

## **MIRACLES.** Technical description

**Scope Setting Meeting** 

### Consorcio ESS-BILBAO

October 20, 2016

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October 20, 2016

1 / 19

## Presentation index

General overview



Chopper cascade





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# **General Overview**

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## MIRACLES

#### General overview

- Located in W5 (between MAGIG and BIFROST).
- Proposal concept updated to adapt to the butterfly moderator. The 3 cm height butterfly is considered for the redesign.

#### Instrument concept and performance

- The guide starts at 2.0 m from the moderator, and transports the beam to the sample, placed at 162.5 m.
- Chopper cascade
  - PWD pair to adapt flux-resolution (cutting the long pulse of ESS to improve the spectral resolution).
  - PS pair to select a single frame per source period.
  - WBD/FO choppers to select the wavelength band for each experiment, and prevent the frame overlap.

• Analyser: spherical, radius 2.5 m. Used to select  $E_f$ : Si 111 ( $\lambda_f$ =6,27Å), Si333 ( $\lambda_f$ =2,08Å), and Si311 as an update ( $\lambda_f$ =3,27Å).

## General overview

### Energy and momentum transfer range. Si 111



Energy transfer-Momentum transfer of MIRACLES

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- 2

## General overview

### Energy and momentum transfer range. Si 111



## General overview

#### Secondary spectrometer

- Spherical analyser with R=2.5m. Near backscattering geometry ( $\theta_B \approx 88^\circ$ ).
- Scattering angle covered by the analyser  $\approx$  155 °. Vertical viewing window of  $\pm$  22 °.
- Beryllium filter to avoid  $S_{i_{333}}$  (2.08Å) while selecting  $\lambda_f = 6.27$ Å with the  $S_{i_{111}}$ . Not decided yet, as in the case of the radial collimator.
- 2 cylindrical arrays with at least 80  $He^3$  detectors each, covering a scattering angle between 10 and 165 °. 2 ° separation between detectors.



Source: N. Tsapatsaris et al Review of Scientific Instruments 87, 085118 (2016); doi: 10.1063/1.4961569 E.Mamontov. A time-of-flight backscattering spectrometer at the SNS, BASIS

# **Chopper cascade**

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## Chopper cascade





Source: N. Tsapatsaris et al Review of Scientific Instruments 87, 085118 (2016); doi: 10.1063/1.4961569

## Chopper cascade

#### Updated concept of the chopper cascade

Chopper Id	Position	Updated parameters. Values are in meters, degrees and Hertz
Ch1	7.0	R=0.35; Freq=252; slit opening= 7.95 (x2) and 40 (x2), window height=0.124
Ch2	7.01	R=0.35; Freq=-252; slit opening= 7.95 (x2) and 40 (x2), window height=0.124
Ch3	7.5	R=0.35; Freq=28 $\rightarrow$ 56; slit opening= 14.3, window height=0.124
Ch4	7.51	R=0.35; Freq=-28 $\rightarrow$ -56; slit opening= 14.3, window height=0.124
Ch5	22.5	R=0.35; Freq=14 $\rightarrow$ 56; slit opening= 55, window height=0.124
Ch6	56.25	R=0.35; Freq=14 $\rightarrow$ 56; slit opening= 125, window height=0.124
Ch7	80.5	R=0.35; Freq=14 $\rightarrow$ 56; slit opening= 167, window height=0.124

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October 20, 2016 10 / 19

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#### General overview

- Maximize brilliance transfer to a 3cm  $\times$  3xm sample, within a solid angle of 7.62  $\cdot$  10<sup>-3</sup> sr.
- The first part of the guide is a double trumpet (converging-diverging guide).
- The last part of the guide is converging.
- Curved section to avoid direct view of sight and avoid neutrons below 1.5 Å.



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### New guide configuration

ID	Description	Values
G1	Guide	$w_1=0.06; h_1=0.076; w_2=0.04; h_2=0.12; m=3; L=4.99$
G2	Guide	$w_1 = 0.04$ ; $h_1 = 0.12$ ; $w_2 = 0.043$ ; $h_2 = 0.12$ ; m=3; L=0.47
G3	Guide	$w_1=0.043; h_1=0.12; w_2=0.11; h_2=0.12; m=3; L=9.04$
G4	Guide	$w_1 = 0.11; h_1 = 0.12; w_2 = 0.11; h_2 = 0.12; m = 1.5; L = 5.94$
G5	Guide	$w_1=0.11; h_1=0.12; w_2=0.11; h_2=0.12; Radius=5000; m=2; L=32.5$
G6	Guide	$w_1 = 0.11; h_1 = 0.12; w_2 = 0.11; h_2 = 0.12; m = 1.5; L = 1.15$
G7	Guide	$w_1$ =0.11; $h_1$ =0.12; $w_2$ =0.11; $h_2$ =0.12; Radius=5000; m=2; L=22.5
G7 exit	Guide	$w_1=0.11; h_1=0.12; w_2=0.11; h_2=0.12; m=1.5; L=1.5$
G8	Guide	$w_1 = 0.11; h_1 = 0.12; w_2 = 0.11; h_2 = 0.12; m = 1.5; L = 75.0$
G9	Guide	$w_1=0.11; h_1=0.12; w_2=0.032; h_2=0.032; m=4; L=6.98$

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#### Results



Brilliance transfer for a 1x1 cm and for a 3x3 cm sample depending on the wavelength

At 6 Å with  $3\times3$  cm sample, we obtain BT=0.142, that is the value predicted by Ken.

#### Results



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#### Results



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#### Results with the updated primary spectrometer



Elastic flux at sample, depending on PWD frequency

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#### Results with the updated primary spectrometer



Elastic resolution of MIRACLES for each PWD frequency

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#### Results with the updated primary spectrometer



Elastic energy resolution against elastic flux

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