

Local 5.1 kick-off

Carsten Søgaard

Introduction

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Local 5.1 kick-off

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Overview

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Plan



GEM principle

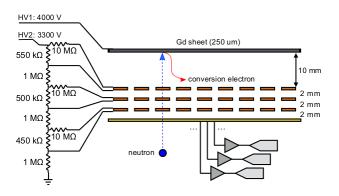
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Tracks

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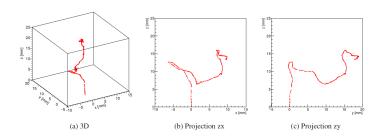
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70 keV electron in Ar/CO_2 (70/30) atmosphere





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Two beam-tests have been performed by Dorothea Pfeiffer et al.

Two different drift volumes have been tried:

- 10 mm drift volume
- 5 mm drift volume

Single events can be displayed graphically using 'uTPCViewer' written by Dorothea. The program also performs an analysis of the tracks and determines the origin. The found origin can occasionally be questioned.



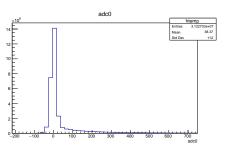
Negative ADC-values

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Status Plan A puzzling discovery is that a substantial part of the ADC-values are negative. Would expect positive integer values.



I have contacted Dorothea about his matter and she has forwarded the question. The answer agrees with my own thoughts - the negative values are a result of subtracting the pedestal.



A look at data and the current analysis

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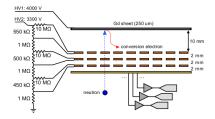
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Two beam tests were performed last year in Oslo. One in February and one in November.

The two datasets have different conditions:

- 10mm drift space no zero suppression
- 2 5 mm drift space zero suppressed





Successful origin

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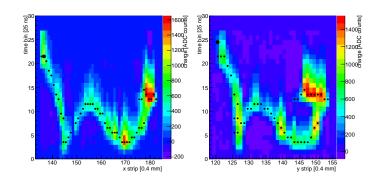
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Run 1 - 10mm drift space - no zero suppression





Questionable origin

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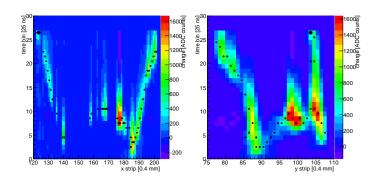
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Run 1 - 10mm drift space - no zero suppression





Wrong origin

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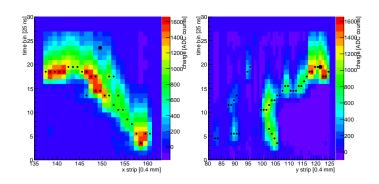
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Run 1 - 10mm drift space - no zero suppression





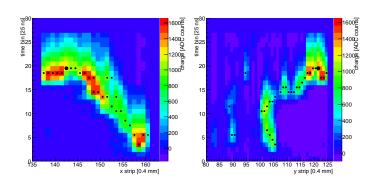
Wrong origin

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Run 1 - 10mm drift space - no zero suppression



Better result by increasing threshold to 200 (150 default)



Gamma events

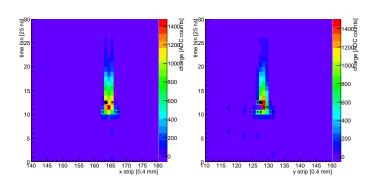
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Status Plan Run 1 - 10mm drift space - no zero suppression

Typical event from run, where gamma's from an Am-source were detected by the GEM.



Gamma's produce 'tracks' which in general are narrow and never 'begin' at high time-bin values.





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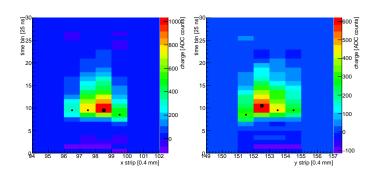
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Run 2 - 5 mm drift space - zero suppressed







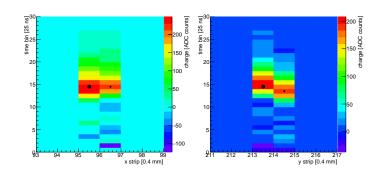
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Run 2 - 5 mm drift space - zero suppressed





Run 2 - Gamma events?

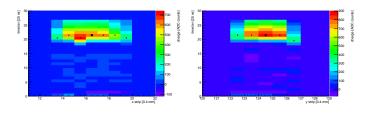
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Run 2 - 5 mm drift space - zero suppressed



Gamma events appear to spread over more strips than neutron events !



The code

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The code is generally doing what we expect. Improvements can probably be made. There should be an investigation of the impact of tuning thresholds etc.

I have started analysing the code provided by Dorothea in order to understand how the algorithm is implemented.

The current code is very monolithic (main analysis function \sim 400 lines long). This has some consequences:

- Difficult to read and get an overview of the code
- Cumbersome to maintain and improve
- To some extent error prone (copy-paste code x-y directions)



Understanding of the code so far

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Status Plan The main analysis function consists of a major loop running over all hit strips (some filtering has happened).

The main function loops over all strips:

- A loop 'backwards' over time bins finds the 'latest' time bin over a threshold
- 2 The local maximas are determined. A maximum must be a maximum over more than 3 time bins
- 3 Then the first time bin over $\frac{1}{2}$ the maximum amplitude is located at present unclear to me for what purpose
- Time bin no, amplitude and other values are stored i relation to the above
- Several quantities are calculated
- 6 A final algorithm does some searching presumably for the entry point



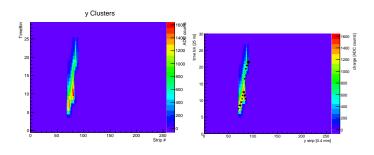
Reading root-files

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Status Plan I have written a class which can read the root-files.





Tentative plan

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Plan

Current and tentative plan - suggestions are more than welcome

- Fully understand the current algorithm and its implementation
- Adjust the algorithm and rewrite the entire code omitting ROOT classes
- Setup test environment and begin running the algorithm in a realistic setup
- Complete code for simulating tracks of low energy electrons in a narrow volume



Thoughts - ouch!

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The two tested conditions of the GEM's provide quite different data.

The VMM3 chip will provide less information - only time-stamp and amplitude per strip (possible a little more information)

Since no final decision on the detector has been made, we should pin point a small number of 'most likely scenarios' and produce the required algorithms and code.