

# The primary Instrument of the BIFROST Spectrometer

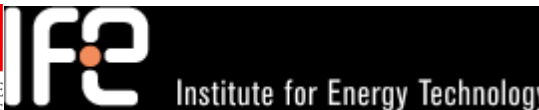


EUROPEAN  
SPALLATION  
SOURCE

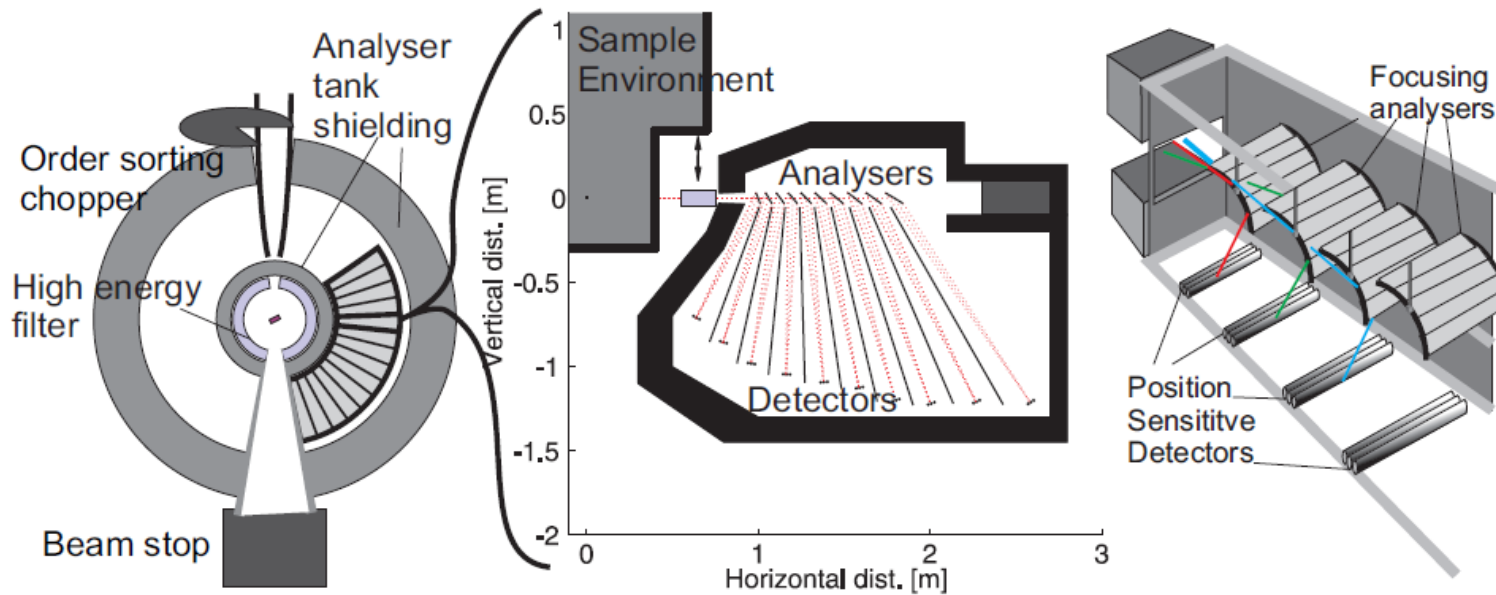
Rasmus Toft-Petersen  
BIFROST  
Technical University of Denmark



PAUL SCHERRER INSTITUT



# BIFROST

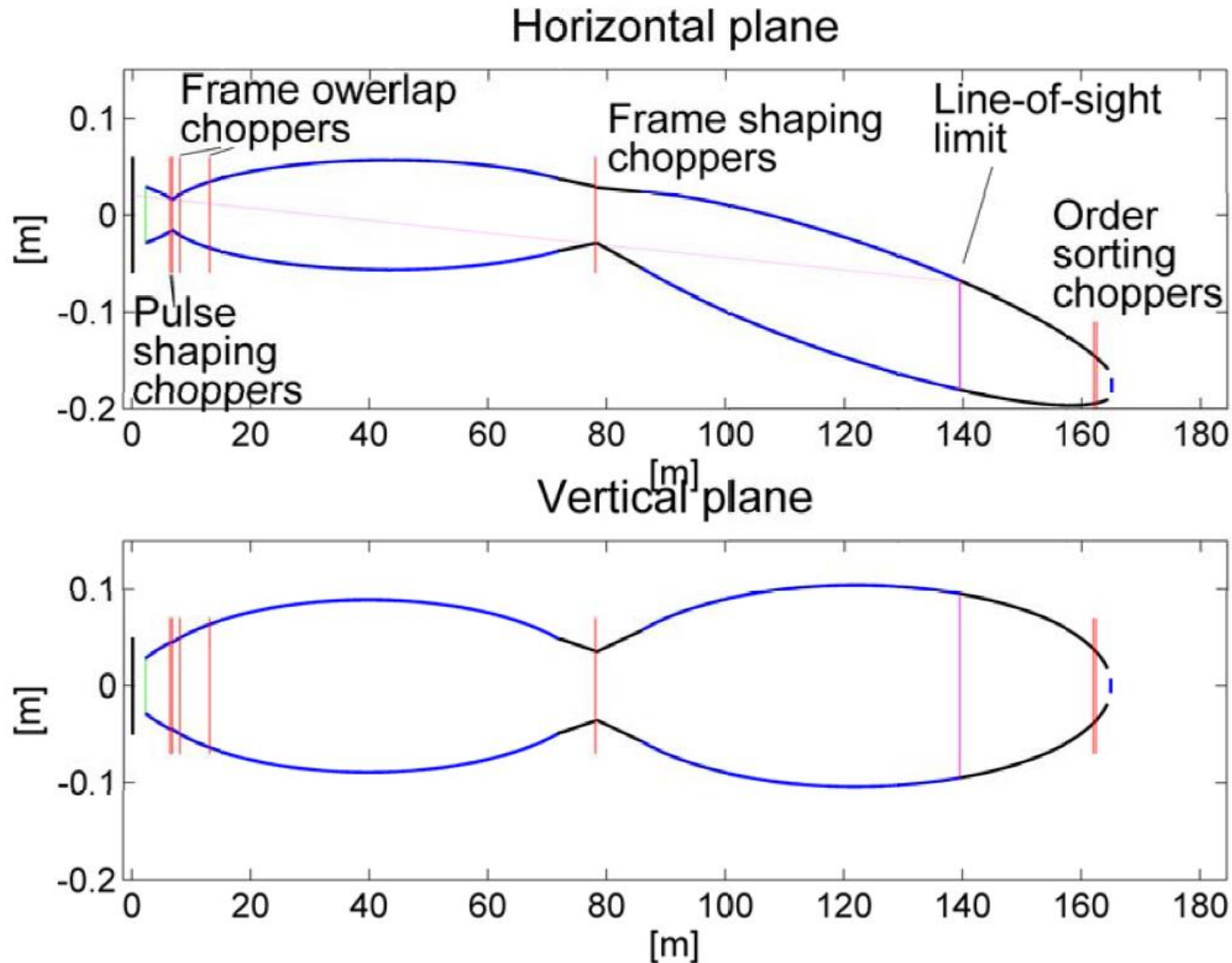


$E_f = 2.3 - 4.7$  meV,  
up to 8.3 meV?

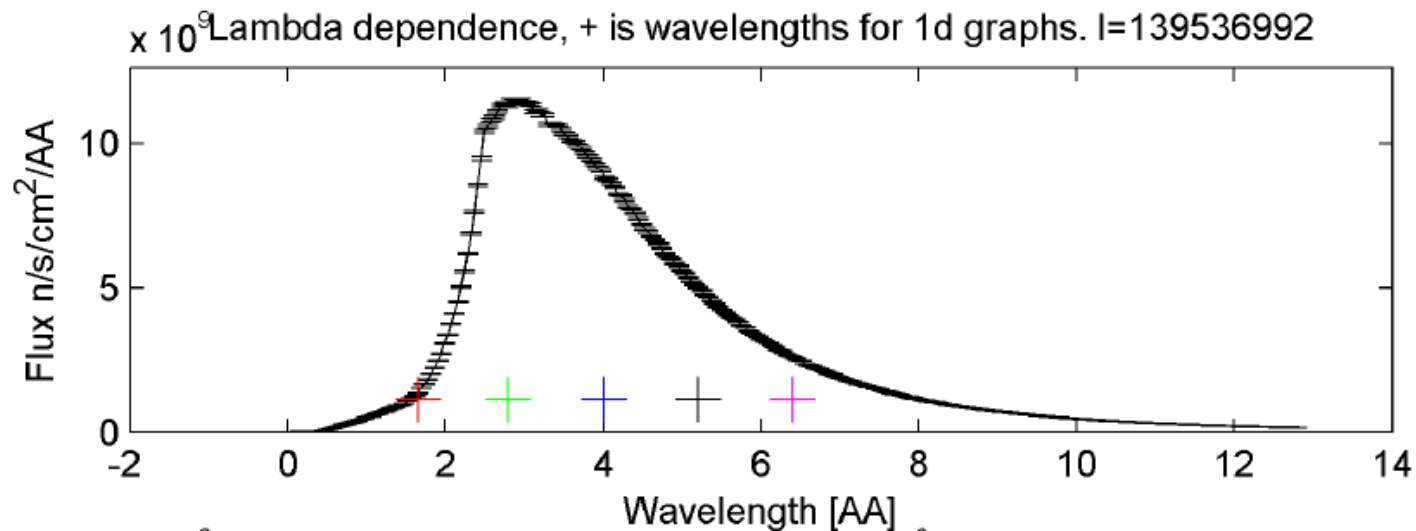
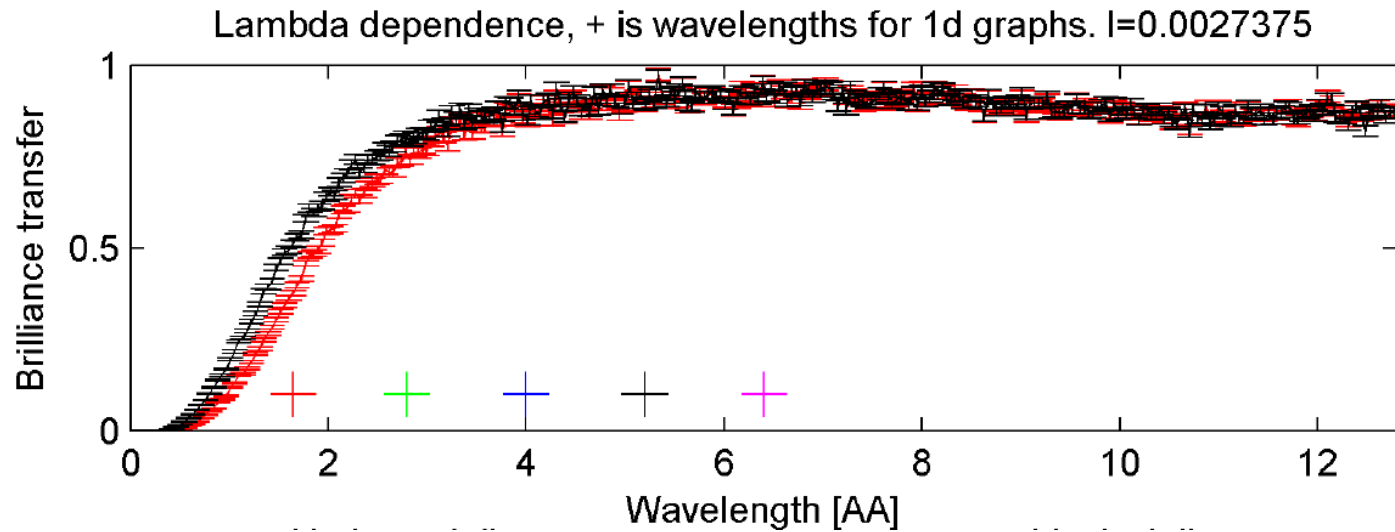
$k_f = 1.05 - 1.5 \text{ \AA}^{-1}$

- Efficient use of pyrolytic graphite and a long primary flight path is a powerful cocktail (at least in one scattering plane)
- Scope is focused on extreme sample environment but there is plenty to do without a 20 T magnet
- Our primary concern is the "order sorting" for the analyzers

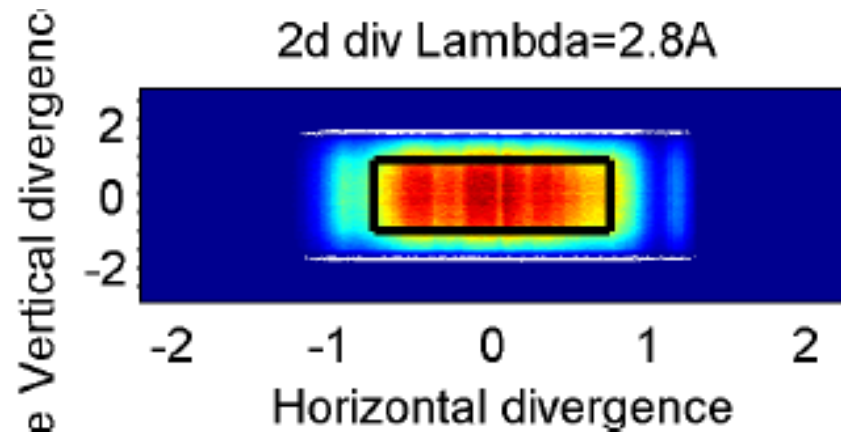
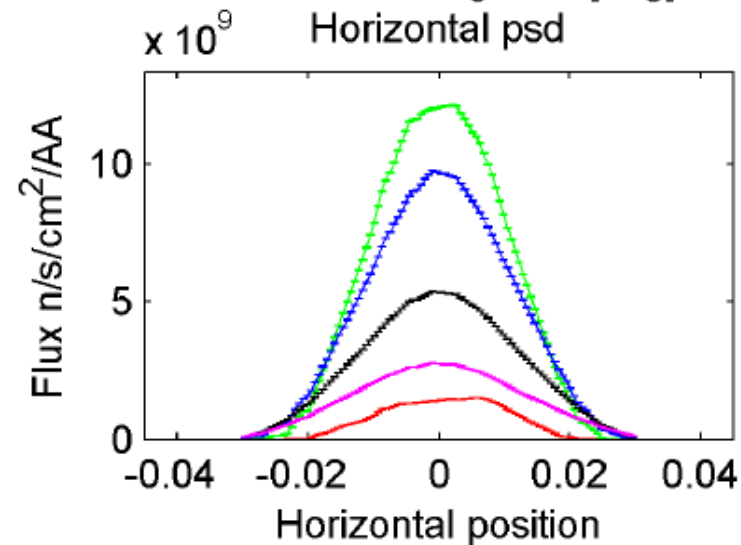
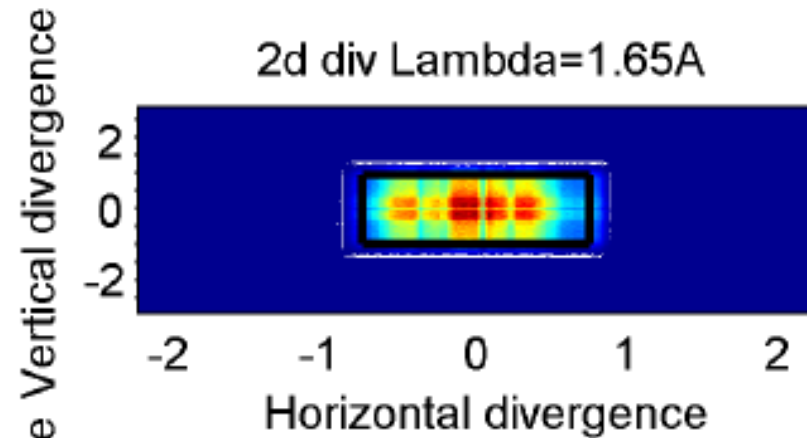
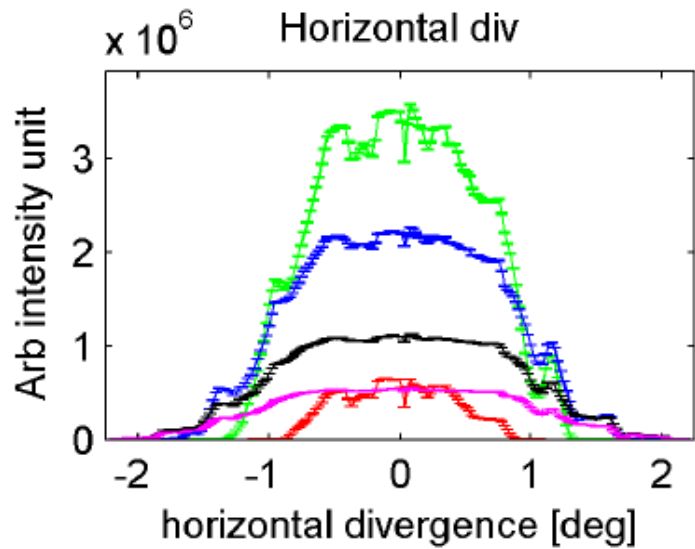
# Guide design



# Flux



# Divergence and spatial distribution



# Acceptance band of the BIFROST back end

The resolution is determined by Bragg's law and acceptance angles. Mosaicity provides the acceptance

$$k_I = \frac{\tau}{2 \sin \theta} \Rightarrow \Delta k_I = -k_I \cot \theta \Delta \theta$$

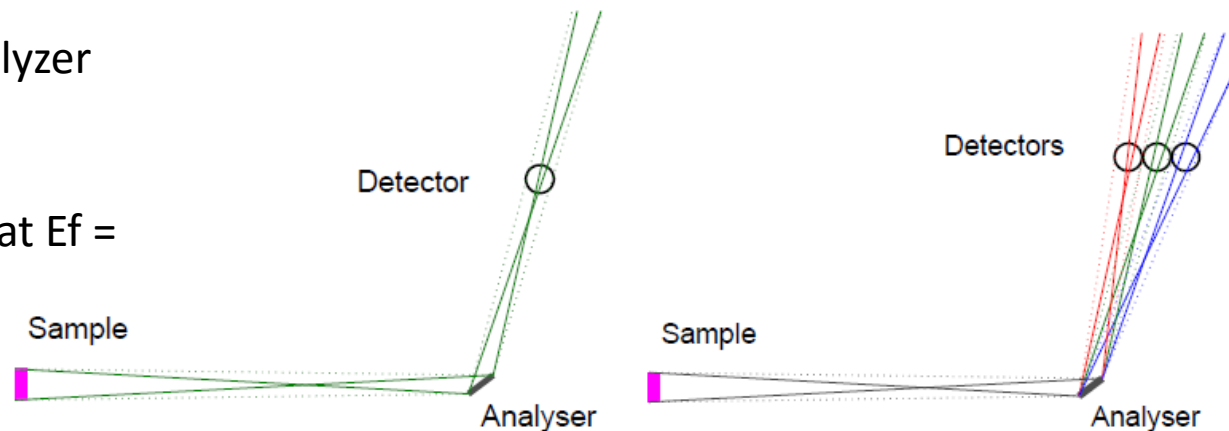
$E_f$	FWHM [ $\mu\text{eV}$ ]	$I_{peak}$	$I_{int}$
2.5 meV	57(2)	3.9(1)	220(9)
3.0 meV	79(1)	4.7(1)	374(7)
3.5 meV	129(2)	3.0(1)	385(10)
4.0 meV	182(4)	1.8(1)	331(15)
4.5 meV	209(5)	1.7(1)	311(20)

The analyzer take-off angle completely dominates the resolution.

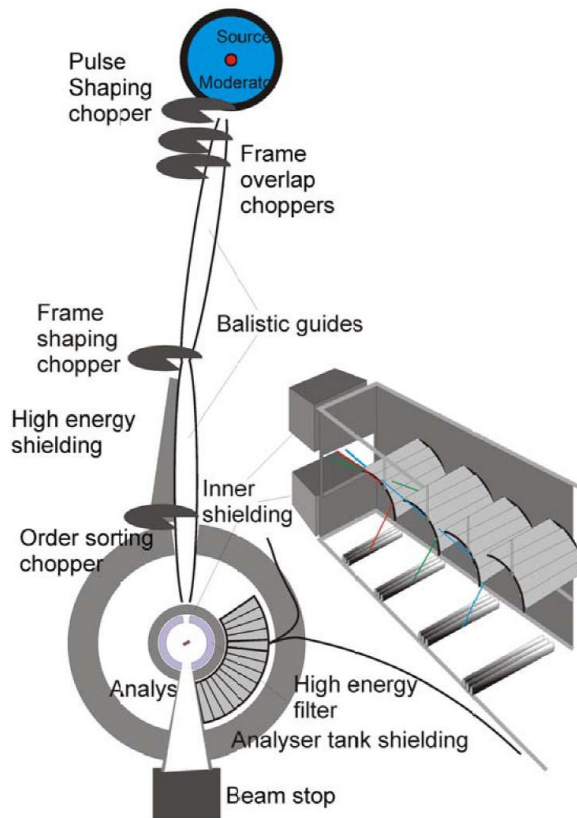
Jonas Okkels Birks prismatic analyzer concept.

Resolution down below 80  $\mu\text{eV}$  at  $E_f = 4.5 \text{ meV}$ .

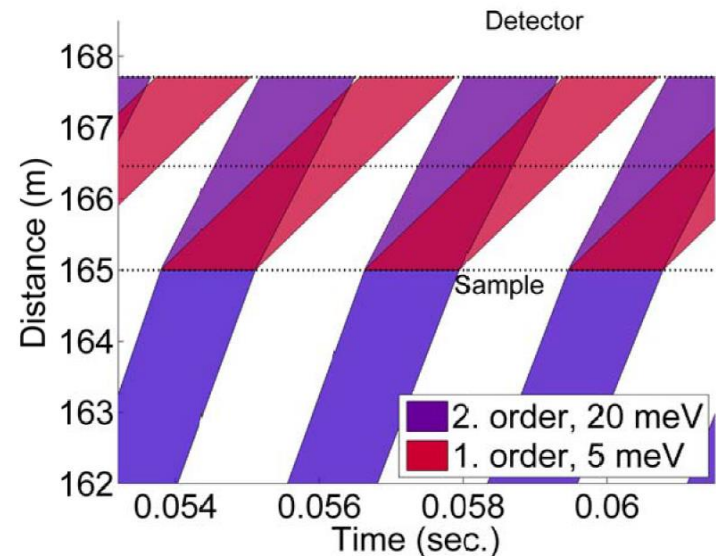
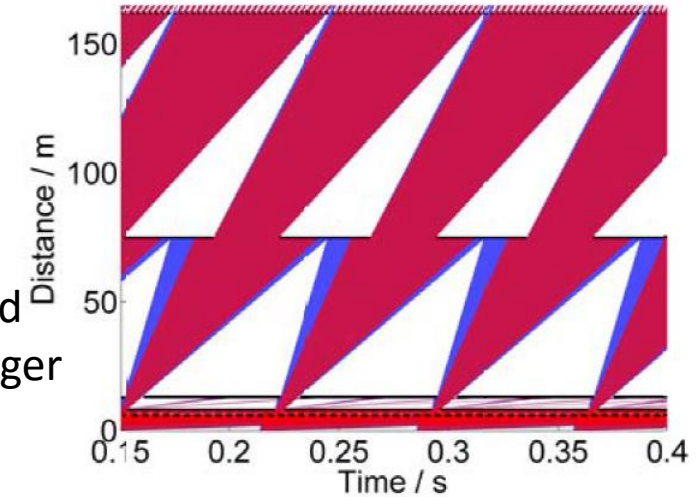
That acceptance bandwidth is a maximum for the setup and it is cheap.



# Chopper system and order sorting



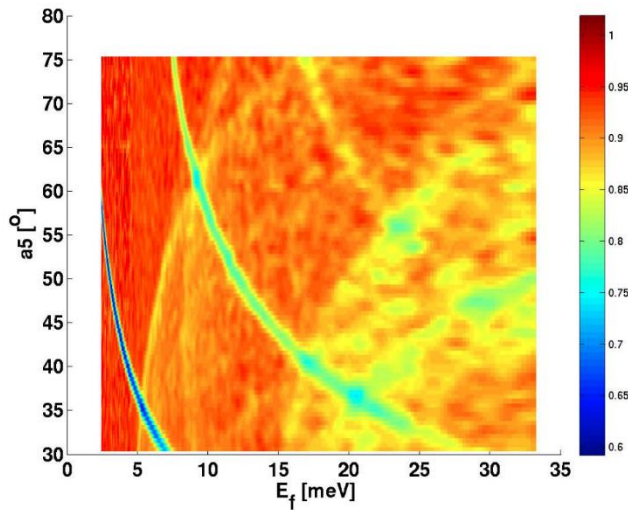
Pulse shaping choppers (210 Hz)  
Allows unprecedented resolution even at larger energy transfers



At larger  $E_i$  and energy transfers, the momentum transfer range become a problem as  $|\mathbf{k}_i - \mathbf{k}_f| < Q < |\mathbf{k}_i + \mathbf{k}_f|$ .  
is simply too weird for the first order setting.

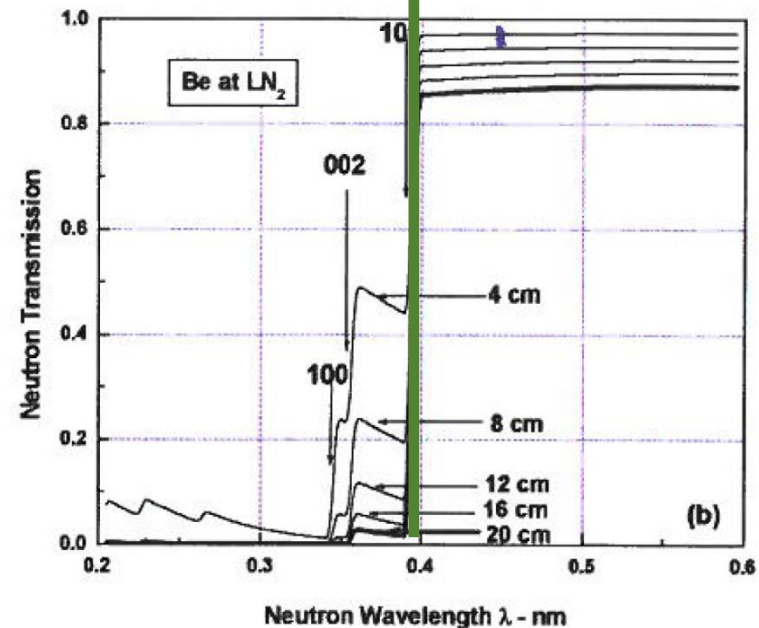
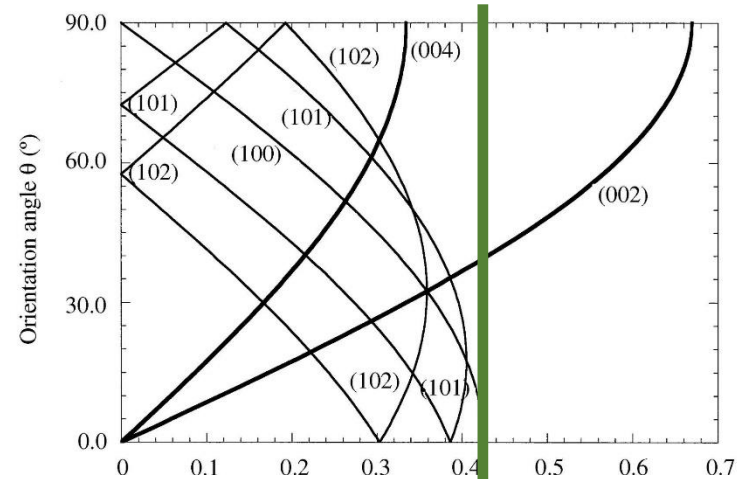


# Problems using 2nd order reflections in HOPG



Reflectivity goes down at short wavelengths. Parasitic reflections become prominent and the intensity has to go somewhere.

We need a translatable filter in any case. With a filter there is no problem at all. Spurious from PG are well understood





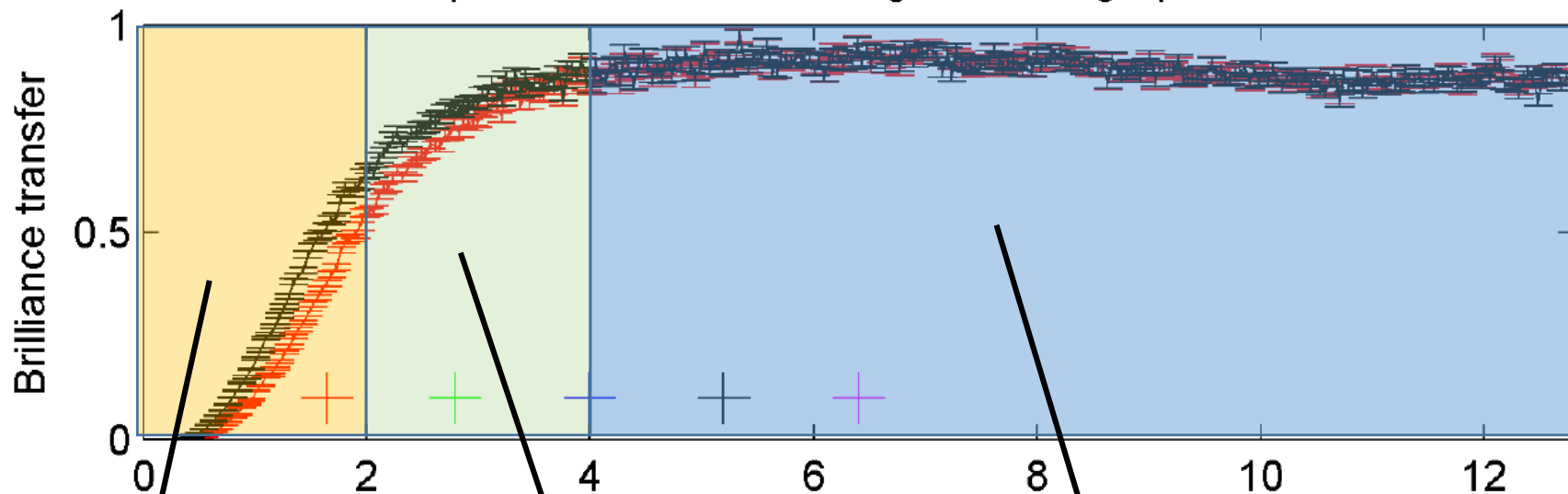
# Utilizing of high incoming energies at BIFROST

$$\Delta k_I = -\frac{k_I}{2} \cot \theta_A (\gamma_1 - \gamma_0)$$

Going second order is cheap – the resolution would only be quadrupled. It opens up a whole new range of stuff to do: Complex high energy dispersions

5  $\mu$ B in 20 T is roughly 15 meV

Lambda dependence, + is wavelengths for 1d graphs.  $l=0.0027375$



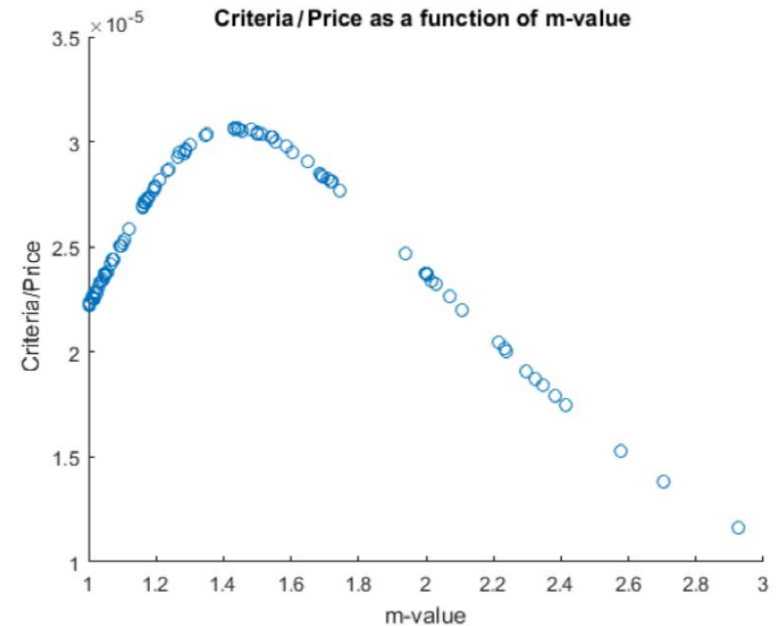
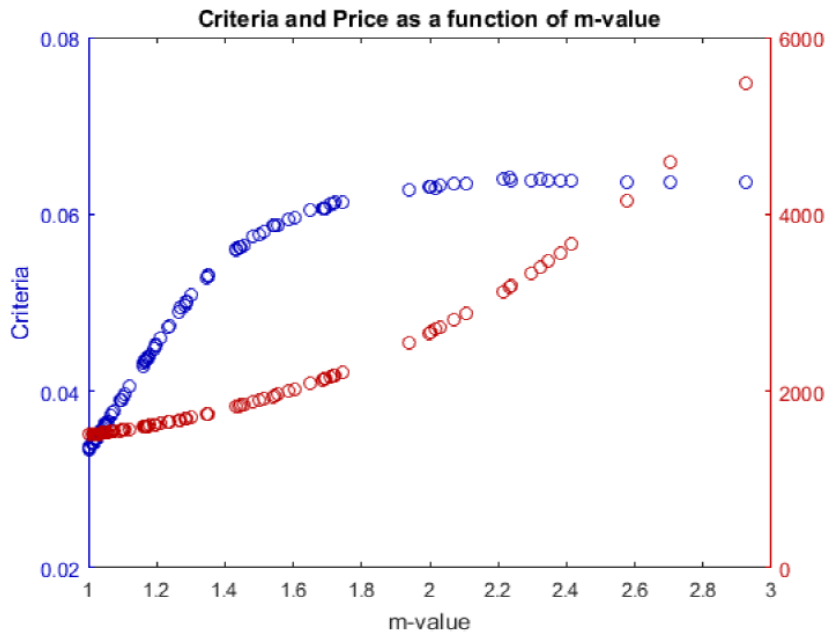
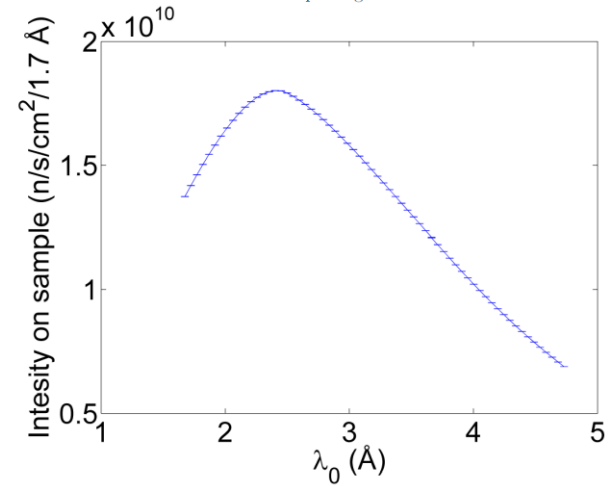
Risky, but with game changer potential

Work horse range

Suboptimal range

# Utilizing of high incoming energies at BIFROST

We need to find out exactly how expensive it is to keep the order sorting option on the table. Can we save money by sacrificing only the short wavelengths ?



# The most expensive posts

- Guide – 1.3 M€
- Shielding – 1.1 - 2.1 M€ (uncertain)
- Second order:  
PG analyzers, chopper,  
end shielding (2 M€)
- Vacuum tank (1.1 M€)
- Polarization  
analysis (2.1 M€)

Optimize with respect to price and examine if short wavelengths can be reasonably sacrificed

Examine the price of heavy shielding at guide end

Restrict ourselves below the Be cutoff, save 1 chopper plus shielding

Stick to argon overpressure?

Keep it as an option, examine discount solutions

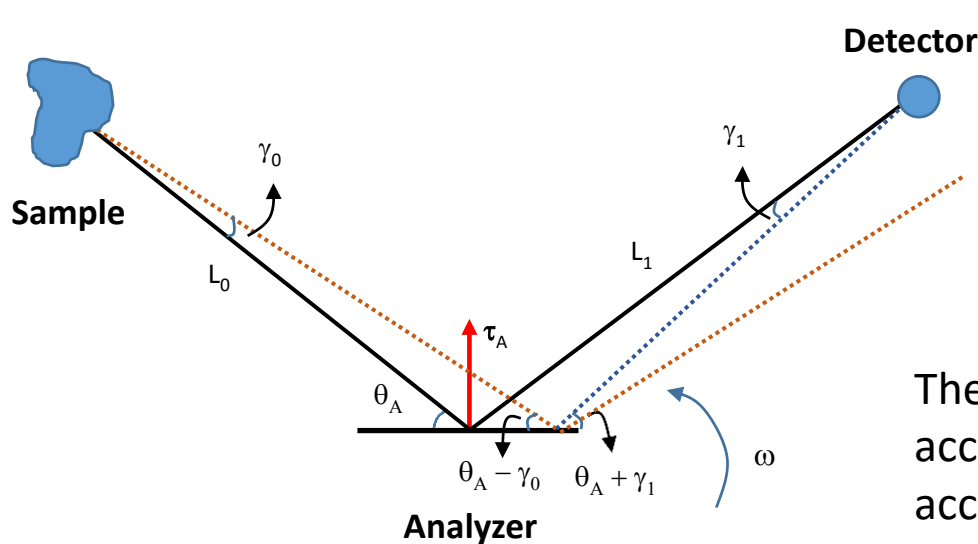
# Conclusions

- The short wavelengths may be expendable
- Removing order sorting can cut costs via less shielding at the endpoint, less PG, a saved chopper and a cheaper guide
- Price of shielding a question mark that we need to address

# Guide design

Component	length of coating segment	coating value	position relative to moderator
Feeder	1.74 m	3	2.16 m - 3.90 m
Feeder	1.74 m	3.5	3.90 m - 5.63 m
Feeder	0.87 m	3	5.63 m - 6.5 m
Ellipse	6.52 m	3.5	6.6 m - 13.12 m
Ellipse	6.52 m	2	13.12 m - 19.64 m
Ellipse	39.12 m	1.5	19.64 m - 58.78 m
Ellipse	6.52 m	2	58.78 m - 65.28 m
Ellipse	6.52 m	3	65.28 m - 71.80 m
Straight	13.94 m	2	71.80 m - 85.74 m
Ellipse	15.73 m	2	85.74 m - 101.47 m
Ellipse	47.20 m	1	101.47 m - 148.67 m
Ellipse	7.87 m	2	148.67 m - 156.53 m
Ellipse	7.87 m	3.5	156.53 m - 164.4 m

# Acceptance band of the BIFROST back end



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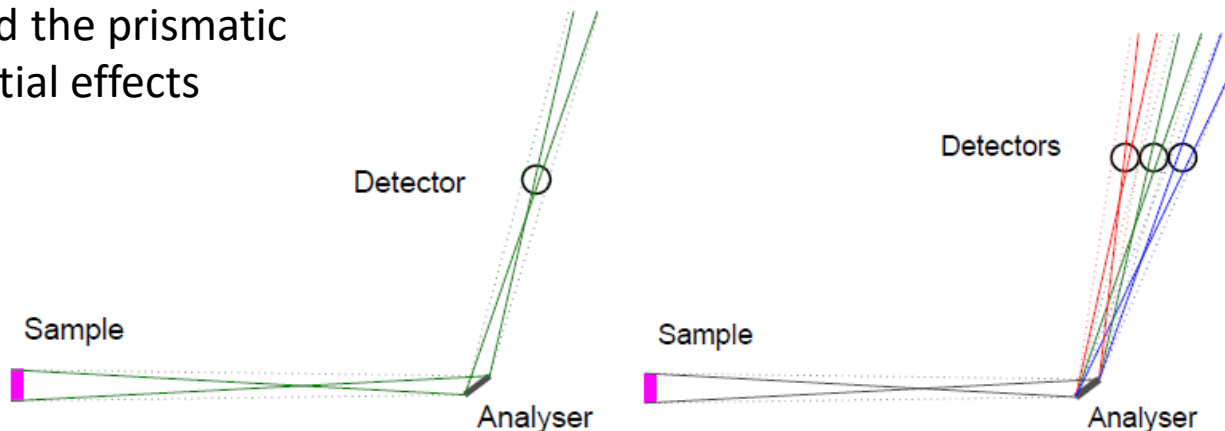
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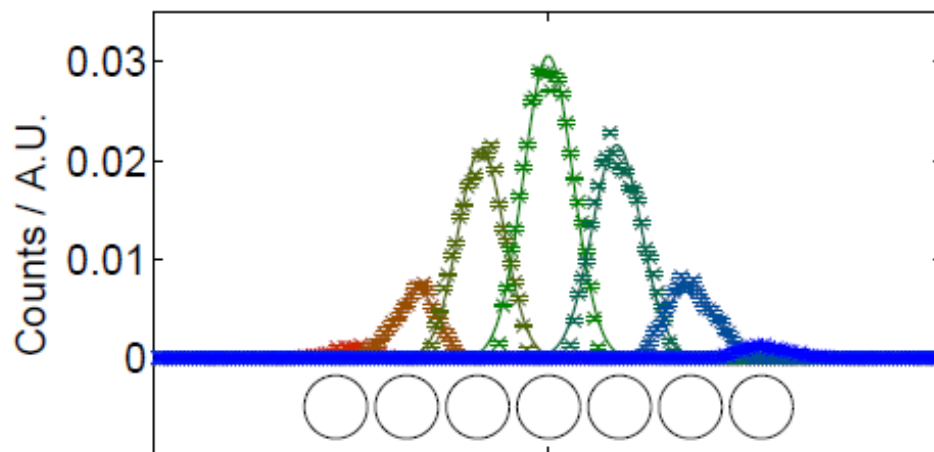
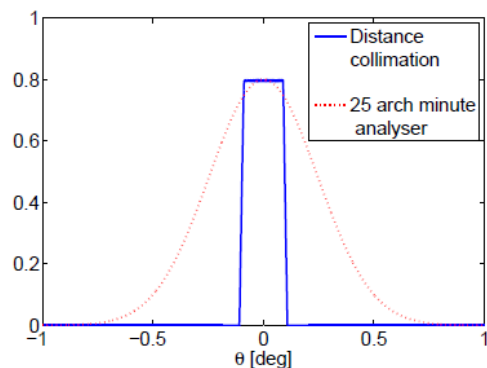
# Improving this acceptance band

Jonas Okkels Birk developed the prismatic analyzer concept when spatial effects are minimized.

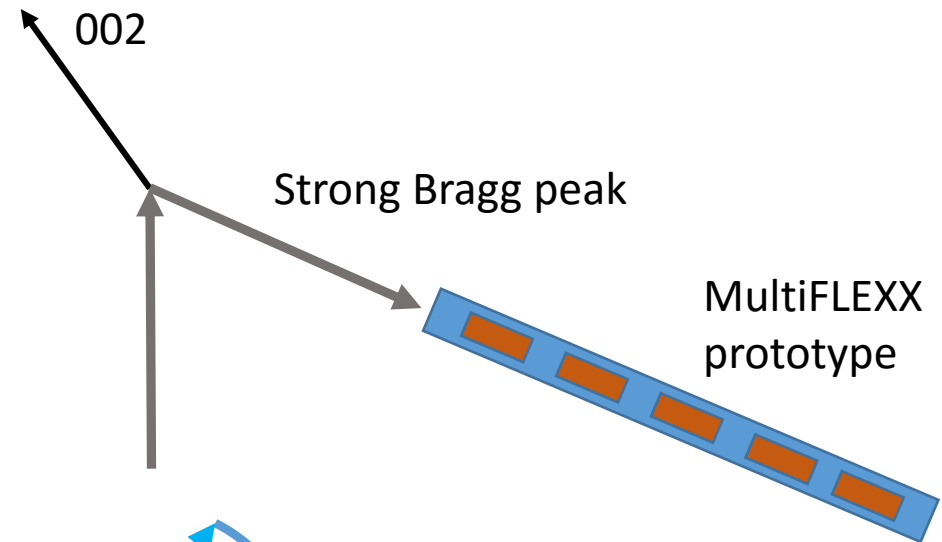
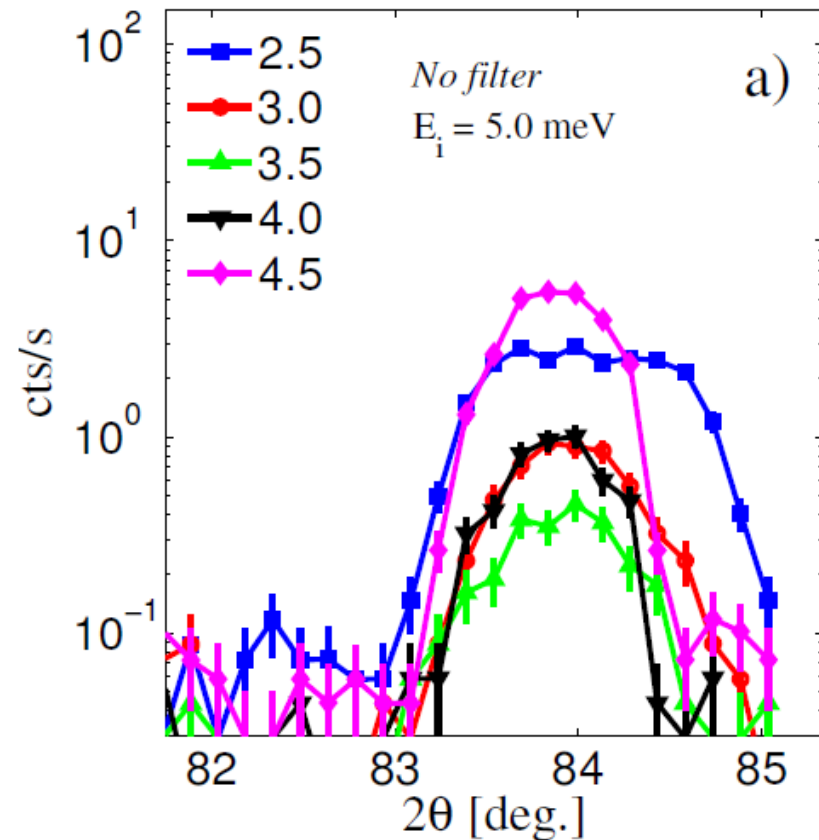
Resolution  
at the  $E_f = 4.5$  meV  
Analyzer  $\rightarrow 80 \mu\text{eV}$ .



That acceptance bandwidth is a maximum for the setup and it is cheap. It gives a lot of perspective.



# Problems on the TAS versions

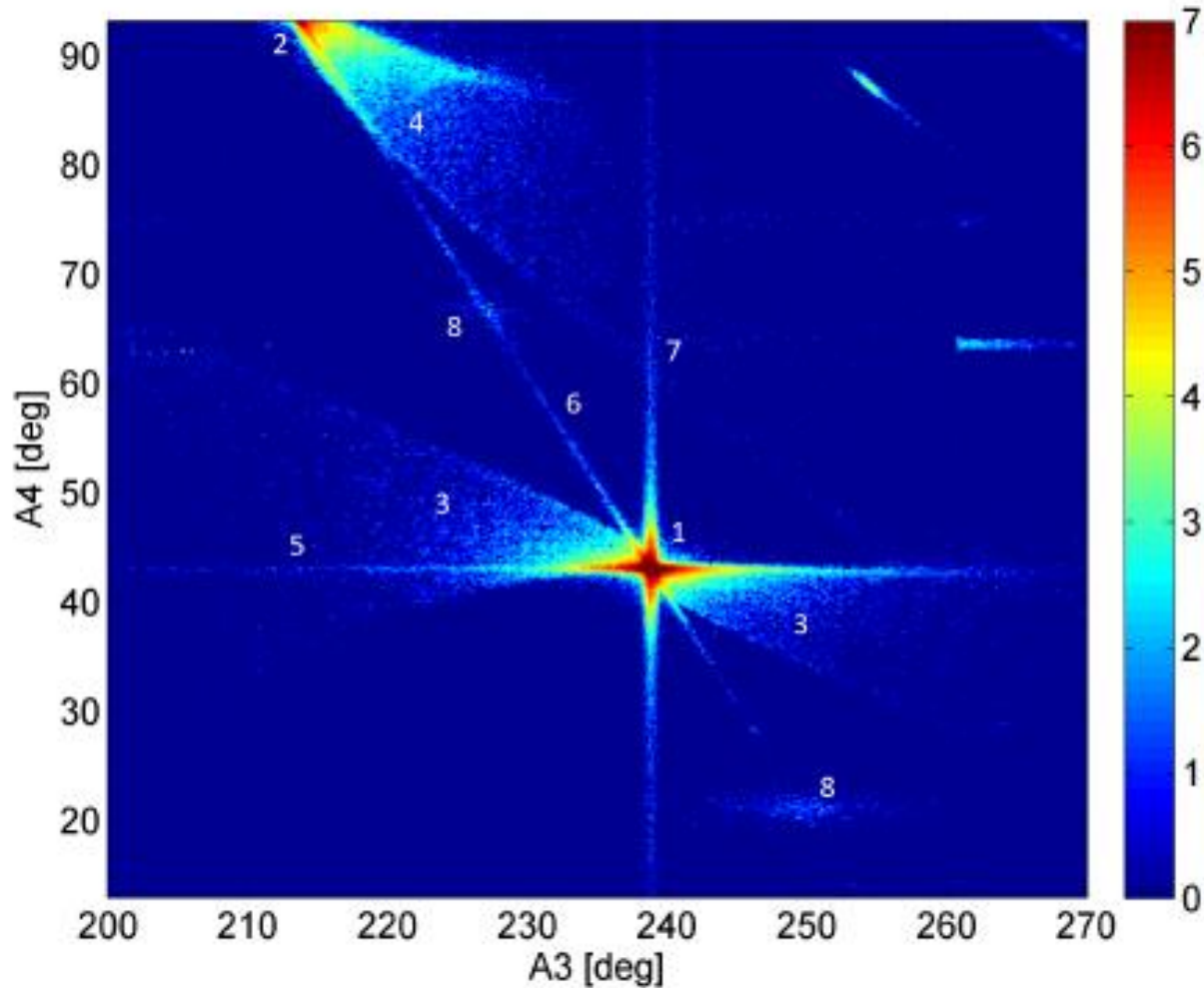


White beam =>  
Bragg peaks become  
Lines in  $\Theta, 2\Theta$  space

Thermal beam =>  
Loads of Bragg peaks

Low symmetry systems  
would generate a universe  
of spurions

# Characterizing HOPG



Beautiful DMC data  
taken by master student  
Rasmus Laurberg Hansen

Well understood by  
McStas simulations

