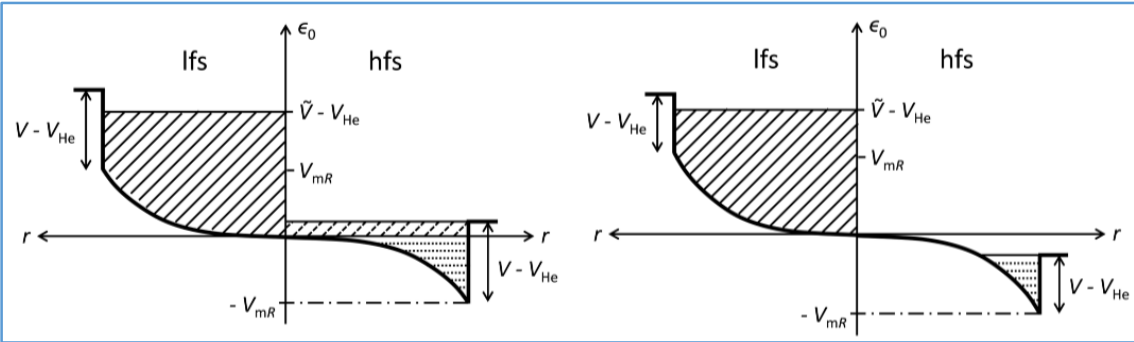
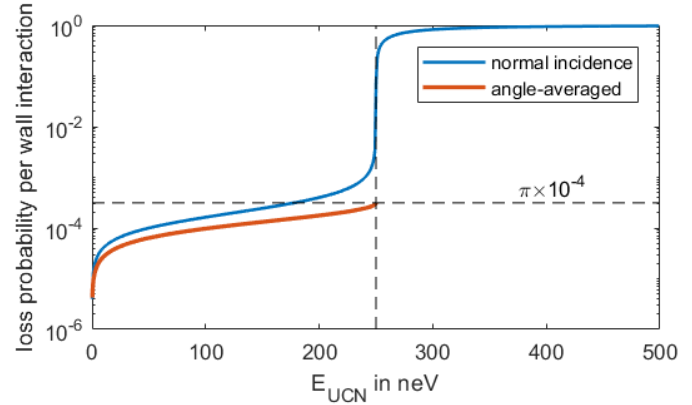
# Optimizing Coatings and Vessels

## SuperSUN phase II: magnetic octupole reflector

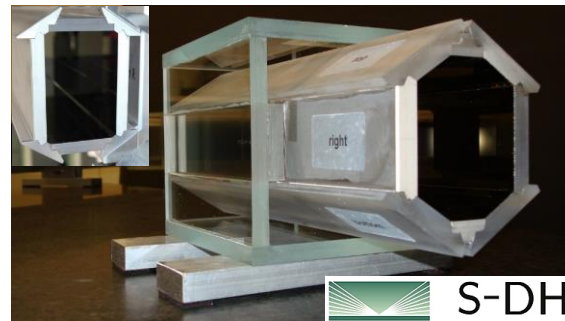
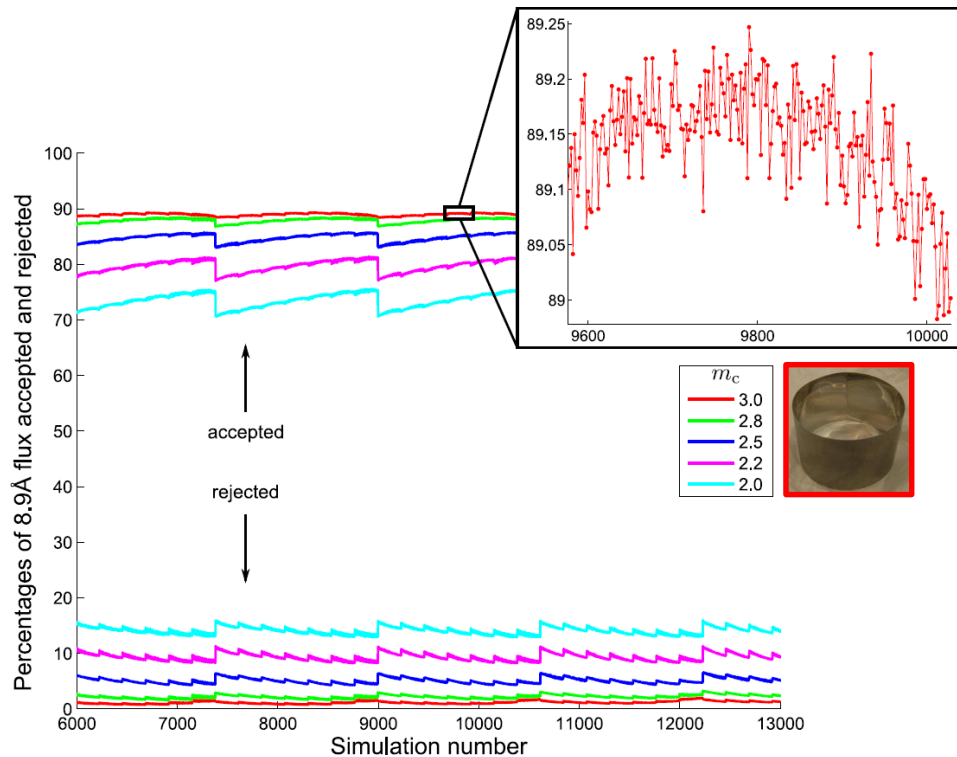
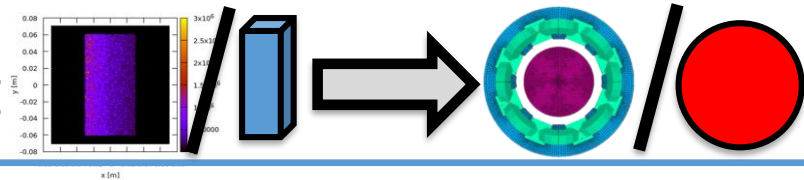
PHYSICAL REVIEW C 92, 015501 (2015)



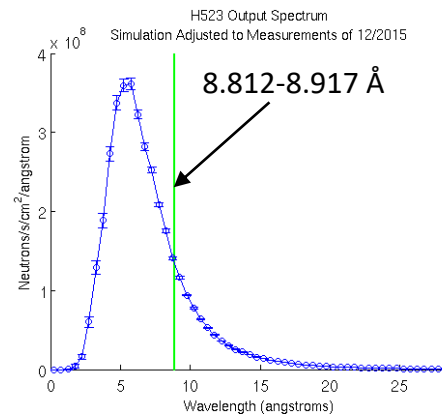
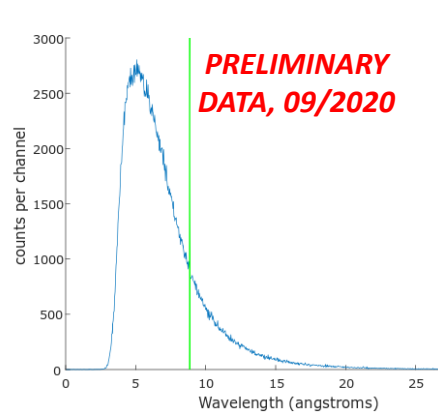
## Material wall potentials



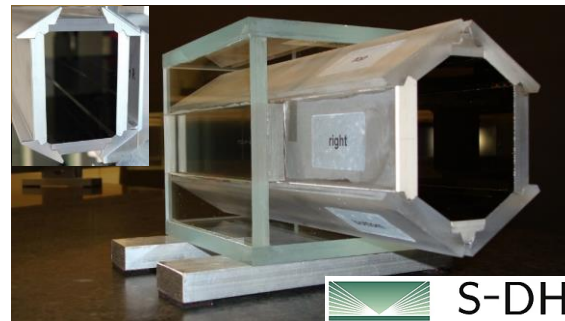
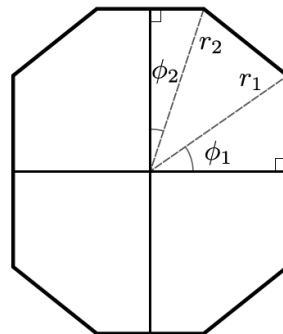
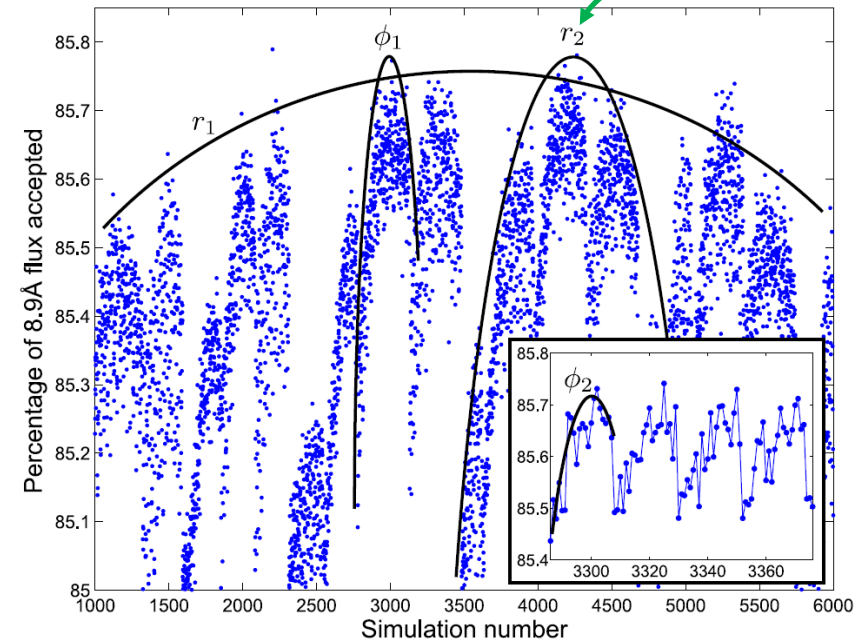
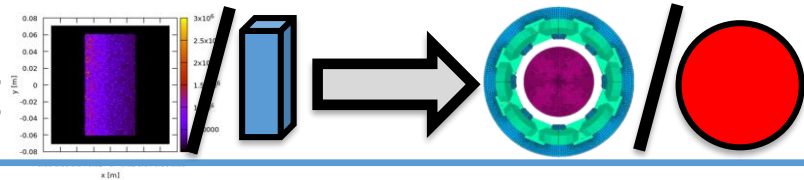
# SuperSUN Cold Neutron Guides:



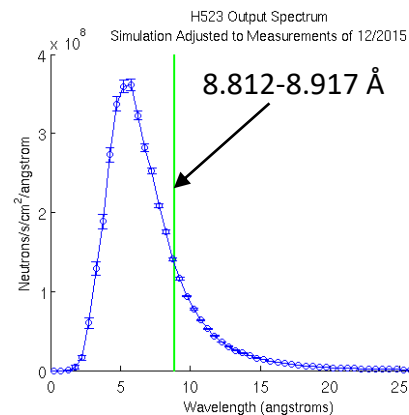
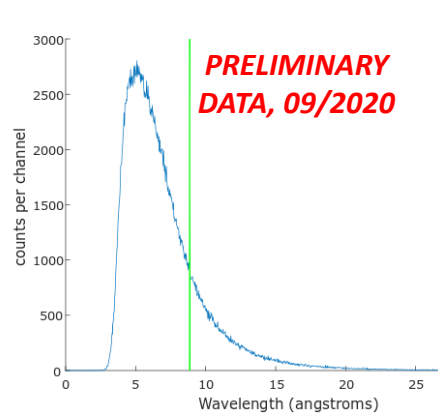
Capture-weighted flux (03/2020):  $2 \times 10^{10}$  n/(cm<sup>2</sup> s)



# SuperSUN Cold Neutron Guides:



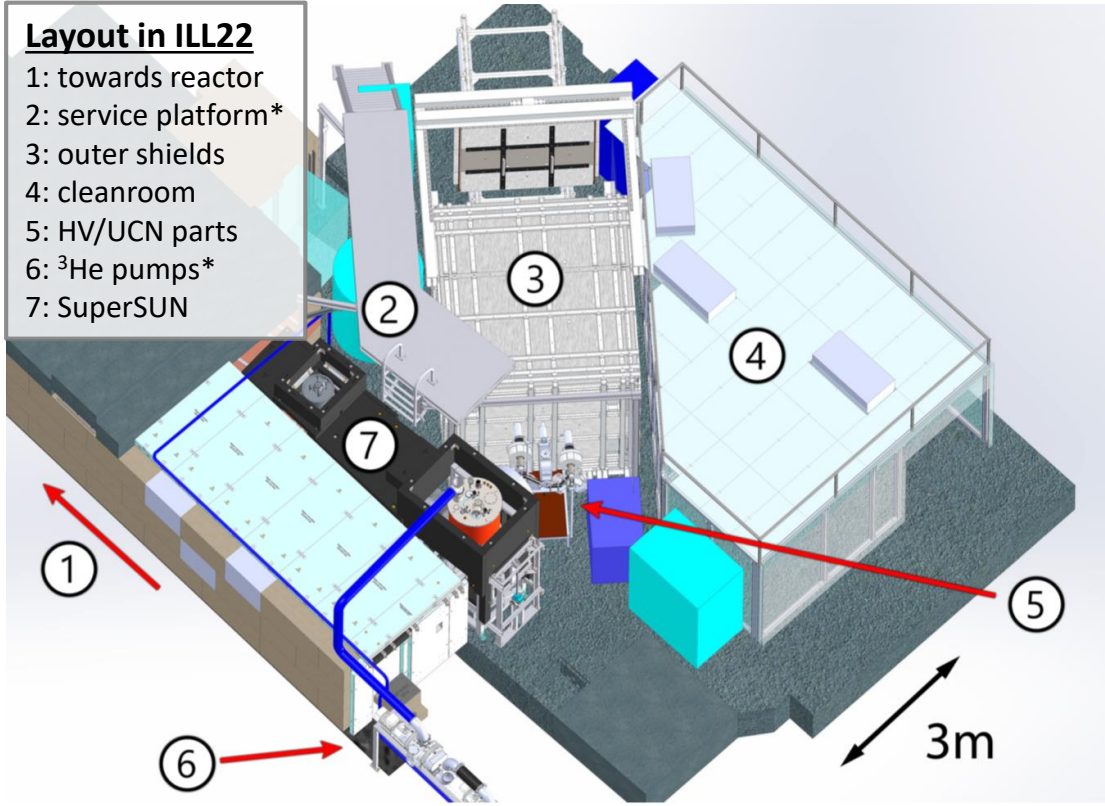
Capture-weighted flux (03/2020):  $2 \times 10^{10}$  n/(cm<sup>2</sup> s)



14/10/2020

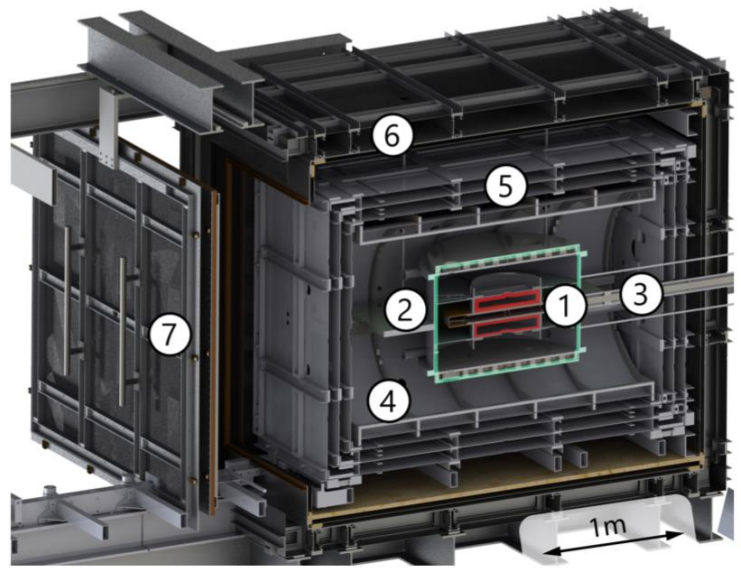
Skyler Degenkolb

# Experimental Zone, PanEDM

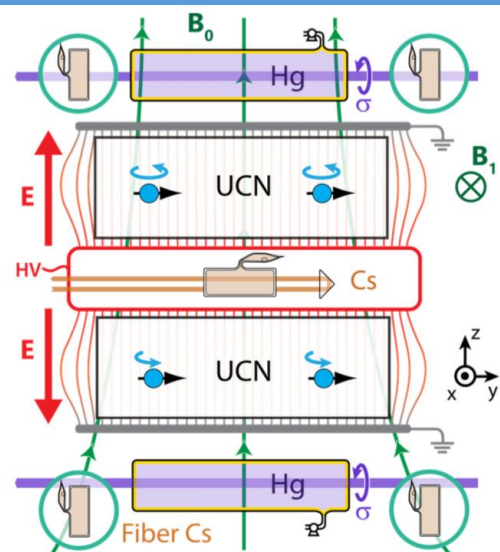


- Layout in ILL22**
- 1: towards reactor
  - 2: service platform\*
  - 3: outer shields
  - 4: cleanroom
  - 5: HV/UCN parts
  - 6: <sup>3</sup>He pumps\*
  - 7: SuperSUN

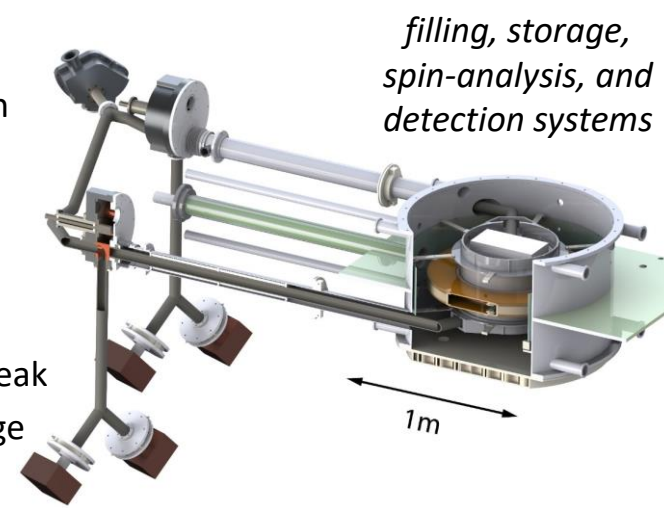
- PanEDM Magnetic and RF Shielding**
- 1: UCN cells
  - 2: vacuum chamber
  - 3: HV insertion
  - 4: cylindrical shield
  - 5: inner shields (3)
  - 6: outer MSR (2+1)
  - 7: MSR door



# SuperSUN-PanEDM: Key Components



- Two chambers, at room temperature
- $^{199}\text{Hg}$  magnetometers with few-fT resolution
- Cs magnetometers (also at HV)
- Magnetic shield (SF:  $6 \times 10^6$  at 1 mHz)
- Simultaneous spin detection
- SuperSUN UCN source at ILL, in 2 phases:
  - Phase I: unpolarized UCN with 80 neV peak
  - Phase II: polarized UCN, magnetic storage
- Ongoing installation/commissioning



$^3\text{He}/^4\text{He}$  Heat exchanger

$m=4$  "replica"

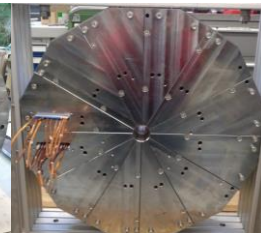
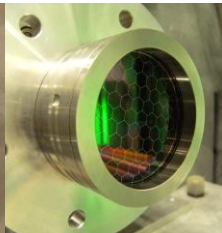
PP foil

3-way switch

$\text{H}_2\text{O}$ -cooled polarizer

vacuum chamber

"Sunioño" test vessel



14/10/2020

Skyler Degenkolb

THE EUROPEAN NEUTRON SOURCE

NEUTRONS FOR SOCIETY

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# Schedule and Sensitivity Estimate

Statistical sensitivity:

<b>SuperSUN</b>	<b>Phase I</b>
Saturated source density [cm <sup>-3</sup> ]	330
Diluted density [cm <sup>-3</sup> ]	63
Density in cells [cm <sup>-3</sup> ]	3.9
<b>PanEDM Sensitivity [1σ, e cm]</b>	
Per run	5.5 × 10 <sup>-25</sup>
Per day	3.8 × 10 <sup>-26</sup>
Per 100 days	3.8 × 10 <sup>-27</sup>

*extraction losses...*

## Systematic effects: magnetic field, soft spectrum

- Nondynamical phases (field control, spectrum)
- No comagnetometer (phase I; magnetic stability)
- Gradiometer stack + Cs sensor(s) in HV electrode
- Optically decoupled leakage current monitor
- On-site magnetic screening

## Phase II: superconducting octupole trap

- Lower source losses, so higher UCN density
- UCN pre-polarized in the source

## Schedule:

- Highly dependent on ILL cycle planning
- Try for first UCN in 2021
- Fortunately, we can do a lot in the meantime...

More details, including phase II:

EPJ Web of Conferences **219**, 02006 (2019)  
<https://doi.org/10.1051/epjconf/201921902006>

# Challenges in the Next Generation

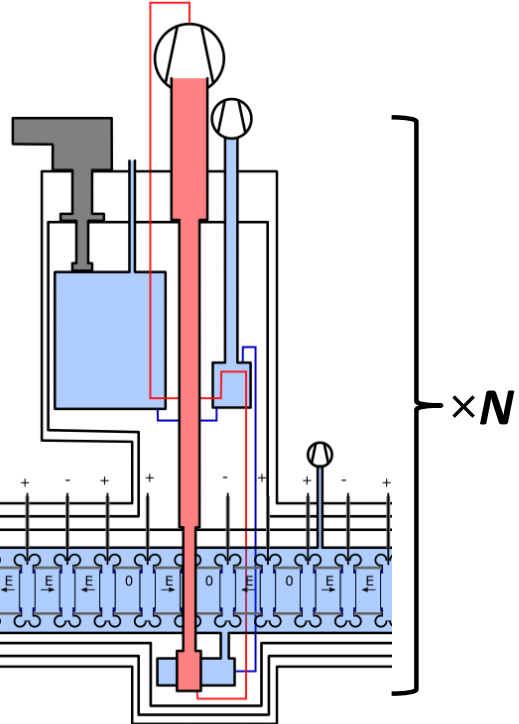
## How to get around the problems of production rate and loss rate?

- Statistical limitations: low source **phase-space** density, inefficient transport
  - *In-situ* experiments are challenging → why?
  - **Modularity and Scalability**
- Systematic limitations: imperfect knowledge/control of experimental conditions
  - Can source and spectrometer use the same prototypes?
  - Can we avoid some challenges already faced by others who tried this?
- Thought experiment: what is the ultimate limit within existing means?



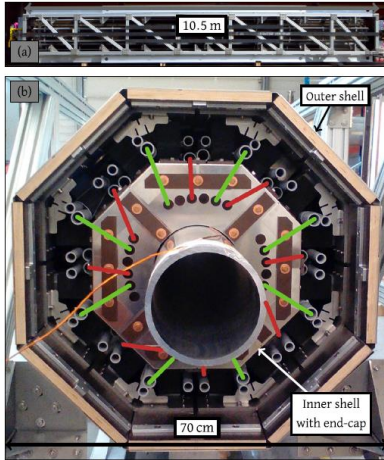
# Where has there already been progress?

## Cryogenics (SuperSUN, SNS)

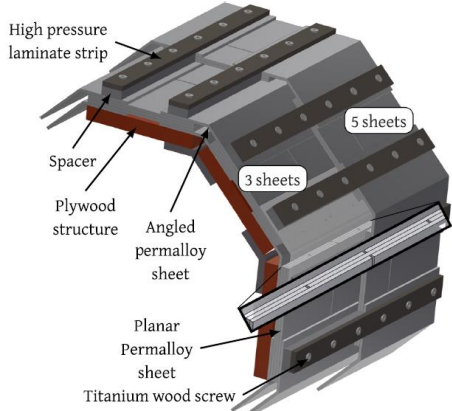


## Magnetic Shielding (pEDM and AMO)

see e.g., Rev. Sci. Instrum. 91, 035117 (2020)



notice we like octagons...



## HV (SNS, CryoEDM)

Image: Brad Filippone



**Missing: in-situ polarization, detection, analysis**

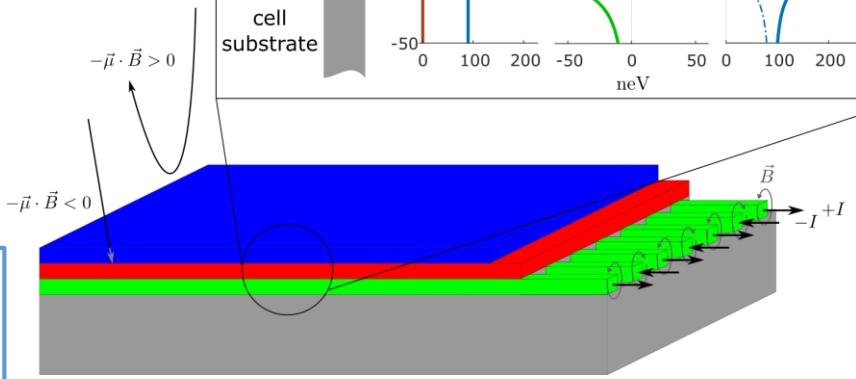
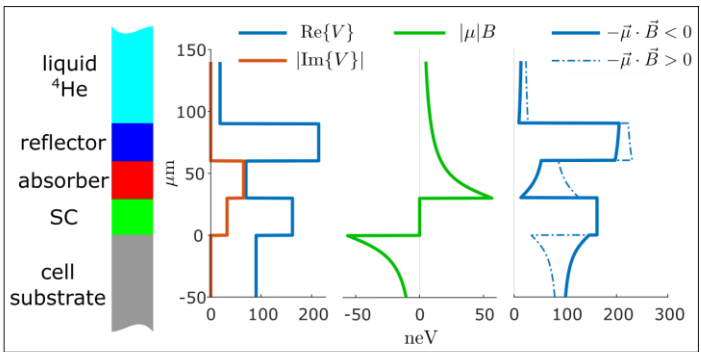
# New Detection Methods (*in-situ*)

- Many chances for UCN to be detected
- Meander field creates strong *local* gradient at wall surface (watch out!)
- Limitations:
  - Slowest UCN never penetrate
  - Fastest UCN always penetrate
  - Cell dimensions
  - Holding time
  - Readout efficiency

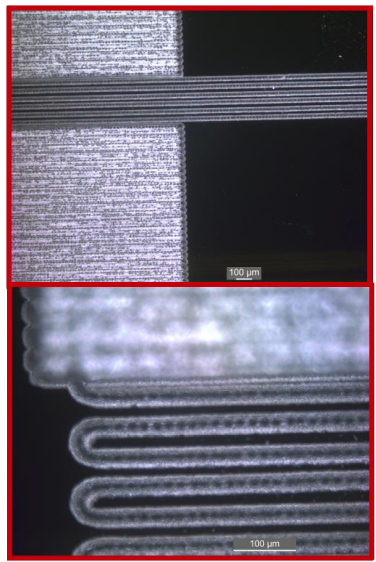
• Remember the themes...

Central element:  
in-situ, **polarization sensitive**  
UCN detectors

Detector concept: SMD, P Fierlinger, O Zimmer



Nb meanders on Si wafers: R Gernhäuser, S Winkler



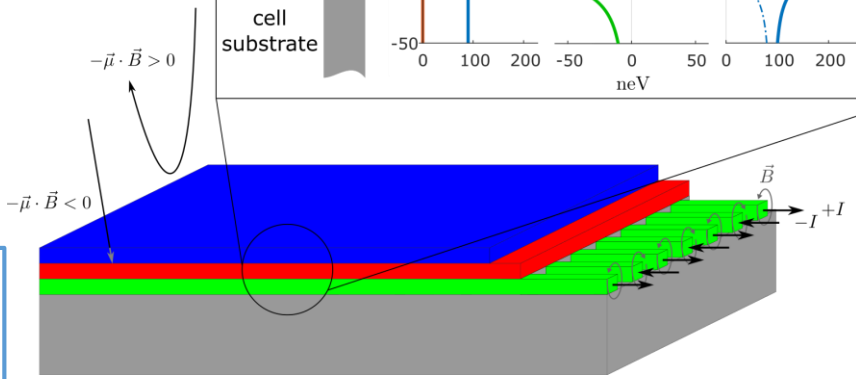
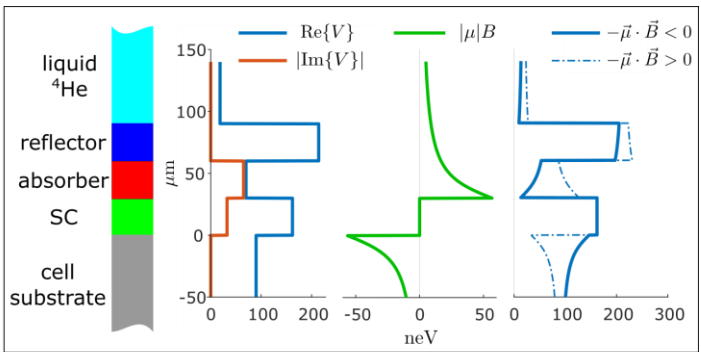
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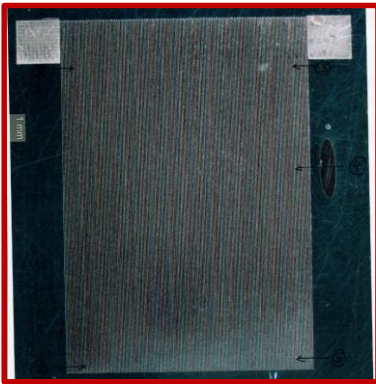
• Remember the themes...

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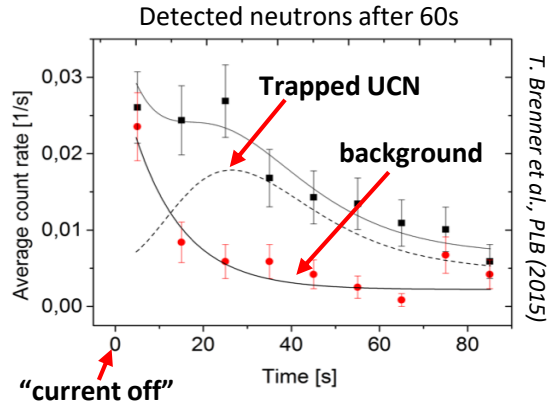


Nb meanders on Si wafers:  
R Gernhäuser, S Winkler

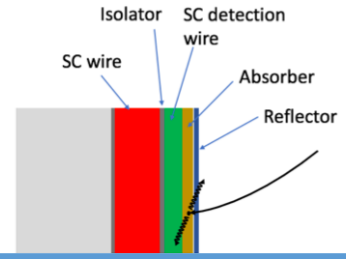


# (EDM)<sup>n</sup>: Further Detector Development

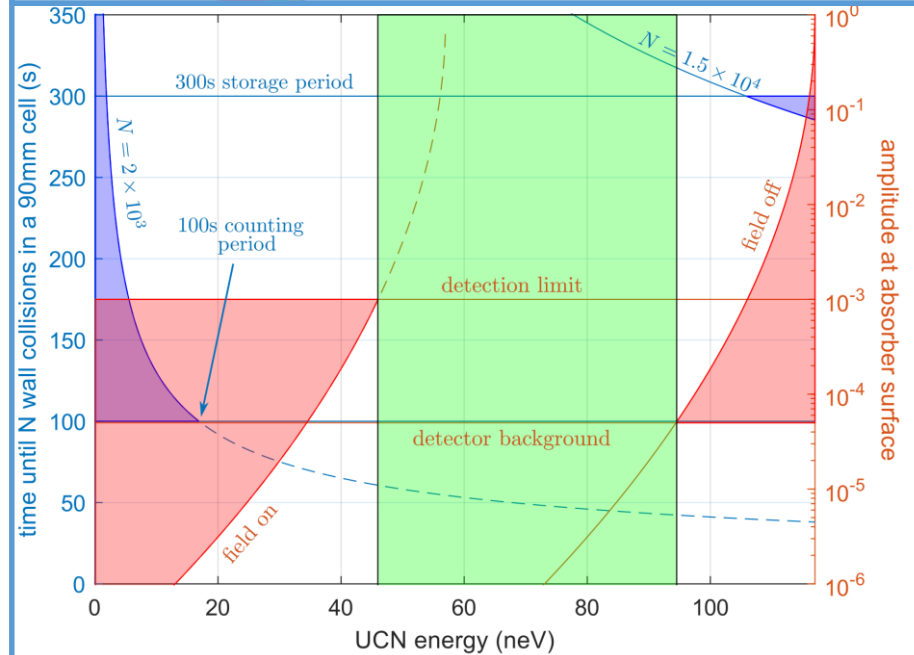
- CB-KID preferred to TES ( $T < T_c$  raises  $J_c$ )
- Testable in small/scalable experiments
  - First: simple cryo environment (dry)
  - Next: ~1K by pumping on LHe
  - Later:  $T < 1K$  w/  $^3He$  or dilution fridge
  - Develop robust cells before scaling up
- Systematically test new materials (cf. Suniño)



**Inspiration:**  
 Trapping/detection  
 of high-field-seekers  
 ...at PF2/TES (flux source)



Concept and a parameter space for in-situ polarized UCN detection, with a fomblin-like spectrum (e.g., the SUN2 converter vessel)



# (EDM)<sup>n</sup>: Ultimate Reach

## Main chances for meaningful improvements:

- Density (via extraction/transfer loss)
  - Requires in-situ polarization, analysis, detection
  - Watch out for systematics...
- Full use of cold neutron beam
  - Also brighter beams, new moderators\*
- Electric field (limited)
- Storage time (limited\*)
  - Also leads to some gain in density, via accumulation
- **Systematics**
- Total measurement time...

	Full Version	Small Scale
E	8.5 MV/m	7 MV/m
T	300s	250s
UCN/cc	1000	55
UCN/cell pair	$4.4 \times 10^6$	$6 \times 10^4$
N(T)/cell pair	$1.6 \times 10^6$	$2 \times 10^4$
M (per day)	$170 \times 144 = 24480$	1440 (10 cells)
$\alpha$	0.85	0.85
$\sigma_d$ (95% CL)	$2.1 \times 10^{-29} e \text{ cm}$	$7 \times 10^{-27} e \text{ cm}$

Special thanks to:

P Fierlinger  
H Filter  
D Wurm

O Zimmer  
X Tonon  
M Kreuz

D Beck  
T Neulinger  
??? (Post-Doc)

R Gernhäuser  
S Winkler  
R Georgii

M Thomas  
Elytt Energy  
S-DH, GmbH



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# Modified targets for 2020 & 2021

Reference reactor schedule (dates are tentative)



# The PanEDM Collaboration

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