#### HELMHOLTZAI ARTIFICIAL INTELLIGENCE COOPERATION UNIT

# SAMPLING NEUTRONS WITH ARTIFICIAL INTELLIGENCE

José I. Robledo

Jülich Supercomputing Centre (JSC)

Jülich Centre for Neutron Science (JCNS)

Forschungszentrum Jülich (FZJ)





## INTRODUCTION



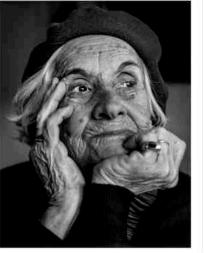


























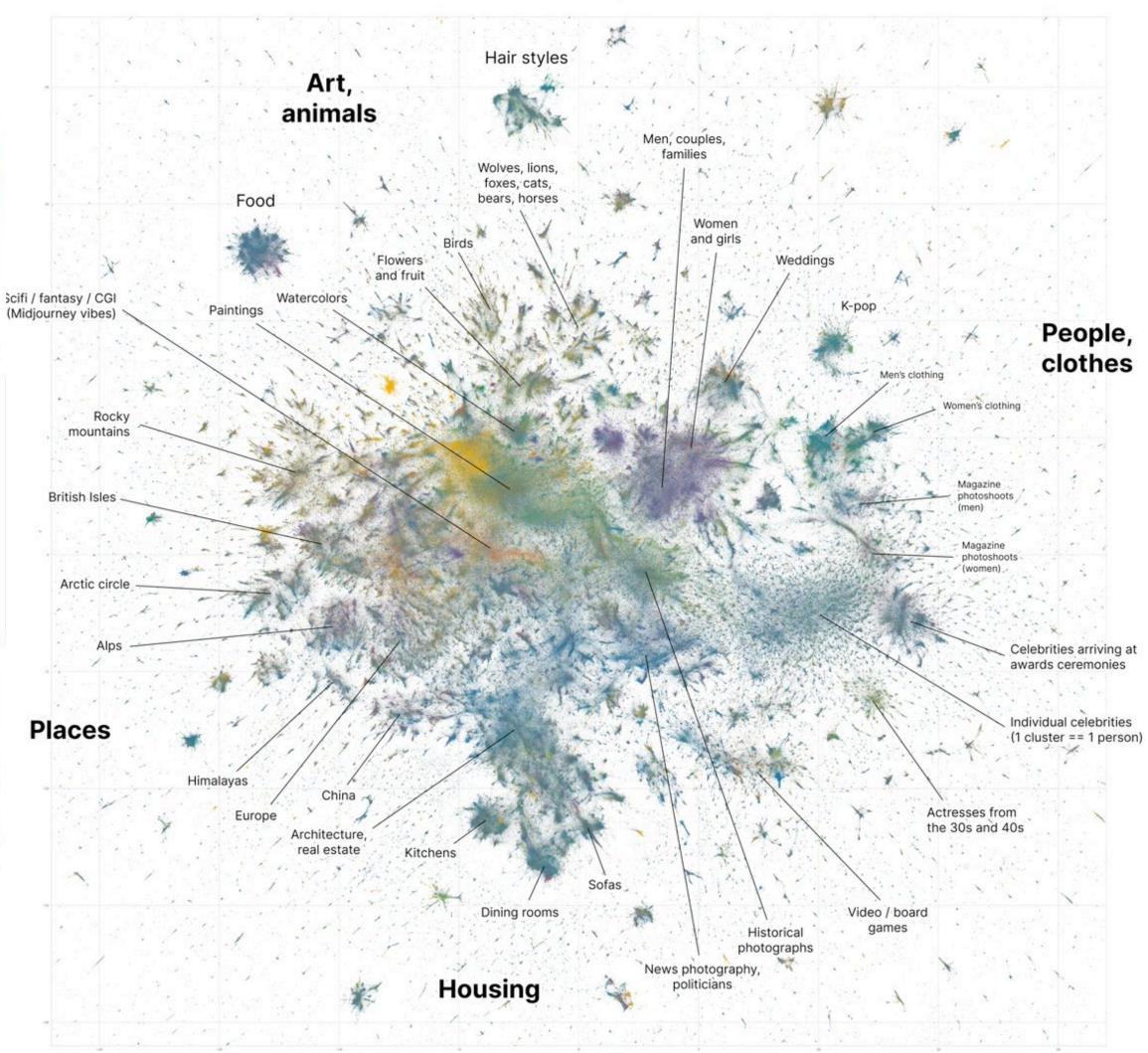






#### All captions from LAION-Aesthetics with score > 6 (n=12M)

Embedded with CLIP, UMAP to 2d

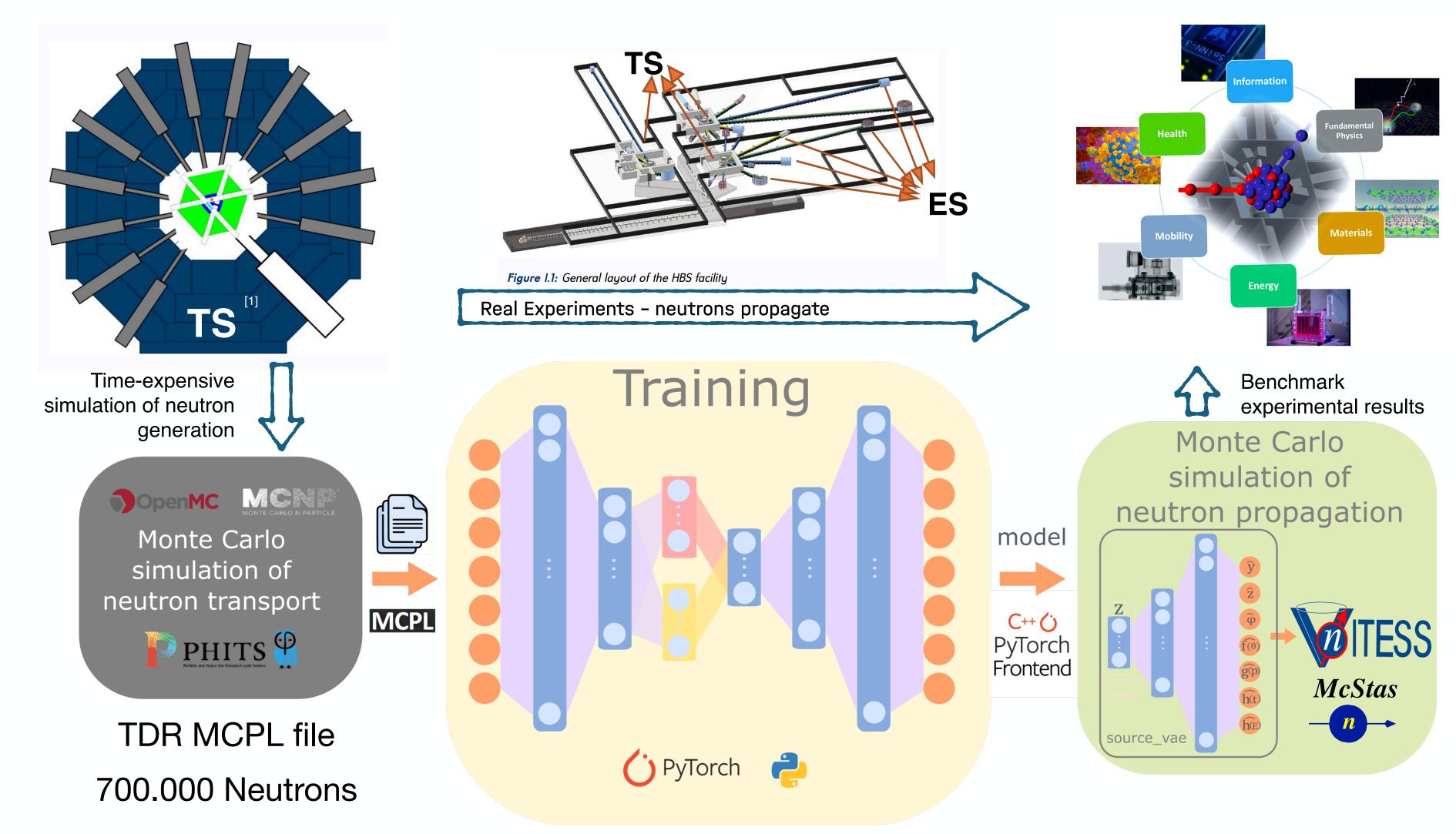


Why can't we generate neutrons?



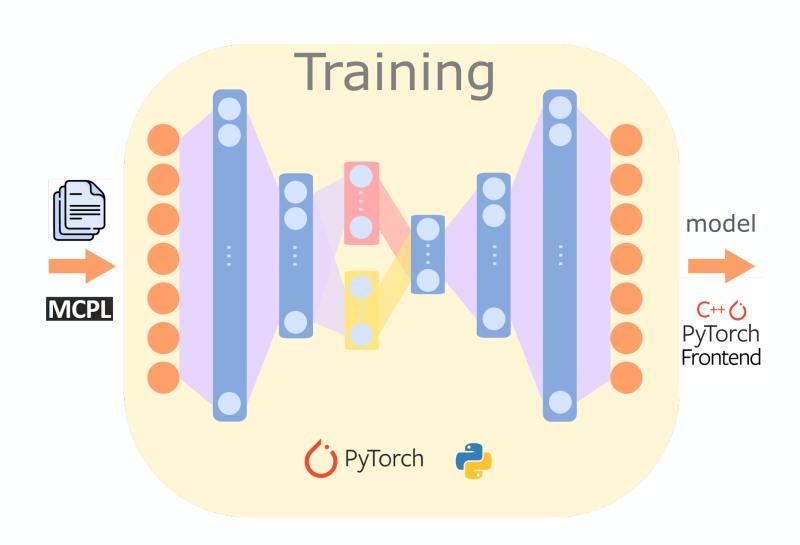


## INTRODUCTION





### TRAINING DATA: MCPL FILES

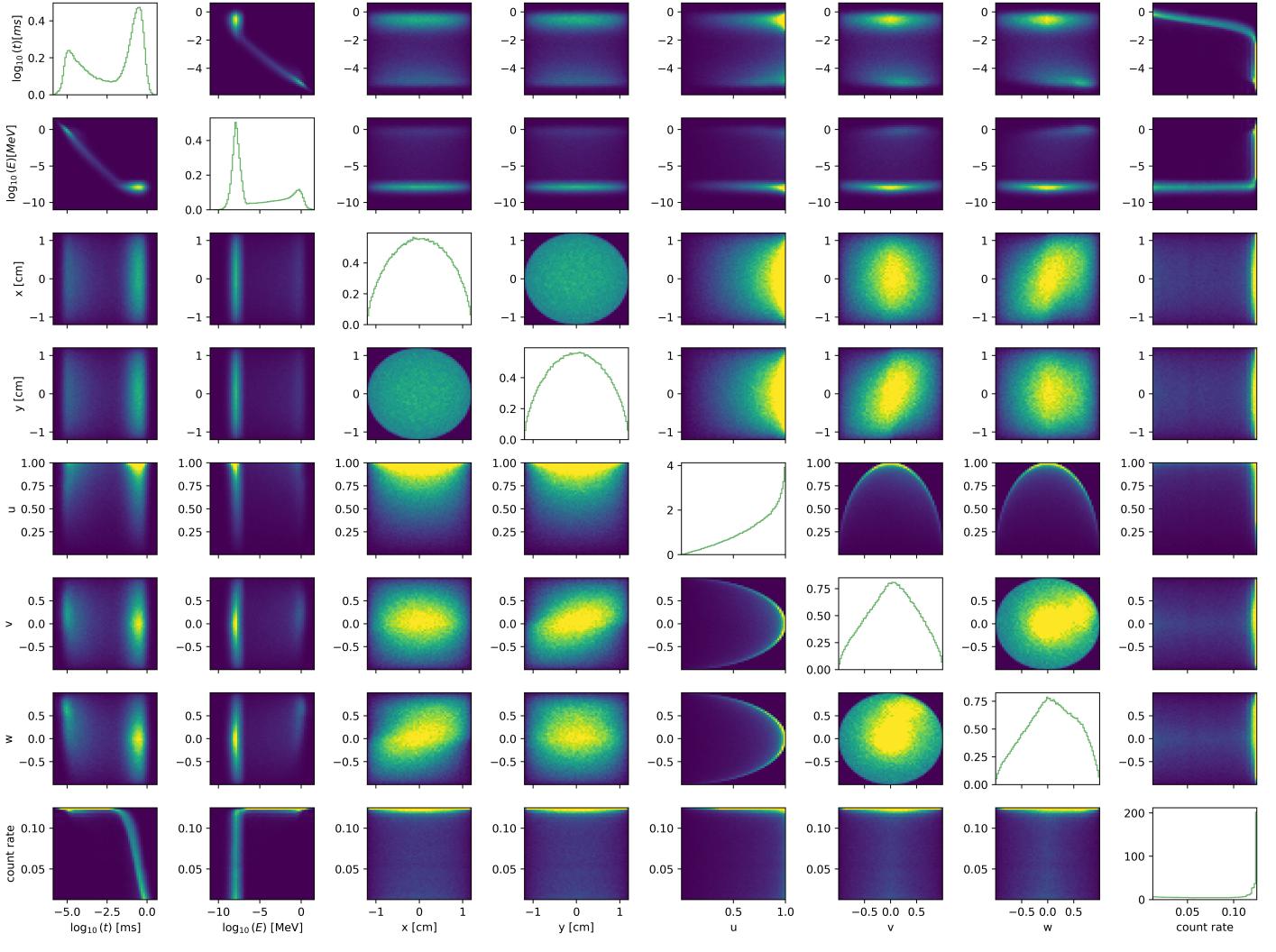


### PyTorch DataLoader Variables

\/	[cm]	and	7	[cm]
У		anu	_	[cm]

- u, v, and w
- count rate (weight)
- $\log_{10}(t)$  [ms]
- $\log_{10}(E)$  [MeV]

	pos_y	pos_z	u	v	w	count_rate	TOF	E
0	-1.157290	-0.156233	0.974585	-0.203838	-0.092924	0.124997	0.000059	9.301730e-02
1	-0.675784	0.331821	0.534119	-0.218695	0.816633	0.074287	0.157121	8.421972e-08
2	0.598687	-0.840697	0.647142	0.740285	-0.182169	0.074748	0.156539	8.232719e-08
3	0.775426	-0.258992	0.508785	0.752556	0.418089	0.075199	0.156472	1.423273e-07
4	-0.778382	0.114899	0.702195	-0.399447	-0.589376	0.122384	0.007535	5.222545e-08
676553	-0.711618	-0.400495	0.916834	-0.103239	-0.385690	0.124989	0.000039	8.381045e-03
676554	-0.052427	0.881021	0.818265	0.572754	-0.048949	0.124998	0.000005	2.565348e-01
676555	-0.042421	0.981967	0.545910	-0.667388	0.506533	0.025299	1.222780	2.869063e-09
676556	0.244454	-0.181313	0.743137	-0.051606	0.667146	0.080170	0.105039	2.322999e-08
676557	-1.063570	-0.350045	0.633484	-0.657062	-0.408617	0.032235	0.657213	6.777769e-09
676558 ro	ws × 8 colum	ne						



T. Kittelmann et al. Monte Carlo particle lists: MCPL. Computer Physics Communications, 218, 17-42.



bw=0.05, N=100

bw=0.2, N=100

bw=0.5, N=100

Possible values

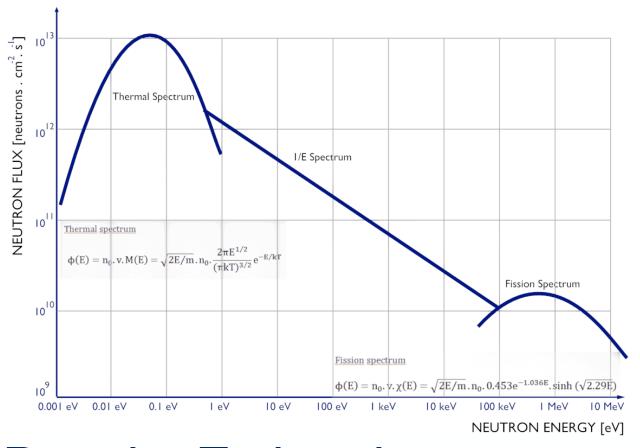
Eveduency 0.5 0.3

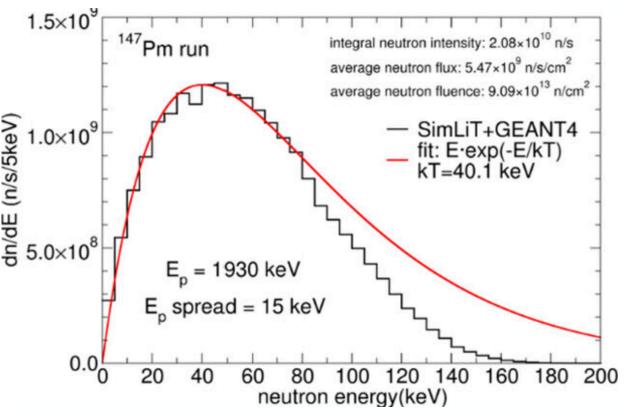
Relative 0.1

Frequency 8.0

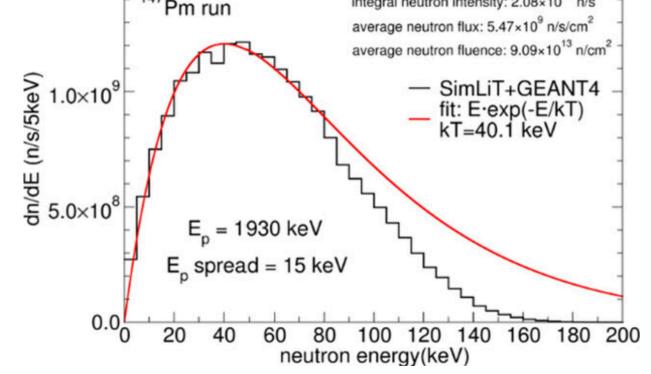
#### CURRENT APPROACHES TO SOURCE ESTIMATION

Analytical approximation based on theory / fitting to observed data









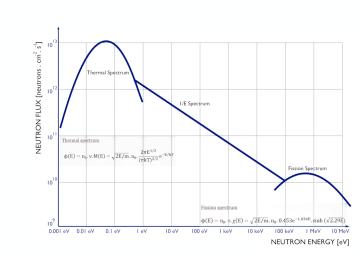
$$\hat{f}(\mathbf{x}) = \hat{f}(x_1, x_2, \dots, x_D) = \sum_{i=1}^{N} w_i \left\{ \prod_{j=1}^{D} \frac{1}{h} K\left(\frac{x_j - (\tilde{p}_i)_j}{h}\right) \right\}$$

Schmidt, N. S. et al. (2022), KDSource, a tool for the generation of Monte Carlo particle sources using kernel density estimation. *Annals of nuclear energy*, 177, 109309.



### PROS & CONS

- Analytical approximation based on theory / fitting to observed data
  - Reliable when assumptions are adequate (Theoretical foundation)
  - It is an approximation and assumptions are needed
  - simple and fast for sampling
  - lack features specific to individual cases
  - Fitting parameters are characteristic of the distribution (interpretability)
- Kernel Density Estimation
  - Non-parametric approach, making it adaptable (flexibility)
  - Although, hyper-parameter: Bandwidth and kernel
  - Data-driven, potentially providing a better representation of the distribution
  - Computational cost for high dimensional spaces
  - Fast sampling
  - Data dependency for sampling





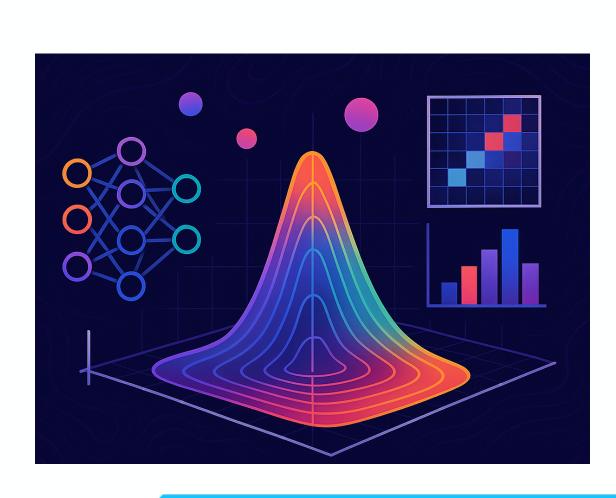


### GENERATIVE MODELS

 Objective: Learn the underlying patterns and distributions of a dataset and generate new data points that resemble the original

# We can use generative models to learn the multivariate phase-space distribution of neutrons from a Monte Carlo Particle List (MCPL file)

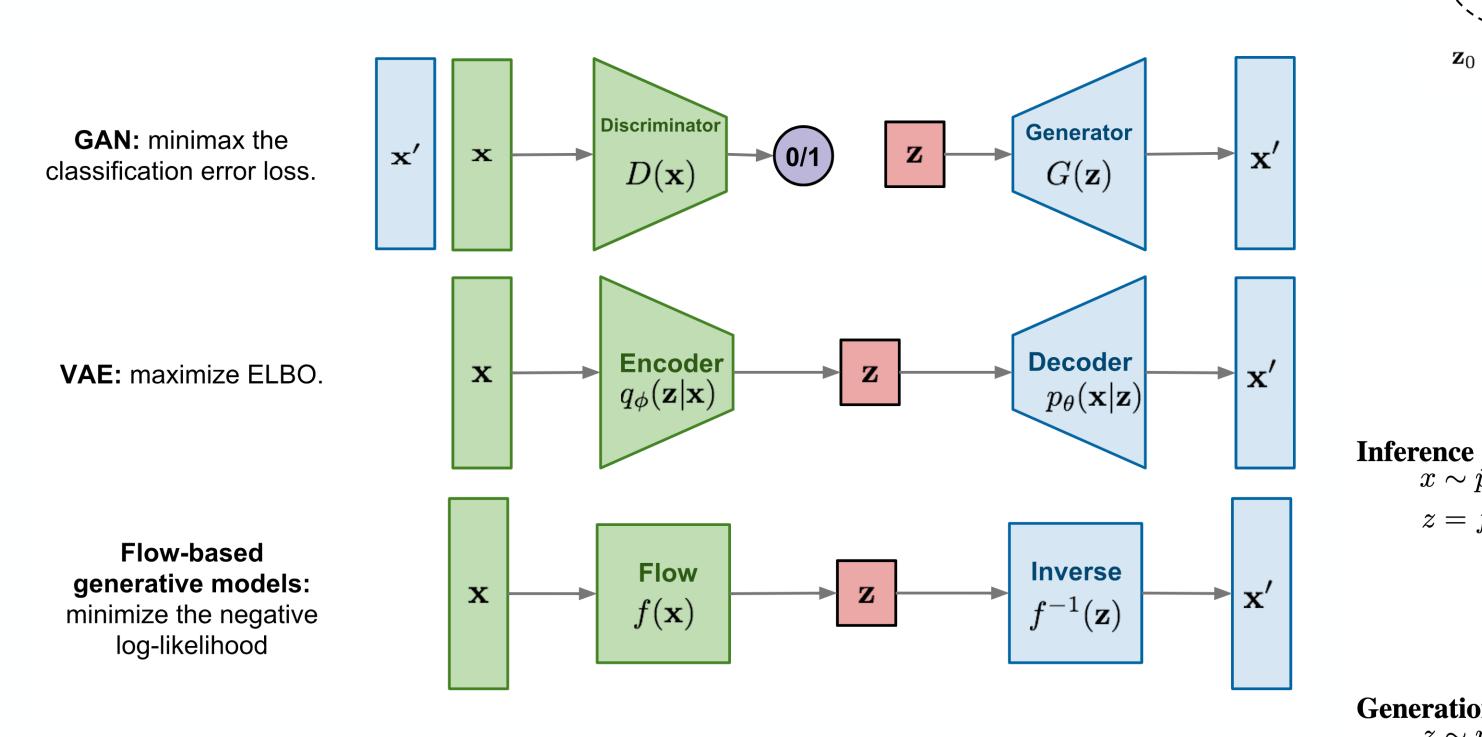
- Easy to generate new neutrons to propagate inside tracing software.
- capable of learning and generating data with complex patterns and structures between phase-space variables! (High fidelity and realism)
- Not limited to specific types of data (Flexibility)
- · Once trained, no need to keep the original dataset used to train it.
- High-computational cost for training
- Large datasets
- Lack of interpretability
- Depending on model size, slow for sampling

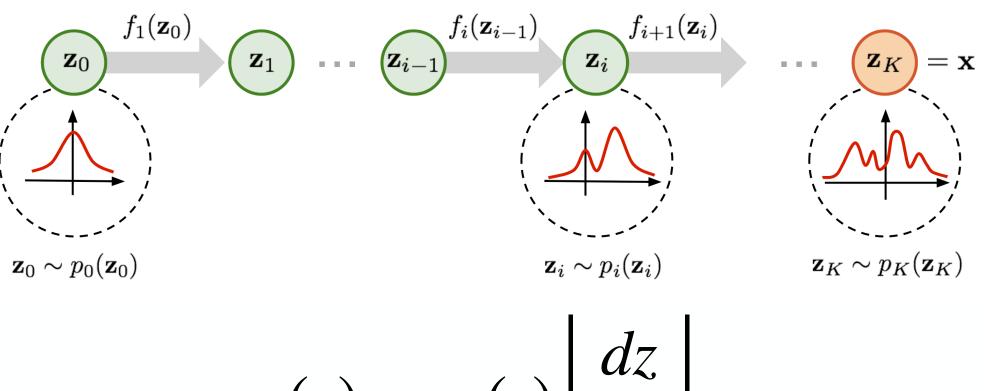


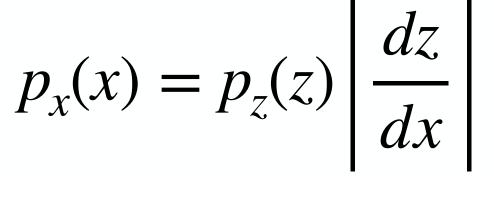


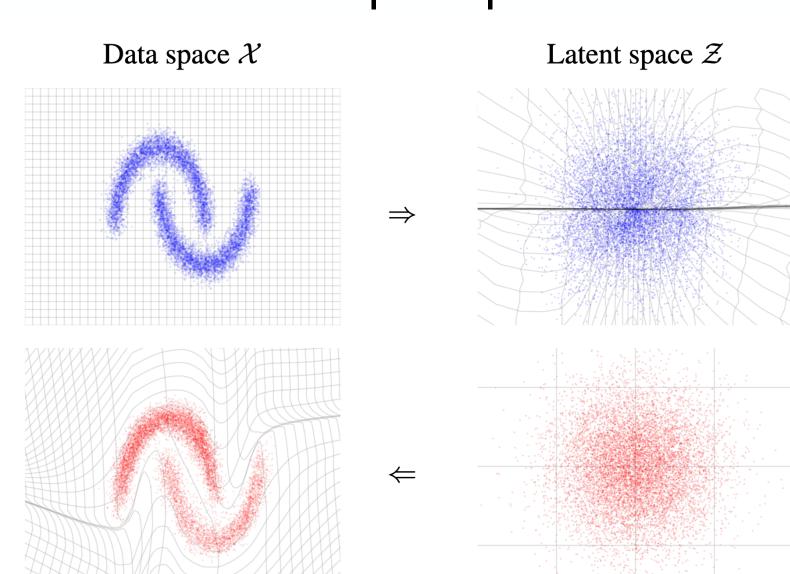
#### **Normalizing Flows**

## GENERATIVE MODELS









Generation

 $x \sim \hat{p}_X$ 

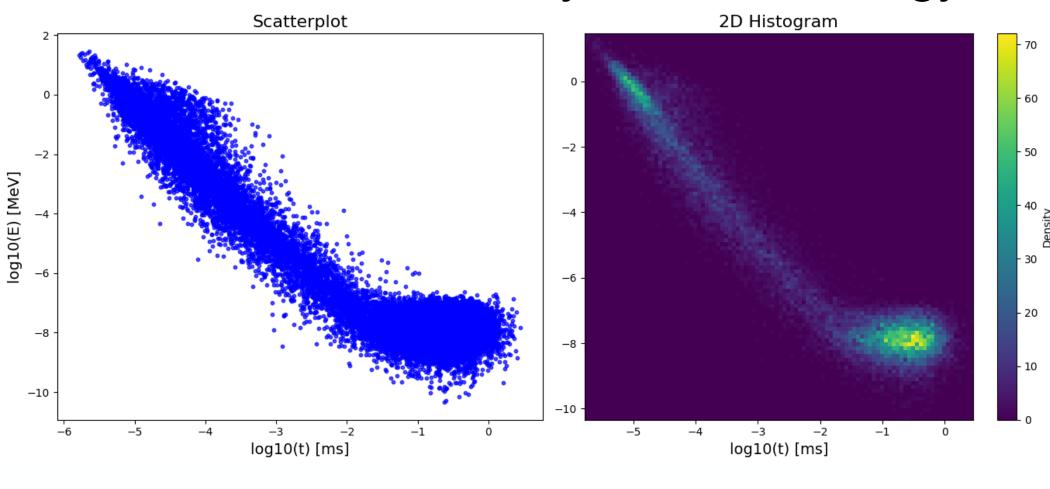
 $z = f\left(x\right)$ 

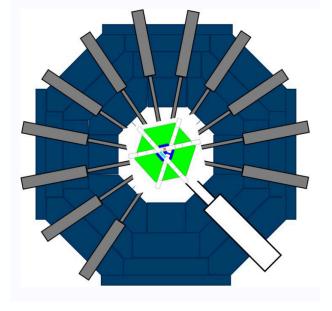
$$z \sim p_Z$$
$$x = f^{-1}(z)$$

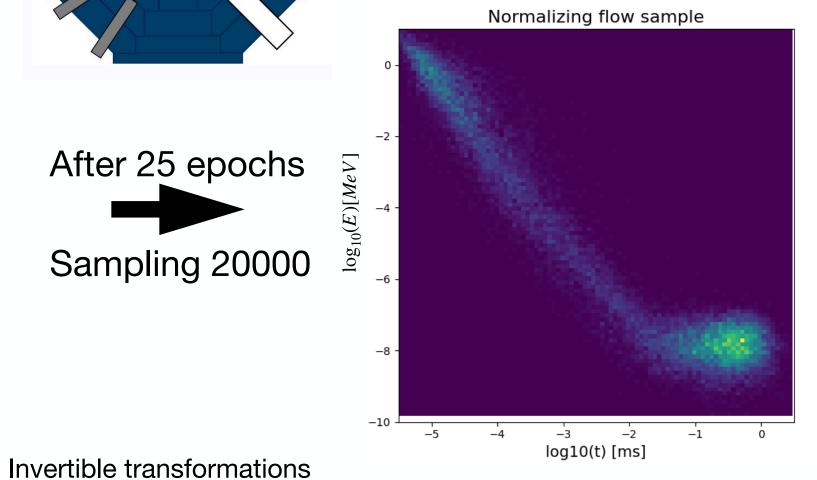
#### HELMHOLTZAI Artificial Intelligence Cooperation Unit

### **TOY EXAMPLE ON 2D**

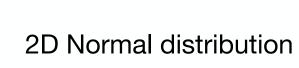
20.000 neutrons, only time and energy



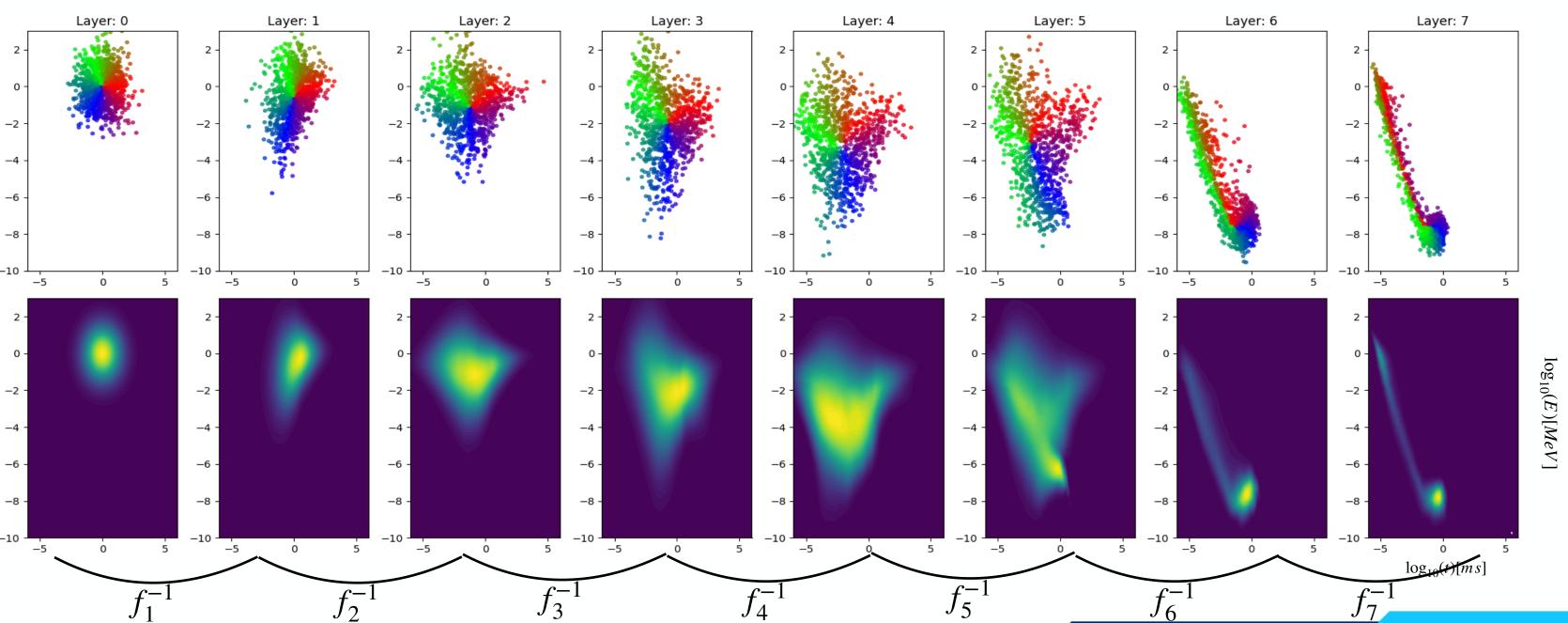








Easy to sample



Artificial sample





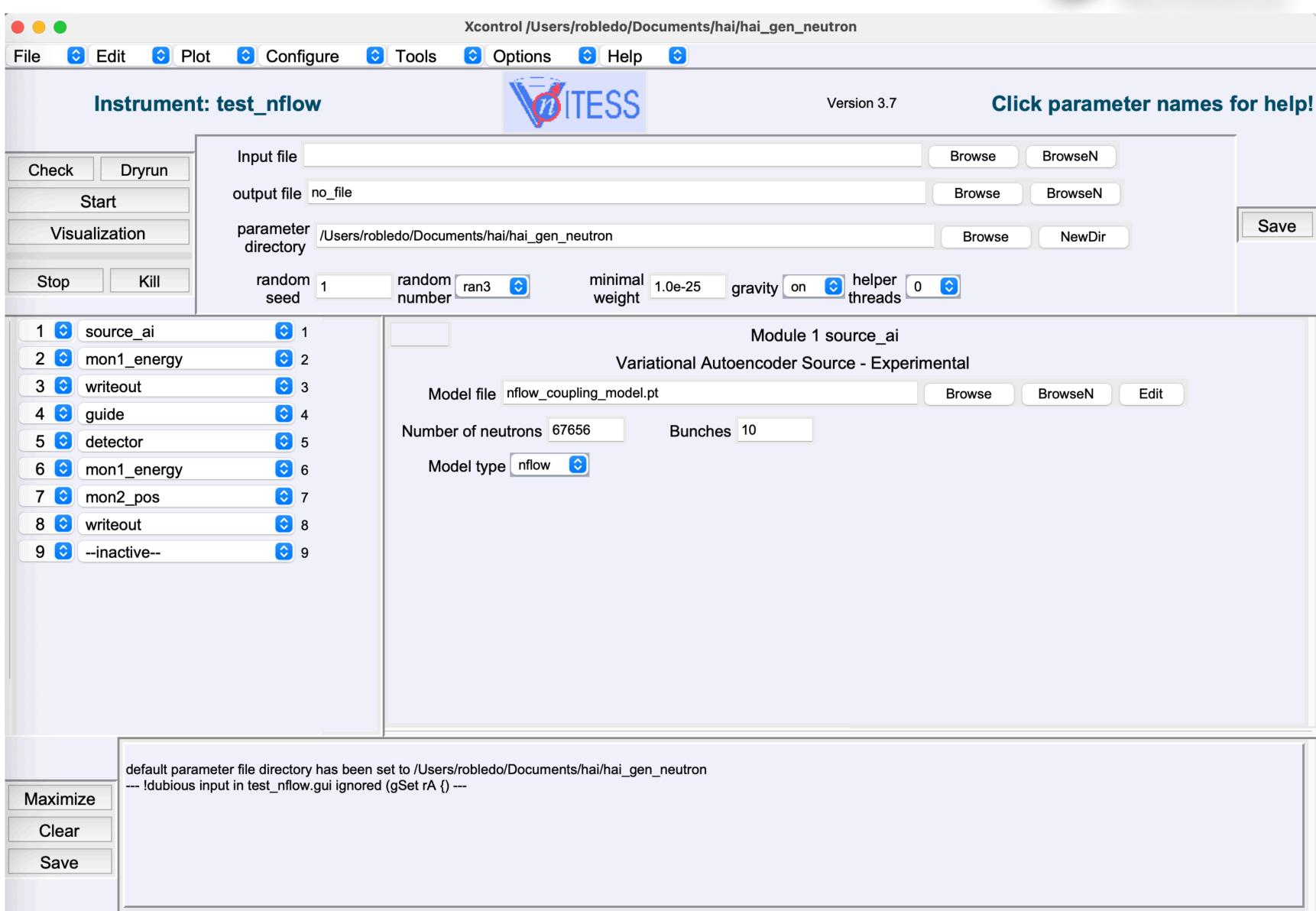
### VITESS

New source\_Al module



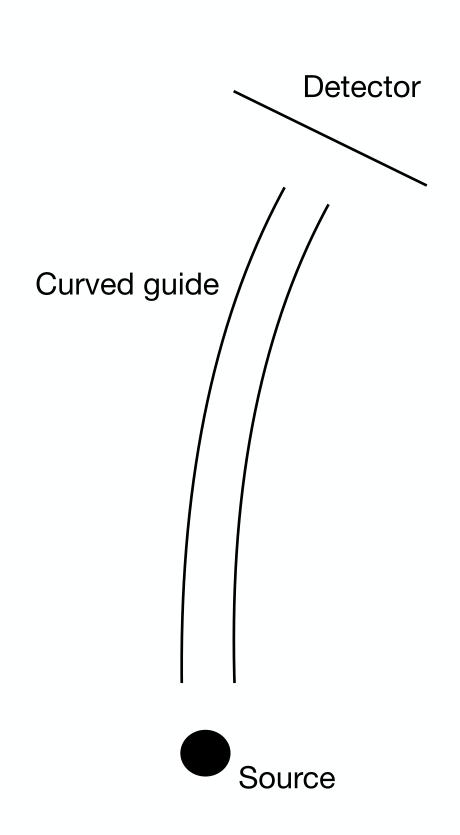
C++ Frontend

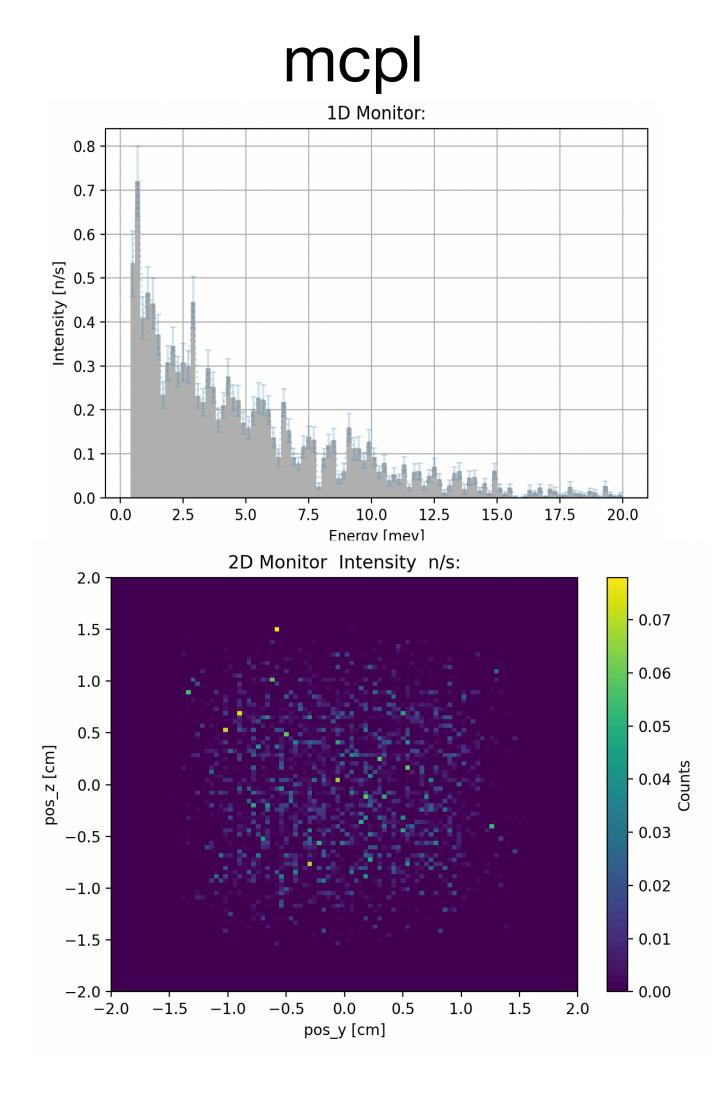
Model needs to be jit compiled



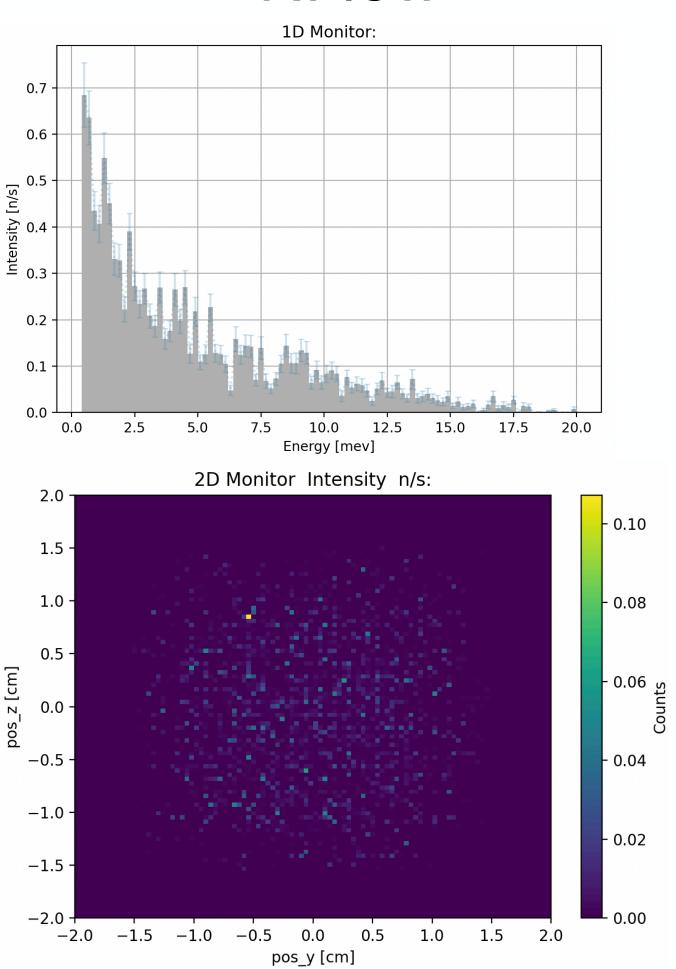


## SAMPLING ON VITESS





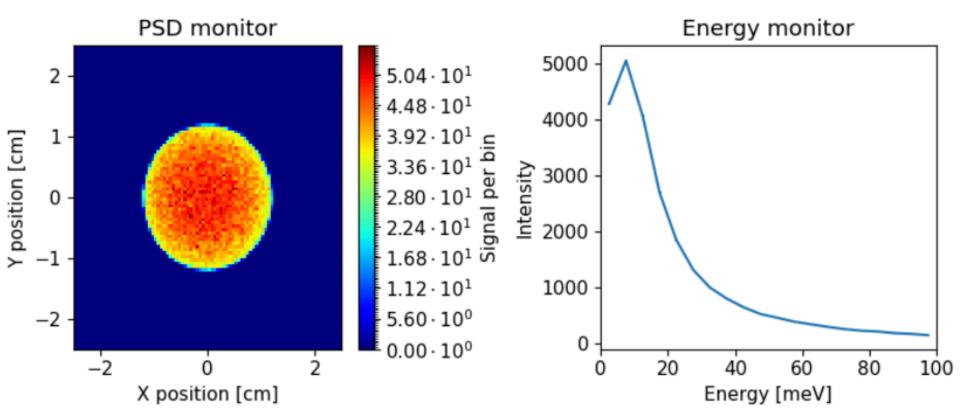
#### NFlow

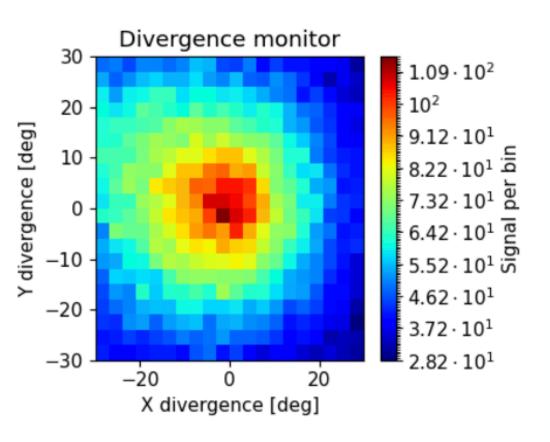




### MCSTAS

```
import np2mcpl
# Load model
model = torch.load("nflows_model.model", map_location=torch.device('cpu'), weights_only=False)
# sample model
samples = model.sample(int(1e7))
# Transform to adapt to mcpl format if necessary
\bullet \bullet \bullet
# Save data
np2mcpl.save("output", samples)
# Load data
import mcstasscript as ms
instrument = ms.McStas_instr("sample_normalizing_flow")
source = instrument.add_component("source", "MCPL_input")
source.filename = '"output.mcpl.gz"'
PSD = instrument.add_component("PSD", "PSD_monitor")
PSD.set_AT([0, 0, 0.2], RELATIVE=source)
PSD.set_parameters(xwidth=1, yheight = 1, filename='"PSD.dat"')
data = instrument.backengine()
ms.make_sub_plot(data)
```





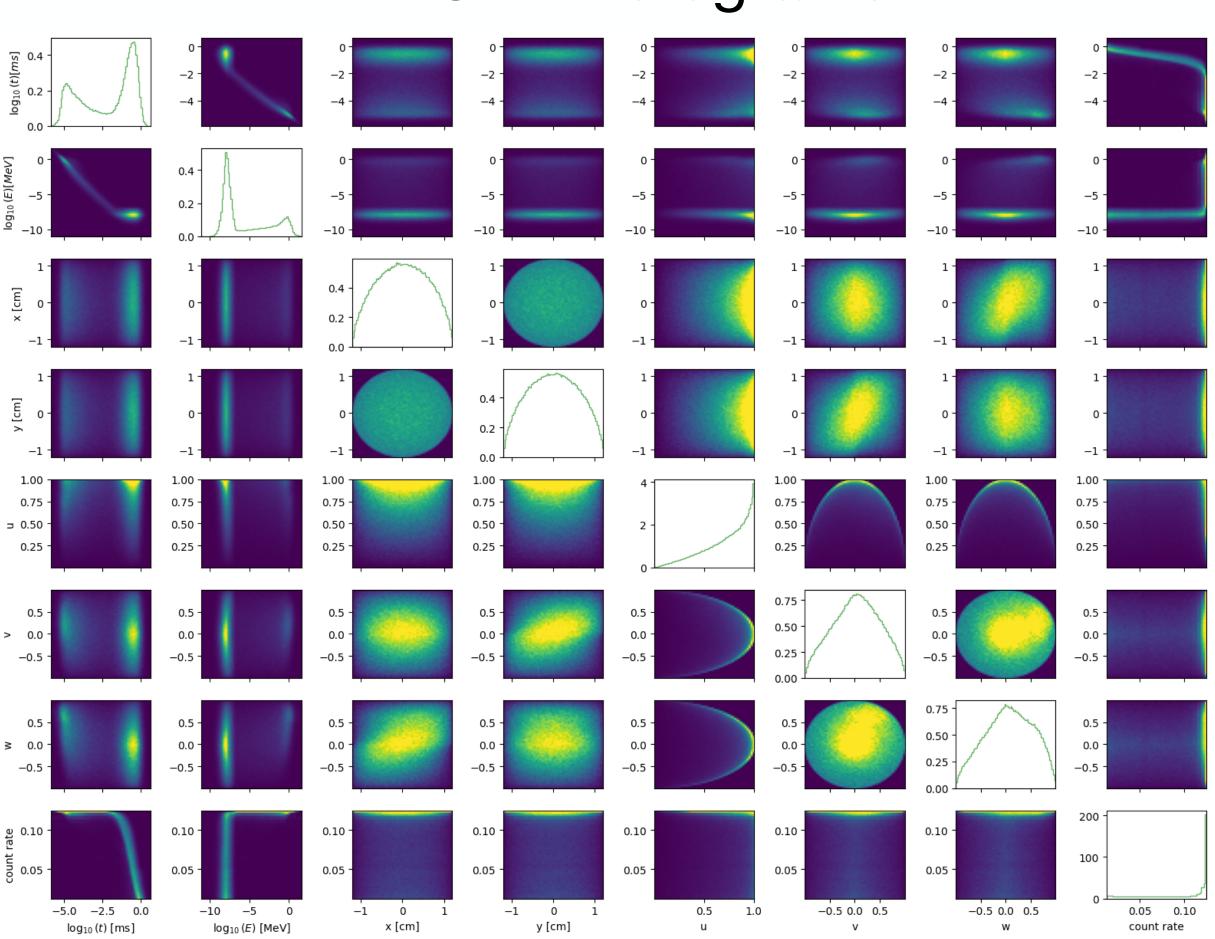
Same procedure can be done in vitess-python

Poster: "Vitess-Python: A Python API for VITESS Instruments" - Fabian Beule

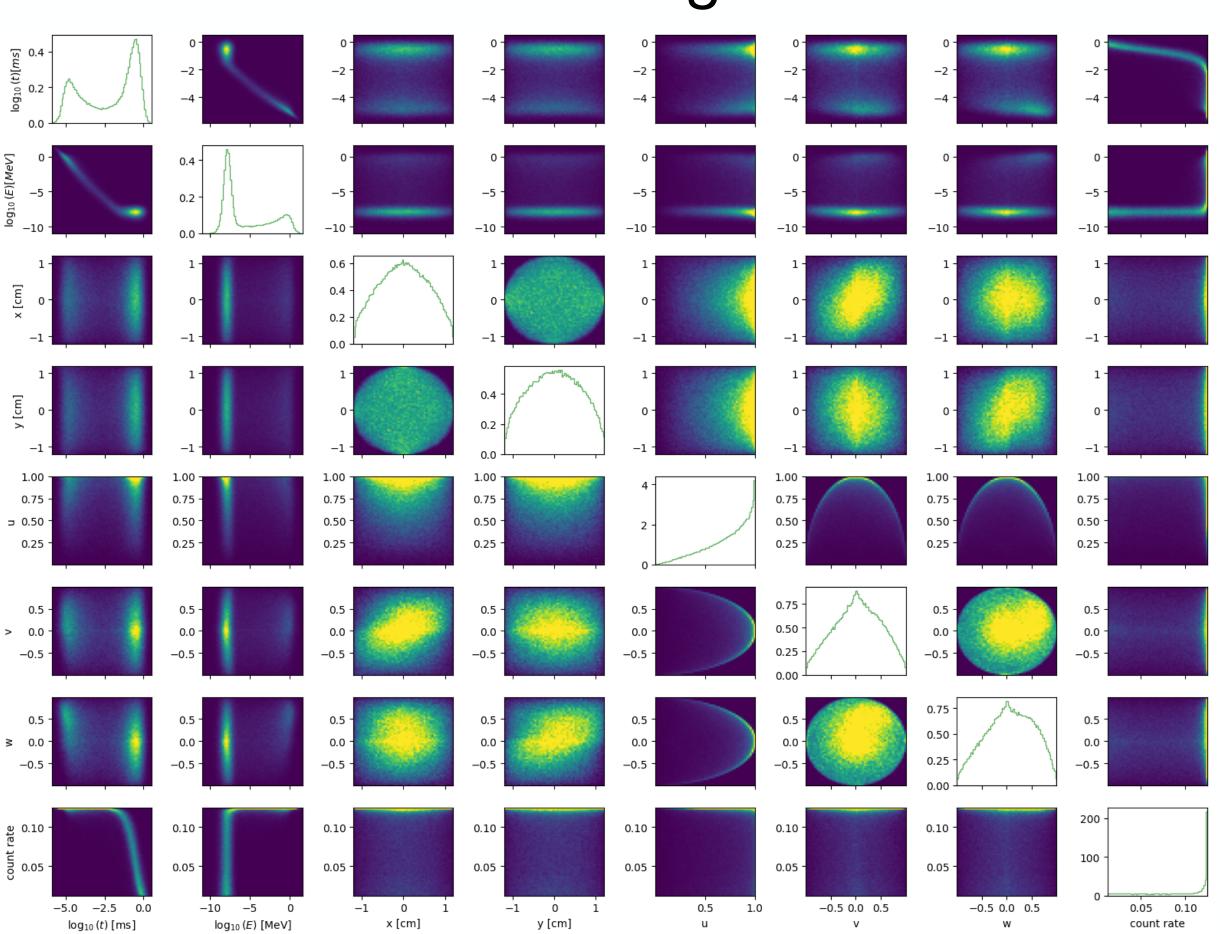


#### Comparison between NFlow and MCPL

### MCPL histograms



#### NFlow histograms



# JÜLICH FORSCHUNGSZENTRUM

0.00100

0.00075

0.00050

0.00025

0.00000

Density

-0.00050

-0.00075

-0.00100

#### Comparison between NFlow and MCPL

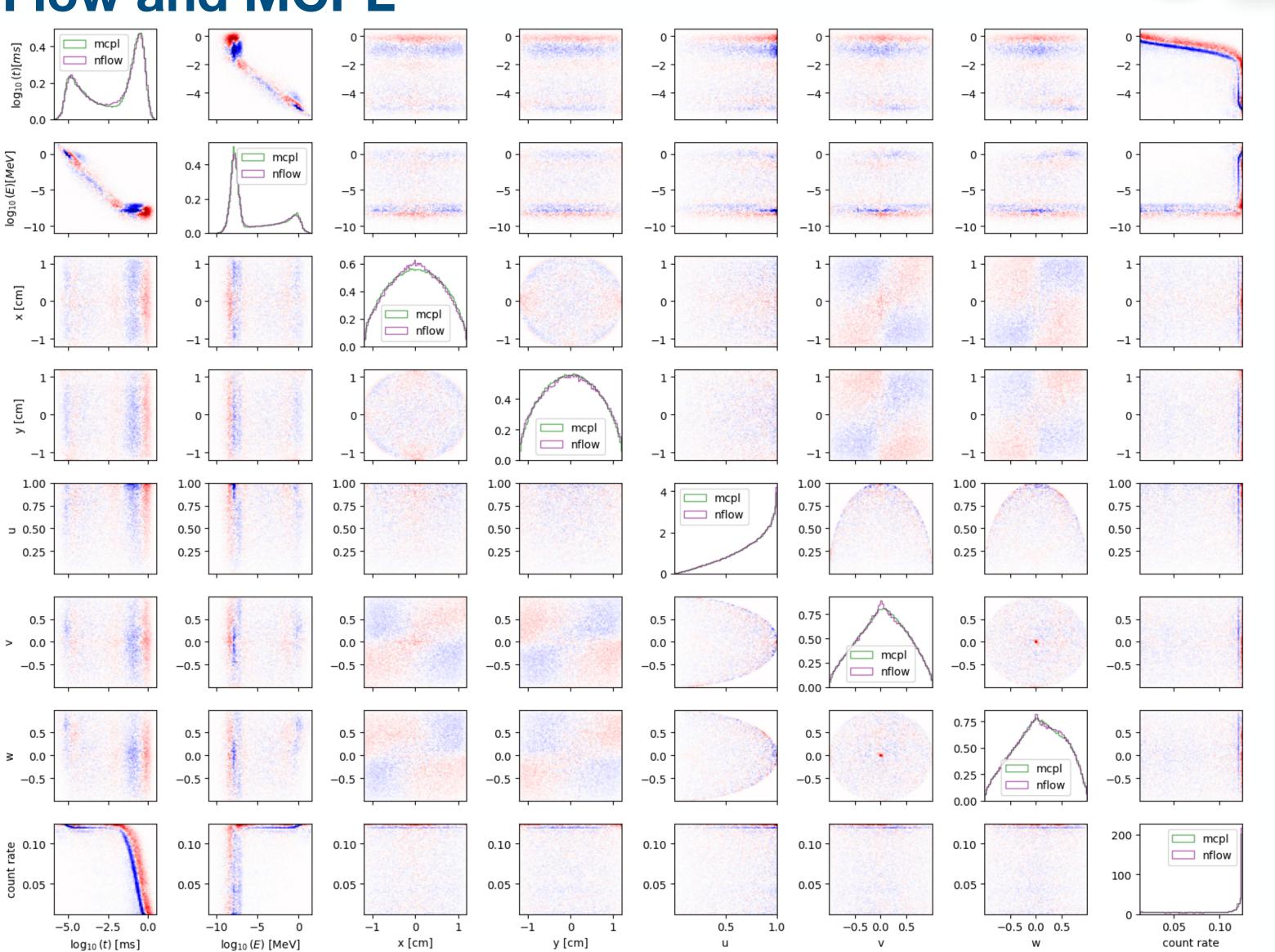
Blue: under-estimate

Red: over-estimate

model can be stored in a file consisting of few KB

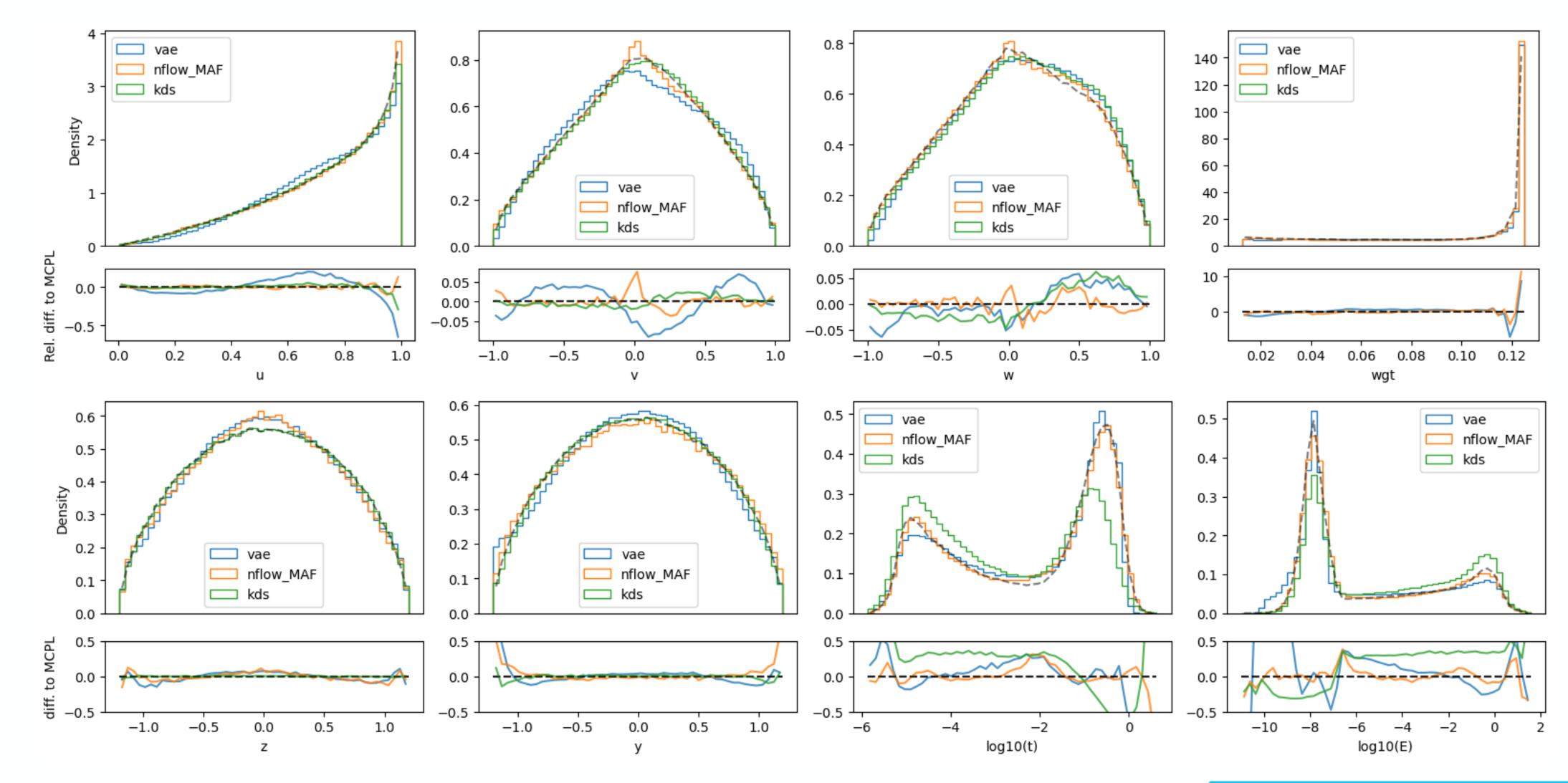
easily loadable through PyTorch API

we can sample from the latent space



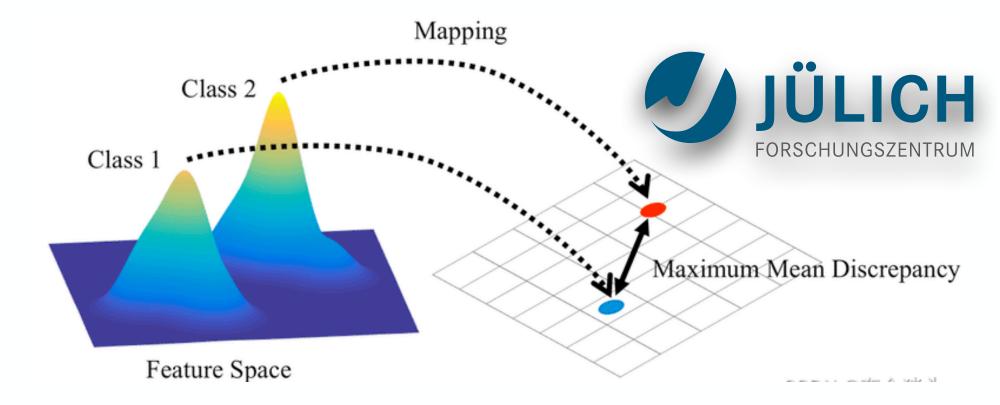


## COMPARISON BETWEEN MODELS









#### Statistical measure: Maximum mean discrepancy (MMD)

Determine if two datasets are likely to have been drawn from the same distribution by embedding probability distributions into a Reproducing Kernel Hilbert Space (RKHS) and then calculating the distance between the means of these embeddings.

Model	Average MMD *
MCPL	0.00015 ± 0.00016
MAF NFlow	0.00053 ± 0.00010
Coupling Flow	0.00171 ± 0.00015
VAE	0.00308 ± 0.00017
KDSource	0.08014 ± 0.00143
Uniform	0.15231 ± 0.00149

<sup>\*</sup> sample size = 10000, average over 10 samples

Model	Sampling time / 700.000 n (s)
KDSource	2
VAE	8
NFlow	7





### SUMMARY

- Generative models can learn multivariate probability distributions from data
- MCPL files make great training data!
- Normalizing Flows show great potential in learning neutron phase-space variable distributions, but they can estimate poorly if distributions have sharp features.
- These sources can already be used in Vitess and McStas, and are easily extensible to other Monte Carlo software.
- There are multiple architectures of NFs, as well as VAEs and GANs. Exploring which model does best is still an art. Physical constraints can be added inside the loss function.
- Metrics for comparing multivariate distributions should be taken into account for model selection.



