

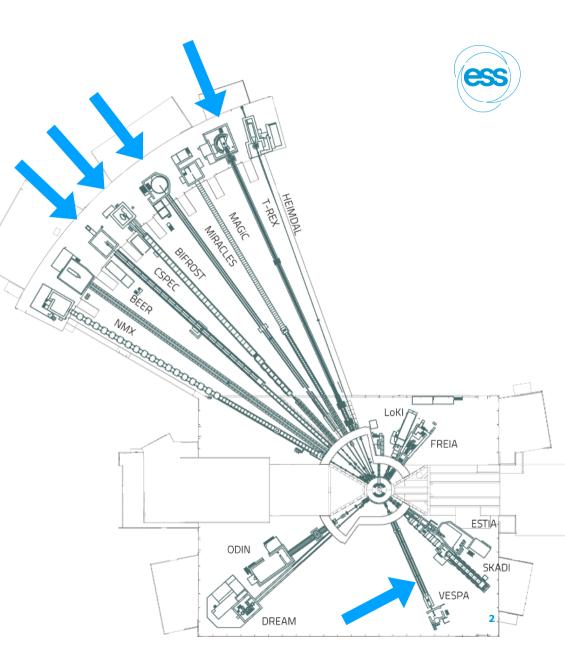
## **ESS Spectroscopy division**

Progress Staffing Rescoping First science

PASCALE DEEN

## Spectroscopy division

# CSPEC cold Chopper SpectrometerImage: SpectrometerIm





## **ESS Spectroscopy division**

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# **ESS Spectroscopy division**

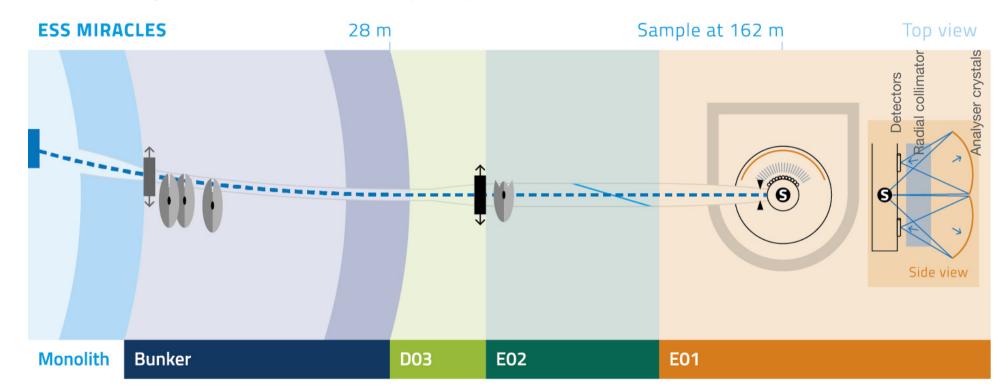
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## MIRACLES: High resolution backscattering spectrometer Ready for hot commisioning TG5: June 2027

Lead scientist: Felix Villacorta (Bilbao), Jose Pereira (Bilbao) Lead instrument engineer: Alex Conde Estebanez (Bilbao)







#### MIRACLES: High resolution backscattering spectrometer Progress Panels and alignment for analyser system Si crystals tested at IN16B Detector developments Detector tank HV Supply (DT5533EP) Charge Sens. SHV-SHV cables Preamplifier (A1422C) CPX400DP (length 15 - 25 m) (12V supply) ab-DB9, Coaxi SHV-SHV cables (length 0.5 m) Digitizer (R5560B 2010.2010.20 RJ45-RJ45 cables (length 5/15/25 m) To PC 128 chann LAA2 100 120 -200 200 -100 100 0 LAA2 Rot

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MIRACLES: High resolution backscattering spectrometer Concerns & actions:



Choppers: Delayed delivery and technical challenges of fast speed choppers.

Common projects not prioritised for Tranche 2/3 instruments: electrical and conventional utilities, monitors, detector scope.

Lack of onsite staff

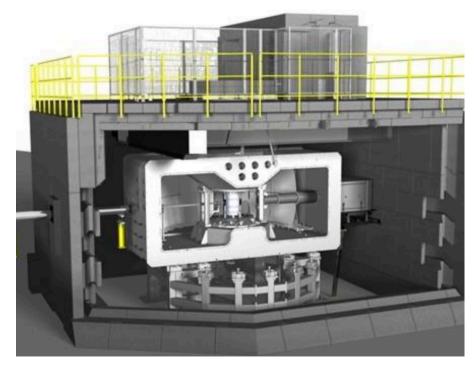
Position offered for instrument scientist based at ESS (October 2024), will need local engineering resources: position open for on-site engineer.

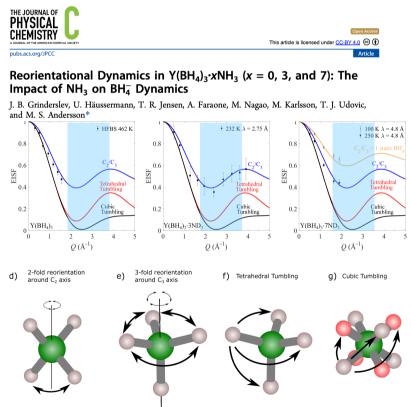
Data chain & workflow: Data scientist recruitment approved.

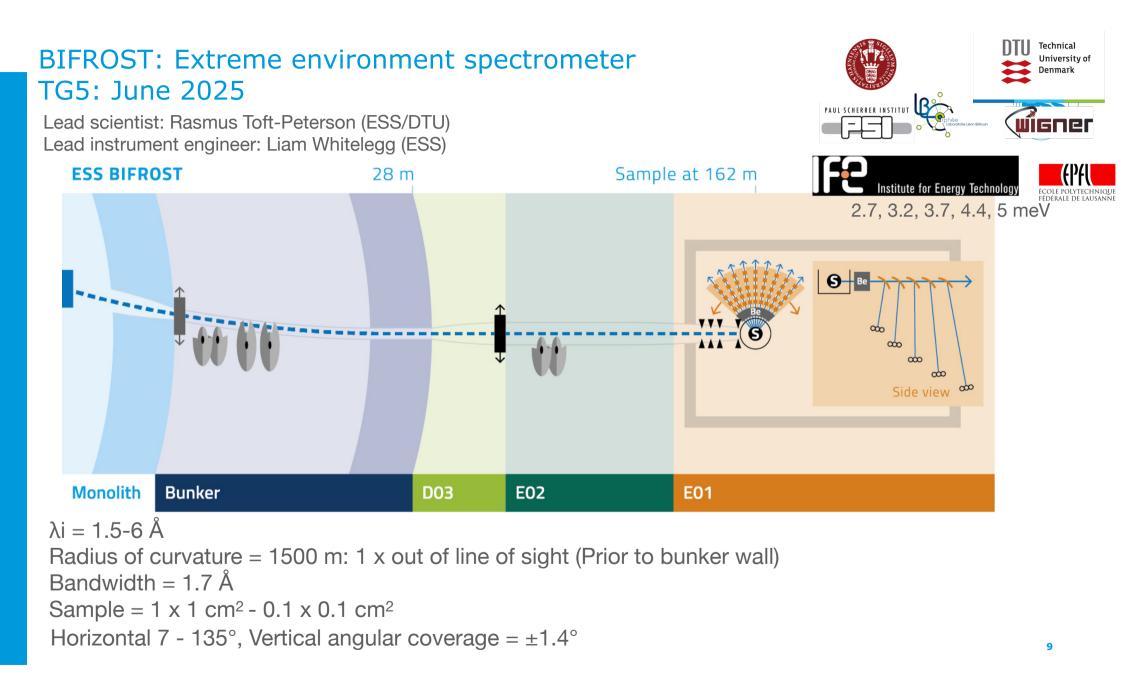
## MIRACLES: Rescoping

The completion of the instrument will include the full coverage of the scattering system: a second analyser, radial collimator and <sup>3</sup>He detectors on the right side.

50% of right-side analyzer with Si(111) crystals (50%+), 50% utilize Si(311) peaks, similar to other backscattering instruments (BASIS, DNA) and extend the Q-range to 3.8 Å. Si(333) reflections from Si(111) crystals.







# BIFROST: Extreme environment spectrometer Progress:



Sample environment testing and installation in cave





Detector testing and read out chain

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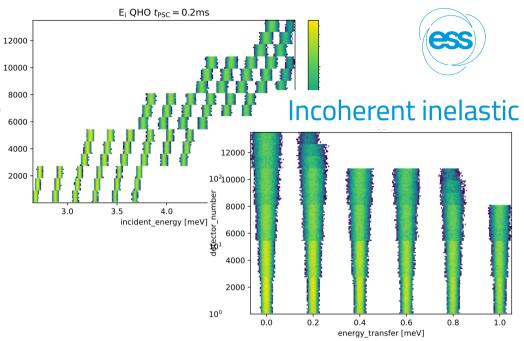
## Bifrost: Extreme environment spectrometer Data pipeline (IDS: Greg Tucker) (current input : McStas data)

#### Workflow requirements

Status	Description
$\checkmark$	propagate frame_time to sample position
$\checkmark$	calculate final (energy, wavelength, wavenumber, wavevector, velocity) from secondary geometry
$\checkmark$	unwrap frame_time at sample to find wall_time at sample
$\checkmark$	calculate incident (energy, wavelength, wavenumber, wavevector, velocity) from chopper cascade
$\checkmark$	calculate energy transfer and lab-frame momentum transfer
$\checkmark$	convert lab-frame momentum transfer to sample-table momentum-transfer
ongoing	normalize by incident wavelength-dependent flux
ongoing	normalize by detector efficiency, analyzer reflectivity, etc.
to-do	scale normalized intensity to absolute intensity units

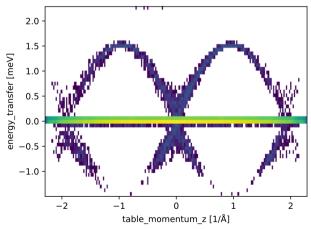


Status	Software	Advantages	Disadvantages				
$\checkmark$	scipp+	event-based data, simple fitting, plotting	no high-level spectroscopy utilities				
ongoing	Horace	extensive tof suite, resolution	detector-trajectory histograms				
to-do	MJOLNIR	Bragg scattering spurions, visualization	flattened (Q <sub>x</sub> , Q <sub>y</sub> , E) histogram				
	Mantid	event-based data					



detector\_number

## Coherent inelastic



BIFROST: Extreme environment spectrometer Concerns & actions

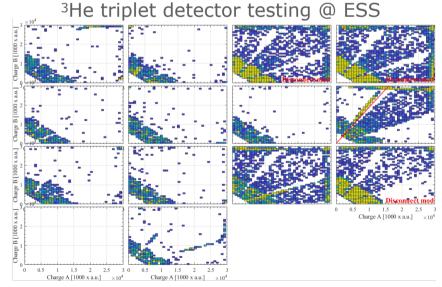


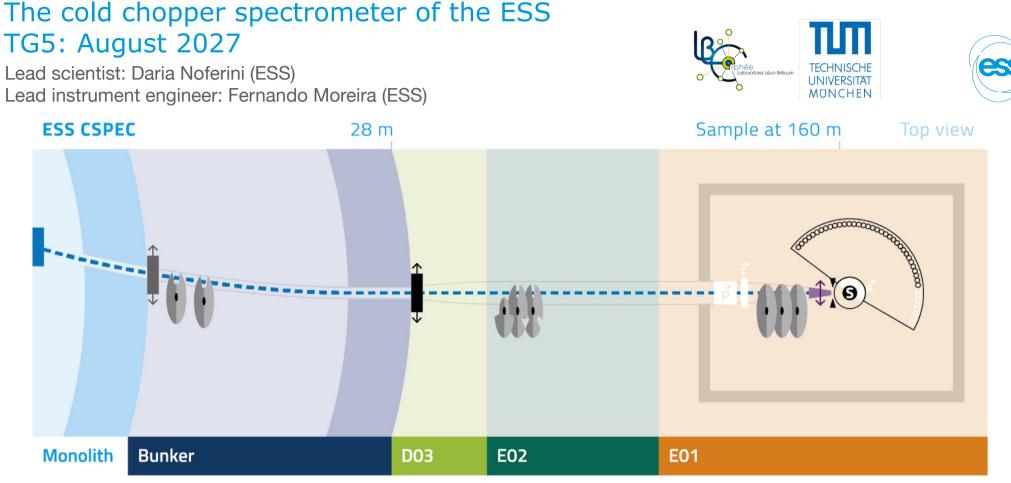
Cold commissioning - network to 1 instrument (Bifrost) Energised only in D01 - not D03/E02/E01.

Without this holistic overview testing of instrument is not possible

Focus on PA is currently a distraction.

First science workshop: 4/5th February 2025. Engagement from community





•Cold neutrons (2-20 Å),  $\Delta E/E = 1.5 \% @ 4 Å$ .

•Focus flux on range of sample areas  $4 \times 2 \text{ cm}^2 \rightarrow 1 \times 1 \text{ cm}^2$ .

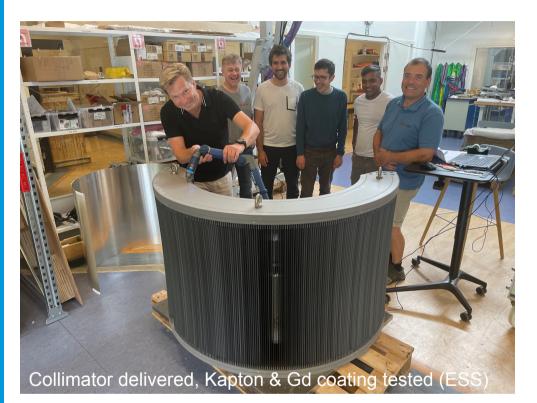
•Signal to noise =  $10^5$  (@5 Å, Vanadium).

# The cold chopper spectrometer of the ESS Progress:

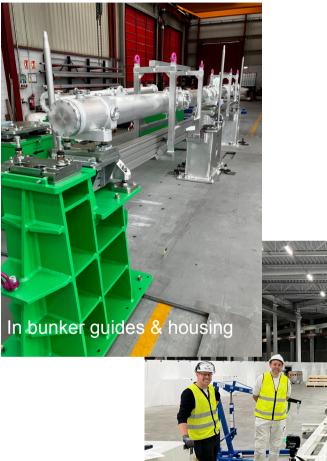




# The cold chopper spectrometer of the ESS Progress:



Installation of cave structure & hutch etc.. (as we speak)

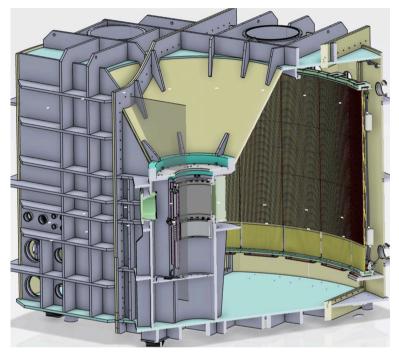






CSPEC Detector kick-off meeting (contract signed): 16th/17th October 2024 ILL-ESS partnership for multitube <sup>3</sup>He detectors.







Engineering work commenced prior to kick off.

Engineering implementation of detector modules nearly complete. Vanes not shown.

## CSPEC: The cold chopper spectrometer of ESS Concerns & actions



Working in a linear manner: Could really do with further engineering resources & technical writer to ensure timeline (beyond the IPL provided).

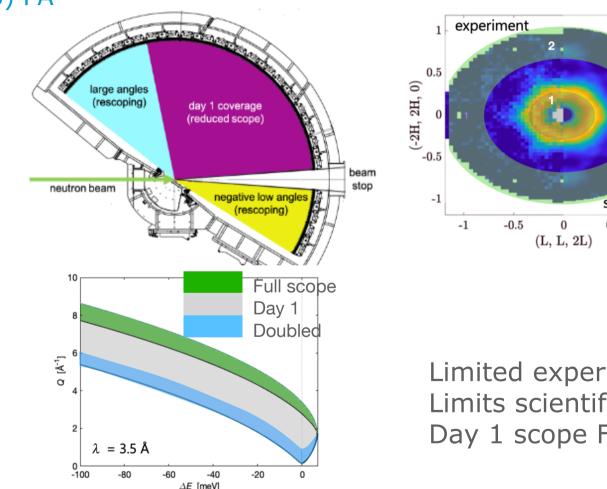
Common projects not prioritised for Tranche 2/3 instruments: electrical and conventional utilities, monitors, detector scope BUT team also do not have resources to focus on these.

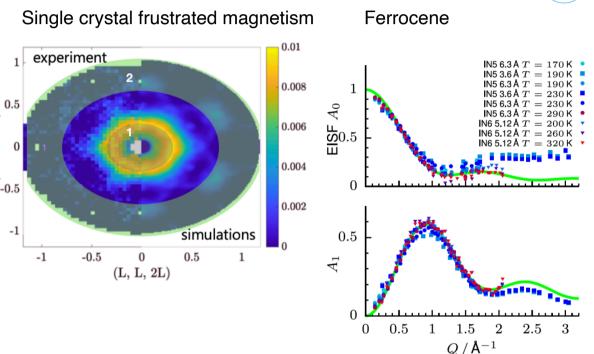
ESS documentation criteria and requirements are difficult for many of our partner companies and laboratories to fulfil - leads to delays. Working actively to readjust.

Critical path is detector delivery for 4 modules at TG5, start with 1.

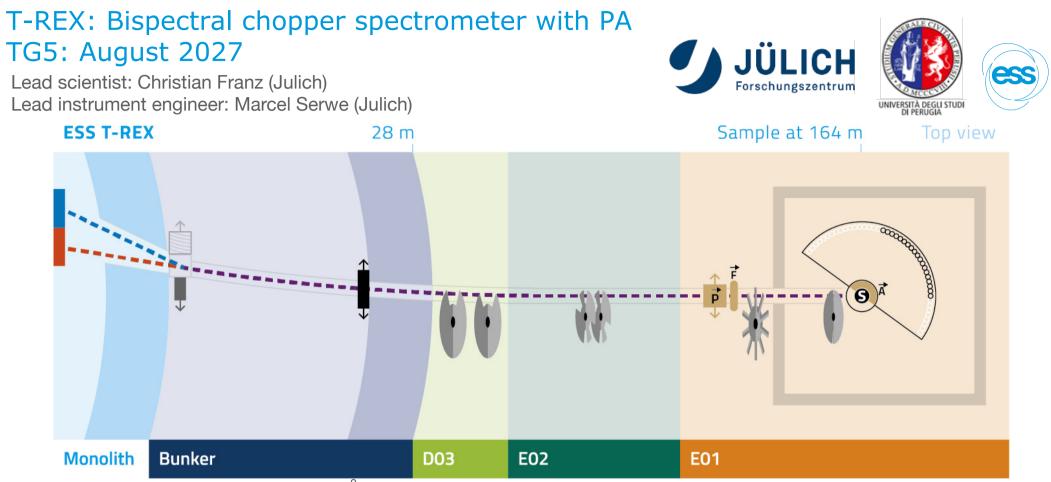
#### Rescoping CSPEC: detector coverage from 58% to 100 % (1) **<sup>3</sup>He gas filling for complete detector** (2.5 bar <sup>3</sup>He pressure)

(2) Extend to 5 bar <sup>3</sup>He pressure(3) PA





Limited experimental signatures & Q range Limits scientific capability Day 1 scope FOM is not world leading.



•Bispectral extraction (0.7-6.5 Å, 2 - 160 meV), 1%–7% energy resolution.

Focus flux on sample area 3 x 1 cm<sup>2</sup> (Div: ±0.25° in the thermal band to ±1° in the cold regime).
Polarisation using PASTIS XYZ set-up.

Novel B10 detector technology

TREX: The bispectral chopper spectrometer with PA. Progress:

Ittel







## TREX: The bispectral chopper spectrometer with PA. **Progress:**

Extensive activity on Multigrid detector

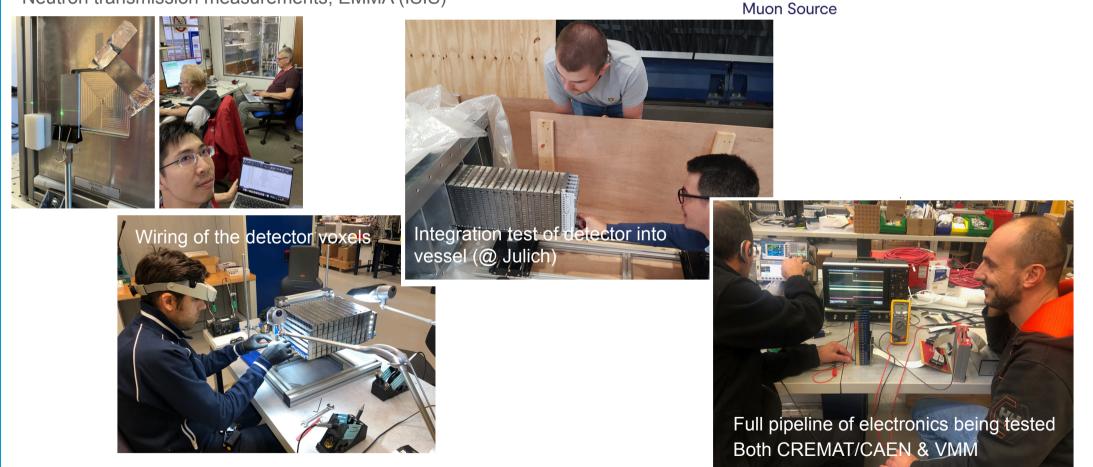
Neutron transmission measurements, EMMA (ISIS)



Science and



Technology **Facilities** Council **ISIS Neutron and** 



TREX: The bispectral chopper spectrometer with PA. Concerns & actions



Common projects not prioritised for Tranche 2/3 instruments: electrical and conventional utilities, monitors, detector scope.

ESS documentation criteria and requirements are difficult for many of our partner companies and laboratories to fulfil - leads to delays. Working actively to readjust.

No staff member on site:

Recruitment completed for instrument scientist & will start 7th January 2025. Concomitant recruitment of engineering resources.

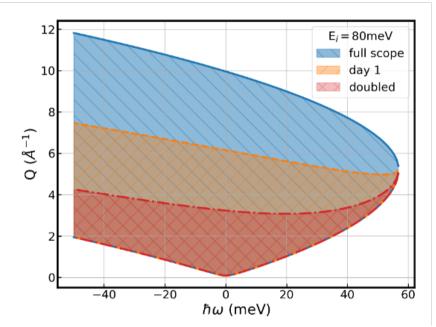
Critical path is detector delivery.

Data chain & workflow: no current focus.

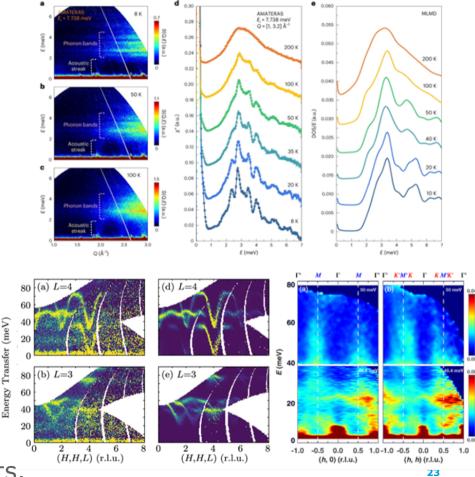
## Rescoping TREX: (1) Detector coverage from 40% to 100%

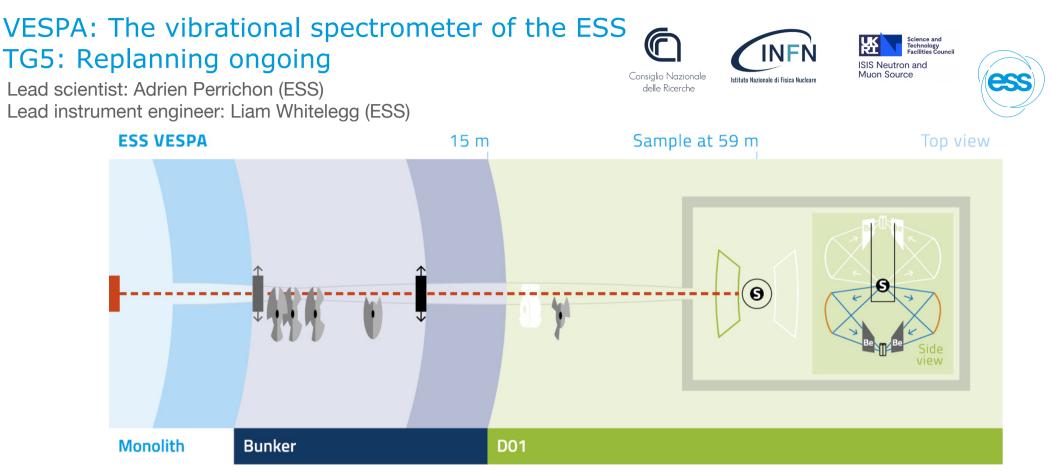


(2) Advanced pulse shaping option with 4 blade pulse shaping chopper
(3) T0 chopper



Limited experimental signatures & Q range Limits scientific capability Day 1 scope FOM is not world leading. Note possible DG loss of competence if one waits.





Direct view of the thermal moderator (Ei = 3 - 1000 meV)

Complex chopper cascade to achieve 3 subframes and variable resolution

Indirect-geometry crystal-analyser

Complex and compact secondary spectrometer:

280 <sup>3</sup>He tubes, 8500 HOPG crystals, 200 kg of cryo-cooled Be, in <2 m<sup>3</sup> vacuum vessel

VESPA: The vibrational spectrometer of the ESS TG5: Replanning ongoing



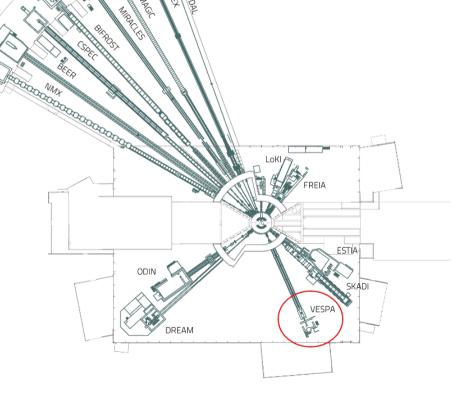
Staff recruitment in 2024

(January) Liam Whitelegg: Lead engineer (50%)(March) Adrien Perrichon: Lead Scientist(September) Rosa Camilleri Lledó: Lead Engineer(October) Gianfranco Belcastro (ISIS/CNR): Engineer[open] Mechanical engineer at ESS

STFC ISIS Facility collaborators S. Parker, R. Bewley, J. Nightingale, D. Raspino

# VESPA: The vibrational spectrometer of the ESS TG5: Replanning ongoing

Review, upgrading and finalisation of VESPA design Space claim for VESPA end-station Neutronics testing of HOPG tiles Rescoping



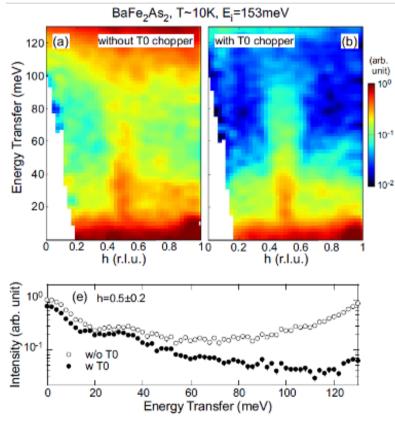
**Primary** Chopper cascade Slit system Guide gaps and substrates

Secondary spectrometer Update to TOSCA+ type geometry

**Cave & sample preparation area** Evaluation of alternative designs



### Rescoping VESPA: (1) T0 chopper (2) Rescoping secondary spectrometer



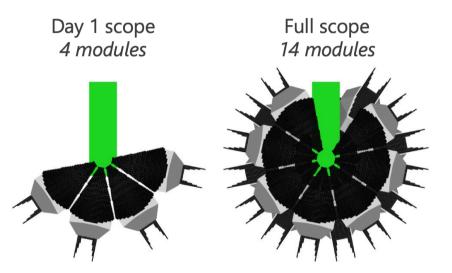


 Kajimoto et al., J. Phys. Soc. Jpn. 80 (2011) SB025



- Direct line-of-sight from moderator to sample
- •Flood of epithermal and high-energy neutrons
- •Secondary scattering events at unpredictable times
- •The lack of T0 chopper would ruin the SNR of VESPA

#### Rescoping VESPA: (1) T0 chopper (2) Rescoping secondary spectrometer



Scope = 11% of nominal performance (Sr × reduced source power)

High-Resolution setting of VESPA only competitive if the instrument is fully rescoped High-Flux setting of VESPA only state-of-the-art if the instrument is fully rescoped





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## Spectroscopy division: Rescoping priorities PA not considered Sample environment considered in the Scientific support division.



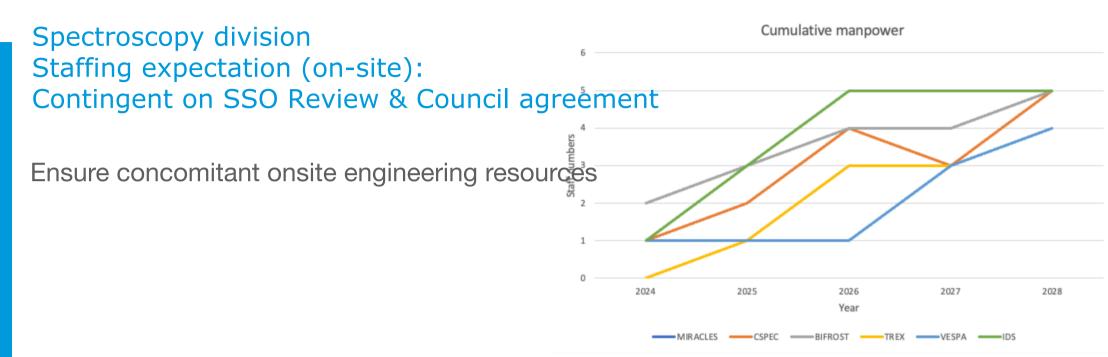
Instrument	Rescoping option	A: Scientifi	B: Manpower (1: high,	C: Technical	B: Cost	FOM: A*B*
		c Impact	5: low)	Risk (1: high,		С
		(5:		5: low)		
		high, 1: low)				
MIRACLES	Complete Analyser/collimat or/detector coverage Si (1 1 1) & (3 1 1)	2	Can be included and delivered within current instrument construction: 2	5	1.67 M€	20
CSPEC	<sup>3</sup> He full detector coverage (2.5 bar)	5	Minimal: 5	5	2.025 M€	125
CSPEC	<sup>3</sup> He full detector coverage (5 bar)	2	Minimal: 5	5	+4.375 M€	50
T-REX	Full detector coverage	5	Significant (Detector group) 2	4	>5 M€	40
T-REX	Extra blades for P chopper	2	Julich chopper group: 3	2	0.5 M€	12
T-REX	T0 chopper	х	x	x	0.85 M€	?
VESPA	T0 chopper	5	Some (Chopper group) 4	4	0.85 M€	80
VESPA	Increase analyser modules	4	Can be included and delivered within current instrument construction: 4	4	3.6 M€ for 10 (0.36 each)	64



## **ESS Spectroscopy division**

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		2024						2028					
	'												
Instrument	TG5 Date	Scientist 1	Scientist 2	Scientist 3	IOE	Post-Doc	Data scientist	Scientist 1	Scientist 2	Scientist 3	IOE	Post-Doc	Data Scientist
MIRACLES	Q2 2027	0	(	ז כ	0 0	1	0 0	1	. 1	. 1	. 1	0	1
CSPEC	Q1 2027	1	. (	) נ	) 0	1	0 0	1	. 1	. 1	. 1	1	. 1
BIFROST	Q3 2025	1	. (	) נ	) 1		0 1	. 1	. 1	. 1	. 1	1	. 1
TREX	Q2 2027	0	(	) נ	) 0	1	0 0	1	. 1	. 1	. 1	0	1
VESPA	Q1 2028	1	. (	0 0	0 0	1	0 0	1	. 1	. 1	. 1	0	1
Spin echo	??												
Bold means C	Q3/4	Red is inkind											
Data Scientist	t												
Total for year	r					Spec Division	4				'	Spec Division	n 22
Note: VESPA Scientist 2 can be a computational material scientist.				IDS	1					IDS	5		
Note: Bifrost scientist 2 is in-kind from KU												32	



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